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A

COLLECTION OF PROBLEMS

IN

ILLUSTRATION OF THE PRINCIPLES

07

THEORETICAL MECHANICS.

BY WILLIAM WALTON, B.A. TRINITY COLLEGE, CAMBRIDGE.



"Examples give a quicker impression than arguments."—BACON.

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PREFACE.

The design of this Work is to facilitate the study of Theoretical Mechanics, by presenting to the student a systematic collection of Problems in illustration of the more important principles of the science. The want of any such treatise, it is believed, has been felt by many as a serious impediment to the acquisition of adequate ideas in this branch of mathematical philosophy. Much importance, it may be observed, was attached by the great discoverers of the mechanical theories to the full discussion of numerous problems, as will be evident from a reference to the works of the three Bernoullis, of Leibnitz, and of D'Alembert, and to the beautiful investigations scattered throughout so long a series of volumes of the St. Petersburgh Transactions, by the liberal hand of Euler.

The author of this volume has endeavoured, as much as possible, to direct the attention of the student to the original memoirs of which he has so largely availed himself. This he has done, partly, to enable the beginner to obtain more detailed information than is compatible with the nature of this work, on particular questions which may excite an interest in his mind: his chief object, however, has been, to offer every facility to those who have already overcome at least the elementary difficulties of the subject, for acquiring a practical familiarity with the historical development of the science. Although it be admitted that useful and exact knowledge

may be obtained from even an exclusive perusal of the concise and methodical treatises which are generally adopted for the purpose of academic instruction: yet it may be asserted with confidence, that an excessive adherence to such a system of study, must deprive the student of much delightful and most valuable information.

In regard to the mode in which the author of this treatise has completed the task which he has proposed to himself, he feels every degree of diffidence, and would willingly that it had been undertaken by an abler hand. In apology for the imperfections, of which either he is himself aware or which may have eluded his observation, he can plead only the fact of engrossing occupations, or of perhaps insufficient preparation for a work requiring greater research than was originally contemplated.

Many of the problems in this volume have been extracted, with appropriate modifications, from the Ancient Transactions of the various Academies and learned Societies of Europe; many have been selected from the Cambridge Senate-House Papers; and for not a few the author is under obligation to the contributions of his friends. In arriving at original sources of information, it is scarcely necessary to state that great assistance has been obtained from the historical matter of Lagrange's Mécanique Analytique, and from Montucla's Histoire des Mathématiques.

Cambridge, October 1842.

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STATICS.

CHAPTER I.

CENTRE OF GRAVITY.

LET dm represent an element of the mass of a body at any point x, y, z, referred to any three co-ordinate axes, rectangular or oblique, and let \overline{x} , \overline{y} , \overline{z} , denote the co-ordinates of the centre of gravity of the body; then the formulæ for finding the values of \overline{x} , \overline{y} , \overline{z} , are

$$\bar{x} = \frac{\int x dm}{\int dm}$$
, $\bar{y} = \frac{\int y dm}{\int dm}$, $\bar{z} = \frac{\int z dm}{\int dm}$,

the limits of the integrations being determined by the form of the body.

If the body be bounded by a surface expressible by a single algebraical equation in x, y, z, the evaluation of each of the expressions $\int xdm$, $\int ydm$, $\int zdm$, will require the performance of the operation of integration on a single function of x, y, z, between appropriate limits; if, however, the body be bounded by discontinuous surfaces, the evaluation of each of these expressions will require the integration between proper limits of several functions of x, y, z, corresponding to the several discontinuous surfaces; the sum of the definite integrals of these functions being the required value of the expression.

The idea of the centre of gravity of material bodies is due to Archimedes, by whom the centres of gravity of various areas were investigated in his treatise, entitled Ἐπιπέδων ἰσορροπικῶν ης κέντρα βαρῶν ἐπιπέδων. He likewise determined the centre of gravity of the parabolic conoid. Among the mathematical successors of Archimedes who have cultivated the science of the centre of gravity, may be mentioned Pappus¹, Guido Ubaldi², Lucas Valerius³, La-Faille⁴, Guldin⁶, Wallis⁶, Carré⊓, Varignon⁶, Clairaut⁶.

SECT. 1. Symmetrical Area.

Let x be the abscissa and y the ordinate of any point in the circumference of a plane area, symmetrical with respect to the axis of x; the axes of co-ordinates being either rectangular or oblique. Then the centre of gravity of any portion of this area intercepted between any assigned pair of double ordinates will lie in the axis of x, and its distance \overline{x} from the origin will be given by the formula

$$\overline{x} = \frac{\int xy \, dx}{\int y \, dx} \dots (I.),$$

where the integrations are to be performed between limits depending upon the positions of the intercepting ordinates.

The value of \bar{x} is sometimes more readily obtained by polar co-ordinates, when the formula will be

$$\overline{x} = \frac{\iint r^2 \cos \theta \ d\theta \ dr}{\iint r \ d\theta \ dr} \dots \dots (II.),$$

where r denotes the distance of any point within the area from the origin, and θ the inclination of r to the axis of x. The nature of the limits in the double integrations will depend upon the form of the area in each particular case.

- 1 Mathemat. Collect., lib. 8, published for the first time in 1588.
- In duos Archimedis Æquiponderantium libros Paraphrasis, 1588.
- De Centro Gravitatis Solidorum, 1604.
- 4 De Centro Gravitatis partium Circuli et Ellipsis Theoremata, 1632.
- Centrobaryca, 1635.
- Opera, tom. 1. cap. 4 et 5, 1670.
- ' Mesure des Surfaces, 1700.
- Mém. de l'Acad. des Sciences de Paris, 1714.
- Mém. de l'Acad. des Sciences de Paris, 1731, p. 159.

(1) To find the centre of gravity of the area of any portion **BAC** (fig. 1) of a parabola cut off by any chord **BC**.

Let Py be a tangent to the parabola at a point P parallel to the chord CB; from P draw Px parallel to the axis of the parabola. Then, Px and Py being taken as the axes of x and y, the equation to the curve will be

$$y^2=4mx,$$

m being the distance of the point P from the focus.

Hence, if PE = a,

$$\overline{x} = \frac{\int_{0}^{a} xy \, dx}{\int_{0}^{a} y \, dx} = \frac{\int_{0}^{a} x^{\frac{3}{4}} dx}{\int_{0}^{a} x^{\frac{1}{4}} dx} = \frac{\frac{2}{5}a^{\frac{3}{4}}}{\frac{2}{3}a^{\frac{3}{4}}} = \frac{3}{5}a.$$

Archimedes, Ἐπιπέδων ἰσορροπικῶν, Lib. 11. Prop. 8; Guldin, Centrobaryca, Lib. 1. cap. 9, p. 121.

(2) To find the centre of gravity of the area of the Cissoid of Diocles, *EAE'*, (fig. 2).

The equation to the curve is

$$y^{2} = \frac{x^{3}}{a - x};$$
hence
$$\bar{x} = \frac{\int xy \, dx}{\int y \, dx} = \frac{\int_{0}^{a} \frac{x^{\frac{1}{2}}}{(a - x)^{\frac{1}{4}}} \, dx}{\int_{0}^{a} \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx} \cdot \dots \cdot (1);$$
but
$$\int \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx = -2x^{\frac{1}{4}} (a - x)^{\frac{1}{4}} + 5 \int x^{\frac{1}{4}} (a - x)^{\frac{1}{4}} \, dx,$$
and therefore
$$\int_{0}^{a} \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx = 5 \int_{0}^{a} x^{\frac{1}{4}} (a - x)^{\frac{1}{4}} \, dx$$

$$= 5a \int_{0}^{a} \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx - 5 \int_{0}^{a} \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx = \frac{5}{6}a \int_{0}^{a} \frac{x^{\frac{1}{4}}}{(a - x)^{\frac{1}{4}}} \, dx;$$

hence from (1) we have

$$\overline{x} = \frac{5}{6}a \frac{\int_{0}^{a} \frac{x^{\frac{3}{4}}}{(a-x)^{\frac{1}{4}}} dx}{\int_{0}^{a} \frac{x^{\frac{3}{4}}}{(a-x)^{\frac{1}{4}}} dx} = \frac{5}{6}a.$$

(3) To find the centre of gravity of the sector ABC (fig. 3) of a circle, of which C is the centre.

From C draw the straight line CEx bisecting the sectorial area; and draw Cy at right angles to Cx. Let CE = a, and $\angle ACx = a$; then Cx, Cy, being the axes of x and y,

$$\overline{x} = \frac{\int_0^a xy \, dx}{\int_a^a y \, dx} \cdot \dots (1).$$

But the equations to the straight line CA, and to the circle of which AEB is an arc, are respectively

$$y = x \tan a$$
, $y^2 = a^2 - x^3$;

also CF is equal to $a \cos a$; hence

$$\int_0^a xy \ dx = \int_0^a \cos a \ \tan a \ x^2 dx + \int_{a \cos a}^a x (a^2 - x^2)^{\frac{1}{4}} \ dx \dots (2),$$

and

$$\int_0^a y \ dx = \int_0^{a \cos a} \tan a \ x \, dx + \int_{a \cos a}^a (a^2 - x^2)^{\frac{1}{4}} \, dx \dots (3).$$

Now by the ordinary processes of the Integral Calculus,

$$\int_0^{a\cos x} \tan a \, x^2 dx = \frac{1}{3}a^3 \sin a \cos^2 a,$$

and

$$\int_{a\cos a}^{a} x(a^{2}-x^{2})^{\frac{1}{2}} dx = \frac{1}{3}a^{3} \sin^{3} a;$$

hence from (2) we have

$$\int_0^a xy \, dx = \frac{1}{3}a^3 \sin a \dots \dots (4).$$

hence from (3) we have

$$\int_0^a y \, dx = \frac{1}{2}a^2a \ldots \ldots (5).$$

From the relations (1), (4), (5),

$$\bar{x} = \frac{\frac{1}{3} a^3 \sin a}{\frac{1}{3} a^2 a} = \frac{2}{3} a \frac{\sin a}{a}.$$

This result however may be obtained more readily by polar co-ordinates: let P be any point in the area of the sector; CP = r, $\angle PCx = \theta$; then

$$\bar{x} = \frac{\int_{-a}^{+a} \int_{0}^{a} r^{2} \cos \theta \ d\theta \ dr}{\int_{-a}^{+a} \int_{0}^{a} r \ d\theta \ dr} = \frac{\frac{1}{3} a^{2} \int_{-a}^{+a} \cos \theta \ d\theta}{\frac{1}{2} a^{2} \int_{-a}^{+a} d\theta} = \frac{2}{3} a \frac{\sin a}{a}.$$

We might have effected the double integration in a different order; thus

$$\bar{x} = \frac{\int_{0}^{a} \int_{-a}^{r_{a}} r^{2} \cos \theta \, dr \, d\theta}{\int_{0}^{a} \int_{-a}^{r_{a}} r \, dr \, d\theta} = \frac{2 \sin a \int_{0}^{a} r^{2} \, dr}{2a \int_{0}^{a} r \, dr}$$
$$= \frac{2 \sin a \cdot \frac{1}{3}a^{3}}{2a \cdot \frac{1}{2}a^{2}} = \frac{2}{3}a \frac{\sin a}{a} = \frac{2}{3} \frac{\text{radius} \times \text{chord}}{\text{arc}}.$$

According to the former order of integration, the sector ACB is conceived to be subdivided into an infinite series of infinitesimal triangles having a common vertex C, their bases being elements of the arc AEB; according to the latter order, we conceive the sector to be made up of a series of circular rings of indefinitely small breadth, having a common centre C.

Carré; Mesure des Surfaces, &c., p. 76.

(4) To find the centre of gravity of the segment AEBF (fig. 3) of a circle.

The construction and notation remaining the same as in the preceding example, produce CP to cut the chord AB in Q and the arc AEB in R.

Then by the formula (II.) for polar co-ordinates, if CQ = r',

$$\overline{x} = \frac{\int_{-a}^{+a} \int_{r'}^{a} r^2 \cos \theta \ d\theta \ dr}{\int_{-a}^{+a} \int_{r'}^{a} r \ d\theta \ dr} \cdots (1).$$

Now, since $r' = a \frac{\cos a}{\cos \theta}$, we have

$$\int_{r}^{a} r^{2} \cos \theta \ d\theta \ dr = \frac{1}{3} \left(a^{3} - r^{3}\right) \cos \theta \ d\theta = \frac{1}{3} a^{3} \left(1 - \frac{\cos^{3} a}{\cos^{3} \theta}\right) \cos \theta \ d\theta;$$

hence
$$\int_{-a}^{a} \int_{r}^{a} r^{3} \cos \theta \ d\theta \ dr = \frac{1}{3} a^{3} \int_{-a}^{a} \left(\cos \theta - \frac{\cos^{3} a}{\cos^{2} \theta} \right) d\theta$$
$$= \frac{1}{3} a^{3} \left(\sin \theta - \cos^{3} a \tan \theta \right), \quad \text{from } \theta = -a \text{ to } \theta = +a,$$
$$= \frac{2}{3} a^{3} \left(\sin a - \cos^{2} a \sin a \right) = \frac{2}{3} a^{3} \sin^{3} a \dots \dots \dots \dots (2).$$

Again,
$$\int_{r}^{a} r \, d\theta \, dr = \frac{1}{2} (a^{2} - r^{2}) \, d\theta = \frac{1}{2} a^{2} \left(1 - \frac{\cos^{2} a}{\cos^{2} \theta} \right) d\theta,$$

and therefore

$$\int_{-a}^{+a} \int_{r}^{a} r \, d\theta \, dr = \frac{1}{2} a^2 \left(\theta - \cos^2 a \tan \theta \right), \text{ from } \theta = -a \text{ to } \theta = +a,$$
$$= a^2 \left(a - \sin a \cos a \right) \dots \dots \dots \dots (3).$$

Hence from (1), (2), (3), we get

but

$$\overline{x} = \frac{2}{3}a \frac{\sin^3 a}{a - \sin a \cos a}.$$

This result may be obtained as easily by rectangular coordinates; thus, putting $a \cos a = a'$,

$$\overline{x} = \frac{\int_{a'}^{a} xy \, dx}{\int_{a}^{a} y \, dx}$$

$$y^{2} = a^{2} - x^{2};$$
(a)

hence
$$\int_{a'}^{a} xy \, dx = \int_{a'}^{a} x (a^{2} - x^{2})^{\frac{1}{4}} \, dx$$

$$= -\frac{1}{3}(a^{2} - x^{2})^{\frac{3}{4}}, \text{ between limits, } = \frac{1}{3}a^{3}\sin^{3}a \dots (b).$$
Again,
$$\int_{a'}^{a} y \, dx = \int_{a'}^{a} (a^{2} - x^{2})^{\frac{1}{4}} \, dx$$

$$= (a^{2} - x^{2})^{\frac{1}{4}} x + \int_{a'}^{x} (a^{2} - x^{2})^{\frac{1}{4}} \, dx, \text{ between limits,}$$

$$= (a^{2} - x^{2})^{\frac{1}{4}} x + a^{2} \int_{a'}^{a} \frac{dx}{(a^{2} - x^{2})^{\frac{1}{4}}} - \int_{a'}^{a} (a^{2} - x^{2})^{\frac{1}{4}} \, dx$$

$$= \frac{1}{2}(a^{2} - x^{2})^{\frac{1}{4}} x + \frac{1}{2}a^{2} \sin^{-1}\frac{x}{a}, \text{ from } x = a' \text{ to } x = a,$$

$$= -\frac{1}{2}a^{2} \sin a \cos a + \frac{1}{2}a^{2} a = \frac{1}{2}a^{2}(a - \sin a \cos a) \dots (c).$$
Hence from (a) , (b) , (c) , there results
$$\overline{x} = \frac{2}{3}a \frac{\sin^{3}a}{a - \sin a \cos a}.$$

In the integrations by polar co-ordinates the segmental area is conceived to be made up of frustums of an infinite number of infinitesimal triangles intercepted by the chord AB, C being the common vertex of the triangles, and a series of elements of the arc AEB being their bases; on the other hand, when rectangular co-ordinates are made use of, the segment is conceived to be made up of an infinite number of indefinitely thin parallelograms parallel to the chord AB.

Guldin; Centrobaryca, Lib. 1. cap. 9, p. 107.

(5) To find the centre of gravity of any portion of a semicubical parabola comprised between the curve and a double ordinate.

The equation to the curve being $ay^2 = x^3$, we shall have

$$\overline{x} = \frac{5}{7}x$$
.

(6) To find the centre of gravity of the whole area of the curve of which the equation is

$$y^2 = b^2 \frac{a - x}{x}.$$

$$\overline{x} = \frac{1}{2} a.$$

(7) To find the centre of gravity of the whole area of a cycloid.

The equation to the cycloid being

$$y = a \text{ vers}^{-1} \frac{x}{a} + (2ax - x^2)^{\frac{1}{2}},$$

we shall have $\bar{x} = \frac{7}{6}a$.

(8) To find the centre of gravity of a semi-ellipse, the bisecting line being any diameter.

If the bisecting diameter be taken as the axis of y, and the conjugate diameter as the axis of x, the equation to the ellipse will be

$$y^2 = \frac{b^2}{a^2}(a^2 - x^2),$$

and we shall have

$$\bar{x}=\frac{4a}{3\pi}$$
.

Guldin; Centrobaryca, Lib. 1. cap. 9, p. 115.

(9) To find the centre of gravity of a loop of the Lemniscata of James Bernoulli.

The equation to the curve being $r^2 = a^2 \cos 2\theta$, we shall have

$$\bar{x}=\frac{2^{\frac{1}{4}}\pi}{8}a.$$

SECT. 2. Area not Symmetrical.

The formulæ for the determination of the co-ordinates of the centre of gravity of an area not symmetrical with respect to either of the axes, are

$$\overline{x} = \frac{\iint x \, dx \, dy}{\iint dx \, dy}$$
, $\overline{y} = \frac{\iint y \, dx \, dy}{\iint dx \, dy}$;

x and y in these expressions are the co-ordinates of any point whatever within the area, and the limits of the double integration will depend upon the form of the bounding curve.

It frequently happens that the method of polar co-ordinates is

more convenient for the determination of \bar{x} and \bar{y} than that by rectangular co-ordinates: the formulæ will be

$$\overline{x} = \frac{\iint r^2 \cos \theta \ d\theta \ dr}{\iint r \ d\theta \ dr} \ , \quad \overline{y} = \frac{\iint r^2 \sin \theta \ d\theta \ dr}{\iint r \ d\theta \ dr} \ .$$

(1) To find the centre of gravity of the area CPD (fig. 4) of an ellipse, where CP, CD, are two semi-conjugate diameters.

If CP = a, CD = b, and CP, CD, produced indefinitely, be taken as the axes of x, y, the equation to the ellipse will be

$$y^2 = \frac{b^2}{a^2}(a^2 - x^2) \dots (1);$$

and for the position of the centre of gravity we have, indicating the limits of integration,

$$x = \frac{\int_{0}^{a} \int_{0}^{y} x \, dx \, dy}{\int_{0}^{a} \int_{0}^{y} dx \, dy}, \quad \overline{y} = \frac{\int_{0}^{a} \int_{0}^{y} y \, dx \, dy}{\int_{0}^{a} \int_{0}^{y} dx \, dy},$$

the value of y in the limit being pm in the figure; in the integration indicated with respect to y, the figure pqnm is considered as being made up of an infinite number of indefinitely small parallelograms p'q'; and in the integration indicated with respect to x, the whole figure CPD is conceived to be composed of an infinite number of indefinitely thin figures such as pqnm.

$$\int_0^a \int_0^y x \ dx \ dy = \int_0^a xy \ dx = \frac{b}{a} \int_0^a x (a^2 - x^2)^{\frac{1}{4}} dx,$$

since the value of y in the limit coincides with the ordinate in the equation (1); hence

$$\int_{0}^{a} \int_{0}^{y} x \, dx \, dy = -\frac{1}{3} \frac{b}{a} (a^{2} - x^{2})^{\frac{3}{4}}, \text{ from } x = 0 \text{ to } x = a,$$

$$= \frac{1}{3} a^{2} b.$$
Again,
$$\int_{0}^{a} \int_{0}^{y} dx \, dy = \int_{0}^{a} y \, dx$$

$$= \frac{b}{a} \int_{0}^{a} (a^{2} - x^{2})^{\frac{3}{4}} dx = \frac{1}{4} \frac{b}{a} \pi a^{2} = \frac{1}{4} \pi a b.$$

Hence by the general formula for \bar{x} we have

$$\overline{x} = \frac{\frac{1}{3}a^2b}{\frac{1}{4}\pi ab} = \frac{4a}{3\pi}.$$

Again,
$$\int_{0}^{a} \int_{0}^{y} y \, dx \, dy = \frac{1}{2} \int_{0}^{a} y^{2} \, dx$$

$$= \frac{1}{2} \frac{b^{2}}{a^{2}} \int_{0}^{a} (a^{2} - x^{2}) \, dx = \frac{1}{2} \frac{b^{2}}{a^{2}} (a^{3} - \frac{1}{3} a^{3}) = \frac{1}{3} a b^{2},$$
and therefore
$$\overline{y} = \frac{\frac{1}{3} a b^{2}}{\frac{1}{3} \pi a b} = \frac{4b}{3\pi},$$

a result which might have been foreseen from the value of \bar{x} .

Instead of the order of the limits which we have chosen, we might equally well have integrated, first with respect to x and then with respect to y, when the formula for \bar{x} and \bar{y} would have been

$$\overline{x} = \frac{\int_{0}^{b} \int_{0}^{x} x \, dy \, dx}{\int_{0}^{b} \int_{0}^{x} dy \, dx}, \quad \overline{y} = \frac{\int_{0}^{b} \int_{0}^{x} y \, dy \, dx}{\int_{0}^{b} \int_{0}^{x} dy \, dx}.$$

(2) To find the centre of gravity of the segment APBp (fig. 5) of an ellipse cut off by a quadrantal chord ApB.

Let CA = a, CB = b, CM = x, PM = y, pM = y'; then the equations to the ellipse and to the chord will be

$$y^2 = \frac{b^2}{a^2}(a^2 - x^2), \quad y' = \frac{b}{a}(a - x) \dots (1).$$

The formula for \bar{x} will be, indicating the limits,

$$\overline{x} = \frac{\int_{0}^{a} \int_{y'}^{y} x \, dx \, dy}{\int_{0}^{a} \int_{y'}^{y} x \, dx \, dy} \dots (2).$$
Now
$$\int_{0}^{a} \int_{y'}^{y} x \, dx \, dy = \int_{0}^{a} (y - y') x \, dx$$

$$= \frac{b}{a} \int_{0}^{a} \left\{ (a^{2} - x^{2})^{\frac{1}{2}} - (a - x) \right\} x \, dx, \text{ by the equations (1),}$$

$$= \frac{b}{a} \left\{ -\frac{1}{3} (a^{2} - x^{2})^{\frac{1}{2}} - \frac{1}{2} ax^{2} + \frac{1}{3} x^{3} \right\}, \text{ from } x = 0 \text{ to } x = a,$$

$$= \frac{1}{6} a^{2} b.$$

Also
$$\int_{0}^{a} \int_{y'}^{y} dx \, dy = \int_{0}^{a} (y - y') \, dx = \frac{b}{a} \int_{0}^{a} \left\{ (a^{2} - x^{2})^{\frac{1}{4}} - (a - x) \right\} \, dx$$
$$= \frac{b}{a} \int_{0}^{a} (a^{2} - x^{2})^{\frac{1}{4}} \, dx - \frac{b}{a} (a^{2} - \frac{1}{2}a^{2})$$
$$= \frac{b}{a} \left(\frac{1}{4} \pi a^{2} - \frac{1}{2} a^{2} \right) = \frac{1}{4} (\pi - 2) \, ab.$$

Hence from (2), $\bar{x} = \frac{2}{3} \frac{a}{\pi - 2}$.

Similarly we should evidently get

$$\ddot{y}=\frac{2}{3}\,\frac{b}{\pi-2}\,.$$

(3) To find the centre of gravity of the area KSL (fig. 6) of a parabola, of which S is the focus, and SK, SL, any two radii.

Take S as the origin of co-ordinates; also, A being the vertex of the parabola, let ASx be the axis of x, and Sy at right angles to Sx the axis of y. Let SP = r, $\angle ASP = \theta$, AS = m. Then for the position of the centre of gravity, if $\angle ASK = a$, $\angle ASL = \beta$,

$$\overline{x} = \frac{\int_a^\beta \int_0^r r^2 \cos(\pi - \theta) \ d\theta \ dr}{\int_a^\beta \int_0^r r \ d\theta \ dr}, \quad \overline{y} = \frac{\int_a^\beta \int_0^r r^2 \sin(\pi - \theta) \ d\theta \ dr}{\int_a^\beta \int_0^r r \ d\theta \ dr}.$$

Now $\int_0^r r^2 \cos(\pi - \theta) d\theta dr = \frac{1}{3} r^3 \cos(\pi - \theta) d\theta = -\frac{1}{3} r^3 \cos\theta d\theta;$

but, by the nature of the parabola,

$$r = \frac{m}{\cos^2 \frac{1}{2} \theta};$$
hence
$$\int_0^r r^2 \cos (\pi - \theta) \, d\theta \, dr = -\frac{1}{3} m^3 \frac{\cos \theta}{\cos^6 \frac{1}{2} \theta} \, d\theta,$$
and therefore
$$\int_a^\beta \int_0^r r^2 \cos (\pi - \theta) \, d\theta \, dr = -\frac{1}{3} m^3 \int_a^\beta \frac{\cos \theta}{\cos^6 \frac{1}{2} \theta} \, d\theta,$$
but
$$\frac{\cos \theta}{\cos^6 \frac{1}{4} \theta} = \frac{1 - \tan^2 \frac{1}{2} \theta}{1 + \tan^2 \frac{1}{2} \theta} \frac{1}{\cos^6 \frac{1}{2} \theta} = \frac{1 - \tan^2 \frac{1}{2} \theta}{\cos^4 \frac{1}{2} \theta} = (1 - \tan^2 \frac{1}{2} \theta) \sec^4 \frac{1}{2} \theta;$$

hence
$$\int_{a}^{\beta} \int_{0}^{r} r^{2} \cos(\pi - \theta) d\theta dr = -\frac{1}{3}m^{3} \int_{a}^{\beta} \sec^{2} \frac{1}{2} \theta (1 - \tan^{2} \frac{1}{2} \theta) \sec^{2} \frac{1}{2} \theta d\theta \\
= -\frac{1}{3}m^{3} \int_{\tan \frac{1}{2}\alpha}^{\tan \frac{1}{2}\beta} (1 - \tan^{4} \frac{1}{2}\theta) 2d \tan \frac{1}{2}\theta \\
= -\frac{2}{3}m^{3} \left\{ \tan \frac{1}{2}\beta - \tan \frac{1}{2}\alpha - \frac{1}{5} (\tan^{5} \frac{1}{2}\beta - \tan^{5} \frac{1}{2}\alpha) \right\}.$$
Also
$$\int_{a}^{\beta} \int_{0}^{r} r d\theta dr = \frac{1}{2} \int_{a}^{\beta} r^{2} d\theta = \frac{1}{2}m^{2} \int_{a}^{\beta} \frac{d\theta}{\cos^{4} \frac{1}{2}\theta} \\
= \frac{1}{2}m^{2} \int_{\tan \frac{1}{2}\alpha}^{\tan \frac{1}{2}\beta} (1 + \tan^{2} \frac{1}{2}\theta) 2d \tan \frac{1}{2}\theta \\
= m^{2} \left\{ \tan \frac{1}{2}\beta - \tan \frac{1}{2}\alpha + \frac{1}{3} (\tan^{3} \frac{1}{2}\beta - \tan^{3} \frac{1}{2}\alpha) \right\}.$$
Hence
$$\overline{x} = \frac{2m}{3} \frac{\frac{1}{5} (\tan^{6} \frac{1}{2}\beta - \tan^{6} \frac{1}{2}\alpha) - (\tan \frac{1}{2}\beta - \tan^{3} \frac{1}{2}\alpha)}{\frac{1}{3} (\tan^{3} \frac{1}{2}\beta - \tan^{3} \frac{1}{2}\alpha) + (\tan \frac{1}{2}\beta - \tan^{3} \frac{1}{2}\alpha)}.$$
Again,
$$\int_{a}^{\beta} \int_{0}^{r} r^{2} \sin(\pi - \theta) d\theta dr = \frac{1}{3} \int_{a}^{\beta} r^{3} \sin \theta d\theta \\
= \frac{1}{3}m^{3} \int_{a}^{\beta} \frac{\sin \theta}{\cos^{6} \frac{1}{2}\theta} d\theta = \frac{2}{3}m^{3} \int_{a}^{\beta} \frac{\sin \frac{1}{2}\theta}{\cos^{6} \frac{1}{2}\theta} d\theta \\
= -\frac{1}{3}m^{3} \int_{\cos \frac{1}{4}\alpha}^{\cos \frac{1}{4}\theta} \frac{d \cos \frac{1}{2}\theta}{\cos^{6} \frac{1}{2}\theta} = \frac{1}{3}m^{3} (\sec^{4} \frac{1}{2}\beta - \sec^{4} \frac{1}{2}\alpha),$$
and therefore
$$\overline{y} = \frac{1}{3}m \frac{\sec^{4} \frac{1}{2}\beta - \tan^{4} \frac{1}{2}\alpha + (\tan^{2} \frac{1}{2}\beta - \tan^{2} \frac{1}{2}\alpha)}{\frac{1}{3} (\tan^{3} \frac{1}{3}\beta - \tan^{3} \frac{1}{2}\alpha) + (\tan^{2} \frac{1}{2}\beta - \tan^{2} \frac{1}{2}\alpha)}.$$

Let SQ be a radius vector very near to SP; and let pq, p'q', be two circular arcs described about S as a centre, with radii Sp, Sp', very nearly equal to each other. In the integrations which we have executed for the determination of the values of \bar{x} and \bar{y} , we have first conceived the indefinitely thin triangle PSQ to be made up of an infinite series of infinitesimal parallelograms pqp'q', and we have then conceived the whole area KSL to be composed of an infinite number of indefinitely thin triangles, such as PSQ: thus the expressions

$$r^2 \cos (\pi - \theta) d\theta dr$$
, $r^2 \sin (\pi - \theta) d\theta dr$,

represent the moments of the area pqp'q' about the axes of y and x; the expressions

$$\int_0^r r^2 \cos (\pi - \theta) \ d\theta \ dr, \quad \int_0^r r^2 \sin (\pi - \theta) \ d\theta \ dr,$$

the moments of the area SPQ about the axes of y and x; and the expressions

$$\int_{a}^{\beta} \int_{0}^{r} r^{2} \cos (\pi - \theta) d\theta dr, \quad \int_{a}^{\beta} \int_{0}^{r} r^{2} \sin (\pi - \theta) d\theta dr,$$

the moments of the whole area KSL about the axes of y and x. Also the expressions

$$r d\theta dr$$
, $\int_0^r r d\theta dr$, $\int_a^\beta \int_0^r r d\theta dr$,

denote respectively the areas pqp'q', PSQ, KSL.

(4) Cx, Cy, are asymptotes to an hyperbola EAF, (fig. 7); PM, QN, are parallel to yC; to find the centre of gravity of the area PMNQ.

If a, b, be the semiaxes of the hyperbola; Cx, Cy, be taken as the axes of x, y; and CM, CN, be denoted by a, a'; then

$$\overline{x} = \frac{a' - a}{\log a' - \log a}$$
, $\overline{y} = \frac{1}{8} \frac{a^2 + b^2}{aa'} \frac{a' - a}{\log a' - \log a}$.

(5) AB (fig. 8) is a parabola, of which the equation is $a^{m-1}y = x^m$; to find the centre of gravity of the area PMNQ, comprised between two ordinates.

If AM = a, AN = a', we shall have

$$\overline{x} = \frac{m+1}{m+2} \frac{a^{\prime m+2} - a^{m+2}}{a^{\prime m+1} - a^{m+1}}, \quad \overline{y} = \frac{m+1}{2(2m+1)a^{m-1}} \frac{a^{\prime 2m+1} - a^{2m+1}}{a^{\prime m+1} - a^{m+1}}.$$

Carré; Mesure des Surfaces, &c. p. 80.

(6) To find the centre of gravity of the area of a quadrant of a circle.

The equation to the circle being

$$x^2+y^2=a^2,$$

we shall have

$$\overline{x}=\frac{4a}{3\pi}\,,\quad \overline{y}=\frac{4a}{3\pi}\,.$$

(7) To find the centre of gravity of the portion of the area of the curve $y = \sin x$, between x = 0 and $x = \pi$,

$$\overline{x} = \frac{1}{2}\pi$$
, $\overline{y} = \frac{1}{8}\pi$.

(8) To find the centre of gravity of the area intercepted between a straight line $y = \beta x$ and a parabola $y^2 = 4mx$.

$$\overline{x}=\frac{8m}{5\beta^2}, \quad \overline{y}=\frac{2m}{\beta}.$$

SECT. 3. Solid of Revolution.

Let a solid of revolution be generated by the rotation of a plane curve about the axis of x; then the centre of gravity will be within the axis of x, its position being given by the formula

$$\bar{x} = \frac{\int \int xy \ dx \ dy}{\int \int y \ dx \ dy} = \frac{\int x (y^2 - y'^2) \ dx}{\int (y^2 - y'^2) \ dx},$$

y, y', being the limiting values of y for any assignable value of x; if y' = 0, we have

$$\bar{x} = \frac{\int xy^2 dx}{\int y^2 dx}.$$

If polar co-ordinates be adopted, which are frequently convenient, the formula will be

$$\bar{x} = \frac{\iint r^3 \sin \theta \cos \theta \, d\theta \, dr}{\iint r^2 \sin \theta \, d\theta \, dr},$$

the pole being taken at the origin of x, and θ being the angle of inclination of the radius vector r to the axis of x.

(1) To find the centre of gravity of the segment of a sphere.

The centre of the generating circle being taken as the origin, its equation will be

$$x^2+y^2=a^2\ldots\ldots(1);$$

and, c being the distance of the centre of the plane face of the segment from the origin,

$$\overline{x} = \frac{\int_{c}^{a} xy^{2} dx}{\int_{c}^{a} y^{2} dx} \qquad (2);$$

but
$$\int_{c}^{a} xy^{3} dx = \int_{c}^{a} (a^{2} - x^{2}) x dx, \text{ from } (1),$$

$$= \frac{1}{2} a^{2}x^{2} - \frac{1}{4} x^{4}, \text{ from } x = c \text{ to } x = a,$$

$$= \frac{1}{4} a^{4} - \frac{1}{3} a^{2}c^{2} + \frac{1}{4} c^{4} = \frac{1}{4} (a^{2} - c^{2})^{2};$$
also
$$\int_{c}^{a} y^{2} dx = \int_{c}^{a} (a^{2} - x^{2}) dx$$

$$= a^{2}x - \frac{1}{3}x^{3}, \text{ from } x = c \text{ to } x = a,$$

$$= \frac{2}{3} a^{3} - a^{2}c + \frac{1}{3}c^{3}:$$

hence from (2),

$$\overline{x} = \frac{3}{4} \frac{(a^2 - c^2)^2}{2a^3 - 3a^2c + c^3} = \frac{3}{4} \frac{(a+c)^2}{2a+c}.$$

If the segment become a semi-circle, then c = 0, and therefore $\overline{x} = \frac{3}{8}a$.

Lucas Valerius; De Centro Gravitatis Solidorum, Lib. 11. Prop. 33, and Lib. 111. Prop. 31. Guldin; Centrobaryca, Lib. 1. cap. 11, p. 130. Wallis; Opera, tom. 1. p. 728.

(2) To find the centre of gravity of the solid formed by the revolution of the sector of a circle about one of its extreme radii.

Let β denote the angle between the extreme radii of the sector; then the centre of the circle being the origin of x, and a the radius,

$$\overline{x} = \frac{\int_0^\beta \int_0^a r^3 \sin \theta \cos \theta \, d\theta \, dr}{\int_0^\beta \int_0^a r^2 \sin \theta \, d\theta \, dr} \dots \dots (1),$$

hut

$$\int_0^\beta \int_0^a r^3 \sin \theta \cos \theta \, d\theta \, dr = \frac{1}{4} a^4 \int_0^\beta \sin \theta \cos \theta \, d\theta$$
$$= \frac{1}{8} a^4 \int_0^\beta \sin 2\theta \, d\theta = \frac{1}{16} a^4 (1 - \cos 2\beta),$$

and $\int_{0}^{\beta} \int_{0}^{a} r^{2} \sin \theta \, d\theta \, dr = \frac{1}{3} a^{3} \int_{0}^{\beta} \sin \theta \, d\theta = \frac{1}{3} a^{3} (1 - \cos \beta);$

hence from (1) we have

$$\bar{x} = \frac{3}{16} a \frac{1 - \cos 2\beta}{1 - \cos \beta} = \frac{3}{8} a (1 + \cos \beta) = \frac{3}{4} a \cos^2 \frac{1}{2} \beta.$$

We might equally well have integrated the numerator and denominator of (1), first with respect to θ , and afterwards with respect to r. In the one order of integration, we conceive the sector to be made up of an infinite number of thin triangles, of which the centre of the circle is the common vertex; in the other order, the sector is conceived to be made up of an infinite number of infinitesimal rings, having the centre of the circle as their common centre.

Wallis; Opera, tom. 1. p. 728.

(3) To find the centre of gravity of the solid generated by the revolution of the parabolic area ABC (fig. 9), about the tangent Ax at the vertex A, BC being at right angles to the axis Ay of the parabola.

Taking Ax, Ay, as the axes of x, y, the equation to the curve will be $x^2 = 4my.$

Let AC = a, BC = b; then

$$\overline{x} = \frac{\int_{0}^{b} \int_{y}^{a} x y \, dx \, dy}{\int_{0}^{b} \int_{y}^{a} y \, dx \, dy} = \frac{\int_{0}^{b} (a^{2} - y^{2}) x \, dx}{\int_{0}^{b} (a^{3} - y^{2}) \, dx}$$

$$= 2m^{\frac{1}{2}} \frac{\int_{0}^{a} (a^{2} - y^{2}) \, dy}{\int_{0}^{a} (a^{2}y^{-\frac{1}{2}} - y^{\frac{1}{2}}) \, dy}, \text{ from (2),}$$

$$= 2m^{\frac{1}{2}} \frac{a^{3} - \frac{1}{3} a^{4}}{2a^{\frac{1}{2}} - \frac{2}{3} a^{\frac{1}{2}}} = \frac{5}{6} m^{\frac{1}{4}} a^{\frac{1}{4}} = \frac{5}{12} b.$$

This is a case of a more general problem given by Carré, Mesure des Surfaces, &c. p. 93.

(4) To find the centre of gravity of the solid formed by the revolution of any parabola, of which the equation is

$$y^{m+n}=a^mx^n.$$

For any portion of the solid from x = 0 to x = b,

$$\overline{x}=\frac{m+3n}{2m+4n}b.$$

(5) To find the centre of gravity of the frustum of a paraboloid.

If a, b, be the radii of the less and of the greater ends, h the length of the frustum, and \bar{x} the distance of the centre of gravity from the smaller end;

$$\bar{x} = \frac{1}{3}h \frac{a^2 + 2b^2}{a^2 + b^2}.$$

(6) To find the centre of gravity of an hyperboloid. If the equation to the generating hyperbola be

$$y^{2} = \frac{b^{2}}{a^{2}}(x^{2} + 2ax),$$

we shall have for the volume between x = 0 and x = c,

$$\overline{x}=\frac{8ac+3c^2}{4\left(3a+c\right)}.$$

Carré; Mesure des Surfaces, &c. p. 97.

(7) ABC (fig. 10) is a portion of the area of a common parabola, where BC is at right angles to the axis Ax of the parabola; to find the centre of gravity of the solid generated by the revolution of the area ABC about BC.

Let BC = b; then, G being the centre of gravity,

$$CG = \frac{5}{16}b.$$

Carré; Ib. p. 90.

(8) AC, BC, (fig. 11) are the semiaxes of an hyperbola, AD being a portion of the curve intercepted by BD drawn parallel to CA; to find the centre of gravity of the solid generated by the revolution of the area ACBD about CB.

If BC = b, then G being the position of the centre of gravity in BC, $CG = \frac{9}{16}b$.

Carré; Ib. p. 97.

(9) To find the centre of gravity of the solid generated by the revolution about the axis of x of the curve corresponding to the equation

 $y=(a-x)\left(\frac{x}{a}\right)^{\frac{m}{n}}.$

For any portion of the solid cut off by a section at a distance c from the origin,

$$\overline{x} = \frac{140a^2c - 240ac^2 + 105c^3}{168a^2 - 280ac + 120c^2},$$

which for the whole solid becomes

$$\bar{x} = \frac{5}{8}a$$
.

Carré; Mesure des Surfaces, &c. p. 99.

(10) To find the position of the centre of gravity of the volume included between the surfaces generated by the revolution of two parabolas, $y^2 = lx$, $y^2 = l'(a-x)$, round the axis of x. $\overline{x} = \frac{1}{3} a \frac{l+2l'}{l+l'}.$

SECT. IV. Any Solid.

Let x, y, z, be the co-ordinates of any point whatever within any assigned solid; let $\bar{x}, \bar{y}, \bar{z}$, be the co-ordinates of the centre of gravity of this solid; then

$$\overline{x} = \frac{\iiint x \, dx \, dy \, dz}{\iiint dx \, dy \, dz}, \quad \overline{y} = \frac{\iiint y \, dx \, dy \, dz}{\iiint dx \, dy \, dz}, \quad \overline{z} = \frac{\iiint z \, dx \, dy \, dz}{\iiint dx \, dy \, dz},$$

where each of the triple integrations is to be performed in accordance with the nature of the bounding surface of the solid.

(1) To find the centre of gravity of a portion of the cone, of which the equation is

 $y^2+z^2=\beta^2x^2,$

which is contained between the planes of xz, xy, and a given plane parallel to that of yz.

Let a be the length of the axis of the portion of the cone:

$$\bar{x} = \frac{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} x \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} dx \, dy \, dz}, \quad \bar{y} = \frac{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} y \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} dx \, dy \, dz},$$

$$\bar{z} = \frac{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} z \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{\beta z} \int_{0}^{s} dx \, dy \, dz}.$$

Now
$$\int_{0}^{a} \int_{0}^{\beta x} \int_{0}^{s} dx \, dy \, dz = \int_{0}^{a} \int_{0}^{\beta x} z \, dx \, dy$$

$$= \int_{0}^{a} \int_{0}^{\beta x} (\beta^{2}x^{2} - y^{2})^{\frac{1}{2}} \, dx \, dy$$

$$= \int_{0}^{a} \frac{1}{4} \pi \beta^{2}x^{2} \, dx = \frac{1}{12} \pi \beta^{2}a^{3}.$$
Also
$$\int_{0}^{a} \int_{0}^{\beta x} \int_{0}^{s} x \, dx \, dy \, dz = \int_{0}^{a} \int_{0}^{\beta x} xz \, dx \, dy$$

$$= \int_{0}^{a} \int_{0}^{\beta x} x \, (\beta^{2}x^{2} - y^{2})^{\frac{1}{2}} \, dx \, dy = \int_{0}^{a} x \cdot \frac{1}{4} \pi \beta^{2}x^{2} \, dx$$

$$= \frac{1}{16} \pi \beta^{2}a^{4};$$
hence
$$\overline{x} = \frac{\frac{1}{16} \pi \beta^{2}a^{4}}{\frac{1}{12} \pi \beta^{2}a^{3}} = \frac{3}{4}a.$$
Again,
$$\int_{0}^{a} \int_{0}^{\beta x} \int_{0}^{s} y \, dx \, dy \, dz = \int_{0}^{a} \int_{0}^{\beta x} yz \, dx \, dy$$

$$= \int_{0}^{a} \int_{0}^{\beta x} (\beta^{2}x^{2} - y^{2})^{\frac{1}{2}} y \, dx \, dy$$

$$= \int_{0}^{a} \frac{1}{3} \beta^{3}x^{3} \, dx = \frac{1}{12} \beta^{3}a^{4},$$
and therefore
$$\overline{y} = \frac{\frac{1}{12} \beta^{3}a^{4}}{\frac{1}{2} \pi \beta^{2}a^{3}} = \frac{\beta}{\pi} a.$$
Similarly,
$$\overline{z} = \frac{\beta}{\pi} a.$$

(2) To find the centre of gravity of half the solid intercepted between the surfaces of a hemisphere and a paraboloid on the same base, the latus-rectum of the paraboloid coinciding with the diameter of the hemisphere, and the solid being bisected by a plane passing through its axis.

Take the centre of the sphere as the origin of co-ordinates, and the axis of the paraboloid as the axis of z; also let the axis of z be so taken that the plane of xz coincides with the bisecting plane; and take the axis of y at right angles to this plane. Then, if a be the radius of the sphere, the equation to the sphere will be $x^2 + y^2 + z^2 = a^2.$

and to the paraboloid,

$$x^2+y^2=a(a-2z).$$

The centre of gravity will be somewhere in the plane of yz, and is to be determined by the formulæ

$$\bar{y} = \frac{\int_{-a}^{+a} \int_{0}^{a} \int_{-a}^{a} y \, dx \, dy \, dz}{\int_{-a}^{+a} \int_{0}^{a} \int_{-a}^{a} \int_{0}^{a} \int_{-a}^{a} \int_{0}^{a} \int_{-a}^{a} z \, dx \, dy \, dz};$$

where x' is taken to represent $(a^2 - x^2)^{\frac{1}{2}}$, and where the limits x', x, of the general value of x, are its values for any assigned values of x and y in the paraboloid and sphere respectively.

Now

$$\int_{a'}^{a} dx \, dy \, dz = dx \, dy \, (z - z') = dx \, dy \, \left\{ (a^{2} - x^{2} - y^{2})^{\frac{1}{2}} - \frac{1}{2a} (a^{2} - x^{2} - y^{2}) \right\};$$

hence
$$\int_{0}^{\omega} \int_{\omega}^{x} dx \, dy \, dx = \int_{0}^{\omega} dx \, dy \, \left\{ (a^{2} - x^{2} - y^{2})^{\frac{1}{2}} - \frac{1}{2a} (a^{2} - x^{2} - y^{2}) \right\}$$

$$= dx \left\{ \frac{1}{4} \pi (a^3 - x^3) - \frac{1}{3a} (a^2 - x^3)^{\frac{3}{4}} \right\},\,$$

and therefore

$$\int_{-a}^{+a} \int_{0}^{a} \int_{a'}^{a} dx \, dy \, dz = \int_{-a}^{+a} dx \left\{ \frac{1}{4} \pi \left(a^{3} - x^{2} \right) - \frac{1}{3a} \left(a^{2} - x^{2} \right)^{\frac{3}{2}} \right\}$$
$$= \frac{1}{3} \pi a^{3} - \frac{1}{3a} \int_{-a}^{+a} \left(a^{3} - x^{2} \right)^{\frac{3}{2}} dx = \frac{1}{3} \pi a^{3} - \frac{1}{8} \pi a^{2} = \frac{5}{24} \pi a^{3}.$$

Again,
$$\int_{u}^{u} y \, dx \, dy \, dz = y \, dx \, dy \, \left\{ (a^{3} - x^{3} - y^{3})^{\frac{1}{4}} - \frac{1}{2a} (a^{2} - x^{3} - y^{3}) \right\},$$

$$\int_{0}^{u} \int_{u}^{u} y \, dx \, dy \, dz = \int_{0}^{u'} y \, dx \, dy \, \left\{ (a^{3} - x^{2} - y^{3})^{\frac{1}{4}} - \frac{1}{2a} (a^{2} - x^{2} - y^{3}) \right\}$$

$$= dx \left\{ -\frac{1}{3} (a^{2} - x^{2} - y^{3})^{\frac{1}{4}} - \frac{1}{4a} (a^{2} - x^{2}) y^{3} + \frac{1}{8a} y^{4} \right\}, \text{ between limits,}$$

$$= dx \left\{ \frac{1}{3} (a^{2} - x^{3})^{\frac{1}{4}} - \frac{1}{9a} (a^{3} - x^{3})^{\frac{3}{4}} \right\};$$

hence

$$\int_{-a}^{+a} \int_{0}^{a} \int_{a}^{a} y \, dx \, dy \, dz = \frac{1}{3} \int_{-a}^{+a} \left(a^{2} - x^{2} \right)^{\frac{3}{2}} \, dx - \frac{1}{8a} \int_{-a}^{+a} \left(a^{4} - 2a^{2}x^{2} + x^{4} \right) \, dx$$
$$= \frac{1}{8} \pi a^{4} - \frac{2}{15} a^{4} = \frac{15\pi - 16}{120} a^{4}.$$

Again,
$$\int_{-\pi}^{\pi} z \, dx \, dy \, dz = \frac{1}{2} \, dx \, dy \, (z^{3} - z^{2}) = \frac{1}{2} \, dx \, dy \, \{(a^{3} - x^{3} - y^{3}) - \frac{1}{4a^{3}}(a^{3} - x^{3} - y^{3})^{3}\},$$

$$\int_{0}^{\pi} \int_{-\pi}^{\pi} z \, dx \, dy \, dz$$

$$= \frac{1}{2} \int_{0}^{\pi} dx \, dy \, \{a^{3} - x^{3} - y^{2} - \frac{1}{4a^{3}}(a^{3} - x^{3})^{3} + \frac{1}{2a^{3}}(a^{2} - x^{3}) \, y^{3} - \frac{1}{4a^{2}} \, y^{4}\}$$

$$= \frac{1}{2} dx \{(a^{3} - x^{3})^{\frac{3}{4}} - \frac{1}{3}(a^{2} - x^{3})^{\frac{3}{4}} - \frac{1}{4a^{3}}(a^{2} - x^{3})^{\frac{3}{4}} + \frac{1}{6a^{3}}(a^{2} - x^{3})^{\frac{3}{4}} - \frac{1}{20a^{3}}(a^{4} - x^{3})^{\frac{3}{4}}\}$$
between limits,
$$= \frac{1}{3} dx \, \{(a^{2} - x^{3})^{\frac{3}{4}} - \frac{1}{5a^{3}}(a^{2} - x^{3})^{\frac{3}{4}}\},$$

$$= \frac{1}{3} dx \left\{ (a^2 - x^2)^{\frac{3}{4}} - \frac{1}{5a^2} (a^2 - x^2)^{\frac{3}{4}} \right\},$$

$$\int_{-a}^{+a} \int_{0}^{a} \int_{a'}^{a} z \, dx \, dy \, dz = \frac{1}{3} \int_{-a}^{+a} dx \left\{ (a^2 - x^2)^{\frac{3}{4}} - \frac{1}{5a^2} (a^2 - x^2)^{\frac{3}{4}} \right\}$$

$$= \frac{1}{3} \int_{-a}^{+a} dx (a^2 - x^2)^{\frac{3}{4}} - \frac{1}{15a^2} \int_{-a}^{+a} (a^2 - x^2)^{\frac{3}{4}} \, dx$$

$$= \frac{1}{8} \pi a^4 - \frac{1}{48} \pi a^4 = \frac{5}{48} \pi a^4.$$

From the formulæ for \overline{y} and \overline{z} then we have

$$\bar{y} = \frac{\frac{1}{120} (15\pi - 16) a^4}{\frac{5}{24} \pi a^3} = \frac{15\pi - 16}{25\pi} a,$$

$$\bar{z} = \frac{\frac{5}{88} \pi a^4}{\frac{5}{24} \pi a^3} = \frac{1}{2} a.$$

(3) AOC (fig. 12) is a right-angled triangle, O being the right angle; AOBD is a rectangle, of which the plane is perpendicular to that of the triangle; from every point R in the line AC a straight line RQ is drawn to meet BD in Q, in a plane at right angles to the areas of the rectangle and triangle; to find the centre of gravity of the volume so generated.

Let OAx, OBy, OCz, be taken as axes of x, y, z; from R draw RM at right angles to OA, join QM, and draw PN at right angles to QM; let OA = a, OB = b, OC = c; OM = x, MN = y, NP = z', z being the distance of any point in the line PN from the point N: then for the determination of the centre of gravity we have

$$\bar{x} = \frac{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} x \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} dx \, dy \, dz}, \qquad \bar{y} = \frac{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} y \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} dx \, dy \, dz},$$

$$\bar{z} = \frac{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} z \, dx \, dy \, dz}{\int_{0}^{a} \int_{0}^{b} \int_{0}^{s'} dx \, dy \, dz}.$$

From the geometry it is evident that

$$z'=c\;\frac{a-x}{a}\;\frac{b-y}{b};$$

hence we have

$$\bar{x} = \frac{\int_{0}^{a} \int_{0}^{b} x (a - x) (b - y) dx dy}{\int_{0}^{a} \int_{0}^{b} (a - x) (b - y) dx dy}$$

$$= \frac{\int_{0}^{a} x (a - x) dx}{\int_{0}^{a} (a - x) dx} = \frac{\frac{1}{6} a^{3}}{\frac{1}{2} a^{2}} = \frac{1}{3} a.$$

$$\bar{y} = \frac{\int_{0}^{a} \int_{0}^{b} (a - x) (b - y) y dx dy}{\int_{0}^{a} \int_{0}^{b} (a - x) (b - y) dx dy} = \frac{\int_{0}^{b} (b - y) y dy}{\int_{0}^{a} (b - y) dy} = \frac{\frac{1}{6} b^{3}}{\frac{1}{2} b^{2}} = \frac{1}{3} b,$$

$$\bar{z} = \frac{\int_{0}^{a} \int_{0}^{b} \frac{1}{2} c^{2} \frac{(a - x)^{2}}{a^{2}} \frac{(b - y)^{2}}{b^{2}} dx dy}{\int_{0}^{a} \int_{0}^{b} (a - x)^{2} (b - y)^{2} dx dy} = \frac{c}{2ab} \int_{0}^{a} \int_{0}^{b} (a - x) (b - y) dx dy$$

$$= \frac{c}{2ab} \frac{\frac{1}{3} a^{3} \cdot \frac{1}{3} b^{3}}{\frac{1}{4} a^{2} \cdot \frac{1}{4} b^{2}} = \frac{2c}{9}.$$

(4) To find the centre of gravity of the portion of the sphere $x^2 + y^2 + z^2 = a^2$,

which is cut off by three planes, x = 0, y = 0, z = 0.

$$\bar{\bar{x}} = \bar{\bar{y}} = \bar{\bar{z}} = \frac{3}{8}a.$$

(5) To find the centre of gravity of a portion of the paraboloid $y^2 + z^2 = 4mx$,

which is cut off by the three planes x = a, y = 0, z = 0.

If b be the radius of the section of the paraboloid made by the plane x = a, then

$$\overline{x}=\frac{2}{3}a, \quad \overline{y}=\overline{z}=\frac{16b}{15\pi}.$$

(6) To find the centre of gravity of a portion of the solid $z^2 = xy$, which is cut off by the five planes x = 0, y = 0, z = 0, z

(7) To find the centre of gravity of the volume of the cylinder $y^2 = 2ax - x^2$, which is cut off between the two planes $z = \beta x$, $z = \beta' x$. $\bar{x} = \frac{5}{8}a$, $\bar{y} = 0$, $\bar{z} = \frac{5}{8}(\beta + \beta')a$.

(8) A solid is generated by a variable rectangle moving parallel to itself along an axis perpendicular to its plane through its centre; one side of the rectangle varies as the distance from a fixed point in the axis, while half the other is the sine of a circular arc, of which this distance is the versed sine; to determine the distance of the centre of gravity of the whole solid from the fixed point.

The required distance is equal to four-fifths of the length of the axis.

SECT. 5. A Plane Curve.

Let x, y, be the co-ordinates of any point of a plane curve, and let ds denote an element of the length of the curve at that point; then \overline{x} , \overline{y} , denoting the co-ordinates of the centre of gravity of any assigned portion of the curve,

$$\overline{x} = \frac{\int x \ ds}{\int ds}$$
, $\overline{y} = \frac{\int y \ ds}{\int ds}$,

the integrations being performed in accordance with the limits of the portion.

The idea of the determination of the centres of gravity of curve lines is due to La-Faille, a Flemish mathematician, by whom it was applied in the instances of portions of the circle and the ellipse, in a work entitled "De centro gravitatis partium circuli et ellipsis theoremata," published in the year 1632. The theorems of La-Faille were afterwards published in a somewhat

more elegant form, and with amplifications, by Guldin; Centro-baryca, lib. 1. cap. 4, 5, 6, 7.

(1) To find the centre of gravity of the arc of the curve $y = \sin x$ from x = 0 to $x = \pi$.

From the equation to the curve we have

$$ds^{2} = dx^{2} + dy^{2} = (1 + \cos^{2} x) dx^{2};$$
hence
$$\bar{y} = \frac{\int y ds}{\int ds} = \frac{\int_{0}^{\pi} \sin x (1 + \cos^{2} x)^{\frac{1}{2}} dx}{\int_{0}^{\pi} (1 + \cos^{2} x)^{\frac{1}{2}} dx}.$$

Now by the ordinary processes of the integral calculus,

$$\int_0^{\pi} \sin x \left(1 + \cos^2 x\right)^{\frac{1}{4}} dx = 2^{\frac{1}{4}} + \log \left(2^{\frac{1}{4}} + 1\right).$$

Also, c denoting the length of the curve from x = 0 to $x = \pi$,

$$c = \int_0^{\pi} (1 + \cos^2 x)^{\frac{1}{4}} dx = 2^{\frac{1}{4}} \int_0^{\pi} (1 - \frac{1}{2} \sin^2 x)^{\frac{1}{4}} dx,$$

an elliptic function, of which $\cos \frac{1}{2}\pi$ is the modulus. Hence \bar{y} is given by the equation

$$c\widetilde{y}=2^{\frac{1}{4}}+\log\left(2^{\frac{1}{4}}+1\right).$$

(2) To find the centre of gravity of any arc of a circle.

Let the centre of the circle be taken as the origin of coordinates, and let the axis of x bisect the arc; then, if a be the radius of the circle, c the chord of the arc, and s the length of the arc,

$$\bar{x} = \frac{ac}{s}$$
, $\bar{y} = 0$.

Guldin; Centrobaryca, lib. 1. cap. 5, p. 59. Wallis; Opera, tom. 1. p. 712.

(3) To find the centre of gravity of the arc of a semicycloid. The equation to the curve being

$$y = a \text{ vers}^{-1} \frac{x}{a} + (2ax - x^3)^{\frac{1}{2}},$$

we shall have

$$\overline{x} = \frac{2}{3}a$$
, $\overline{y} = (\pi - \frac{4}{3})a$.
Wallis; Opera, tom 1. p. 520.

(4) To find the centre of gravity of the arc of a catenary

$$y=\frac{1}{2}a\left(e^{\frac{x}{a}}+e^{-\frac{x}{a}}\right),$$

cut off by any assigned double ordinate.

If 2s be the whole length of the intercepted arc,

$$\overline{x}=0, \qquad \overline{y}=\frac{ax+sy}{2s}.$$

(5) To find the centre of gravity of the arc of a parabola $y^2 = 4mx$, cut off by the latus rectum.

$$\bar{x} = \frac{1}{4}m \frac{3 \cdot 2^{\frac{1}{4}} - \log(1 + 2^{\frac{1}{4}})}{2^{\frac{1}{4}} + \log(1 + 2^{\frac{1}{4}})}, \quad \bar{y} = 0.$$

(6) To find the centre of gravity of the semi-arc of a loop of the Lemniscata of James Bernoulli.

If the axis of the loop be taken as the axis of x, the node being the origin; then, a being the length of the axis and l of the semi-arc,

$$\bar{x} = \frac{a^2}{a^{\frac{1}{4}}}, \quad \bar{y} = \frac{(2^{\frac{1}{4}} - 1) a^2}{a^{\frac{1}{4}}}.$$

SECT. 6. Curve of Double Curvature.

The formulæ for the determination of the centre of gravity of a curve of double curvature, are

$$\bar{x} = \frac{\int x \, ds}{\int ds}$$
, $\bar{y} = \frac{\int y \, ds}{\int ds}$, $\bar{z} = \frac{\int z \, ds}{\int ds}$;

where x, y, z, are the co-ordinates of any point in the curve, and ds an element of the arc at that point: the limits of the integrations will depend upon the positions of the ends of that portion of the curve of which the centre of gravity is required.

Ex. To find the centre of gravity of the Helix.

The equations to the curve are

$$x^2 + y^2 = a^2,$$
 $z = b \cos^{-1} \frac{x}{a};$

and for the centre of gravity of any length of the curve, beginning at the origin of co-ordinates,

$$\overline{x} = \frac{by}{z}$$
, $\overline{y} = \frac{b(a-x)}{z}$, $\overline{z} = \frac{1}{2}z$.

SECT. 7. Surface of Revolution.

Let x, y, be the co-ordinates of any point of a curve, by the revolution of which about the axis of x a surface is supposed to be generated; then, if ds denote an element of the generating curve at the point, we have for the position of the centre of gravity of the surface of revolution in the axis of x,

$$\bar{x} = \frac{\int xy \ ds}{\int y \ ds};$$

the integrations being performed between limits depending upon the magnitude of the surface.

(1) To find the centre of gravity of the surface of a segment of a sphere.

If the equation to the generating circle be

$$y = (2ax - x^{2})^{\frac{1}{4}},$$

$$dy = \frac{a - x}{(2ax - x^{2})^{\frac{1}{4}}} dx,$$

we shall have

and therefore

$$ds^2 = dx^2 + dy^2 = \frac{a^2}{2ax - x^2} dx^2 = \frac{a^2 dx^2}{y^2}$$
, or $y ds = a dx$;

hence for any segment, of which the limiting abscissa is c,

$$\bar{x} = \frac{\int_0^c ax \, dx}{\int_0^c a \, dx} = \frac{\frac{1}{2}c^2}{c} = \frac{1}{2}c.$$

(2) To find the centre of gravity of the surface of a cone.

Let the equation to the generating straight line be y = ax; then, c being the length of the axis of the cone,

$$\overline{x} = \frac{2}{7}c$$
.

Guldin; Centrobaryca, lib. 1. cap. 10, prop. 3.

(3) To find the centre of gravity of the surface generated by the revolution of a semicycloid about its axis.

The equation to the curve being

$$y = a \text{ vers}^{-1} \frac{x}{a} + (2ax - x^2)^{\frac{1}{4}},$$

we shall have

$$\overline{x} = \frac{2}{15} a \frac{15\pi - 8}{3\pi - 4}.$$

(4) To find the centre of gravity of the surface generated by the revolution of the parabola $y^2 = 4mx$ about the axis of x.

$$\overline{x} = \frac{1}{5}(3x - 2m).$$

SECT. 8. Any Surface.

Let x, y, z, be the co-ordinates of any point of a surface referred to three rectangular axes; and let $\frac{dz}{dx} = p$, $\frac{dz}{dy} = q$; then for the centre of gravity of any portion of the surface

$$\bar{x} = \frac{\iint x (1 + p^{3} + q^{2})^{\frac{1}{4}} dx dy}{\iint (1 + p^{2} + q^{2})^{\frac{1}{4}} dx dy},$$

$$\bar{y} = \frac{\iint y (1 + p^{2} + q^{2})^{\frac{1}{4}} dx dy}{\iint (1 + p^{2} + q^{2})^{\frac{1}{4}} dx dy},$$

$$\bar{z} = \frac{\iint z (1 + p^{2} + q^{2})^{\frac{1}{4}} dx dy}{\iint (1 + p^{3} + q^{2})^{\frac{1}{4}} dx dy},$$

the integrations being performed between limits corresponding to the boundary of the surface.

(1) Suppose the surface to be any portion of the superficies of a sphere, of which the equation is

$$x^2 + y^2 + z^3 = r^2.$$

Then clearly we shall have

$$p=-\frac{x}{z}, \qquad q=-\frac{y}{z};$$

and therefore
$$\bar{z} = r$$

$$\overline{z} = r \frac{\iint dx \, dy}{\iint dx \, dy \, (1 + p^2 + q^2)^{\frac{1}{4}}}.$$

Now it is evident that, the integrations being performed within the given limits, the denominator of this expression for \overline{z} represents the area of the given portion of the surface, while the numerator represents the area of the projection of this same portion upon the plane of x and y. Hence in general language: the distance of the centre of gravity of any portion whatever of the surface of a sphere from the plane of any one of its great circles, is a fourth proportional to the area of the portion itself, the area of its projection on this plane, and the radius of the sphere.

The truth of this proposition depends solely upon the property expressed by the equation

$$z(1+p^3+q^2)^{\frac{1}{3}}=r;$$

but this equation holds good for the whole class of surfaces generated by the motion of a sphere of invariable radius, of which the centre describes a plane curve traced arbitrarily in the plane of x and y; hence we may evidently extend the preceding proposition to all these surfaces under the following enunciation:—

"Upon any surface whatever, generated by the motion of a sphere of which the centre never departs from a given plane, let any portion S be taken, and let S' be the projection of S upon the given plane; then the distance of the centre of gravity of S from this plane will be a fourth proportional to S, S', and the radius of the generating sphere."

(2) To find the centre of gravity of any spherical triangle formed by three great circles.

Let ABC (fig. 13) be any spherical triangle, O the centre of the sphere; and OA, OB, OC, the three radii at the angles A, B, C, of the triangle. Let Z_a , Z_b , Z_c , denote the distances of the centre of gravity of this triangle from the three planes BOC, AOC, AOB; then, by the proposition of the preceding article, if r be the radius of the sphere, S the area of the

spherical triangle ABC, and S' of its projection upon the plane BOC, $Z_a = \frac{S'}{S'} r.$

But by the principles of spherical trigonometry,

$$S = \frac{\pi r^3}{180} (A + B + C - 180);$$

also it is clear that, a, b, c, being the number of degrees subtended at the centre of the sphere by the sides of the spherical triangle opposite to the angles A, B, C,

$$S' = \text{area } BOC - \text{area } AOB \times \cos B - \text{area } AOC \times \cos C$$

$$= \frac{\pi r^3}{360} (a - c \cos B - b \cos C),$$
and therefore
$$Z_a = \frac{1}{2}r \frac{a - b \cos C - c \cos B}{A + B + C - 180}.$$
Similarly,
$$Z_b = \frac{1}{2}r \frac{b - c \cos A - a \cos C}{A + B + C - 180}.$$

$$Z_c = \frac{1}{2}r \frac{c - a \cos B - b \cos A}{A + B + C - 180}.$$

The position of the centre of gravity of the spherical triangle may be elegantly expressed likewise in terms of its distances from three great circles of the sphere, at right angles to the three edges OA, OB, OC, of the spherical pyramid ABCO. Let D_{e} , D_{b} , D_{c} , denote these distances; then by Art. (1) we have

$$D_a = r \frac{S_a}{S}, \quad D_b = r \frac{S_b}{S}, \quad D_c = r \frac{S_c}{S},$$

where S denotes the spherical area ABC, and S_a , S_b , S_c , its projections upon the three great circles at right angles to OA, OB, OC.

Now it is evident that the projections of the spherical area ABC, and of the sector BOC, upon the great circle which is at right angles to OA, are identically the same, and therefore, if the arc Aa be drawn at right angles to BC, we have

$$S_a = \text{area of sector } BOC \times \cos\left(\frac{\pi}{2} - \frac{Aa}{r}\right)$$

= $\frac{\pi}{360} ar^2 \sin\frac{Aa}{r} = \frac{\pi}{360} ar^2 \sin B \cdot \sin c$:

but
$$S = \frac{\pi r^2}{180} (A + B + C - 180);$$

hence $D_a = \frac{1}{2} r \frac{a \sin B \sin c}{A + B + C - 180}.$
Similarly, $D_b = \frac{1}{2} r \frac{b \sin C \sin a}{A + B + C - 180},$
 $D_c = \frac{1}{2} r \frac{c \sin A \sin b}{A + B + C - 180}.$

If we desire to determine the position of the centre of gravity of the triangle by means of three rectangular co-ordinates x, y, z, let the plane of the side c be taken as the plane of x and y, and let the radius OA be taken to coincide with the axis of x. Then from the preceding results we have at once

$$x = \frac{1}{2}r \frac{a \sin B \sin c}{A + B + C - 180}, \qquad z = \frac{1}{2}r \frac{c - b \cos A - a \cos B}{A + B + C - 180}.$$

Again, let the great circle, of which BC is an arc, meet the plane of x, z, in the point D, as in fig. 14; join A and D by an arc of a great circle. Then clearly the projection of the spherical triangle ABC upon the plane of x and z, is equal to the difference of the projections of the sectors AOC, BOC, upon this plane, and therefore to the expression

$$\frac{\pi}{180} r^2 b \cos CAD - \frac{\pi}{180} r^2 a \cos D$$

$$= \frac{\pi r^2}{180} (b \sin A - a \sin B \cos c);$$

hence, by the principles of Art. (1), we have

$$y = \frac{1}{2} \frac{b \sin A - a \sin B \cos c}{A + B + C - 180}.$$

(3) The general formula

$$\bar{z} = \frac{\iint z (1 + p^2 + q^2)^{\frac{1}{4}} dx dy}{\iint (1 + p^2 + q^2)^{\frac{1}{4}} dx dy}$$

furnishes us with the following general proposition:-

"Upon the surface (A), generated by the revolution of the curve of equilibrium of a homogeneous catenary about the

vertical line which passes through its lowest point, trace arbitrarily a perimeter enclosing a portion S of the surface; project this perimeter upon a horizontal plane which intersects the axis of revolution at a distance c below the lowest point of the surface, where c is equal to the horizontal tension of the catenary divided by the mass of a unit of its length; let V be the volume contained between the surface S, its projection, and the cylindrical surface formed by the perpendiculars from the perimeter of S upon the plane of projection. Then the altitude of the centre of gravity of S above this plane will be double of that of the centre of gravity of V."

In fact, the plane touching the surface (A) in a point situated at an altitude z above the plane of projection, which we shall take for the plane of x and y, makes with this plane an angle, of which the cosine is $\frac{c}{z}$; and therefore we have the equation

$$(1+p^3+q^2)^{\frac{1}{4}}=\frac{z}{c};$$

hence, by the formula for \bar{z} , we obtain

$$\overline{z} = \frac{\iint z^3 \, dx \, dy}{\iint z \, dx \, dy} .$$

But calling z the altitude of the centre of gravity of V above the same plane, we have

$$\underline{z} = \frac{\int \int \frac{1}{2} z z \, dx \, dy}{\int \int z \, dx \, dy},$$

and the limits of the integrations being the same in the expressions for \bar{z} and z, we see that $\bar{z} = 2z$.

The property expressed by the partial differential equation

$$(1 + p^2 + q^2)^{\frac{1}{4}} = \frac{z}{c},$$

being common to all the surfaces which can be generated by the surface (A) when it moves in such a manner that its axis always remains vertical, and that one of its points describes a plane curve traced arbitrarily upon a horizontal plane, the proposition which we have demonstrated is susceptible of the same extension as that of (1).

The illustrations of the general formulæ for the determination of the centre of gravity of any surface, which we have given in (1), (2), (3), are extracted from a memoir by Professor Giulio, of Turin, which may be seen in Liouville's *Journal de Mathematiques*, tom. IV. p. 386.

(4) To find the centre of gravity of the surface of a cone

$$y^2+z^2=\beta^2x^2,$$

intercepted by the planes x = 0, x = a, y = 0, z = 0.

$$\overline{x} = \frac{2}{3}a$$
, $\overline{y} = z = \frac{4}{3}\frac{\beta a}{\pi}$.

SECT. 9. Heterogeneous Bodies.

In the preceding sections we have computed the centres of gravity of various classes of homogeneous bodies; we will now give a few examples of the determination of the centre of gravity when the density is variable.

(1) To find the centre of gravity of the surface of a hemisphere, when the density of each point in the surface varies as its perpendicular distance from the circular base of the hemisphere.

Let the equation to the quadrantal arc, by the revolution of which the hemispherical surface may be generated, be

$$x^2 + y^2 = a^2 \dots (1),$$

the axis of x being the axis of revolution.

The area of the strip of the surface which is generated by the element ds of the arc, will be equal to $2\pi y ds$; and, if ρ be its density, its mass will be equal to $2\pi \rho y ds$: hence

$$\bar{x} = \frac{\int 2\pi\rho \ y \ ds \cdot x}{\int 2\pi\rho \ y \ ds} = \frac{\int \rho \ xy \ ds}{\int \rho \ y \ ds};$$

$$e \qquad \qquad \dot{\bar{x}} = \frac{\int x^2 \ y \ ds}{\int xy \ ds};$$

but $\rho \propto x$; hence

and therefore, since from (1) it is easily seen that

$$y ds = a dx$$

we have
$$\bar{x} = \frac{\int_0^a x^3 dx}{\int_0^a x dx} = \frac{\frac{1}{3}a^3}{\frac{1}{2}a^2} = \frac{2}{3}a.$$

(2) To find the centre of gravity of a physical line, the density of which at any point varies as the n^{th} power of its distance from a given point in the line produced.

Let a, b, be the distances of the given point from the two extremities, and \bar{x} its distance from the centre of gravity of the physical line; then

$$\bar{x} = \frac{n+1}{n+2} \frac{b^{n+2}-a^{n+2}}{b^{n+1}-a^{n+1}}.$$

(3) To find the centre of gravity of a quadrant of a circle, the density at any point of which varies as the nth power of its distance from the centre.

Let a denote the radius of the circle, and \bar{x} , \bar{y} , the coordinates of the centre of gravity of the quadrant referred to its two extreme radii as axes; then

$$\overline{x}=\frac{n+2}{n+3}\cdot\frac{2a}{\pi}=\overline{y}.$$

SECT. 10. Centre of Parallel Forces.

When any number of parallel forces act on a system of rigidly connected points, they generally have a single resultant acting on a point of which the position is invariable while their common direction is changed in every possible way. This point is called the Centre of the Parallel Forces: the Centre of Gravity of a body is a particular case of this. Let x, y, z, denote the co-ordinates of the point of application of any force P of the system referred to any axes, rectangular or oblique; and let \overline{x} , \overline{y} , \overline{z} , be the co-ordinates of the Centre of Parallel Forces. Then, R representing the resultant,

$$R = \Sigma(P), \quad \overline{x} = \frac{\Sigma(Px)}{\Sigma(P)}, \quad \overline{y} = \frac{\Sigma(Py)}{\Sigma(P)}, \quad \overline{z} = \frac{\Sigma(Pz)}{\Sigma(P)}.$$

Whenever $\Sigma(P)$ is equal to zero, these formulæ cease to be applicable, there being in this case no single resultant; the forces will be reducible to a resultant couple. For the complete development of the theory of Statical Couples, the reader is referred to Poinsot's beautiful work entitled *Elémens de Statique*.

The formulæ for \bar{x} , \bar{y} , \bar{z} , were first given by Varignon, in the *Mémoires de l'Académie des Sciences de Paris* for the year 1714.

(1) Three parallel forces acting at the angular points A, B, C, of a plane triangle, are respectively proportional to the opposite sides a, b, c; to find the distance of the centre of parallel forces from A.

Produce AB, AC, indefinitely to points x, y, and let Ax, Ay, be taken as co-ordinate axes.

Let μa , μb , μc , be the forces applied at A, B, C, where a, b, c, denote the opposite sides of the triangle. The coordinates of the points of application of these three forces are 0, 0; c, 0; 0, b; hence by the two first of the formulæ (B) we have

$$\bar{x} = \frac{c \cdot \mu b}{\mu a + \mu b + \mu c} = \frac{bc}{a + b + c},$$

$$\bar{y} = \frac{b \cdot \mu c}{\mu a + \mu b + \mu c} = \frac{bc}{a + b + c} = \bar{x}.$$

Let r be the distance of the centre of parallel forces from A; then

$$r^3 = \bar{x}^2 + \bar{y}^2 + 2\bar{x}\bar{y}\cos A = 2\bar{x}^2(1 + \cos A) = 4\bar{x}^2\cos^2\frac{1}{2}A,$$
 and therefore

$$r = 2\bar{x} \cos \frac{1}{2}A = \frac{2bc \cos \frac{1}{2}A}{a+b+c}$$
.

(2) Three parallel forces P, Q, R, act at the angles A, B, C, of a given triangle, and are to each other as the reciprocals of the opposite sides a, b, c; to determine the distance of their centre from A.

Required distance =
$$a \frac{(b^4 + 2b^2c^2 \cos A + c^4)^{\frac{1}{4}}}{ab + ac + bc}$$
.

(3) At the corners B, C, D, of a quadrilateral pyramid ABCD, three parallel forces P, Q, R, are applied; to find the distance of their centre from the corner A.

Let AB=b, AC=c, AD=d; $\angle BAC=(b, c)$, $\angle BAD=(b, d)$, $\angle CAD=(c, d)$; r= the required distance; then

$$r^{2}(P+Q+R)^{3} = P^{2}b^{2} + Q^{2}c^{3} + R^{2}d^{2} + 2PQbc \cos(b, c) + 2PRbd \cos(b, d) + 2QRcd \cos(c, d).$$

(4) At three fixed points (a, b), (a', b'), (a', b'), in the plane of x, y, are applied three parallel forces p, p', p'; supposing the magnitude of p' to vary in every possible way, to find the locus of the centre of parallel forces.

The locus will be a straight line of which the equation is

$$(ap + a'p')b'' + \{(a'' - a)p + (a'' - a')p'\}y$$

= $(bp + b'p')a'' + \{(b'' - b)p + (b'' - b')p'\}x$.

SECT. 11. The Properties of Pappus.

- I. If any plane area revolve about any axis in its own plane through any assigned angle, the volume of the surface generated by the motion of the area will be equal to a prism, of which the base is equal to the revolving area, and the altitude to the length of the path described by the centre of gravity of the area during its revolution.
- II. If any plane area revolve through any angle about any axis in its own plane, the area of the surface generated by its perimeter will be equal to a rectangle, of which one side is the length of the perimeter, and the other the length of the path described by the centre of gravity of the perimeter.

The enunciation of these properties, which are generally called Guldin's properties, is due to Pappus, and may be seen

¹ The words of Pappus in the Latin translation are: "Perfectorum utrorumque ordinum proportio composita est ex proportione amphismatum, et rectarum linearum similiter ad axes ductarum a punctis, quæ in ipsis gravitatis centra sunt. Imperfectorum autem proportio composita est ex proportione amphismatum, et circumferentiarum a punctis quæ in ipsis sunt centra gravitatis, factarum, &c." In the former case he is alluding to those solids which are formed by the entire revolution of the generating figures through 360°; in the latter, to those which are formed by revolution through any smaller angle.

at the end of the Preface to the seventh book of his Mathematical Collections, of which the first edition appeared in the form of a Latin translation in the year 1588. They were afterwards published, with various applications, by Guldin, in his treatise De Centro Gravitatis, lib. 2 and 3, which appeared for the first time in the year 1635. Cavalieri, in reply to objections advanced by Guldin against his method of indivisibles, gave a demonstration of these properties by this method; stating likewise, in allusion to Guldin's claims as a discoverer, that they had been communicated to him, long before the publication of Guldin's work, by a pupil of his, Antonio Roccha. Elegant demonstrations of these properties were given also by Varignon in the Mémoires de l'Académie des Sciences de Paris for the year 1714, p. 77.

(1) From any point P (fig. 15) in a parabola, is drawn a straight line PM at right angles to the axis, and meeting it in the point M; to find the content of the solid generated by the complete revolution of the area APM about PM.

Let AM = x, PM = y; V = the required volume; and $\overline{x} =$ the distance of the centre of gravity of the area APM from PM. Then the whole path described by the centre of gravity will be equal to $2\pi \overline{x}$; hence, by (I.),

$$V = 2\pi \bar{x} \times \text{area of } PAM$$
:

but $\bar{x} = \frac{2}{5}x$, and area of $PAM = \frac{2}{3}xy$; and therefore

$$V = 2\pi \cdot \frac{2}{5}x \cdot \frac{2}{3}xy = \frac{8}{15}\pi x^2y.$$

Complete the parallelogram MPmA; then the area of this parallelogram will be equal to xy, and the distance of its centre of gravity from PM will be equal to $\frac{1}{2}x$. Conceive this parallelogram to make an entire revolution about PM; then the path of its centre of gravity will be equal to

$$2\pi \cdot \frac{1}{2}x = \pi x$$
;

and therefore, if U denote the volume of the cylinder which is generated by the revolution,

$$U=\pi x\cdot xy=\pi x^2y.$$

Hence U:V::15:8.

This is one of the problems proposed in Kepler's Stereometria.

Guldinus; De Centro Gravitatis, lib. 11. cap. 12, prop. 6.

1 Exercitationes Geometricæ Sex, Exercit. 1 & 2; Bononiæ, 1647.

(2) To find the surface of a sphere.

Let BAb (fig. 16) be a semicircle, by the revolution of which about the diameter BCb the sphere is generated.

Let CA be at right angles to Bb, C being the centre of the circle, and let G be the centre of gravity of the semicircular arc BAb. Let CA = a, surface required = S; then, by (II.),

$$2\pi CG$$
 . arc $BAb = S$;

but $CG = \frac{2a}{\pi}$, and arc $BAb = \pi a$; hence

$$S=2\pi\cdot\frac{2a}{\pi}\cdot\pi a=4\pi a^2.$$

Now πa^2 is the area of a great circle of the sphere; and thus we find that the whole surface of a sphere is four times as great as that of one of its great circles. This proposition was first proved by Archimedes, $\Pi_{\epsilon\rho}$ Equipag kal Kulivõpou, Bi $\beta\lambda$. A, $\pi\rho\delta\tau a$. A; and afterwards, according to the method which we have given, by Guldin, De Centro Gravitatis, lib. IV. cap. 1, prop. 7.

(3) To find the volume and the surface of the solid ring generated by the complete revolution of a circle about any external line in its own plane.

Let b be the distance of the centre of the circle from the axis of revolution, and a the radius of the circle; then

volume =
$$2\pi^2 a^2 b$$
, and surface = $4\pi^2 ab$.

(4) To find the volume of the solid ring generated by the revolution of an ellipse about an external axis in its own plane through an angle of 180°.

If a, b, be the semiaxes of the ellipse, and c the distance of its centre from the axis of rotation, then

volume =
$$\pi^2 abc$$
.

(5) To find the volume generated by the revolution through a given angle of a portion APM (fig. 15) of a parabola about a tangent at its vertex A, PM being parallel to the tangent, and AM at right angles to it.

If AM = x, PM = y, and β be the angle through which the revolution takes place; then

volume =
$$\frac{1}{3}\beta x^2y$$
.

(6) To find the volume and the surface of the solid generated by the complete revolution of a cycloid about its axis.

If a be the radius of the generating circle,

volume =
$$\pi a^3 (\frac{3}{2} \pi^2 - \frac{8}{3})$$
, surface = $8\pi a^2 (\pi - \frac{4}{3})$.

(7) To find the volume and the surface of the solid generated by the complete revolution of a cycloid about its base.

Volume =
$$\frac{5}{2}\pi^2a^3$$
, surface = $\frac{32}{3}\pi a^2$.

(8) To find the content of the solid generated by the complete revolution of a right-angled triangle about its hypothenuse.

If a, b, denote the two sides of the triangle, the content will be equal to

$$\frac{\pi a^2 b^2}{3 \left(a^2 + b^2\right)^{\frac{1}{4}}}.$$

CHAPTER II.

EQUILIBRIUM OF A PARTICLE.

Let P denote any one of a system of forces acting on a particle; and let a, β , γ , be the angles which the direction of this force makes with any three proposed straight lines, no two of which are parallel; then the sufficient and necessary conditions for the equilibrium of the particle are expressed by the three following equations,

$$\Sigma(P\cos a) = 0$$
, $\Sigma(P\cos \beta) = 0$, $\Sigma(P\cos \gamma) = 0$,

where the Σ represents the summation of all such quantities as $P\cos a$, $P\cos \beta$, $P\cos \gamma$, for all the different forces of the system; or the algebraical sum of the resolved parts of all the forces of the system estimated parallel to each of the three straight lines must be equal to zero. If all the forces acting on the particle lie within a single plane, then two of the three straight lines being taken in this plane, the three equations of equilibrium will evidently be reduced to two.

The conditions for the equilibrium of a particle acted on by oblique forces, appear to have been first distinctly conceived by Stevin of Bruges.¹ He establishes by reasoning, which although indirect is satisfactory and ingenious, the ratio which the weight of a particle supported on an inclined plane bears to the force by which it is sustained, the force being supposed to act along the plane. He then announces generally, without however supplying a corresponding extension of demonstration, that the condition of equilibrium of any three forces acting on a particle, consists in the proportionality of the forces to the sides

Beghinselen der Waaghconst, 1586. I. Livre de la Statique, prop. 19.

of a triangle to which they are parallel. The first rigorous demonstration of Stevin's theorem in its general form, was obtained by Roberval¹ from the nature of the lever. The idea of a triangle of equilibrium had occurred indeed somewhat earlier to Michael Varro, of Geneva, in application to the equilibrium of forces acting on the sides of a right-angledtriangular wedge: it does not appear, however, that Varro's notion was based upon any very distinct conception of the nature of statical pressure. The Principle of the Parallelogram of Forces, which is in fact a mere modification of Stevin's theorem, was announced almost simultaneously by Newton³ and Varignon; by whom it was inferred from the consideration of the composition of motions. In the same year was published by Lami, in a little treatise entitled Nouvelle manière de démontrer les principaux Théorèmes des élémens des Mécaniques, a theorem in which it is asserted, that if three forces P, Q, R, keep a particle at rest, then

$$P: Q: R: \sin(Q, R): \sin(P, R): \sin(P, Q),$$

where (Q, R), (P, R), (P, Q), denote the angles between the directions of Q and R, P and R, P and Q, respectively. The virtual coincidence of this theorem with the Principle of the Parallelogram of Forces, subjected Lami to the imputation of plagiarism, an aspersion cast upon him by the author of the Histoire des Ouvrages des Savans, (April 1688). Lami combatted this insinuation in a letter published in the Journal des Savans, (Sept. 13, 1688), to which the Journalist replied in the following December, when the controversy appears to have terminated. The first unexceptionable demonstration of the Parallelogram of Forces on pure statical principles, without the introduction of the idea of motion, was given by Daniel Bernoulli. Many other proofs of the proposition have been

¹ Traité de Mécanique, printed in 1636, in the Harmonie Universelle de Mersenne, and in a work also by Mersenne, entitled Cogitata Physico-Mathematica, published in 1644.

^{*} Tractatus de Motu, 1584.

Principia, lex iii. cor. 2, 1687.

¹ Projet de la Nouvelle Mécanique, 1687.

⁸ Comment. Petrop., tom. 1. p. 126, 1726.

since given. Eighteen demonstrations have been collected and examined by Jacobi', by the following authors: D. Bernoulli, 1726; Lambert, 1771; Scarella, 1756; Venini, 1764; Araldi, 1806; Wachter, 1815; Kæstner, Marini, Eytelwein, Salimbeni, Duchayla; two different proofs by Foncenex, 1760; three by D'Alembert; and those of Laplace and Poisson.

SECT. 1. No Friction.

1. P and W (fig. 17) are two heavy particles; W is attached to the end of a fine thread, and P is suspended from a fixed point C of the thread; the thread has one extremity attached to a fixed point A, and passes through a smooth small ring at B in the same horizontal line with A; to find the ratio between P and W, that the vertical line through C may bisect AB in D.

From the supposition it is evident that $\angle ACD = \angle BCD$; let each of these angles be denoted by θ : let T = the tension of the string CA; CA = b, AB = a; the ring B being perfectly smooth, W will be the tension of the string BC.

Hence for the equilibrium of the point C we have, resolving vertically the forces which act on it,

$$(T+W)\cos\theta=P$$
,

and resolving horizontally,

$$T \sin \theta = W \sin \theta$$
, or $T = W$;

hence

$$2 W \cos \theta = P, \qquad \cos \theta = \frac{P}{2 W} \dots (1);$$

but from the geometry we see that

$$b \sin \theta = \frac{1}{2}a, \quad \sin \theta = \frac{a}{2b} \dots (2).$$

Squaring the equations (1) and (2), and adding the resulting equations, we have

$$1 = \frac{P^2}{4W^2} + \frac{a^2}{4b^2}, \qquad \frac{P^2}{4W^2} = \frac{4b^2 - a^2}{4b^2},$$

^{&#}x27; Whewell's Philosophy of the Inductive Sciences, vol. 1. p. 197.

and therefore
$$\frac{P}{W} = \frac{(4b^2 - a^2)^{\frac{1}{4}}}{b}$$
,

which determines the required ratio.

2. A particle P (fig. 18) is placed on the surface of a smooth prolate spheroid, and is attracted towards the foci S and H with forces varying as SP^m and $HP^{m'}$; to find the position of equilibrium.

Draw a tangent KPL at the point P in the plane passing through the three points S, H, P; let μ , μ' , be the absolute forces towards S, H; SP = r, HP = r'. Then for the equilibrium of the particle we have, resolving forces parallel to the line KPL,

$$\mu \cdot r^m \cos \angle SPK = \mu' r'^{m'} \cos \angle HPL$$
;

but $\angle SPK = \angle HPL$, by the nature of ellipses; hence

$$\mu r^m = \mu' r'^{m'}$$
;

also 2a denoting the axis of the spheroid, 2a = r + r'; hence for the determination of r and r' we have the two equations

$$\mu r^{m} = \mu' (2a - r)^{m'}, \qquad \mu (2a - r')^{m} = \mu' r'^{m'}.$$

3. Two weights m, m', are attached to the points O, O', (fig. 19) of a string AOO'A', suspended from two tacks at A and A' in the same horizontal line; to find the positions of the points that their vertical distances from the horizontal line through A and A' may have given equal values.

Draw OE, O'E', vertical; let OE = a = O'E', AA' = b, c = the length of the string; $\angle AOE = \theta$, $\angle A'O'E = \theta'$; T = the tension of the string OO'.

Then for the equilibrium of O we have, by Lami's Principle,

$$\frac{T}{m} = \frac{\sin (\pi - \theta)}{\sin (\frac{1}{2}\pi + \theta)} = \tan \theta;$$

and for the equilibrium of O,

$$\frac{m'}{T} = \frac{\sin\left(\frac{1}{2}\pi + \theta'\right)}{\sin\left(\pi - \theta'\right)} = \cot \theta'.$$

From these two equations we get

$$m \tan \theta = m' \tan \theta' \dots (1)$$

Again, from the geometry,

$$EE' = AA' - AE - A'E'$$

$$= b - a (\tan \theta + \tan \theta');$$

but we have also, from the geometry,

$$EE' = OO' = c - AO - A'O'$$
$$= c - a (\sec \theta + \sec \theta');$$

hence
$$a(\sec \theta - \tan \theta + \sec \theta' - \tan \theta') = c - b \dots (2)$$
.

From the equations (1) and (2) the values of θ , θ' , are to be determined, and then, EO and E'O' being given, AO, A'O' will be known.

Diarian Repository, p. 627.

4. To determine that point in the axis of a hemispherical body, the particles of which attract inversely as the square of the distance, where a corpuscle must be placed so as to remain in equilibrium by the equal and contrary action of the matter of the hemisphere surrounding it.

Let CA (fig. 20) be the axis of the hemisphere, DCD' a diameter of its base, and O the required position of the corpuscle; DAD' the intersection of the plane through CA, DCD', with the surface of the hemisphere; draw BOB' at right angles to CA, join OD; take any points P, p, in the arcs AB, BD, join PO, po, and draw PM, pm, at right angles to CA. Let CA = a = CD, CO = c, CB = b, CD = b', CD = r, CDM = x, CDM

The attraction of a thin slice of the hemisphere at right angles to its axis at the point M, and having a thickness dx, will be

 $2\pi\mu\rho \ dx\left(1-\frac{x}{r}\right)$

as may be seen in elementary treatises on attraction; hence

$$A=2\pi\mu\rho\int_0^{a-c}dx\left(1-\frac{x}{r}\right),$$

$$\frac{A}{2\pi\mu\rho}=a-c-\int_0^{a-c}\frac{x\,dx}{r}\ldots(1);$$

similarly we have

$$B = 2\pi\mu\rho \int_0^c dx' \left(1 - \frac{x'}{r'}\right),$$

$$\frac{B}{2\pi\mu\rho} = c - \int_0^c \frac{x' dx'}{r'} \dots (2).$$

Now from the geometry we see that

$$r^2 = x^2 + y^2 = x^2 + a^2 - (x + c)^2 = a^2 - c^2 - 2cx = b^2 - 2cx,$$
hence $2cx = b^2 - r^3, c dx = -r dr,$
and therefore $\frac{x dx}{r} = -\frac{b^2 - r^2}{2c^2} dr;$

hence from (1), it being observed that r is equal to a-c, b, when x is equal to a-c, 0, we have

$$\frac{A}{2\pi\mu\rho} = a - c + \frac{1}{2c^2} \int_b^{a-c} (b^2 - r^2) dr \dots (3).$$

Again, from the geometry,

$$r'^2 = x'^2 + y'^2 = x'^2 + a^2 - (c - x')^2 = a^2 - c^2 + 2cx' = b^2 + 2cx';$$

hence $2cx' = r'^2 - b^2, cdx' = r'dr',$
and therefore $\frac{x'dx'}{r'} = -\frac{b^2 - r'^2}{2c^2}dr';$

hence from (2), since r' is equal to b', b, when x' is equal to c, 0,

$$\frac{B}{2\pi\mu\rho}=c+\frac{1}{2c^2}\int_b^{b'}(b^2-r'^2)\ dr';$$

but it is evident that

$$\int_{b}^{b'} (b^2 - r'^2) dr' = \int_{b}^{b'} (b^2 - r^2) dr;$$
hence
$$\frac{B}{2\pi\mu\rho} = c + \frac{1}{2c^2} \int_{b}^{b'} (b^2 - r^2) dr \dots (4).$$

But since the corpuscle is in equilibrium we must have A = B, and therefore by (3) and (4),

$$a-c+\frac{1}{2c^2}\int_b^{a-c}(b^2-r^2)\ dr=c+\frac{1}{2c^2}\int_b^{b'}(b^2-r^2)\ dr;$$

hence

$$a-2c=\frac{1}{2c^2}\int_{a-c}^{b'}(b^2-r^2)\,dr$$
:

performing the integration, and putting for b' its value $(a^2 + c^2)^{\frac{1}{4}}$, we shall get, after certain obvious simplifications,

$$a^3 - 4c^3 = (a^2 - 2c^2)(a^2 + c^2)^{\frac{1}{2}};$$

squaring both sides and simplifying,

$$12c^4 - 8a^3c + 3a^4 = 0,$$

an equation from which the value of c is to be determined; as an approximation $c = \frac{3}{7}a$.

Diarian Repository, p. 629.

5. A weight W is sustained upon a smooth inclined plane AB (fig. 21), by three forces, each equal to $\frac{1}{3}W$, one acting vertically upwards, another along AB, and the third parallel to the horizontal line AC; to find the inclination of AB to the horizon.

Required angle of inclination = $2 \tan^{-1} \frac{1}{2}$.

6. Two forces P, Q, of known magnitudes, acting respectively parallel to the base and length of an inclined plane, will each of them singly sustain upon it a particle of weight W; to determine the magnitude of W.

$$W = \frac{PQ}{\left(P^2 - Q^2\right)^{\frac{1}{2}}}.$$

7. Two heavy particles, P and Q, (fig. 22), are connected together by a fine thread passing over a smooth pulley at C; P rests on a smooth inclined plane AB, and Q hangs freely; to determine the position of equilibrium and the pressure on the inclined plane.

Let a = the inclination of the plane to the horizon, R = the pressure, and $\theta =$ the angle CPB; then

$$\cos \theta = \frac{P \sin a}{Q}, \qquad R = P \cos a - (Q^2 - P^2 \sin^2 a)^{\frac{1}{2}}.$$

8. A particle P is placed within a thin parabolic tube AP, (fig. 23), the axis Ax of the parabola being vertical; the particle

is acted on by gravity and by a force μ . PM tending from Ax, to which PM is perpendicular; to find the conditions of equilibrium.

There will be no position of equilibrium unless the latus rectum of the parabola be equal to $\frac{2g}{\mu}$; and under this condition every point of the tube will be a position of rest.

9. The sides of an isosceles triangle are formed of slender uniform prisms, attracting with forces which vary inversely as the square of the distance; to determine the vertical angle in order that a particle may remain at rest in a point which divides the perpendicular from the vertex in a given ratio.

If a be the distance of the particle from the vertex, and b from the base, then

vertical angle =
$$2 \sin^{-1} \left(\frac{b}{a} \right)$$
.

10. Two straight lines AB, AC, at right angles to one another, attract a particle P placed at the point where the perpendicular AP meets BC; to find the direction and magnitude of the force necessary to keep the particle at rest, the law of attraction being that of the inverse square.

Let AB = a, AC = b, BC = c, $\mu =$ the absolute force of a unit of length of the attracting lines condensed into a point; then the direction of the required force will make an angle of 45° with AB, and its magnitude will be equal to

$$\frac{2^{\frac{1}{4}}\mu c^3}{c^2b^2}$$
.

11. The perpendiculars from a point in a triangle upon the sides are a, b, c, and the angles which the sides subtend at the point are a, β , γ ; a particle is placed at this point, and is acted on by the attraction of the sides, the law of attraction being that of the inverse square; to find the relations between a, b, c, and a, β , γ , that the particle may be at rest.

The required relations are

$$\frac{1}{a}:\frac{1}{b}:\frac{1}{c}::\frac{\sin\frac{1}{2}a}{a}:\frac{\sin\frac{1}{2}\beta}{\beta}:\frac{\sin\frac{1}{2}\gamma}{\gamma}.$$

SECT. 2. Friction.

1. Two heavy particles P and P' (fig. 24) rest on two inclined planes CA, C'A, and are connected together by a fine string passing over a smooth pulley at O in the vertical line through A; to determine the positions of P and P' when P is only just sustained.

Let a be the length of the string POP, T its tension, which will be the same throughout; W, W', the weights of the particles P, P'; μ , μ' , the coefficients of friction on the planes CA, C'A, and R, R', their reactions; a, a', the inclinations of the two planes to the vertical.

Then, since by hypothesis the friction on P will be exerted up CA, and that on P down AC', we have for the equilibrium of P, resolving forces parallel and perpendicular to CA,

$$\mu R + T \cos (a - \theta) = W \cos a \dots (1),$$

$$R + T \sin (a - \theta) = W \sin a \dots (2);$$

and in the same way for the equilibrium of P,

$$\mu'R' + W' \cos \alpha' = T \cos (\alpha' - \theta') \dots (3),$$

$$R' + T \sin (\alpha' - \theta') = W \sin \alpha' \dots (4).$$

From (1) and (2),

$$T\left\{\cos\left(a-\theta\right)-\mu\sin\left(a-\theta\right)\right\}=W\left(\cos\,a-\mu\sin\,a\right);$$
 and from (3) and (4),

$$T\left\{\cos\left(\alpha'-\theta'\right)+\mu'\sin\left(\alpha'-\theta'\right)\right\}=W'(\cos\alpha'+\mu'\sin\alpha').$$

Eliminating T between these two last equations, we obtain

$$W'(\cos a' + \mu' \sin a') \{\cos (a - \theta) - \mu \sin (a - \theta)\}$$
= $W(\cos a - \mu \sin a) \{\cos (a' - \theta') + \mu' \sin (a' - \theta')\}.$

Assume $\mu = \tan \epsilon$, $\mu' = \tan \epsilon'$; then this equation becomes

$$W' \cos(a'-\epsilon')\cos(a-\theta+\epsilon) = W\cos(a+\epsilon)\cos(a'-\theta'-\epsilon')...(5).$$

Again, from the geometry, if CA = k,

$$OP = \frac{k \sin a}{\sin (a - \theta)}, \qquad OP' = \frac{k \sin a'}{\sin (a' - \theta')},$$

and therefore, since OP + OP' = a,

$$a = \frac{k \sin \alpha}{\sin (\alpha - \theta)} + \frac{k \sin \alpha'}{\sin (\alpha' - \theta')} \dots (6).$$

The angles θ , θ' , are to be determined from the equations (5) and (6).

2. Given the semi-sum and semi-difference of the greatest and least angles which the direction of a force, supporting a heavy particle on a rough inclined plane, may make with the plane, and the least elevation of the plane when the particle would, without being supported, slide down it; to determine the angle at which the same force, when inclined to a smooth plane of the same elevation, would support the same particle.

Let ε denote the least angle which the force may make with the rough plane to support the particle, P the magnitude of the force, R the reaction of the plane at right angles to itself, μ the coefficient of friction, a the inclination of the plane to the horizon, W the weight of the particle. Then, since the friction must in this case act down the plane, we have for the equilibrium of the particle, resolving forces parallel to the inclined plane, $P \cos \varepsilon = \mu R + W \sin a$:

and resolving forces at right angles to the plane,

$$P \sin \epsilon + R = W \cos \alpha$$
.

Eliminating R between these two equations, we get

$$P(\cos \epsilon + \mu \sin \epsilon) = W(\mu \cos a + \sin a) \dots (1).$$

Let ϕ be the least elevation of the plane for the particle without support to slide down it, then tan ϕ will be equal to μ ; hence from (1),

$$P = W \frac{\sin (a + \phi)}{\cos (\varepsilon - \phi)} \dots (2).$$

If ϵ' denote the greatest angle which P may make with the inclined plane consistently with the equilibrium of the particle, then the friction will act with the greatest force it can exert up

the plane; hence, making μ negative, or putting $-\phi$ for ϕ , we shall have from (2), ϵ' replacing ϵ ,

$$P = W \frac{\sin (a - \phi)}{\cos (\epsilon' + \phi)} \dots (3).$$

Also, if ϵ denote the angle of P's inclination in the case of a smooth plane of the same elevation, we have, putting $\phi = 0$ in (2), and replacing ϵ by ϵ ,

$$P = W \frac{\sin \alpha}{\cos \varepsilon} \dots \dots \dots (4).$$

From (2) and (3)

$$\cos(\varepsilon-\phi)+\cos(\varepsilon'+\phi)=\frac{W}{P}\{\sin(\alpha+\phi)+\sin(\alpha-\phi)\},$$

and therefore if $S = \frac{1}{2}(\varepsilon' + \varepsilon)$ and $D = \frac{1}{2}(\varepsilon' - \varepsilon)$, we get

$$2\cos S\cos (D+\phi)=2\frac{W}{P}\sin a\cos \phi$$

=
$$2 \cos \epsilon' \cos \phi$$
, by (4);

hence
$$\cos \varepsilon' = \frac{\cos S}{\cos \phi} \cos (D + \phi).$$

3. P is the lowest point on the rough circumference of a circle in a vertical plane at which a particle can rest, friction being equal to pressure; to determine the inclination of the radius through P to the horizon.

Required angle =
$$\frac{\pi}{4}$$
.

CHAPTER III.

EQUILIBRIUM OF A SINGLE BODY.

Let any system of forces act upon a body consisting of a system of points rigidly connected together. Take any three straight lines OA, OB, OC, (fig. 25), in space, no two of which are parallel to each other. Let A, B, C, denote the whole resolved part, parallel to each of these straight lines, of any one of the forces of the system; A, B, C, being positive or negative quantities according as they act in the directions OA, OB, OC, or the opposite ones. Then, $\Sigma(A)$, $\Sigma(B)$, $\Sigma(C)$, denoting the algebraical sums of the resolved parts of all the forces of the system parallel to these three straight lines, it is necessary for the equilibrium of the body that we have

$$\Sigma(A) = 0$$
, $\Sigma(B) = 0$, $\Sigma(C) = 0$(I.)

Again, let OA', OB', OC', be any three straight lines in space, no two of which are parallel to each other. Let any force of the system be resolved into two parts, the one at right angles to OA, and the other parallel to it; let A' be the magnitude of the part which is at right angles to OA, and a' the perpendicular distance between OA' and the direction of A'; then A'a' is called the moment of A' about O'A', and the sum of all such moments for all the forces of the system will be denoted by $\Sigma(A'a')$, those moments being considered positive which tend to twist the body about OA' in one direction, and those which tend to twist it in the opposite direction being considered negative. Similarly the algebraical sums of the moments about O'B', O'C', respectively, will be denoted by $\Sigma(Bb')$, $\Sigma(C'c')$. Then for the equilibrium of the body it is necessary that we have

$$\Sigma(A'a')=0$$
, $\Sigma(B'b')=0$, $\Sigma(C'c')=0$...(II.)

The three equations (I.), together with the three equations (II.), are universally sufficient and universally necessary for the equilibrium of any body. It may be proper to remark, that any of the lines OA', OB', OC', may be taken to coincide with any of the lines OA, OB, OC, according to convenience.

If all the forces of the system lie in one plane, then the lines OA, OB, OA', OB', being taken within this plane, and the line OC at right angles to it, the six equations of equilibrium will be reduced to the three following,

$$\Sigma(A) = 0$$
, $\Sigma(B) = 0$, $\Sigma(C'c') = 0$ (III.);

for it is evident that the three other equations will be identically satisfied.

The basis of the general equations of equilibrium consists in the Theory of the Composition and Resolution of Forces, of which we have treated in the preceding chapter, and in the Theory of Moments. The latter theory, in the case of weights acting at right angles to the arms of a straight lever, was established by Archimedes. In the year 1499, the condition of equilibrium of a force acting obliquely on a lever, and supporting a weight suspended from it, was correctly stated by Leonardo Da Vinci, the celebrated painter, to whom must therefore be ascribed the discovery of the theory of oblique action, investigated at a later date by Stevin, in application to the Equilibrium of a Particle. The following elegant geometrical proposition, the application of which to the general Theory of Moments depends upon the Principle of the Parallelogram of Forces, was given by Varignon in his Nouvelle Mécanique, sect. I, lemma xvI: If from any point whatever in the plane of a parallelogram we let fall perpendiculars upon the diagonal and upon the two sides which comprehend this diagonal, the product of the diagonal by its perpendicular is equal to the sum of the products of the two sides by their

¹ 'Αρχιμήδους 'Επιπέδων Ισορροπικών η κέντρα βαρών έπιπέδων το Α. Πρότ. στ. Ζ.

⁸ Venturi; Essai sur les Ouvrages Physico-Mathématiques de Léonard da Vinci, avec des Fragmens tirés de ses Manuscrits apportés d'Italie, Paris, 1797; quoted in Whewell's History of the Inductive Sciences, vol. 11, p. 122.

respective perpendiculars, if the point lie without the parallelogram, or to their difference, if it lie within the parallelogram. The six general conditions of equilibrium of a system of rigidly connected points acted on by any forces whatever, were first laid down by D'Alembert, in the second chapter of his Recherches sur la Precession des Equinoxes, published in the year 1749.

SECT. 1. No Friction.

(1) A beam AB (fig. 26) rests with one end against a horizontal plane in a point A, and with the other against a vertical one in the point B; the vertical plane passing through the beam intersects at right angles the former plane in the line AC, and the latter in the line BC; the beam is attached to the point C by a string EC without weight: to find the tension of the string, E being any assigned point in the beam.

The actions of the horizontal and vertical planes upon the beam at A and B, will be in the directions AR' and BR, which are parallel respectively to CB and CA; let them be denoted by R' and R. Again, let T denote the tension of the string EC. Let G be the centre of gravity of the beam, and W its weight; then instead of supposing the beam to have weight, we may suppose it to be a rigid rod without weight, provided that we apply the force W vertically downwards at G. Thus we have four forces, R, R', T, W, acting at four points B, A, E, G, rigidly connected together. We proceed to express the equations of equilibrium. Let $\angle ECA = \varepsilon$, $\angle BAC = a$, AG = BG = a. Then resolving the forces parallel to the line CA, we have

$$R - T \cos \epsilon = 0 \ldots (1);$$

resolving the forces parallel to CB, we have

$$R'-W-T\sin \epsilon=0\ldots(2);$$

and taking moments about the point C,

$$R \cdot 2a \sin a + Wa \cos a - R' \cdot 2a \cos a = 0,$$

or
$$2R \sin a + W \cos a = 2R' \cos a \dots$$
 (3).

From (1), (2), (3), there is $2T\cos \varepsilon \sin \alpha + W\cos \alpha = 2W\cos \alpha + 2T\sin \varepsilon \cos \alpha,$ $2T(\cos \varepsilon \sin \alpha - \cos \alpha \sin \varepsilon) = W\cos \alpha,$ $2T\sin (\alpha - \varepsilon) = W\cos \alpha,$ and therefore $T = \frac{W\cos \alpha}{2\sin (\alpha - \varepsilon)}.$

If ε be equal to α we have $T = \infty$, which shews that no tension, however great, can sustain the beam in a position of equilibrium. It is easily seen that in this case E coincides with G; and that the length of CE is sufficient to allow the beam to descend continually.

If ε be greater than a, T will clearly be negative; and since the string can pull but not push, the equilibrium is impossible. Thus for the possibility of the equilibrium we must have a greater than ε .

(2) A smooth beam AB, (fig. 27), rests against two horizontal bars which pierce the vertical plane through the beam at right angles in the points A', A''; the beam passes under the lower and over the higher bar, its lower extremity A being sustained upon a smooth horizontal plane: to determine the pressures upon the two bars and upon the horizontal plane.

The pressures upon the bars and upon the horizontal plane will be equal to their reactions upon the beam; the reactions of the bars upon the beam will be two forces R', R'', at right angles to the beam; and the reaction of the horizontal plane will be a vertical force R. Let G be the centre of gravity of the beam; then if we suspend its weight W from G, we may, without affecting the circumstances of the equilibrium, conceive the beam to be a rigid rod without weight. Thus we have four forces R, R', R', W, acting respectively at four points A, A', A', G, rigidly connected together, so as to produce equilibrium. Let AG = a, A'A'' = b, and a = the inclination of the beam to the horizon.

Then resolving forces parallel to the beam, we have

$$W \sin a - R \sin a = 0$$
, and therefore $R = W \dots (1)$.

Remains times at right angles to the beam,

$$R' + W' \cos a - R' - R \cos a = 0,$$

and therefore by (1),

$$R' = R \ldots (2)$$

Igain, taking moments about A,

$$R \cdot AA - R' \cdot AA' - Wa \cos a = 0$$

and therefore by (2),

$$Rb = Wa \cos a$$
;

whence
$$R' = R' = \frac{Wa \cos a}{b}$$
.

(3) A rigid rod AB, fig. (28), rests upon a fixed point E, while its lower extremity A presses against a vertical line FF; to find its position of equilibrium and also the pressures at A and E.

Let G be the centre of gravity of the rod, and W its weight; we suppose the whole weight of the rod to be collected at its centre of gravity. Let R be the reaction of the vertical line FF' upon the rod, which will be at right angles to FF'; also let R' be the reaction of the fixed point E which will be at right angles to the rod. Let EF be at right angles to FF''; and let EF = c, AG = a, $\angle AEF = \theta$.

Then, resolving forces parallel to the rod,

$$R \cos \theta = W \sin \theta \dots (1);$$

resolving forces at right angles to the rod,

$$R' = W \cos \theta + R \sin \theta \dots (2);$$

and taking moments about E,

$$R. AE \sin \theta = W. EG \cos \theta$$

= $W(AG - AE) \cos \theta$
= $W'(a \cos \theta - c)$,

and therefore $Rc \sin \theta = W(a \cos^2 \theta - c \cos \theta) \dots (3);$ hence from (1) and (3),

$$Wc \frac{\sin^2 \theta}{\cos \theta} = W(a \cos^2 \theta - c \cos \theta)$$

$$\frac{c}{\cos \theta} = a \cos^3 \theta, \cos \theta = \left(\frac{c}{a}\right)^{\frac{1}{3}} \dots \dots (4),$$

which gives the value of θ , and defines the position of the beam. From (1) and (4) we have

$$R = W \tan \theta = W \frac{\left\{1 - \left(\frac{c}{a}\right)^{\frac{3}{2}}\right\}^{\frac{1}{4}}}{\left(\frac{c}{a}\right)^{\frac{1}{3}}} = W \frac{\left(a^{\frac{1}{3}} - c^{\frac{3}{3}}\right)^{\frac{1}{4}}}{c^{\frac{1}{3}}},$$

which determines the pressures on the vertical line.

Also, from (1) and (2),

$$R' = W \cos \theta + W \frac{\sin^2 \theta}{\cos \theta} = \frac{W}{\cos \theta},$$

and therefore by (4), $R' = W\left(\frac{a}{c}\right)^{\frac{1}{3}}$,

which determines the presure on the fixed point.

If c be greater than a, then we see by (4) that $\cos \theta$ would be greater than unity, which is impossible; thus equilibrium is impossible unless a be at least equal to c.

Fontana, Memorie della Societa Italiana, 1802; p. 626, &c.

(4) One end A of a beam AB, (fig. 29), is connected to a fixed point by a hinge, about which the beam is capable of revolving in a vertical plane; the other end B is attached to a weight P by means of a string passing over a pulley C in the same vertical plane; to find the position of equilibrium.

Let a horizontal line AD through A meet a vertical line through C in the point D. Let G be the centre of gravity of the beam, at which we shall suppose its whole weight W to be collected. Produce CB to meet AE at right angles to it.

AG = a, BG = b, AD = k, CD = l, $\angle BAD = \theta$, ϕ = the inclination of CE to the horizon.

Then taking moments about A,

$$P \cdot AE = W \cdot AF$$
 or $P(a+b)\sin(\phi-\theta) = Wa\cos\theta...................(1);$ again, from the geometry,

$$(a + b) \sin \theta + BC \sin \phi = l,$$

$$(a + b) \cos \theta + BC \cos \phi = k,$$

and therefore, eliminating BC,

$$\frac{(a+b)\sin\theta-l}{(a+b)\cos\theta-k}=\tan\phi\ldots(2).$$

The equations (1) and (2) are sufficient for the determination of θ and ϕ , or of the position of equilibrium.

(5) A uniform beam AB, (fig. 30), of which one end A is placed upon a smooth horizontal plane OA, and of which the other end B presses against a vertical plane OB, is attracted by a centre of force situated in the point O, the intensity of the force varying directly as the distance; to determine the position of equilibrium.

Conceive the beam to be inclined at an angle ω to the horizon. Take P any point in the beam and join OP.

Let OP = r, AP = s, AG = BG = a, $\angle POA = \theta$, $\mu =$ absolute force of attraction, R, R', the reactions of the planes at A, B. Then, resolving forces horizontally, we have

$$R' = \int \mu r ds \cos \theta = \mu \int_0^{2a} ds (2a - s) \cos \omega = 2\mu a^2 \cos \omega \dots (1).$$

Resolving forces vertically,

$$R - W = \int \mu r ds \sin \theta = \mu \int_0^{2a} \sin \omega s ds = 2\mu a^2 \sin \omega \dots (2).$$

Taking moments about O,

$$R \cdot 2a \cos \omega = Wa \cos \omega + R' \cdot 2a \sin \omega$$

 $2R \cos \omega = W \cos \omega + 2R' \sin \omega \dots (3),$

and therefore, substituting in this equation the values of R' and R from (1) and (2), we have

$$2 \cos \omega (W + 2\mu a^2 \sin \omega) = W \cos \omega + 2 \sin \omega \cdot 2\mu a^2 \cos \omega,$$

and therefore
$$W \cos \omega = 0, \quad \omega = \frac{1}{2}\omega,$$

or the beam lies in contact with the vertical plane OB.

It is evident, however, that this is not the only position of equilibrium; the beam will plainly remain at rest if it be placed in contact with the horizontal plane OA with one extremity at O. In writing down the equations (1), (2), (3), it is tacitly assumed that the beam receives no pressure from the planes ex-

cepting at its extremities, an hypothesis which holds good in the former position of equilibrium while it evidently does not in the latter: it is for this reason that, in our analytical investigation, out of the two admissible values 0 and $\frac{1}{2}\pi$ for ω we obtained only the latter.

(6) A uniform bar AB (fig. 31) is placed in the straight line joining two centres of force K, L, which attract with forces varying directly as the distance; to find the position in which the bar will rest.

Let μ , μ' , be the absolute forces of the centres K, L; let P be any point in the bar AB; KA = x, LB = y, KP = s, BP = s', AB = 2a, KL = l; ρ = the density of the bar, κ = the area of a transverse section. Then for the equilibrium of the bar we must have

$$\int_{a}^{a+2a} \kappa \rho \mu \ s \ ds = \int_{y}^{y+2a} \kappa \rho \mu' \ s' \ ds' \ ;$$

or, since κ and ρ are supposed to be the same for all points of the bar,

$$\mu \int_{a}^{x+2a} s \, ds = \mu' \int_{y}^{y+2a} s' \, ds',$$

$$\mu \left\{ (x+2a)^{2} - x^{2} \right\} = \mu' \left\{ (y+2a)^{2} - y^{2} \right\},$$

$$\mu \left(x+a \right) = \mu' \left(y+a \right) = \mu' \left(l-a-x \right),$$

$$(\mu + \mu') x = \mu' l - (\mu + \mu') a,$$

$$x = \frac{\mu' l}{\mu + \mu'} - a;$$
similarly
$$y = \frac{\mu l}{\mu + \mu'} - a.$$

The value of x, or of y, determines the position of equilibrium.

(7) To find the force requisite to keep a door in a given position, the post being inclined at a given angle to the vertical; neglecting friction.

Let AB(fig. 32) be the door-post, ABCD the door; Az a vertical line through A; Ax at right angles to Az and in the plane of BAz; E the intersection of the line CD produced with the horizontal plane through A; join AE. With A as a centre

describe a sphere cutting AB, Ax, AE, in the points p, q, r; and join these points by great circles of the sphere.

Let $\angle BAz = \beta$, and a = the inclination of the plane of the door to the plane xAz; W = the weight of the door.

Then, since the angle pqr of the spherical triangle pqr is a right angle, we have, by Napier's rules,

$$\cos prq = \sin qpr \cos pq = \sin a \sin \beta;$$

but if ϕ denote the angle which W's direction makes with the plane of the door, it is clear that

$$\sin \phi = \cos prq$$
;

hence, a denoting the distance of the centre of gravity of the door from the post, the moment of W about AB will be equal to

Wa sin a sin
$$\beta$$
;

let P be the force applied at right angles to the door, at a point distant from the door-post by a space b, sufficient to keep it in its present position; then, by the equation of moments, we have

$$Pb = Wa \sin a \sin \beta$$
.

(8) A rigid rod AB (fig. 33) rests against a smooth vertical wall CD, and has its lower extremity A attached to a hinge about which it can revolve freely; to find the pressure on the wall and upon the hinge.

Let G be the centre of gravity of the rod at which we may suppose its whole weight W to be collected; let AG = b, AB = a, $\angle BAC = a$. Also let R denote the reaction of the wall against the rod, which will take place at right angles to CD; and let R, S, be the vertical and horizontal parts of the reaction of the hinge upon the rod. Then

$$R = \frac{Wb}{a \tan a} = S, R' = W.$$

This problem was first proposed under a vicious form in a work by Stone; where the author proposes to determine the position of AB corresponding to a maximum value of R. In a review of Stone's work by John Bernoulli', the solution given by Stone was declared to be erroneous, and a different one was

¹ Opera, tom. IV. p. 189.

offered by the reviewer. Bernoulli's solution is, however, essentially vicious. The problem was correctly solved for the first time by Couplet'. The opinions however, both of mathematicians and of architects, were for many years divided as to the respective merits of the solutions given by Bernoulli and by Couplet, and even down to very late years numerous memoirs have appeared on the subject by different mathematicians, with various conclusions; several of whom have arrived at results at variance with the solutions both of Bernoulli and of Couplet. The reader who may be curious to examine the various solutions of this problem, which by the aberrations of the learned rather than by any intrinsic difficulty has obtained considerable celebrity, is referred to a memoir by Franchini in the Memorie della Societa Italiana, tom. xvi. parte 1, p. 228; 1813.

(9) A ladder of uniform thickness rests with its lower end upon a smooth horizontal plane, and its upper end on a slope inclined at an angle of 60° to the horizon; the ladder makes an angle of 30° with the horizon: to find the force which must act horizontally at the foot of the ladder to prevent sliding.

If W denote the weight of the ladder;

required force =
$$\frac{3^{\frac{1}{4}}}{4}$$
 W.

(10) A sphere rests upon two inclined planes; to find the pressure experienced by each.

Let W be the weight of the sphere; a, a', the inclinations of the inclined planes to the horizon; and R, R', their respective pressures. Then

$$R = \frac{W \sin a'}{\sin (a + a')}, R' = \frac{W \sin a}{\sin (a + a')}$$

Leibnitz; Opera, tom. III. p. 176.

(11) A uniform beam rests upon two perfectly smooth inclined planes; to find its position and its pressure upon the two planes.

Let a, a', be the inclinations of the two planes to the horizon; R, R', the pressures which they experience; then, supposing the end of the beam which rests against the former plane to be the

² Mémoires de l'Académie de Paris, 1781, p. 69.

lower one, and θ to be the inclination of the beam to the horizon, we shall have, W being the weight of the beam,

$$\tan \theta = \frac{\sin (a'-a)}{2 \sin a' \sin a}, \quad R = \frac{W \sin a'}{\sin (a+a')}, \quad R' = \frac{W \sin a}{\sin (a+a')}.$$

(12) A beam AB (fig. 34) leans against a smooth vertical prop CD, the end A being prevented from sliding along the horizontal plane AD by a string AD fastened at D; to find the tension of the string.

Let G be the centre of gravity of the beam; AG = a, CD = b, AD = c, W = the weight of the beam, T = the tension; then

$$T = \frac{abc}{\left(b^2 + c^2\right)^{\frac{3}{2}}} W.$$

(13) A uniform rigid rod AB (fig. 28) rests upon a fixed point E, while its lower end A presses against a vertical line FF'; a weight P is suspended from the extremity B; to find its position of equilibrium.

Let W = the weight of the rod, b = the perpendicular distance of E from the line FF', AE = x, a = the length of the rod; then

$$x = \left(ab^2 \frac{P + \frac{1}{2} W}{P + W}\right)^{\frac{1}{3}}.$$

Fontana, Memorie della Societa Italiana, 1802, p. 630.

If we suppose W = 0, then we shall have $x = (ab^3)^{\frac{1}{3}}$, whatever be the magnitude of P. This problem is discussed by Euler, Acad. des Sciences de Berlin, tom. VII. p. 196, in illustration of Maupertuis' Principle of Rest.

(14) A sphere, of which C is the centre, is supported on an inclined plane AB by a string CB which is horizontal; to find the tension of CB.

If W be the weight of the sphere, and a the inclination of the plane to the horizon,

(15) A uniform rod of given length rests against a peg at the focus of a parabola, its lower extremity being supported on the curve; to determine the angle which it makes with the axis of the parabola which is vertical.

If a be the length of the rod, and 4m the latus rectum of the parabola; then

required angle =
$$2 \cos^{-1} \left(\frac{m}{a}\right)^{\frac{1}{4}}$$
.

(16) A weight W hangs from a rod BC, (fig. 35), which rests on a fulcrum at B, and is supported by a string DA at right angles to the rod, D being the middle point of BC; to determine the magnitude and direction of the pressure on the fulcrum, the rod being inclined to the horizon at an angle of 30° , and being without weight.

Let BD = CD = a; and let X, Y, represent the vertical and horizontal pressure exerted by the rod on the fulcrum; then

$$X = \frac{1}{2} W, \quad Y = \frac{3^{\frac{1}{4}}}{2} W;$$

and, if ϕ denote the inclination of the resultant pressure to the vertical, and R its magnitude,

$$R=W, \quad \phi=\frac{1}{3}\pi.$$

(17) One end of a beam is connected with a horizontal plane by a hinge about which the beam can revolve freely in a vertical plane; the other end is attached to a weight by means of a string passing over a pulley in the same vertical plane; to find the position of equilibrium.

Let a, b, be the distances of the centre of gravity of the beam from its lower and its higher extremities, W its weight, and θ its inclination to the horizon; let ϕ be the inclination of the string to the horizon, and P the weight attached to its extremity; let l be the distance of the pulley from the horizontal and k from the vertical line through the hinge. Then the position of equilibrium will depend upon the equations

$$P(a+b)\sin(\phi-\theta) = Wa\cos\theta,$$

$$(a+b)\sin(\phi-\theta) = k\sin\phi - l\cos\phi.$$

(18) To find the position of equilibrium of a uniform beam, one end of which rests against a vertical plane, and the other on the interior surface of a given hemisphere.

Let r be the radius of the hemisphere, c the distance of its centre from the vertical plane, 2a the length of the beam; θ the inclination of the beam to the horizon, and ϕ of the radius at the

point where the beam presses against the hemisphere. Then the position of equilibrium will depend upon the equations

$$\tan \phi = 2 \tan \theta$$
, $2a \cos \theta = r \cos \phi + c$.

(19) A uniform beam rests with one end upon a given inclined plane, the other end being suspended by a string from a fixed point above the plane; to determine the position of equilibrium, the tension of the string, and the pressure on the plane.

Let 2a be the length of the beam, θ its inclination to the inclined plane, W its weight, and R the pressure which it exerts on the inclined plane; let T be the tension of the string, c its length, and ϕ its inclination to the inclined plane; also let b be the distance of the fixed point from the plane; and a the inclination of the plane to the horizon.

Then the position of the beam will depend upon the two equations

$$2 \sin (\phi - \theta) \sin a = \cos \phi \cos (\theta + a),$$

$$c \sin \phi + 2a \sin \theta = b;$$

and thence R and T will be given by

$$R = \frac{W\cos(a + \phi)}{\cos\phi}, \ T = \frac{W\sin a}{\cos\phi}.$$

(20) A uniform beam rests with one end against a smooth vertical plane, its other end being supported by a string attached to a fixed point in the plane; to determine the position of the beam, its pressure against the plane, and the tension of the string.

Let b be the length and T the tension of the string; 2a the length of the beam, W its weight, and R its pressure against the vertical plane; also let ϕ , θ , be the inclinations of the beam and of the string to the vertical. Then

$$\sin \theta = \left(\frac{16a^2 - b^2}{3b^2}\right)^{\frac{1}{2}}, \sin \phi = \left(\frac{16a^2 - b^2}{12a^2}\right)^{\frac{1}{2}},$$

$$T = \frac{3^{\frac{1}{2}}b W}{(4b^2 - 16a^2)^{\frac{1}{2}}}, \quad R = \left(\frac{16a^2 - b^2}{4b^2 - 16a^2}\right)^{\frac{1}{2}}. W.$$

(21) A uniform beam AB (fig. 36) moveable in a vertical plane about a hinge at A, leans upon a prop CD situated in the same plane; to determine the strain upon the prop CD.

Let AB = 2a, CD = b, $\angle BAC = a$, $\angle ACD = \beta$. Then the resolved part of the pressure of AB on CD at right angles to CD, which measures the strain on the prop, will be equal to

$$\frac{\mathbf{W}a\,\sin\,2a\,\cos\,(a+\beta)}{2b\,\sin\,\beta}.$$

(22) A uniform isosceles triangle is placed within a smooth hemispherical bowl, its three angles touching the bowl; to find the position in which it will rest.

Let a = the length of each of the equal sides, h = the altitude of the triangle, r = the radius of the hemisphere, $\theta =$ the inclination of the triangle to the vertical; then

$$\tan \theta = \frac{3(4r^3h^2 - a^4)^{\frac{1}{4}}}{4h^2 - 3a^2}.$$

(23) A uniform beam ABC (fig. 37) is placed with one end A in a hemispherical bowl, and, being of greater length than the diameter of the bowl, rests upon the rim of the bowl at the point B; to find the position in which the beam will rest, the radius OB of the bowl being horizontal.

If r be the radius of the bowl, 2a the length of the beam, and θ its angle of inclination to the horizon; then

$$4r\cos^2\theta-a\cos\theta-2r=0.$$

(24) A given weight P is suspended from the rim of a uniform hemispherical bowl placed on a horizontal plane; to find the position in which the bowl will rest.

If W denote the weight of the bowl, c the distance between its centre and its centre of gravity, and θ the inclination of the axis of the bowl to the vertical,

$$\tan \theta = \frac{Pr}{Wc}.$$

(25) An isosceles right-angled triangle rests in a vertical plane with the right angle downwards, between two pegs at a distance a from each other in the same horizontal line; to determine its positions of equilibrium.

Let h = the perpendicular from the right angle on the base, and θ = the inclination of the base to the horizon; then

$$\theta = 0$$
, or $\theta = \cos^{-1}\left(\frac{h}{3a}\right)$.

(26) A uniform circular lamina is placed with its centre upon a prop; to find at what points on its circumference three weights w_1 , w_2 , w_3 , must be attached that it may remain at rest in a horizontal position.

Let (w_1, w_2) , (w_1, w_3) , (w_2, w_3) , denote the angles at the centre of the lamina between the distances of w_1, w_2 ; w_1, w_3 ; respectively. Then

$$\begin{split} \cos\left(w_{_{1}},\,w_{_{2}}\right) &= -\,\frac{w_{_{1}}^{\,2} + w_{_{2}}^{\,2} - w_{_{3}}^{\,2}}{2w_{_{1}}w_{_{3}}}\,\,,\,\,\cos\left(w_{_{1}},\,w_{_{3}}\right) &= -\,\frac{w_{_{1}}^{\,2} + w_{_{3}}^{\,2} - w_{_{2}}^{\,2}}{2w_{_{1}}w_{_{3}}}\,,\\ &\cos\left(w_{_{2}},w_{_{3}}\right) &= -\,\frac{w_{_{2}}^{\,2} + w_{_{3}}^{\,2} - w_{_{1}}^{\,2}}{2w_{_{2}}w_{_{3}}}\,. \end{split}$$

SECT. 2. Friction.

Statical friction consists in the resistance arising from mutual roughness, which is opposed to the production of relative motion between two substances in contact. If the substances were perfectly smooth, their mutual pressure at every point of the surfaces of contact would take place in some determinate straight line depending upon the forms of the surfaces; if the consideration of roughness be introduced, the force of friction when called into play will exert itself at each point in a direction at right angles to the mutual pressure corresponding to perfect smoothness. The estimation of the magnitude of friction for assigned substances and for given surfaces of contact, can be effected solely by experiment.

Suppose R to denote the total pressure of two substances, of which the surfaces of contact are two planes, and let F be the greatest force which friction can exert in the prevention of relative motion; then F is taken as the measure of the statical friction. After the performance of numerous experiments,

Amontons, who was the first to discuss scientifically the subject of friction, was led to conclude that, so long as the substances remain the same, F varies directly as R, and is independent of the magnitude of the area of contact. Thus, μ denoting some constant quantity, the magnitude of which is to be obtained by experiment, we should have for any assigned substances

$$F = \mu R$$
,

where u is called the coefficient of friction. This relation however, although generally adopted by mathematicians, is probably not quite accurate. Muschenbroek,2 and the Abbé Nolet, concluded from experiments that the value of \u03c4 depends in some degree upon the magnitude of the area of contact, and that for an assigned area of contact it does not remain invariable for all values of R. Bossut' agreed with Amontons in supposing u to be independent of the area of contact, but considered that its value decreases as R increases. Various experimenters have exerted their labours on the same subject with very different conclusions. Professor Vince⁵ concluded, after the performance of very careful experiments, that the coefficient of friction does really diminish with the increase of R, and that for a given pressure it decreases when the area of contact is diminished. It would appear however, from the valuable experiments of Coulomb6 and Ximenes7, that the variation of u, owing to any change in the magnitude of the area of contact, is extremely small and of an irregular character, and that it decreases very slightly as R increases. Bossut has remarked, that the statical friction between two substances becomes greater by allowing them to remain in contact for some time before it is called into play, an observation which has been fully confirmed by the experiments of Coulomb.

^{*} Mémoires de l'Académie des Sciences de Paris, 1699, p. 206.

^{*} Introduct. ad Phil. Nat. tom. 1. cap. 9, 1762. Lect. Phys. Exp. tom. 1. p. 241.

^{*} Leçons de Physique Expérimentale, tom. 1. p. 230 ; 1754.

^{*} Traité de Mécanique, Part 1. chap 4, sect. 1, p. 178.

^{*} Philosophical Transactions, 1785, Part 1. p. 165.

^{*} Mémoires Present, a l'Academie, tom. x. 1785.

I Terria e Pentica delle Resist. de' Sol. ne' loro Attr. Pisa, 1782.

If the surfaces of contact be not plane areas, the coefficient of friction will on this account receive a change of value; and generally it will depend upon the forms of the surfaces of contact, as well as upon the nature of the substances. The friction of a solid cylinder against a hollow one has been considered by Coulomb and Ximenes, who have found it to be much smaller than between two plane surfaces of the same substance; the coefficient of friction is, however, approximately constant, as in the case of plane surfaces of contact.

The friction of which we have been speaking, is the friction called into play by the *rubbing* of two substances against each other; the roughness of substances, however, exerts force to interrupt the production of relative motion also in the case when one body is urged to *roll* along another without *rubbing*; this may be called the friction of cohesion, depending probably upon the mutual tenacity of the particles of the two bodies. This species of friction was first noticed by Bossut, and afterwards carefully investigated by Ximenes and Coulomb: in the case of a cylinder rolling along a plane, the friction of cohesion is found to vary inversely as the diameter.

The friction which exists between two substances in motion, which may be called their dynamical friction, is very considerably less than their statical friction. The dynamical friction is measured by the force necessary to keep the bodies in motion; the statical friction by the force necessary to set them originally in motion. The difference of the magnitudes of statical and dynamical friction was noticed by Camus' and Desaguliers², and afterwards by various other experimenters. Professor Vince ascertained by experiments, that dynamical friction is a constant force for hard substances, whatever be the velocity of the relative motion; but that in the case of softer bodies it increases considerably with an increase of velocity. The friction of pivots has been fully considered by Coulomb in the Mémoires de l'Acad. des Sciences de Paris, 1790. The friction and rigidity of ropes was first investigated experimentally by Amontons in the memoir to which we have alluded above, and afterwards by Coulomb and Ximenes.

¹ Traité des Forces Mouvantes.

² Cours de Physique.

(1) A uniform beam AB, (fig. 38), resting with one end A upon a rough horizontal plane KL, has its other end B attached to a string which passes over a smooth pulley E, and supports a weight P; to determine the range of positions in which the beam may be placed consistently with equilibrium.

Let G be the centre of gravity of the beam, and W its weight; θ , ϕ , the angles of inclination of AB, BE, to the horizon for any position of equilibrium; AG = BG = a, F the friction, estimated along LK, which is called into play at A, and which will be at right angles to R the vertical reaction of the plane on the beam. Suppose the whole weight of the beam to be collected at its centre of gravity.

Then for the equilibrium of the beam we have, resolving the forces horizontally,

$$F = P \cos \phi \dots (1);$$

resolving vertically,

$$R + P \sin \phi = W \dots (2);$$

and taking moments about A,

$$Wa \cos \theta = P \cdot 2a \sin (\phi - \theta)$$
or
$$W \cos \theta = 2P \sin (\phi - \theta) \cdot \dots \cdot (3).$$

Assume $F = \lambda R$, where, if μ denote the coefficient of friction between the end of the beam and the plane, λ may have any value from zero up to μ . Then by (1) we have

$$\lambda R = P \cos \phi \dots (4).$$

From (2) and (4), we obtain

$$P\cos\phi + \lambda P\sin\phi = \lambda W$$

or, putting $\lambda = \tan \epsilon$,

$$P\cos(\phi-\epsilon)=W\sin\epsilon$$

which determines the angle ϕ in terms of W, P, ε ; and then θ may be determined from (3). By giving then to ε any values from zero up to $\tan^{-1}\mu$, we shall obtain a series of positions of equilibrium.

Suppose for instance λ to be equal to zero; then from (4)

$$P\cos\phi=0$$
, and therefore $\phi=\frac{1}{2}\pi$;

hence by (3), $W \cos \theta = 2P \cos \theta$,

and therefore either W = 2P, in which case θ remains indeterminate and may have any value whatever, or $\theta = \frac{1}{2}\pi$. Again from (2), since $\phi = \frac{1}{2}\pi$, we have

and therefore, if θ be not equal to $\frac{1}{2}\pi$, we must have $R = P = \frac{1}{2}W$.

Thus we see, that the end B of the beam must be in the vertical line through E; and that, unless AB be placed vertically, the weight P must be equal to half the weight of the beam. If the beam be placed vertically, it is clear from (5) that P may have any value from 0 up to W, but no greater value, because R cannot be negative.

If instead of taking $\lambda = 0$, we were to give it any other value between 0 and μ , we should have to determine the values of θ and ϕ as in the present case.

(2) A beam AB (fig. 39) is supported on a prop CD by a given force P acting at a given angle of inclination to the horizon; to find the position of the beam when it is upon the point of sliding past the point C from A towards B, the prop and beam being relatively rough.

Produce BA, PA, to meet the horizontal line KL in the points F, E; let G be the centre of gravity of the beam. Let AC = a, CG = x, $\angle PEL = a$, $\angle AFE = \theta$, R = the reaction of the prop at right angles to AB, and μ the coefficient of friction; then μR will be the friction, of which BA is the direction.

Then for the equilibrium of the beam we have, resolving forces vertically,

$$P \sin a + R \cos \theta = W + \mu R \sin \theta \dots (1);$$
 resolving horizontally,

$$P \cos a = R \sin \theta + \mu R \cos \theta \dots (2);$$

and taking moments about C ,

$$\mathbf{W}\mathbf{x}\cos\,\boldsymbol{\theta}=\mathbf{P}(\mathbf{a}+\mathbf{x})\sin(\mathbf{a}-\boldsymbol{\theta})\ldots(3).$$

From the equations (1) and (2) there is

$$\frac{\cos \theta - \mu \sin \theta}{\sin \theta + \mu \cos \theta} = \frac{W - P \sin \alpha}{P \cos \alpha},$$

and therefore

$$P \cos a (1 - \mu \tan \theta) = (W - P \sin a)(\tan \theta + \mu),$$

 $P(\cos a + \mu \sin a) - \mu W = \{W + P(\mu \cos a - \sin a)\} \tan \theta;$ assume $\mu = \tan \varepsilon$; then, multiplying both sides of the equation by $\cos \varepsilon$,

$$P\cos(\epsilon - a) - W\sin\epsilon = \{P\sin(\epsilon - a) + W\cos\epsilon\} \tan\theta,$$

$$\tan\theta = \frac{P\cos(\epsilon - a) - W\sin\epsilon}{P\sin(\epsilon - a) + W\cos\epsilon},$$

which determines the inclination of the beam to the horizon.

Knowing θ we may determine x from the equation (3); and thus the position of the beam will be completely ascertained.

If the beam be on the point of sliding in a direction opposite to that which we have supposed, the quantity μ must be replaced by $-\mu$, or ε by $-\varepsilon$; and the formulæ for the former case will all become adapted to the latter.

(3) A uniform rectangular board *KLMN*, (fig. 40), is placed upon a rough inclined plane *AB*; supposing the inclination of the plane *AB* to the horizon to be gradually increased, to find whether the equilibrium of the board will be disturbed by the commencement of a rolling or of a sliding motion.

First suppose that the board begins to slide; let R be the whole of the reaction of the plane at right angles to itself on the board, μ the coefficient of friction, and ϕ the inclination of the plane at the commencement of sliding. Then, resolving forces parallel to the inclined plane,

$$\mu R = W \sin \phi$$
;

and resolving forces at right angles to it,

$$R = W \cos \phi$$
;

hence, eliminating R,

$$\tan \theta = \mu$$
.

Next suppose that the board tumbles over the corner K before the commencement of sliding; then the vertical through G will pass through K when ϕ has received the proper value; draw GH at right angles to the plane, let HK = a, GH = b; then

$$\tan \phi = \tan \angle KGH = \frac{a}{b}.$$

Hence, if μ be less than $\frac{a}{b}$, sliding will take place before rolling; on the contrary, if μ be greater than $\frac{a}{b}$, rolling will take place before sliding; if μ be equal to $\frac{a}{b}$, rolling and sliding will take place simultaneously.

(4) A beam PQ, (fig. 41), which is capable of free motion in every direction about a smooth hinge at P, rests with its end Q against a rough vertical plane ABC; to determine the position of the beam when it is bordering on motion.

From P draw PO at right angles to the plane ABC; join OQ; the locus of Q will be a circle in the vertical plane having O for its centre; let G be the centre of gravity of the beam; PHV be the projection on the horizontal plane through PO of the line PGQ, H and V being the projections of G and Q; draw HK at right angles to PO; let W be the weight of the beam, μ the coefficient of friction between the beam and the vertical plane, and R their mutual pressure; μR will act in the tangent to the locus of Q at the point Q, that is, at right angles to OQ and in the plane ABC, and from A towards B;

let
$$PG = a$$
, $QG = b$, $\angle QPO = a$, $\angle QOA = \theta$.

Then for the equilibrium of the beam we have, taking moments about PO,

$$W. HK = \mu R. OQ.....(1);$$

and taking moments about the horizontal line through P, which is at right angles to PO, it being observed that the vertical resolved part of μR is $\mu R \cos \angle QOV$,

$$W \cdot PK = R \cdot QV + \mu R \cdot PO \cos \angle QOV \cdot \ldots (2)$$

Now from the geometry,

$$HK = GK \cos \theta = a \sin a \cos \theta$$
,

$$OQ = (a + b) \sin a$$
, $PO \cos \angle QOV = (a + b) \cos a \cos \theta$,

$$PK = a \cos a$$
, $QV = OQ \sin \theta = (a + b) \sin a \sin \theta$;

hence from the equations (1) and (2),

Wa sin a cos
$$\theta = \mu R (a + b) \sin a$$
,

and $Wa \cos a = R(a + b) \sin a \sin \theta + \mu R(a + b) \cos a \cos \theta$; dividing the latter of these equations by the former,

$$\frac{\cos a}{\sin a \cos \theta} = \frac{\sin a \sin \theta + \mu \cos a \cos \theta}{\mu \sin a},$$

$$\mu \cos a = \cos \theta (\sin a \sin \theta + \mu \cos a \cos \theta),$$

$$\mu \cos a \sin^{3} \theta = \sin a \sin \theta \cos \theta,$$

$$\mu \tan \theta = \tan a,$$

$$\tan \theta = \frac{1}{\mu} \tan a.$$

We may solve this problem also in the following manner: taking moments about the vertical line through P we have, since $\mu R \sin \theta$ is the horizontal resolved part of μR ,

$$R \cdot OV = \mu R \cdot \sin \theta \cdot PO,$$
 and therefore $OQ \cos \theta = \mu \sin \theta \cdot PO;$ but $OQ = OP \tan a,$ hence $\tan a \cos \theta = \mu \sin \theta, \tan \theta = \frac{1}{\mu} \tan a.$

(5) A beam AB (fig. 42) is placed with one end upon a rough horizontal plane Ox, and rests against a rough plane curve KPL at any point P; supposing that, whatever be the point P against which the beam leans, it is always in an equilibrium bordering on motion, and that the coefficient of friction is the same both for the curve and for the horizontal plane, to find the nature of the curve.

Draw PM at right angles to Ox; let G be the centre of gravity of the beam, W its weight, AG = a, $\angle BAx = \theta$, OM = x, PM = y, $\mu =$ the coefficient of friction; let R and R' be the normal reactions of the curve and of the plane against the beam; in consequence of friction the curve will exert on the beam a force μR along PB, and the horizontal plane a force $\mu R'$ along Ax.

Hence for the equilibrium of the beam, resolving forces parallel to Ox,

$$R \sin \theta = \mu R \cos \theta + \mu R',$$

$$R (\sin \theta - \mu \cos \theta) = \mu R' \dots (1);$$

resolving forces perpendicularly to Ox,

$$R\cos\theta + \mu R\sin\theta + R' = W,$$

$$R(\cos\theta + \mu\sin\theta) + R' = W \dots (2);$$

and taking moments about A,

$$R \cdot AP = Wa \cos \theta$$
, or $R \cdot AM = Wa \cos^2 \theta \cdot \dots (3)$.

From (1) and (2) we get

$$(1+\mu^2) R \sin \theta = \mu W,$$

and therefore from (3)

$$(1 + \mu^2) Wa \sin \theta \cos^2 \theta = \mu W \cdot AM,$$

$$(1 + \mu^2) a \sin \theta \cos^2 \theta = \mu \cdot AM;$$

but
$$\sin \theta = \frac{dy}{ds}$$
, $\cos \theta = \frac{dx}{ds}$, $AM = y \frac{dx}{dy}$;

hence we have,

$$(1 + \mu^2) a \frac{dy}{ds} \frac{dx^2}{ds^2} = \mu y \frac{dx}{dy},$$

$$a(1 + \mu^2) \frac{dy^2}{dx^2} = \mu y \frac{ds^3}{dx^3};$$

put $\mu = \tan \epsilon$, and this equation becomes

$$\frac{2a}{\sin 2s} \frac{dy^3}{dx^3} = y \frac{ds^3}{dx^3};$$

which is the differential equation to the curve.

If the friction of the curve and the plane be different, we may obtain the differential equation to the curve with equal ease.

(6) A beam rests with its lower extremity on a horizontal, and its higher against a vertical plane; having given its length, the position of its centre of gravity, and the coefficients of the friction of the horizontal and of the vertical plane, to find its position when in a state bordering on motion.

If a, b, be the distances of the centre of gravity of the beam from its lower and higher extremity; μ , μ' , the coefficients of friction between the beam and the horizontal, and between

the beam and the vertical plane; and θ the inclination of the beam to the horizon; then

$$\tan \theta = \frac{a - \mu \mu' b}{\mu (a + b)}.$$

(7) A straight uniform beam is placed upon two rough planes, of which the inclinations to the horizon are a and a', and the coefficients of friction $\tan \lambda$ and $\tan \lambda'$; to find the limiting value of the angle of inclination of the beam to the horizon at which it will rest, and the relation between the weight of the beam and each of the two perpendicular pressures upon the planes.

Let θ be the required limiting angle; R, R', the pressures on the planes; and W the weight of the beam. Then

$$\frac{2 \tan \theta = \cot (\alpha' + \lambda') - \cot (\alpha + \lambda),}{R} \frac{R'}{\cos \lambda \sin (\alpha' + \lambda')} = \frac{W}{\sin (\alpha - \lambda + \alpha' + \lambda')} = \frac{R'}{\cos \lambda' \sin (\alpha - \lambda)}.$$

(8) A uniform and straight plank rests with its middle point upon a rough horizontal cylinder, their directions being perpendicular to each other; to find the greatest weight which can be suspended from one end of the plank without its sliding off the cylinder.

Let W be the weight of the plank, and P the attached weight; r the radius of the cylinder, a the length of the plank, tan λ the coefficient of friction. Then P will be given by the relation

$$\frac{P}{W} = \frac{r\lambda}{a - r\lambda}$$
.

(9) A uniform beam AB, (fig. 43), of which the end B presses against a rough vertical plane CD, is supported by a fine string AC attached to a fixed point C in the plane; to find the position of the beam when bordering upon motion.

Let the point B be on the point of ascending; μ = the coefficient of friction; a = the length of the beam, CA = l, $\angle ACB$ = θ , ϕ = $\angle ABD$. Then θ may be found from the equation

$$(4a^2-4l^2-\mu^2l^2)\tan^2\theta-2\mu l^2\tan\theta+4a^2-l^2=0$$
;

and then ϕ may be determined by the equation $a \sin \phi = l \sin \theta$.

- If B be on the point of sliding downwards, μ must be replaced by $-\mu$.
- (10) An elliptical cylinder, placed between a smooth vertical plane and a rough horizontal one, with the major axis of the ellipse inclined at an angle of 45° to the horizon, is just prevented by friction from sliding; to find the coefficient of friction.

If e be the eccentricity of the ellipse, the coefficient of friction will be equal to $\frac{1}{2}e^2$.

- (11) A homogeneous solid hemisphere is capable of rolling on its curve surface upon a horizontal plane, the friction being such as to prevent all sliding; to find the moment of a couple which may keep it at rest with its base inclined at an angle of 30° to the horizon.
- If W be the weight and a the radius of the hemisphere, the moment of the couple will be equal to $\frac{3}{16}$ Wa.

CHAPTER IV.

EQUILIBRIUM OF SEVERAL BODIES.

If there be a system of bodies mutually acting on each other by contact, by connecting rods, or in any conceivable way, it will be necessary, in the determination of the circumstances of equilibrium, to represent the unknown actions and reactions by appropriate symbols. We shall then have to write down the equations of equilibrium for each body separately, including among the known forces to which it is subject, the unknown actions which it experiences from its connection with the other bodies of the system. From these different sets of equations, taken conjointly, we shall have to determine the circumstances of equilibrium.

SECT. 1. No Friction.

(1) AB (fig. 44) is a uniform beam, capable of motion about its middle point D; CE is a beam, moveable about a hinge C in the vertical line through D, and pressing against the beam AB from the extremity B of which a weight P is suspended; to determine the positions of the beams for equilibrium, having given that CD is equal to AD or BD.

Let AD = CD = BD = a, $\angle ACD = \theta$; GC = b, G being the centre of gravity of the beam CE; R = the action and reaction of the two beams at A; W = the weight of the beam CE. Then for the equilibrium of CE, taking moments about C, we have

 $R \cdot 2a \cos \theta = W \cdot b \sin \theta$;

and for the equilibrium of AB, taking moments about D,

 $R.a \cos \theta = P.a \sin 2\theta$, or $R = 2P \sin \theta$;

from these two equations, by the elimination of R, we get $Wb \sin \theta = 2Pa \sin 2\theta = 4Pa \sin \theta \cos \theta$,

and therefore
$$\theta = 0$$
, or $\cos \theta = \frac{bW}{4aP}$;

which determine the required positions of the beams.

(2) Two spheres O and O', (fig. 45), rest upon two smooth inclined planes AC and AC', and press against each other; to determine their position.

Let W, W', be the weights of the spheres O, O'; R their mutual action and reaction; a, a', the inclinations of the planes AC, AC', to the horizon; θ the inclination of the line OO', joining the centres of the spheres, to the horizon.

Then for the equilibrium of the sphere O, resolving forces parallel to AC,

$$R\cos(a+\theta)=W\sin a$$
;

and for the equilibrium of the sphere O', resolving forces parallel to AC',

$$R\cos(a'-\theta)=W'\sin a'$$
.

Eliminating R between these two equations,

$$W \sin \alpha \cos (\alpha' - \theta) = W' \sin \alpha' \cos (\alpha + \theta)$$

 $W \tan a (1 + \tan a' \tan \theta) = W' \tan a' (1 - \tan a \tan \theta),$

and therefore
$$\tan \theta = \frac{W' \tan \alpha' - W \tan \alpha}{(W' + W) \tan \alpha' \tan \alpha}$$
.

(3) Three spheres O, O', O', (fig. 46), are placed in contact within a hollow sphere; a vertical plane through the centre of the hollow sphere being supposed to contain the centres of the three solid spheres; to find their positions of equilibrium.

Let C be the centre of the hollow sphere; O, O', O', the centres of the solid spheres; join OC, O'C, O'C'; let W, W', W'', be the weights of the three spheres; CO = r, CO' = r', CO'' = r''; CO'' = a''; $\theta = \text{the inclination of } O'C'$ to the horizon.

Then since the actions of the hollow sphere on the solid ones all three pass through the point C, we have for the equilibrium of the solid spheres, taking moments about C, and observing

that if each of the spheres be in equilibrium singly they would likewise be at rest connectedly,

$$Wr \cos (\theta - a) + W'r' \cos \theta + W''r' \cos (\theta + a'') = 0,$$

$$Wr (\cos a + \sin a \tan \theta) + W'r' + W''r' (\cos a'' - \sin a'' \tan \theta) = 0;$$
and therefore
$$\tan \theta = \frac{W''r' \cos a'' + W'r' + Wr \cos a}{W''r' \sin a'' - Wr \sin a}.$$

(4) A sphere and cone of given weights are placed in contact on two inclined planes, the intersection of which is a horizontal line; to determine the circumstances of equilibrium.

Let W, W', be the weights of the sphere and the cone, which we may suppose to be applied at their centres of gravity G, G', (fig. 47). Let R be the action of the plane AB upon the sphere, and S the mutual action of the sphere and cone: if ϕ denote the semiangle of the cone, then evidently the line of action of S will make an angle ϕ with the plane AB'. The plane AB' will exert at right angles to itself an action upon every element of the base of the cone; the resultant of all these actions will be some force R' applied at some point E of the base of the cone in the line AB'. Let a, a', be the inclinations of the two planes to the horizon.

For the equilibrium of the sphere we have, resolving forces parallel to the plane AB,

$$W \sin a = S \cos (a + a' - \phi) \dots (1),$$

and resolving forces at right angles to the plane,

$$R = W \cos a + S \sin (a + a' - \phi) \dots (2);$$

the equation of moments is an identical equation, since all the forces which act upon the sphere pass through its centre.

Again, for the equilibrium of the cone, resolving the forces which act upon it parallel to the plane AB',

$$W' \sin a' = S \cos \phi \ldots (3);$$

resolving forces at right angles to the plane AB',

$$R' = W' \cos a' + S \sin \phi \dots (4),$$

and taking moments about G', the lines EH, mG', being represented by x, y,

$$R'x = S_V \cos \phi \dots (5).$$

From the equations (1) and (3),

$$\frac{W \sin a}{W' \sin a'} = \frac{\cos (a + a' - \phi)}{\cos \phi} \dots (6),$$

from which $\tan \phi$ may be readily determined: this relation is the only condition to which the cone and sphere are subject to secure equilibrium; as will be evident when it is observed that the three equations (2), (4), (5), introduce four unknown quantities R, R', x, y, each of the three equations at least one, which are not involved in (1) and (3). From this it is evident that there will be an infinite number of positions of equilibrium, or that if ϕ only have the value given by (6), the cone and sphere will rest in contact in whatever manner they may be placed on the two planes, and whatever be their magnitudes.

The values of ϕ being determined by (6), S will be determined by (1) or (3), and therefore R, R', from (2), (4), respectively. Then from the equation (5) we may determine x, provided that y be given; and y can be given only by our knowing the magnitudes of the cone and sphere, and the particular position of equilibrium in which we may choose to place them.

(5) Two uniform rods AC, A'C, of which the lower extremities are situated in the same horizontal plane, and prevented from sliding, lean against each other at the point C, and are in equilibrium; to determine the relation between their angles of inclination to the horizon, the small area of mutual contact at C being vertical.

Let W, W, be the weights of the rods AC, A'C, respectively, and ϕ , ϕ' , their angles of inclination to the horizon; then

$$W \tan \phi = W' \tan \phi'$$
.

Franchini, Memorie della Societa Italiana, tom. xvi. P. i. p. 237; 1813.

(6) An inextensible string binds tightly together two smooth cylinders of given radii; to find the ratio of the mutual pressure between the cylinders to the tension by which it is produced.

If R be the mutual pressure, T the tension of the string, r, r', the radii of the cylinders; then

$$\frac{R}{T}=\frac{4(rr')^{\frac{1}{2}}}{r+r'}.$$

(7) A sphere of given weight and radius is suspended by a string of given length from a fixed point, to which point also is attached another given weight by a string so long that the weight hangs below the sphere; to find the angle which the string, to which the sphere is attached, makes with the vertical.

If P denote the weight, Q the weight and sphere together, a the radius of the sphere, and b the distance of its centre from the point of suspension; then the required angle will be equal to

$$\sin^{-1}\left\{rac{Pa}{Qb}
ight\}^{rac{1}{4}}.$$

(8) A rod AB (fig. 48) is fixed at a given angle of inclination to the vertical; a rod CD is attached to AB by connections at the points B, C, a weight W being suspended from the extremity D; to determine the pressures exerted by AB upon CD, the weight of CD being neglected.

Let F, G, denote the resolved parts of the pressures at B, C, on CD, estimated along its length; and R, S, the pressures at right angles to the former; let CD = b, CB = c, then, a being the inclination of the rods to the vertical,

$$R = \frac{b}{c} W \sin a$$
, $S = \frac{b-c}{c} W \sin a$,
 $F + G = W \cos a$,

the single value of F or G being indeterminate.

SECT. 2. Friction.

(1) Two equal uniform beams AK, AK', (fig. 49), which are capable of revolving in a vertical plane about a point A to which their lower extremities are attached, have their upper extremities connected by a string KK'; a heavy sphere is

placed between the two beams; supposing the string to contract, to determine its tension when the sphere is just going to be forced upwards, the friction between the sphere and each of the beams being given.

It is plain that the two beams must make equal angles with the vertical line AL which passes through A, because the centre of gravity of the system consisting of the two beams and the sphere must lie in this line.

Let R, R', denote the actions of the beams upon the sphere at right angles to their lengths, and F, F', their actions along their lengths which are due to roughness. Let 2a be the angle at which the two beams are inclined to each other, T the tension of the string KK'; W the weight of the sphere, W' of each of the beams, and 2a the length of each.

Then for the equilibrium of the sphere we have, resolving forces parallel to LA,

$$(F + F') \cos a + W = (R + R') \sin a \dots (1);$$

resolving at right angles to LA,

$$(F' - F) \sin a = (R - R') \cos a \dots (2);$$

and taking moments about O, the centre of the sphere,

$$F. OE = F'. OE', \text{ or } F = F'......(3).$$

From (2) and (3) we have

$$R' = R \ldots (4)$$

Now supposing the sphere to be on the point of being disturbed by the contraction of the string, one or both of the points E, E', of the sphere must be on the point of sliding along the corresponding beams. Suppose that sliding is on the point of taking place at E.

Then μ being the coefficient of friction between the sphere and the beam AK, we have

$$F = \mu R$$
;

and therefore from (1), (3), (4),

$$2\mu R\cos a + W = 2R\sin a,$$

and therefore, putting $\mu = \tan \epsilon$,

$$R = \frac{W}{2(\sin a - \mu \cos a)} = \frac{W \cos \varepsilon}{2 \sin (a - \varepsilon)} \dots (5).$$

Also from (3) and (4),

$$\frac{F'}{R'}=\frac{F}{R}=\mu,$$

and therefore $F' = \mu R'$.

Hence we see that, if μ' be the coefficient of friction between the sphere and the beam AK', μ is not greater than μ' , since the greatest value of F' will be $\mu'R'$. If μ be less than μ' , the sphere would, with the slightest increase in the tension of KK', begin to roll along AK' without sliding; and if μ be equal to μ' , the sphere would begin to slide at both points simultaneously.

Again, for the equilibrium of AK we have, taking moments about A, it being remembered that the actions and reactions between the sphere and the beams are equal and opposite,

$$R \cdot AE + W' \cdot a \sin a = T \cdot 2a \cos a$$
;

and therefore, r being the radius of the sphere,

$$Rr \cot a + W'a \sin a = 2Ta \cos a$$
:

hence, putting for R its value given in (5),

$$\frac{Wr\cos\epsilon\cos\alpha}{2\sin\alpha\sin(\alpha-\epsilon)} + W'a\sin\alpha = 2Ta\cos\alpha,$$

and therefore

$$T = \frac{Wr\cos\varepsilon}{4a\sin\alpha\sin(\alpha-\varepsilon)} + \frac{1}{2}W'\tan\alpha.$$

(2) AB (fig. 50) is a uniform beam, capable of motion about its middle point D; CE is a beam, moveable about a hinge C in the vertical line through D, and pressing against the beam AB from the extremity B of which a weight P is suspended; CD, AD, BD, are equal lines; from observing the magnitude of the angle ACD when the end A of the beam AB is on the point of sliding in the direction CE, to find the coefficient of friction between the two beams.

Let G be the centre of gravity of the beam CE; μ the coefficient of friction; R the mutual action of the two beams at right angles to CE; $\angle ACD = \beta = \angle CAD$; AD = a = BD; CG = b, 4Q the weight of the beam CE.

Then for the equilibrium of CE we have, taking moments about C,

$$R \cdot 2a \cos \beta = 4Q \cdot b \sin \beta$$
,
or $aR \cos \beta = 2bQ \sin \beta \cdot \dots (1)$;

and for the equilibrium of AB, taking moments about D,

$$R.a \cos \beta = \mu R.a \sin \beta + P.a \sin 2\beta$$
,

$$R(\cos\beta - \mu \sin\beta) = P \sin 2\beta \dots (2).$$

From (1) and (2) there is

$$\frac{2bQ\sin\beta}{a\cos\beta}(\cos\beta-\mu\sin\beta)=2P\sin\beta\cos\beta;$$

and therefore $bQ(1 - \mu \tan \beta) = aP \cos \beta$, $\mu = \frac{bQ - aP \cos \beta}{bQ \tan \beta}.$

(3) A weight W(fig. 51) is suspended from the middle point of a rigid rod without weight, connecting the centres O, O', of two equal heavy wheels, which rest on a rough inclined plane; the wheel O is locked: to find the greatest inclination of the plane which is consistent with the equilibrium of the carriage.

Let P be the weight, and r the radius of each of the wheels; OO' = 2a, $\phi =$ the inclination of the plane to the horizon; R, R', the reactions of the plane on the wheels at right angles to itself; μR the friction on the wheel O, μ being the coefficient of friction; F the action of the plane on the wheel O' at right angles to R'; X, Y, the resolved parts, parallel and perpendicular to the plane, of the action of the wheel O on the rod OO'; and X', Y', the similarly resolved parts of the reaction.

For the equilibrium of the wheel O and the rod OO, regarded as one system, we have, resolving forces parallel to the inclined plane,

$$\mu R = X + (P + W) \sin \phi \dots (1);$$

resolving forces at right angles to the plane,

$$R + Y = (P + W) \cos \phi \dots (2);$$

and taking moments about O,

$$\mu Rr + 2a Y = Wa \cos \phi \dots (3).$$

Again, for the equilibrium of the wheel O', we have, taking moments about the point of contact of this wheel with the plane,

$$X'r = Pr \sin \phi$$
, or $X' = P \sin \phi \dots (4)$.

From the equations (1) and (4), observing that X' is by the nature of action and reaction equal to X, we get

$$\mu R = (2P + W) \sin \phi \ldots (5).$$

Again, from (2) and (3),

$$\mu r R + 2a (P + W) \cos \phi - 2aR = Wa \cos \phi,$$

$$(2a - \mu r) R = a (2P + W) \cos \phi \dots (6).$$

From (5) and (6) we obtain for the required inclination of the plane,

 $\tan \phi = \frac{\mu a}{2a - \mu r}.$

- Cor. Having ascertained ϕ , we know R from (5) and X' or X from (4), and therefore Y from (2); also F being the only force acting on the wheel O' which does not pass through its centre, it is evident that F must be equal to zero.
- (4) Two equal beams AC, BC, are connected by a smooth hinge at C, and are placed in a vertical plane with their lower extremities A and B resting on a rough horizontal plane; from observing the greatest value of the angle ACB for which equilibrium is possible, to determine the coefficient of friction at the ends A and B.
- If β be the greatest value of $\angle ACB$, and μ be the coefficient of friction at each of the ends; then

$$\mu = \frac{1}{2} \tan \frac{1}{2} \beta.$$

Sect. 3. Systems of Beams.

(1) At the middle points of the sides of any polygon ABCDE.... (fig. 52), and at right angles to them, are applied a series of forces P, Q, R,, respectively proportional to the sides; the sides of the polygon are perfectly rigid, and

capable of moving freely about the angular points A, B, C, D,...; to determine the form of the polygon that it may be in equilibrium, the lengths of the sides being given.

Let p, q, r, s, \ldots denote the mutual actions of the sides of the polygon at the angles A, B, C, D, \ldots , of which the directions will lie in certain straight lines $bB\beta$, $cC\gamma$, $dD\delta$, ...

For the equilibrium of the side BC we have, resolving forces at right angles to it,

$$Q = q \sin CB\beta + r \sin BCc \dots (1);$$

resolving forces parallel to BC,

$$q \cos CB\beta = r \cos BCc \dots (2);$$

and taking moments about the middle point of BC,

$$q \sin CB\beta = r \sin BCc \dots (3)$$
.

Dividing (3) by (2), we have

$$\tan CB\beta = \tan BCc,$$

and therefore
$$\angle CB\beta = \angle BCc \ldots (4)$$
;

hence also, from (2) or (3),
$$q = r \dots (5)$$
.

Again, from (1) and (3), we have

$$Q = 2r \sin \angle BCc$$
;

in precisely the same manner we may find that

$$R = 2r \sin \angle DC\gamma,$$

and therefore

$$\frac{Q}{R} = \frac{\sin \angle BCc}{\sin \angle DC\gamma};$$

but by hypothesis

$$\frac{Q}{R} = \frac{BC}{DC} = \frac{\sin \angle BDC}{\sin \angle CBD};$$

hence
$$\sin \angle BCc = \frac{\sin \angle BDC}{\sin \angle DC\gamma}$$
;

but from the geometry it is evident that

$$\angle B(Y + \angle DCY = \angle BDC + \angle CBD;$$

hence we readily see that

$$\angle BCc = \angle BDC \dots (6).$$

In just the same way we might prove that

$$\angle CB\beta = \angle BAC$$
,

and therefore by (4)

$$\angle BDC = \angle BAC \dots (7)$$

From this relation (7) it is plain that a circle passing through the three points A, B, C, must pass likewise through the point D; similarly we might shew that this circle, since it passes through B, C, D, must likewise pass through E, and so on indefinitely; hence we see that when the sides of the polygon are arranged consistently with equilibrium, all its angular points must be situated in the circumference of a single circle.

From (5) we gather that

$$p=q=r=s=\ldots$$

or that the mutual pressures at all the angular points are equal. It is evident also from the relation (6), that all the lines $aa, b\beta, c\gamma, d\delta, \ldots$ are tangents to the circle passing through A, B, C, D, \ldots

The value of the mutual pressure at each of the angular points is easily obtained: thus, as we have shewn,

$$Q = 2r \sin \angle BCc$$
;

but since $\angle BCc$ is equal to half the angle subtended by BC at the centre of the circle circumscribing the polygon, it is clear that

 $\sin \angle BCc = \frac{\frac{1}{2}BC}{\text{radius}};$

hence
$$r = \text{radius} \times \frac{Q}{BC}$$
,

and therefore $p = q = r = s \dots = k\rho$,

where ρ denotes the radius and k the ratio between any one of the forces and the corresponding side of the polygon.

The following is a different solution of the same problem:— Let the forces P, Q, R,... be represented in magnitude by the lines 2AB, 2BC, 2CD,..., to which they are proportional. Instead of the force 2AB acting at the middle point of the side AB, apply two forces, each equal to AB, one at the end A and the other at the end B of the side AB; each of these forces being at right angles to the side AB. Again, instead of the force 2BC acting at the middle point of BC, apply a force BC at C, and a force BC at the extremity B of the side AB, (which we are at liberty to do, because the point B of AB is rigidly attached to the point B of BC,) each of these forces being at right angles to BC. Now, according to this distribution of the forces, the only force which could twist BC about C, is the action of the rod AB upon the end B of BC; and therefore for the equilibrium of BC it is necessary that this action should take place exactly along BC. Hence conversely the action of CB upon BA will take place entirely in the direction CB. Let this action be denoted by R.

Thus, the line AB is acted upon at the point B by a force AB at right angles to AB, a force BC at right angles to BC, and a force R in the direction CB: but, by the principle of the parallelogram of forces, the forces AB and BC at B are equivalent to a single force AC acting at right angles to AC; hence for the equilibrium of AB we have, taking moments about A,

$$R \cdot AB \cdot \sin \angle ABC = AC \cdot AB \cdot \cos \angle BAC$$

or
$$R \sin \angle ABC = AC \cos \angle BAC$$
.

Similarly for the equilibrium of the side CD,

$$R \sin \angle BCD = BD \cos \angle BDC$$
;

and therefore

$$\frac{\sin \angle ABC}{\sin \angle BCD} = \frac{AC}{BD} \frac{\cos \angle BAC}{\cos \angle BDC}.$$

But by the geometry,

$$\frac{\sin \angle BAC}{\sin \angle BDC} = \frac{\frac{BC}{AC}\sin \angle ABC}{\frac{BC}{BD}\sin \angle BCD} = \frac{BD}{AC} \frac{\sin \angle ABC}{\sin \angle BCD}.$$

Hence from these two relations we have

$$\frac{\sin \angle BAC}{\sin \angle BDC} = \frac{\cos \angle BAC}{\cos \angle BDC},$$

 $\tan \angle BAC = \tan \angle BDC$, $\angle BAC = \angle BDC$;

which shews, as in the former solution, that the sides of the polygon must be so arranged that its angular points may all lie in the circumference of a single circle.

- (2) A quadrilateral ABCD, (fig. 53), consists of four rigid rods, which are capable of free motion about the angular points A, B, C, D; supposing the points A, C, and B, D, to be attached together by strings AC and BD in given states of tension, to determine the geometrical conditions necessary for the equilibrium of the quadrilateral.
- Let P, Q, represent the tensions of the strings AC, BD. Let K, L, M, N, denote the actions and reactions between the four pairs of points (A, B), (B, C), (C, D), (D, A).

The force P acting upon the point A in the direction AC, is equivalent to a force, in the direction AB,

$$= P \frac{\sin CAD}{\sin BAD} = P \frac{\sin ADB}{\sin BAD} \cdot \frac{DO}{AO} = P \frac{OD \cdot AB}{BD \cdot OA};$$

and to some force (F suppose) in AD.

Similarly, the force Q acting upon the point B in the direction BD, is equivalent to

a force, in
$$BA$$
, = $Q \frac{OC.AB}{AC.OB}$,

and some force (G suppose) in BC.

Hence clearly the point A is solicited by a force F-N in AD, and a force

$$P \stackrel{OD.AB}{RD.OA} - K \text{ in } AB \dots (1);$$

and therefore for its equilibrium we have

$$F - N = 0$$
, and $P \frac{OD \cdot AB}{RD \cdot OA} - K = 0$.

Similarly for the equilibrium of the point B there is

$$G-L=0$$
, and $Q\frac{OC \cdot AB}{AC \cdot OB}-K=0 \cdot \dots (2)$.

From (1) and (2) we have

$$P\frac{OD \cdot AB}{RD \cdot OA} = Q\frac{OC \cdot AB}{AC \cdot OB}$$

and therefore

$$\frac{P \cdot OD}{BD \cdot OA} = \frac{Q \cdot OC}{AC \cdot OB};$$

which is the condition for the equilibrium of the quadrilateral.

Euler; Act. Acad. Petrop. 1779, P. 11. p. 106.

The following is a different solution of the same problem:

For the equilibrium of the rod AB there is, taking moments about B,

$$N. BD \cdot \sin \angle BDA = P \cdot BO \cdot \sin \angle BOC;$$

and for the equilibrium of the rod CD, taking moments about C,

$$N. CA. \sin \angle CAD = Q. CO. \sin \angle BOC;$$

hence obviously

$$\frac{BD \sin \angle ODA}{CA \sin \angle OAD}$$
, or $\frac{BD \cdot AO}{AC \cdot DO} = \frac{P \cdot BO}{Q \cdot CO}$.

(3) Four rigid rods AB, BC, CD, DA, (fig. 54), are so joined together that they are capable of revolving freely about the angular points of the quadrilateral which they form; these rods are attached together, two and two, viz. those which are contiguous, by strings aa, $b\beta$, $c\gamma$, $d\delta$, in given states of tension; to determine the form of the quadrilateral which shall correspond to the equilibrium of the rods.

Let A, B, C, D, denote the tensions of the strings aa, $b\beta$, $c\gamma$, $d\delta$. Then the force A in aa upon the point a is equivalent to a force, in BA,

$$=A\frac{\sin aaD}{\sin AaD}=A\frac{\sin aDa \cdot \frac{Da}{aa}}{\sin ADa \cdot \frac{DA}{Aa}}=A\frac{Aa \cdot Da}{aa \cdot DA};$$

and to a force, in aD,

$$= A \frac{\sin Aaa}{\sin AaD} = A \frac{\sin aAa \cdot \frac{Aa}{aa}}{\sin aAD \cdot \frac{AD}{aD}} = A \frac{Aa \cdot Da}{aa \cdot DA}$$

= A' suppose.

But the force A' in aD is equivalent to

a force in
$$AD$$
, = $A' \frac{\sin aDB}{\sin ADB} = A' \frac{\sin aBD \cdot \frac{Ba}{Da}}{\sin ABD \cdot \frac{BA}{DA}}$
= $A' \frac{Ba \cdot DA}{Da \cdot BA} = A \frac{Aa \cdot Ba}{aa \cdot BA}$;

and to a force in BD, = $A' \frac{\sin ADa}{\sin ADB}$

$$=A'\frac{\sin DAa \cdot \frac{Aa}{Da}}{\sin DAB \cdot \frac{AB}{DB}} = A'\frac{Aa \cdot BD}{AB \cdot Da} = A\frac{Aa \cdot Aa \cdot BD}{aa \cdot DA \cdot BA}.$$

Thus we see that the force A, acting upon the point a in the direction aa, is equivalent to the three forces

$$A \frac{aD \cdot Aa}{AD \cdot aa}$$
 in BA upon A , $A \frac{Aa \cdot aB}{AB \cdot aa}$ in AD upon A , and $A \frac{aA \cdot Aa \cdot BD}{AB \cdot AD \cdot aa}$ in BD upon B .

Similarly, the force A acting upon the point a in the direction aa, is equivalent to

$$A \frac{aB \cdot Aa}{AB \cdot aa}$$
 in DA upon A , $A \frac{Aa \cdot aD}{AD \cdot aa}$ in AB upon A , and $A \frac{aA \cdot Aa \cdot DB}{AD \cdot AB \cdot aa}$ in DB upon D .

Now these three forces are equal and opposite to the three former, and therefore the string aa with a tension A produces the same effect, and may therefore be replaced by a string BD with a tension

$$A \frac{aA \cdot Aa \cdot BD}{AB \cdot AD \cdot aa}$$
.

In the same way we may shew, that the tension of $c\gamma$ is equivalent to a string BD, of which the tension is equal to

$$C \frac{cC \cdot C\gamma \cdot BD}{CB \cdot CD \cdot c\gamma}$$
.

Hence the tensions of aa, $c\gamma$, together, are equivalent to a string BD with a tension

$$A \frac{Aa \cdot Aa \cdot BD}{BA \cdot DA \cdot aa} + C \frac{Cc \cdot C\gamma \cdot BD}{BC \cdot DC \cdot c\gamma}$$

Similarly it may be shewn, that the tensions $b\beta$, $d\delta$, are equivalent to a string AC with a tension

$$B \frac{Bb \cdot B\beta \cdot CA}{AB \cdot CB \cdot b\beta} + D \frac{D\delta \cdot Dd \cdot AC}{AD \cdot CD \cdot d\delta}$$

Hence, by the result of the preceding problem, the condition of equilibrium is expressed by the relation

$$\frac{OB \cdot OD}{BD^2} \left(\frac{B \cdot Bb \cdot B\beta}{AB \cdot CB \cdot b\beta} + \frac{D \cdot D\delta \cdot Dd}{AD \cdot CD \cdot d\delta} \right)$$

$$=\frac{OA \cdot OC}{AC^{\alpha}} \left(\frac{A \cdot Aa \cdot Aa}{BA \cdot DA \cdot aa} + \frac{C \cdot Cc \cdot C\gamma}{BC \cdot DC \cdot c\gamma} \right).$$

Euler; Act. Acad. Petrop. 1779, P. 2, p. 106.

(4) Three uniform beams AB, BC, CD, of the same thickness, and of lengths l, 2l, l, respectively, are connected by hinges at B and C, and rest on a perfectly smooth sphere, the radius of which is equal to 2l, so that the middle point of BC and the extremities of AB, CD, are in contact with the sphere; to compare the pressure at the middle point of BC, and the pressures at A and D, with the weight of the three beams.

Let W be the weight of the three beams taken together; R the pressure at each of the points A and D; and R' the pressure at the middle point of BC. Then

$$\frac{R}{W} = \frac{3}{40}, \qquad \frac{R'}{W} = \frac{91}{100}.$$

(5) Four equal uniform beams AB, BC, CD, DE, (fig. 55), connected together by joints at their extremities, rest in equilibrium in a vertical plane; the distances AE and CF, of which the latter is perpendicular to AE and vertical, are given; to determine the conditions of equilibrium.

If a, β , be the inclinations of AB and ED, BC and DC, to the horizon; we must have

$$\tan a = 3 \tan \beta$$
.

Draw BK at right angles to AE; let CF = a, AF = b, FK = x, BK = y; then from the equation in a and β , and the geometry of the figure, we may get

$$x = \frac{a^2 + 2b^2 - (a^4 + a^2b^2 + b^4)^{\frac{1}{2}}}{2b}, \quad y = \frac{2a^2 + b^2 - (a^4 + a^2b^2 + b^4)^{\frac{1}{2}}}{2a}.$$

These values of x and y are obtained by Couplet in his Recherches sur la Construction des Combles de Charpente, in the Mémoires de l'Académie des Sciences de Paris, 1731, p. 69.

(6) A and C (fig. 56) in the same vertical line are fixed points, about which beams AB, CD, are freely moveable by hinge joints; AB is supported in a horizontal position by CD, with which it is connected by a hinge joint at D, and has a weight suspended at B: to find the pressure at C, the weights of the beams being neglected.

Let H and V be the horizontal and vertical pressures at C, and P the weight suspended from B. Then

$$H = P \cdot \frac{AB}{AC}, \quad V = P \cdot \frac{AB}{AD},$$

and therefore the whole pressure at C is equal to

$$P \cdot AB \cdot \left(\frac{1}{AC^2} + \frac{1}{AD^2}\right)^{\frac{1}{2}}$$

(7) Two equal uniform beams AB, AC, moveable about a hinge at A, are placed upon the convex circumference of a circle in a vertical plane; to find their inclination to each other when they are in their position of equilibrium.

Let 2a = the length of each beam, 2θ = their inclination to each other, and r = the radius of the circle. Then θ will be determined by the equation

$$r \cos \theta = a \sin^3 \theta$$
.

CHAPTER V.

EQUILIBRIUM OF FLEXIBLE STRINGS.

THE form of equilibrium assumed by a uniform flexible string sustained at its two extremities and acted on by gravity, attracted the attention of Galileo¹, who, from a want of sufficient examination, concluded it to be a parabola; this mistake may have arisen from the fact, that in the immediate neighbourhood of its lowest point it approximates very nearly to the parabolic The inaccuracy of Galileo's conclusion was experimentally ascertained by Joachim Jungius². This subject having been at last successfully investigated by James Bernoulli3, he proposed the problem of the chaînette, the name which he gave to the required curve, as a trial of skill to the mathematicians of the day. The four mathematicians who succeeded in arriving at correct solutions of the problem were, James Bernoulli, by whom it had been proposed, his brother John, Leibnitz, and Huyghens: their four solutions appeared without analysis in the Acta Eruditorum for the year 1691, Jun. pp. 273-282 demonstration of the results of these four illustrious mathematicians was first published by David Gregory, in the Philosophical Transactions for the year 1697.

The form of equilibrium of the chaînette or catenary, of which the thickness is supposed to be uniform, having been thoroughly discussed, James Bernoulli next directed his attention to more complicated problems of the same character; he

^{&#}x27; Mechanica; Dialogo 2, p. 131.

² Geometria Empyrica.

² Acta Eruditorum, Lips. 1690, Mai. p. 217; Opera, tom. 1. p. 424.

⁴ Acta Eruditorum, Lips. 1691, Jun. p. 289; Opera, tom. I. p. 449.

investigated the form of equilibrium when the thickness varies from point to point according to any assigned law, and, conversely, determined the law of its variation that the string may hang in assigned curves: he likewise considered the problem of the catenary when the string is extensible, the extension of each element being assumed according to the law established experimentally by Hooke¹ to vary as the tension. The analysis of these problems, of which the solutions only were published by James Bernoulli, was supplied by John Bernoulli². The consideration of the general conditions of the equilibrium of flexible strings was first attempted by Hermann³, whose investigations, however, were not free from error; a more accurate analysis was furnished by John Bernoulli⁴, who has particularly examined various cases of the equilibrium of strings acted on by central forces.

Among the numerous mathematicians who afterwards discussed the theory of the equilibrium of flexible strings, may be mentioned Euler⁵, Clairaut⁶, Krafft⁷, Legendre⁸, Fuss⁹, Venturoli¹⁰, and Poisson¹¹.

SECT. 1. Free Inextensible String; general Conditions of Equilibrium.

To investigate the conditions for the equilibrium of an inextensible string, of which the density and thickness vary from point to point according to any assigned law; the accelerating forces which act upon the string being any whatever.

- 1 De Potentia Restitutiva, or Spring.
- ² Lectiones Mathematicæ in usum Hospitalii, Opera, tom. IV. p. 387.
- Phoronomia, lib. 1. cap. 3, and Append. §. v.
- 4 Opera, tom. IV. p. 234.
- * Comment. Petrop. tom. III.; Nov. Comment. Petrop. tom. xv. and tom. xx.
- ⁶ Miscellanea Berolinensia, tom. vii. p. 270, 1743.
- ⁹ Nov. Comment. Petrop. tom. v. p. 143; 1754 and 1755.
- ^a Mém. Acad. Par. 1786, p. 20.
- Nova Acta Petrop. tom. xII. p. 145, 1794.
- 10 Elements of Mechanics, by Cresswell, Part 1. p. 62.
- 11 Traité de Mécanique, tom. 1. p. 564, seconde édition.

Let APB (fig. 57) be any portion of the string in a position of rest; Pp being a small element of its length; x, y, z, and $x + \delta x$, $y + \delta y$, $z + \delta z$, the co-ordinates of P and p respectively; s the length of the string reckoned from some assigned point up to P, and $s + \delta s$ the length up to p; t the tension of the string at P.

The resolved parts, parallel to the axes of x, y, z, of the force exerted upon the point P of the element Pp by the portion AP of the string, will evidently be

$$-t\frac{dx}{ds}$$
, $-t\frac{dy}{ds}$, $-t\frac{dz}{ds}$;

and therefore, since each of these three forces must be some function of s, it is plain by Taylor's theorem that the resolved parts of the force exerted on the element Pp by the portion qB of the string, will be

$$t \frac{dx}{ds} + \frac{d}{ds} \left(t \frac{dx}{ds} \right) \delta s,$$

$$t \frac{dy}{ds} + \frac{d}{ds} \left(t \frac{dy}{ds} \right) \delta s,$$

$$t \frac{dz}{ds} + \frac{d}{ds} \left(t \frac{dz}{ds} \right) \delta s.$$

Again, let X, Y, Z, be the sums of the resolved parts of the accelerating forces which act upon the element Pp; ρ the density of the string at P, and k the area of a section at right angles to its length at that point. Then clearly the mass of the portion Pp of the string will be $k\rho\delta s$, which therefore for a constant value of δs will vary as $k\rho$; hence evidently the product $k\rho$, which we will call m, may be taken to measure the massiveness of the string at the point P; it will be convenient to call it the unit of mass at the point P. The impressed moving force then of the element Pp, will have for its resolved parts parallel to the co-ordinate axes,

$$mX\delta s$$
, $mY\delta s$, $mZ\delta s$.

Hence clearly for the equilibrium of Pp we must have, equating to zero the sum of the resolved forces which act upon

it parallel to each of the three axes, and dividing the three resulting equations by &s,

$$\frac{d}{ds}\left(t\frac{dx}{ds}\right) + mX = 0,$$

$$\frac{d}{ds}\left(t\frac{dy}{ds}\right) + mY = 0,$$

$$\frac{d}{ds}\left(t\frac{dz}{ds}\right) + mZ = 0,$$

$$\ldots \ldots (a);$$

which three equations constitute the conditions of equilibrium of the entire string.

By the elimination of t we readily obtain the three following equations,

$$dx \int m Y ds = dy \int m X ds,$$

$$dy \int m Z ds = dz \int m Y ds,$$

$$dz \int m X ds = dx \int m Z ds;$$

any two of which will be differential equations to the required curve of equilibrium.

Cor. 1. From the equations (a) we have also

$$t\frac{dx}{ds} = -\int mXds$$
, $t\frac{dy}{ds} = -\int mYds$, $t\frac{dz}{ds} = -\int mZds$;

squaring and adding these equations, and observing that

$$\frac{dx^2}{ds^2} + \frac{dy^2}{ds^2} + \frac{dz^2}{ds^2} = 1 \dots (b),$$

we obtain for the value of the tension at any point,

$$t^2 = (\int m X ds)^2 + (\int m Y ds)^2 + (\int m Z ds)^2.$$

We may obtain also another expression for the tension: differentiating (b) with respect to s, we get

$$\frac{dx}{ds}\frac{d^2x}{ds^2}+\frac{dy}{ds}\frac{d^2y}{ds^2}+\frac{dz}{ds}\frac{d^2z}{ds^2}=0 \ldots (c);$$

hence, multiplying the three equations (a) by dx, dy, dz, in order, and adding the resulting equations, we have, by the aid of (b) and (c),

$$t = C - \int m (Xdx + Ydy + Zdz),$$

where C is an arbitrary constant.

Cor. 2. If the whole string lie entirely within one plane, let the plane of xy be so chosen as to coincide with this plane; then the three differential equations to the string will be reduced to the single one

$$dx \int m Y ds = dy \int m X ds \dots (d);$$

and the two formulæ for the tension will become

$$t^{2} = (\int mX ds)^{2} + (\int mY ds)^{2},$$

$$t = C - \int m(X dx + Y dy).$$

These two formulæ for the tension, and also the differential equation (d) to the string, coincide with those given by Fuss, *Mémoires de St. Pétersbourg*, 1794, p. 150, 151.

SECT. 2. Parallel Forces.

(1) A flexible string fixed at any two points A and B, (fig. 58), is acted on by gravity; supposing the unit of mass to vary according to any assigned law as we pass from one point to another, to find the equation to the catenary of rest; and conversely, the curve being known, to determine the law of the unit of mass.

Let the axis of y extend vertically upwards, and let the axis of x be horizontal, the plane xOy coinciding with the plane which contains the catenary. Then since

$$X = 0$$
, $Y = -q$,

we have, by the first two of the equations (a) of section (1),

$$\frac{d}{ds}\left(t\,\frac{dx}{ds}\right)=0\,\ldots\ldots\,(a),$$

$$\frac{d}{ds}\left(t\frac{dy}{ds}\right) = mg \dots (b).$$

Integrating the equation (a), we get

$$t\frac{dx}{ds}=C,$$

where C is a constant quantity: let τ denote the tension at the lowest point of the curve, then evidently $\tau = C$, and therefore

$$t\frac{dx}{ds}=\tau\ldots\ldots(c).$$

From (b) and (c), we have

$$\tau \frac{d}{ds} \frac{dy}{dx} = mg,$$

and therefore

$$\tau \frac{dy}{dx} = \int mg \ ds \; ;$$

but evidently at the lowest point of the catenary $\frac{dy}{dx} = 0$, and therefore, supposing a to be the value of s at the lowest point,

$$0 = \int_a^a mg \, ds ;$$

$$\tau \, \frac{dy}{dx} = g \int_a^s m \, ds \dots \dots \dots \dots \dots (d).$$

hence

If m be given in terms of the variables x, y, s, the form of the catenary may be determined from (d).

Again, differentiating (d), we obtain

$$m = \frac{r}{g} \frac{d^3y}{ds} \dots (e),$$

a formula by which m may be computed for every point of the string when the form of the catenary is given. Also from (c) we get

$$t = r \frac{ds}{dx} \cdot \dots \cdot (f),$$

which gives the tension at any point of the catenary when its form is known.

John Bernoulli; Lectiones Mathematicæ, Lect. 38, 39, 40; Opera, tom. 111.

(2) A flexible string AOB, (fig. 60), fixed at two points A and B, is acted on by gravity; the unit of mass at any point P varies inversely as the square root of the length OP measured from the lowest point O; to find the equation to the catenary.

Let the origin of co-ordinates be taken at O, x being horizontal, and y vertical, and the plane of xy coinciding with the plane of the catenary; also let O be the origin of s.

Then if μ be the unit of mass at a length c from the lowest point,

 $m=\mu\,\frac{c^{\frac{1}{4}}}{s^{\frac{1}{4}}},$

and therefore by (1, d), a being in the present case zero, we have

$$\tau \frac{dy}{dx} = g\mu c^{\frac{1}{2}} \int_0^s \frac{ds}{s^{\frac{1}{2}}} = 2g\mu c^{\frac{1}{2}} s^{\frac{1}{2}};$$

hence, putting for the sake of brevity

$$\frac{2g\mu c^{\frac{1}{4}}}{\tau} = \frac{1}{\beta^{\frac{1}{4}}},$$
we get
$$\frac{dy}{dx} = \left(\frac{s}{\beta}\right)^{\frac{1}{4}}, \quad \frac{dy^2}{dx^2} = \frac{s}{\beta},$$

$$\beta \frac{d}{dx} \frac{dy^2}{dx^2} = \frac{ds}{dx} = \left(1 + \frac{dy^2}{dx^2}\right)^{\frac{1}{4}},$$

$$\frac{\beta \frac{d}{dx} \frac{dy^2}{dx^2}}{\left(1 + \frac{dy^2}{dx^2}\right)^{\frac{1}{4}}} = 1;$$

integrating with respect to x we obtain

$$2\beta\left(1+\frac{dy^2}{dx^2}\right)^{\frac{1}{4}}=x+C;$$

but x = 0, $\frac{dy}{dx} = 0$, simultaneously; hence $C = 2\beta$, and therefore

$$2\beta\left(1+\frac{dy^2}{dx^3}\right)^{\frac{1}{2}}=x+2\beta.\ldots\ldots(a);$$

squaring and transposing,

$$4\beta^{2} \frac{dy^{2}}{dx^{2}} = (x + 2\beta)^{2} - 4\beta^{2},$$
$$2\beta dy = \{(x + 2\beta)^{2} - 4\beta^{2}\}^{\frac{1}{2}} dx;$$

integrating we have

$$C + 2\beta y = \frac{1}{2}(x + 2\beta)(x^2 + 4\beta x)^{\frac{1}{2}} - 2\beta^2 \log \{x + 2\beta + (x^2 + 4\beta x)^{\frac{1}{2}}\};$$

but $x = 0$, $y = 0$, simultaneously; hence

$$C = -2\beta^2 \log (2\beta);$$

hence, eliminating C_{i}

$$2\beta y = \frac{1}{2}(x+2\beta)(x^2+4\beta x)^{\frac{1}{4}} - 2\beta^2 \log \frac{x+2\beta+(x^2+4\beta x)^{\frac{1}{4}}}{2\beta},$$

which is the required equation to the catenary.

Cor. From (a) we get

$$\frac{ds}{dx}=\frac{x+2\beta}{2\beta},$$

and therefore by (1, f)

$$t = r \frac{ds}{dx} = \frac{r}{2\beta}(x + 2\beta),$$

which gives the tension at any point of the curve.

John Bernoulli; Lect. Math., Opera, tom. 111. p. 497.

(3) To find the law of variation of the unit of mass of a catenary acted on by gravity that it may hang in the form of a semicircle with its diameter horizontal.

The notation remaining the same as in (2), the equation to the catenary will be

$$x^2 = 2ay - y^2,$$

where a denotes the radius of the semicircle: hence

$$a^{2} - x^{2} = (a - y)^{2}, \quad y = a - (a^{2} - x^{2})^{\frac{1}{2}};$$

$$\frac{dy}{dx} = \frac{x}{(a^{2} - x^{2})^{\frac{1}{2}}}, \quad \frac{d^{2}y}{dx^{2}} = \frac{a^{2}}{(a^{2} - x^{2})^{\frac{1}{2}}};$$

also

$$\frac{ds^2}{dx^2} = 1 + \frac{dy^2}{dx^2} = \frac{a^2}{a^2 - x^2}, \quad \frac{ds}{dx} = \frac{a}{(a^2 - x^2)^{\frac{1}{2}}};$$

and therefore by (1, e)

$$m = \frac{r}{g} \frac{\frac{d^2y}{dx^2}}{\frac{ds}{dx}} = \frac{r}{g} \frac{a}{a^2 - x^2} = \frac{ra}{g(a - y)^2};$$

or the unit of mass at any point varies inversely as the square of its depth below the horizontal diameter of the semicircle.

Cor. By (1, f) we have for the tension at any point

$$t=\tau\frac{ds}{dx}=\frac{\tau a}{\left(a^2-x^2\right)^{\frac{1}{2}}}=\frac{\tau a}{a-y}.$$

John Bernoulli; Opera, tom. 111. p. 502.

(4) To find the length of a uniform chain ALB, (fig. 61), suspended from two points A and B in the same horizontal line, when the stress on each point of support is equal to the whole weight of the chain; to find also the depth of the lowest point L of the chain below the line AB, and the direction of its tangent at A or B.

Let yCLO be vertical, OL being equal to a length of the chain of which the weight is equal to the tension of the lowest point L, Ox horizontal; PM at right angles to Ox. OM = x, PM = y, OL = c, ALB = l, AC = BC = a.

Then the equation to the curve will be

$$y = \frac{1}{2}c \left(\epsilon^{c} + \epsilon^{-\frac{a}{c}} \right) \dots (1),$$

$$l = c \left(\epsilon^{c} - \epsilon^{-\frac{a}{c}} \right) \dots (2).$$

and also

Let m denote the unit of mass of the chain, which will be the same at all its points; then the tension at P will be equal to

$$mgy = \frac{1}{2}mcg\left(\varepsilon^{\frac{z}{c}} + \varepsilon^{-\frac{z}{c}}\right),$$

and therefore at B to $\frac{1}{2}mcg(\epsilon^{\frac{a}{c}} + \epsilon^{-\frac{a}{c}});$

but by the hypothesis the tension at B is equal to mgl, and therefore by (2) to

hence
$$mcg\left(\varepsilon^{\frac{a}{c}} - \varepsilon^{-\frac{a}{c}}\right);$$

$$\frac{1}{2}mcg\left(\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}}\right) = mcg\left(\varepsilon^{\frac{a}{c}} - \varepsilon^{-\frac{a}{c}}\right);$$

$$\frac{1}{2}\varepsilon^{\frac{a}{c}} = \frac{3}{2}\varepsilon^{-\frac{a}{c}},$$

$$\varepsilon^{\frac{2a}{c}} = 3, \quad \frac{2a}{c} = \log_{\varepsilon} 3, \quad \frac{a}{c} = \frac{1}{2}\log_{\varepsilon} 3 \dots (3).$$

Hence from (2) we have

$$l = \frac{2a}{\log_e 3} \left(3^{\frac{1}{4}} - \frac{1}{3^{\frac{1}{4}}} \right) = \frac{4a}{3^{\frac{1}{4}} \log_e 3},$$

which gives the length of the chain.

Again, putting x = a, we have from (1),

$$OC = \frac{1}{2} c \left(\epsilon^{\frac{a}{c}} + \epsilon^{-\frac{a}{c}} \right),$$

$$CL = \left(\frac{1}{2} \epsilon^{\frac{a}{c}} + \frac{1}{2} \epsilon^{-\frac{a}{c}} - 1 \right) c$$

$$= \left(\frac{3^{\frac{1}{4}} + 3^{-\frac{1}{2}}}{2} - 1 \right) \frac{2a}{\log_{\epsilon} 3}, \text{ from (3),}$$

$$= \left(\frac{2}{a^{\frac{1}{4}}} - 1 \right) \frac{2a}{\log_{\epsilon} 3} = \frac{2a}{\log_{\epsilon} 3} \frac{2 - 3^{\frac{1}{4}}}{a^{\frac{1}{4}}},$$

and therefore

which gives the depth of the lowest point of the chain below the line AB.

Again, from (1) we have

$$\frac{dy}{dx} = \frac{1}{2} \left(\varepsilon^{\frac{x}{c}} - \varepsilon^{-\frac{x}{c}} \right),$$

and therefore, ϕ denoting the inclination of the chain at B to the horizon,

$$\tan \phi = \frac{1}{2} \left(\epsilon^{\frac{a}{o}} - \epsilon^{-\frac{a}{o}} \right) = \frac{1}{2} \left(3^{\frac{1}{4}} - \frac{1}{3^{\frac{1}{4}}} \right) = \frac{1}{3^{\frac{1}{4}}};$$
hence $\phi = \frac{\pi}{6}.$

(5) A uniform string A'ALBB' (fig. 62) is placed over two supports A and B in the same horizontal line, so as to remain in equilibrium; having given the length of the string, and the distance of the points of support, to find the pressure which they have to bear.

Let L be the lowest point of the curve ALB, OLy a vertical line through L, where OL is equal to a length of the chain, the weight of which is equal to the tension at L; Ox horizontal.

Then Ox, Oy, being taken as the axes of co-ordinates, we shall have for the equation to the curve ALB, putting OL = c,

$$y = \frac{1}{2} c \left(\frac{x}{\epsilon^c} + \epsilon^{-\frac{x}{c}} \right) \dots (1);$$

and if m be the unit of mass at each point of the string, the tension at P will be equal to

$$mgy \text{ or } \frac{1}{2}mcg\left(\varepsilon^{\frac{x}{c}} + \varepsilon^{-\frac{x}{c}}\right);$$

hence, if AC = BC = a, the tension at B will be equal to

$$\frac{1}{2}mcg\left(\varepsilon^{\frac{a}{c}}+\varepsilon^{-\frac{a}{c}}\right).....(2).$$

But the tension at B is evidently equal to the weight of BB', and therefore, if BB' = s, to the expression mgs; hence

$$mgs = \frac{1}{2} mcg \left(\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}} \right),$$
or
$$s = \frac{1}{2} c \left(\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}} \right) \dots (3).$$

Suppose that the length of the whole string A'ALBB' is 2l; then the length of the portion LBB' will be l, and l-s will be the length of BL. Hence, by the nature of the catenary,

$$l-s=\frac{1}{2}c\left(\varepsilon^{\frac{a}{c}}-\varepsilon^{-\frac{a}{c}}\right).....(4).$$

Adding together the equations (3) and (4), we obtain

$$l=c\,\epsilon^{\frac{a}{c}},$$

whence c is made to depend upon the known quantities a and l: hence the expression (2) for the tension at B is known.

Differentiating (1), we get

$$\frac{dy}{dx} = \frac{1}{2} \left(\epsilon^{\frac{x}{c}} - \epsilon^{-\frac{x}{c}} \right):$$

but, if LP = s',

$$s' = \frac{1}{2}c\left(\overline{\epsilon^{o}} - \overline{\epsilon^{-\frac{x}{o}}}\right), \quad \frac{ds'}{dx} = \frac{1}{2}\left(\overline{\epsilon^{c}} + \overline{\epsilon^{-\frac{x}{c}}}\right);$$

hence evidently
$$\frac{dy}{ds'} = \frac{\sum_{\epsilon^{c} - \epsilon}^{x} - \sum_{\epsilon^{c} + \epsilon}^{x}}{\sum_{\epsilon^{c} + \epsilon}^{x} - \sum_{\epsilon}^{x}},$$

and therefore if ϕ denote the angle between the line BB' and the curve BL at B,

$$\cos \phi = \frac{\xi^{\frac{a}{c}} - \xi^{\frac{a}{c}}}{\xi^{\frac{a}{c}} + \xi^{\frac{a}{c}}} \cdot \dots (5).$$

Let P denote the pressure on the point B, and r the tension of the string at B; then

$$P^2 = 2r^2 - 2r^2 \cos \phi$$

= $2r^2 (1 - \cos \phi);$

and therefore from (2) and (5),

$$P^{2} = \frac{1}{2}m^{2}c^{2}g^{2}\left(\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}}\right)^{2} \left\{1 - \frac{\varepsilon^{\frac{a}{c}} - \varepsilon^{-\frac{a}{c}}}{\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}}}\right\}$$

$$= m^{2}c^{2}g^{2}\left(\varepsilon^{\frac{a}{c}} + \varepsilon^{-\frac{a}{c}}\right) \varepsilon^{-\frac{a}{c}}$$

$$= m^{2}c^{2}g^{2}\left(1 + \varepsilon^{-\frac{2a}{c}}\right)$$

$$P = mcg\left(1 + \varepsilon^{-\frac{2a}{c}}\right)^{\frac{1}{2}}$$

which gives the required value of the pressure, c having been previously determined.

(6) A uniform chain ABC (fig. 63) is suspended from a point A above an inclined plane RS; having given the angle which the chain at the point of suspension and which the plane makes with the horizon, and also the length of the whole chain, to find the length of the portion BC which is in contact with the plane.

Let ABLA' denote the catenary, of which AB is an arc, L being the lowest point. Let P be any point in the curve AL; ϕ the inclination of the curve at P to the horizon, t the tension at B; a, β , the values of ϕ at A, B respectively; c the length of chain of which the weight is equ l to the tension at L; m the unit of mass of the chain; LP = s, ABC = l, BC = l'.

Then, by the nature of the catenary,

$$t \cos \beta = mcg \dots (1),$$

 $s = c \tan \phi \dots (2).$

Now it is evident that the tension at B is equal to $mgl \sin \beta$; hence from (1),

$$mgl' \sin \beta \cos \beta = mcq$$
, $c = l' \sin \beta \cos \beta \dots (3)$.

Again, from (2) we have

$$LBA = c \tan a$$
, $LB = c \tan \beta$,

and therefore

$$l-l'=c\;(\tan\;a-\tan\;\beta);$$

hence, from (3),

$$l - l' = l' \sin \beta \cos \beta (\tan \alpha - \tan \beta),$$

$$l \cos \alpha = l' (\cos \alpha + \sin \alpha \sin \beta \cos \beta - \sin^2 \beta \cos \alpha)$$

$$= l' \cos \beta \cdot \cos (\alpha - \beta)$$

$$l' = \frac{l \cos a}{\cos \beta \cos (a - \beta)}.$$

(7) AOB (fig. 60) is a flexible string acted on by gravity, and is in a position of rest; the unit of mass at any point varies as the cosine of the angle at which an element of the curve at the point is inclined to the horizon; to find the equation to the catenary.

Assuming $m = \beta \frac{dx}{ds}$, where β is some constant quantity, the equation to the catenary will be

$$x^2 = \frac{2\tau}{\beta a} y;$$

which shows that the catenary is the common parabola.

James Bernoulli; Act. Erudit. Lips. Jun.; Opera, tom. 1. p. 449. John Bernoulli; Opera, tom. 111. p. 501.

(8) To find the equation to the catenary when the unit of mass varies as $x \cos \phi$, where ϕ is the angle of inclination of the element of the curve at any point to the horizon.

Assuming $m = \beta x \frac{dx}{ds}$, the required equation will be

$$6\tau y = g\beta x^3,$$

which belongs to a cubical parabola.

James Bernoulli; Ib. John Bernoulli; Ib.

(9) To find the equation to the catenary when the unit of mass varies as $x^{\frac{1}{4}} \cos \phi$.

Assuming
$$m = \beta x^{\frac{1}{4}} \frac{dx}{ds}$$
, the equation will be

$$16g^2\beta^2x^5 = 225 \tau^2y^2.$$

James Bernoulli; Ib. John Bernoulli; Ib.

(10) To find the equation to the catenary when the unit of mass varies as $y^n \sin \phi$, where n is any positive quantity.

If the origin of co-ordinates be so chosen that the axis of x passes through the lowest point of the catenary, and that $y = \infty$ when x = 0, the required equation will be

$$xy^n=-\frac{(n+1)\tau}{ng\beta}.$$

James Bernoulli; Ib. John Bernoulli; Ib.

(11) To find the law of the variation of the unit of mass when the catenary is the common parabola.

The construction and notation being the same as in (2),

$$m=\frac{2\tau}{g\left(a^2+4x^2\right)^{\frac{1}{4}}},$$

a being the latus rectum of the parabola.

John Bernoulli; Opera, tom. III. p. 504.

(12) A chain suspended at its extremities from two tacks in the same horizontal line, forms itself into a cycloid; to find the unit of mass at any point of the string and the weight of the arc between this and the lowest point.

Let w denote the weight of the arc; then taking the ordinary equations to the cycloid

$$x = a(\theta + \sin \theta), y = a(1 - \cos \theta),$$

we shall have

$$m = \frac{\tau \left(\sec \frac{1}{2}\theta\right)^3}{4ag}, w = \tau \tan \frac{1}{2}\theta.$$

(13) One end of a heavy chain is attached to a fixed point A, and the other to a weight which is placed on a rough horizontal

plane passing through A, and the chain hangs through a slit in the horizontal plane; to find the greatest distance of the weight from A, at which equilibrium is possible.

If a be the length of the chain, x the greatest distance of the weight from A, μ the coefficient of friction, and n twice the ratio between the given weight and that of the chain,

$$\varepsilon^{\frac{\pi}{a}} = \left\{ \frac{1 + \left\{ 1 + \mu^2 \left(1 + n \right)^2 \right\}^{\frac{1}{6}}}{\mu \left(1 + n \right)} \right\}^{\mu(1+n)}.$$

(14) A uniform chain is suspended from two tacks in the same horizontal line at a distance 2a from each other; to determine the length of the chain that the stress on the tacks may be a minimum.

Let c denote a length of the chain of which the weight is equal to the tension at the lowest point; and let l denote the required length of the chain. Then

$$\frac{a}{\epsilon^{r}} = \left\{ \frac{a}{c} + 1 \atop \frac{a}{c} - 1 \right\}^{\frac{1}{4}}, \quad l = c \left(\epsilon^{\frac{a}{r}} - \epsilon^{-\frac{a}{r}} \right);$$

from the former equation $\frac{a}{c}$ and therefore c is to be determined, and then l will be given by the latter.

If for instance 2a = 10 feet, then c = 4.168 feet nearly, and l = 12.578 feet nearly.

Diarian Repository, p. 644.

(15) A chain acted on by gravity hangs in the form of a curve, of which $a^3y = x^4$ is the equation; to find where the unit of mass is a maximum, and its maximum value.

When m is a maximum, x and y being the co-ordinates of the point,

$$x = \frac{a}{2^{\frac{1}{4}}}, y = \frac{1}{4} a, m = \frac{2 \cdot 3^{\frac{1}{4}} \tau}{ag}.$$

The law of the mass of the chain is erroneously investigated in the Lady's and Gentleman's Diary for the year 1745; see also Diarian Repository, p. 435.

SECT. 3. Central Forces.

(1) To find the equation to a flexible string in a position of equilibrium under the action of any central attractive force.

Let APB (fig. 64) be any portion of the string; S the centre of force; P any point in the string, PT a tangent at this point; SY a perpendicular from S upon PT; p a point of the string indefinitely near to P, and pk a tangent at p. Also let OP, Op, be the normals at P, p, O being therefore the centre and OP the radius of curvature at P; let OP produced meet pk in k.

 $OP = \rho$, SP = r, SY = p, $\angle SPT = \phi$, $\angle kOp = \psi$, m = the unitof mass at P; t = the tension at P and t + dt at p; Pp = ds, F =central force at P.

Then for the equilibrium of the element Pp we have, resolving the forces which act upon it at right angles to PT,

Finds $\sin \phi = (t + dt) \cos pkO = (t + dt) \sin \psi$, or retaining infinitesimals of the first order,

Fmds
$$\sin \phi = t\psi = t \frac{ds}{\rho}$$
;

and therefore

$$Fm \sin \phi = \frac{t}{\rho} \dots (a);$$

and resolving forces parallel to PT we have

Finds
$$\cos \phi = (t + dt) \sin Okp - t$$

= $(t + dt) \cos \psi - t$,

or, retaining infinitesimals of the first order only,

$$Fmds\cos\phi=dt;$$

and therefore, $ds \cos \phi$ being equal to dr,

$$Fmdr = dt \dots (b).$$

From the equation (a), since

$$\rho = -r \frac{dr}{dp} \text{ and } \sin \phi = \frac{p}{r},$$

we have

$$Fmdr + \frac{dp}{p} t = 0;$$

and therefore from (b)

$$\frac{dp}{p}+\frac{dt}{t}=0,$$

$$\log (pt) = \log C,$$

where C is an arbitrary constant; and therefore

$$p=\frac{C}{t}=\frac{C}{\int Fmdr}, \ldots (c),$$

which is the equation to the catenary in p and r when the form of F is known.

Let θ be the angle between SP and any fixed line; then

$$p = \frac{r^2 d\theta}{\left(dr^2 + r^2 d\theta^2\right)^{\frac{1}{4}}},$$

and therefore from (c), putting $\int Fmdr = R$,

$$Rr^{2} d\theta = C(dr^{2} + r^{2} d\theta^{2})^{\frac{1}{2}},$$

$$R^{2}r^{4}d\theta^{2} = C^{2}(dr^{2} + r^{2} d\theta^{2}).$$

and therefore

the differential equation to the catenary between r and θ . This the form in which the solution is given by John Bernoulli.

The value of the tension at any point of the catenary is given by (b), when the expression for F in terms of r is known.

The relations at which we have arrived may be deduced from the general equations of equilibrium of section (1); the method however of the tangential and normal resolution is more convenient in the case of central forces.

If the central force be repulsive instead of attractive, we must replace F by -F, wherever it occurs in the above formulæ.

(2). To find the form of the catenary when the central force is attractive and varies inversely as the square of the distance; the unit of mass being invariable.

¹ Opera, tom. IV. p. 238.

Let AOB (fig. 65) be the catenary; S the centre of force; SO the radius vector which meets the curve at right angles.

$$r =$$
the tension at O , and $SO = c$.

Then if k denote the attraction at the distance c.

$$R = \int Fmdr = \int k \frac{c^2}{r^2} mdr = C' - \frac{mkc^2}{r},$$

where C' is an arbitrary constant: but, by (1, b), t = R, and therefore

$$\tau = C' - mkc;$$

hence

$$t = R = r + mkc - \frac{mkc^2}{r} \dots (a).$$

Hence, from (1, d), we have

$$d\theta = \frac{Cdr}{r\left\{\left(r + mkc - \frac{mkc^2}{r}\right)^2 r^2 - C^2\right\}^{\frac{1}{2}}};$$

but from (1, c), since p = c and t = r at the point O, we see that $C = c\tau$; therefore

$$d\theta = \frac{c\tau dr}{r\left\{\left(r + mkc - \frac{mkc^2}{r}\right)^2 r^2 - c^2r^2\right\}^{\frac{1}{6}}}.$$

For the sake of simplicity put r = nmkc; then

$$d\theta = \frac{ncdr}{r\left\{\left(n+1-\frac{c}{r}\right)^{2}r^{2}-n^{2}c^{2}\right\}^{\frac{1}{4}}},$$

$$= \frac{ncdr}{r\left\{(n+1)^{2}r^{2}-2(n+1)cr+c^{2}-n^{2}c^{2}\right\}^{\frac{1}{4}}};$$

the equation to the catenary resulting from the integration of this differential equation will be of three different forms according as n is greater than, equal to, or less than unity.

First, suppose that n is greater than unity; then the integral of the equation will be, supposing that $\theta = 0$ when r = c,

$$r = \frac{(n-1)c}{n\cos\left\{\left(n^2-1\right)^{\frac{1}{2}}\frac{\theta}{n}\right\}-1}.$$

Secondly, suppose that n=1; then the equation to the catenary will be, if $\theta=0$ when r=c,

$$r=\frac{c}{1-\theta^2}.$$

Thirdly, let n be less than unity; then, if as before $\theta = 0$ when r = c, the equation will be

$$\varepsilon^{(1-n^2)\frac{\theta}{n}}+\varepsilon^{-(1-n^2)\frac{\theta}{n}}=\frac{2}{nr}\left\{r-(1-n)c\right\}.$$

Again from (1, c) we have, since $C = c\tau$,

$$p=\frac{c\tau}{t},$$

and therefore by (a)

$$p = \frac{c\tau}{\tau + mkc - \frac{mkc^2}{r}} = \frac{nc}{n+1-\frac{c}{r}};$$

hence putting $r = \infty$ we have $p = \frac{nc}{n+1}$, which shews that the three catenaries, corresponding to the three values of n, have all of them asymptotes passing within a distance $\frac{nc}{n+1}$ from the centre of force. Put $r = \infty$ in the equations to the three curves, and we get for the inclinations of the pair of asymptotes of each to the line SO,

$$\frac{n}{(n^2-1)^{\frac{1}{4}}}\cos^{-1}\frac{1}{n}, \text{ and } \frac{n}{(1-n^2)^{\frac{1}{4}}}\log\frac{1+(1-n^2)^{\frac{1}{4}}}{n}.$$
John Bernoulli; *Opera*, tom. IV, p. 240.
Whewell's *Mechanics*, 3rd edit. p. 183.

(3) To find the equation to a uniform catenary AOB, (fig. 65), acted on by a central force tending to S, the intensity of which varies as the μ^{th} power of the distance; the tension at O being $(1 + \mu)^{th}$ of the weight of a length SO of the string, each element of which length is supposed to be acted on by a constant force equal to that at O and towards S.

The notation remaining the same as in (2), the equation to the catenary will be

$$\left(\frac{c}{r}\right)^{\mu,2}=\cos\left(\mu+2\right)\theta.$$

(4) To find the equation to a uniform catenary SAOB, (fig. 66), acted on by a central repulsive force emanating from S, at which the two ends of the string are fastened, the intensity of this force varying inversely as the μ^{th} power of the distance; the tension at O being $(\mu - 1)^{th}$ of the weight of a length SO of the string, each element of which length is supposed to be acted on by a constant force equal to that at O and from S.

The notation remaining the same as before, the equation to the curve will be

$$\left(\frac{r}{c}\right)^{\mu-2}=\cos\left(\mu-2\right)\theta.$$

SECT. 4. Constrained Equilibrium.

(1) A flexible string ab, (fig. 67), acted on by gravity, rests on the arc of a curve APB in a vertical plane; to find the tension of the string and the pressure on the curve at any point.

Let P, p, be any two points of the curve very near to each other; PO, pO, normals at these points, the point O being the centre of curvature when p approaches indefinitely near to P: let ax, ay, be the axes of x, y, the former being horizontal, the latter vertical; aP = s, Pp = ds; t = the tension at P and t + dt at p; R = the unit of pressure on the curve at P, m = the unit of mass of the string, $\angle POp = \phi$, $PO = \rho$.

Then, resolving forces which act on the element Pp of the string, parallel to the tangent at P, we have

$$(t+dt)\cos\phi-t=mgds.\frac{dy}{ds},$$

or, neglecting infinitesimals of higher orders than the first,

$$dt = mgdy$$
;

integrating and observing that t is equal to zero when y = 0, we get

$$t = mgy \dots (1),$$

which gives the tension at any point of the string.

Again, resolving the forces on the element Pp parallel to the normal OP,

$$mgds. \frac{dx}{ds} + (t + dt) \sin \phi = Rds,$$

or, neglecting infinitesimals of orders higher than the first,

$$mg \frac{dx}{ds} + t \frac{\phi}{ds} = R;$$

but $\frac{\phi}{ds}$ is equal to $\frac{1}{\rho}$; hence we have for the pressure on the curve at any point

$$R = mg \frac{dx}{ds} + \frac{t}{\rho}$$

$$= mg\left(\frac{dx}{ds} + \frac{y}{\rho}\right).$$

(2) Two equal weights Q, Q, are suspended at the extremities of a flexible string hanging over a smooth curve in a vertical plane; to find the pressure at any point of the curve, the weight of the string being reckoned inconsiderable.

Let APB (fig. 68) be the curve; OP, Op, normals at two consecutive points P, p; θ the inclination of OP to some assigned line in the plane of the curve, $POp = d\theta$; $PO = \rho$, AP = s, Pp = ds; p = the unit of pressure on the curve at the point P; t = the tension of the string at P, t + dt = the tension at p.

Then for the equilibrium of the element Pp of the string we have, resolving forces at right angles to the tangent at P,

$$(t+dt)\sin d\theta = pds = p\rho d\theta,$$

and therefore, retaining infinitesimals of the first order,

Again, resolving forces parallel to the tangent at P,

$$(t+dt)\cos d\theta - t = 0,$$

and therefore, retaining infinitesimals of the first order,

$$dt = 0$$
, $t = constant$;

but evidently at A the tension is equal to Q; hence t = Q. Hence from (1) we have

$$Q = p\rho$$
, $p = \frac{Q}{\rho}$.

Cor. The whole pressure on the curve AB is equal to

$$\int pds = Q \int \frac{ds}{\rho} = Q \int_{\theta_1}^{\theta_2} d\theta = Q(\theta_2 - \theta_1),$$

If the tangents at the points where the string leaves the curve be vertical, we have πQ for the whole pressure along the curve; if they be not vertical there will of course be pressures at the points A, B, in addition to the pressure along the curve.

Euler; Nov. Comment. Petrop. 1775, p. 307. Poisson; Traité de Mécanique, tom. 1. ch. 3.

(3) To find the pressure on a curve AB, (fig. 69), when two weights Q, R, balance each other over it by means of a string of negligible weight, the friction between the string and the curve being taken into account; and the weight Q being considered as much greater than R as is consistent with equilibrium.

Let μ be the coefficient of friction; the rest of the notation being the same as in the preceding problem. Then the friction on the element Pp will be μpds , and will act nearly in the direction of the tangent at P. Hence, resolving forces on the element Pp parallel to PO, we have

$$(t + dt) \sin d\theta = pds = p\rho d\theta$$
;

and therefore in the limit

$$td\theta = p\rho d\theta, \quad t = p\rho \dots (1);$$

again, resolving forces parallel to the tangent at P,

$$(t+dt)\cos d\theta - t + \mu p ds = 0,$$

and therefore in the limit

$$dt + \mu p ds = 0,$$

and consequently by (1)

$$dt + \frac{\mu t}{\rho} ds = 0;$$

integrating, we get

$$\log t = -\mu \int \frac{ds}{\rho} = -\mu \int d\theta = C - \mu \theta \dots (2);$$

hence, the values of t at A and B being Q and R,

$$\log Q = C - \mu \theta_1, \quad \log R = C - \mu \theta_2 \dots (3),$$

and therefore
$$\log \frac{Q}{R} = \mu (\theta_2 - \theta_1), \quad \frac{Q}{R} = \epsilon^{\mu (\theta_2 - \theta_1)},$$

which expresses the relation which must subsist between Q and R under the circumstances of the problem.

1

Also, from (2) and (3),

$$\log \frac{t}{R} = \mu (\theta_2 - \theta), \quad t = R \epsilon^{\mu (\theta_1 - \theta)} \dots (4);$$

hence the whole pressure along the curve is equal to

$$\int pds = \int \frac{ds}{\rho} t, \text{ from (1)},$$

$$= \int td\theta = R \int \varepsilon^{\mu(\theta_{n}-\theta)} d\theta = C - \frac{R}{\mu} \varepsilon^{\mu(\theta_{n}-\theta)};$$

but when $\theta = \theta_1$, it is clear that the pressure along the curve is zero; hence

$$0 = C - \frac{R}{\mu} \, \epsilon^{\mu(\theta_2 - \theta_1)},$$

and therefore the whole pressure from θ_1 to θ_2 , is equal to

$$\frac{R}{\mu} \left\{ \varepsilon^{\mu(\theta_{\mathfrak{g}}^-\theta_1)} - 1 \right\}.$$

In addition to this pressure along the curve there are the pressures at the extremities A and B.

Cor. If the curve be a semicircle $\theta_2 - \theta_1 = \pi$, and we have

$$\frac{Q}{R}=\epsilon^{\mu \pi}.$$

Euler; Nov. Comment. Petrop. 1775, p. 316. Poisson; Traité de Mécanique, tom. 1. ch. 3.

SECT. 5. Extensible Strings.

If an extensible string of given length be stretched by any force, it is found by experiment that the extension of the string beyond its natural length is proportional to the force. From this it is easily seen that, if the string be of variable length, the extension will vary as the product of the force and the natural length of the string. Hence if a denote the natural length of the string, and a' the length under the action of a stretching force P, we shall have

$$a'=a(1+\lambda P),$$

where λ is a constant quantity depending upon the quality of the string.

This theory was first announced by Hooke, in the form of an anagram, among a list of inventions at the end of his Descriptions of Helioscopes, published in the year 1676. The anagram is ceiiinosssttuu, from which may be extracted the proposition, "ut tensio sic vis." He afterwards published a work entitled De Potentia Restitutiva or Spring, in which the theory was developed at large with experimental illustrations. Hooke's theory forms the basis of a memoir by Leibnitz, in the Acta Eruditorum for the year 1684, entitled Demonstrationes Novae de Resistentia Solidorum. For additional information on the subject the student is referred to s'Gravesande's Element. Physic. lib. 1. c. 26.

(1) An elastic string AC (fig. 70) is suspended from its extremity A, and has a weight attached to it at a point B; the natural lengths of AB, BC, being given, to find the length of the string AC in its present circumstances.

Let m denote the unit of mass of the string in its natural state; a, b, the natural lengths of AB, BC, and a', b', their lengths under the circumstances of the problem; c the length of a portion of the natural string, the weight of which is equal to the weight attached to B; let P be any point in AB, and p very near to it, AP = x, Pp = dx; t = the tension at P and t + dt at p.

Then, since by Hooke's Principle the unit of mass of the element dx must evidently be diminished in the ratio of $1 + \lambda t : 1$, the weight of Pp will be

$$\frac{mgdx}{1+\lambda t}$$
,

and therefore for the equilibrium of Pp

$$t+dt+\frac{mgdx}{1+\lambda t}-t=0,$$

$$(1 + \lambda t) dt + mgdx = 0;$$

integrating we get

$$t\left(1+\tfrac{1}{2}\,\lambda t\right)+mgx=C;$$

but it is evident that

$$t = mg(a + b + c)$$
, when $x = 0$,
 $t = mg(b + c)$, when $x = a'$;

and

hence we obtain

 $(a+b+c)\{1+\frac{1}{2}\lambda mg(a+b+c)\}=(b+c)\{1+\frac{1}{2}\lambda mg(b+c)\}+a',$ and therefore

$$a' = a + \frac{1}{2} \lambda mg \left\{ (a + b + c)^2 - (b + c)^2 \right\}$$

$$= a + \frac{1}{2} \lambda mg \left(a^2 + 2ab + 2ac \right)$$

$$= a \left\{ 1 + \frac{1}{2} \lambda mg \left(a + 2b + 2c \right) \right\} \dots (1).$$

Again, if Q be any point in BC, BQ = y, and r = the tension at Q, we shall have, as before,

but evidently
$$r = mgb, \quad \text{when } y = C;$$
 and
$$r = 0, \quad \text{when } y = b';$$
 hence we have
$$b' = b \left(1 + \frac{1}{2} \lambda mgb\right) \dots \dots \dots (2).$$

Hence from (1) and (2), if l' denote the whole length of the string AC, we find that

$$l' = a + b + \frac{1}{2} \lambda mg \{ a (a + 2b + 2c) + b^2 \}.$$

(2) An elastic string, of which the unstretched length is a, is placed on a smooth inclined plane the length of which is also equal to a; to find the length which will hang over the plane, the string being stretched by its own weight.

Let ACD (fig. 71) be the string hanging from the point A in the inclined plane AC; P any point in AC, and p a point near to P; t = the tension at P, T at A, and r at C; AP = x, Pp = dx; m = the unit of mass of the string when unstretched, a = the inclination of AC to the horizon.

Then, by virtue of Hooke's Principle, the mass of Pp will be

$$\frac{mdx}{1+\lambda t}$$
;

and therefore, t + dt being the tension at p, we have for the equilibrium of Pp

$$dt + \frac{mg \sin a}{1 + \lambda t} dx = 0,$$

$$(1 + \lambda t) dt + mg \sin a dx = 0;$$

integrating we obtain

$$t\left(1+\tfrac{1}{2}\lambda t\right)+mgx\sin a=C;$$

hence τ being the value of t when x = a and T when x = 0, we have

$$\tau \left(1 + \frac{1}{2}\lambda \tau\right) + mga \sin a = T\left(1 + \frac{1}{2}\lambda T\right),$$

$$(\mathbf{r}-\mathbf{T})\left\{1+\tfrac{1}{2}\lambda\left(\mathbf{r}+\mathbf{T}\right)\right\}+mga\,\sin\,a=0\,\ldots\ldots\,(1)$$

Let s be the natural length of CD; then a - s will be the natural length of AC; hence clearly

$$r = mgs$$
, $T = mg \{s + (a - s) \sin a\}$;

we have then, by (1),

$$(a-s)[1+\frac{1}{2}\lambda mg \{2s+(a-s)\sin a\}]=a,$$

 $\frac{1}{2}\lambda mg (a-s)\{2s+(a-s)\sin a\}=s,$

whence s may be determined by the solution of a quadratic equation.

If s' be the actual length of the portion CD of the string, we may shew that

$$s' = s \left(1 + \frac{1}{2} \lambda m s\right),$$

and therefore s and s' are both known.

(3) A slightly extensible string Aa (fig. 72) is attached to the upper extremity A of the vertical radius AO of a circular arc AB along which it rests; having given its natural length, to find its length as it rests on the arc.

Let P, p, be any two points very near together in the string Aa; draw the lines PO, pO; let AP = s, Pp = ds, $\angle AOP = \phi$, $\angle POp = d\phi$, AO = a; m = the unit of mass of the string when unstretched; t, t + dt, the tensions at P, p; s', ds', the lengths of AP, Pp, without stretching.

Then, by Hooke's Principle,

$$ds = (1 + \lambda t) ds' \dots (1).$$

Again, for the equilibrium of the portion Pp of the string, we have, resolving forces parallel to the tangent at P,

$$(t+dt)\cos(d\phi)+mgds'\sin\phi=t;$$

and therefore, retaining infinitesimals of the first order,

$$dt + mqds' \sin \phi = 0$$
;

hence by the aid of (1), we have

hence by the aid of (1), we have
$$(1 + \lambda t) dt + mg \sin \phi ds = 0;$$
or, since $s = a\phi$,

$$(1+\lambda t) dt + mag \sin \phi d\phi = 0;$$

integrating we get

$$t(1+\frac{1}{3}\lambda t)-mag\cos\phi=C;$$

let β be the angle subtended at O by the arc $A\alpha$; then it is clear that when $\phi = \beta$, t will be equal to zero; hence

- mag
$$\cos \beta = C$$
,

and therefore $t(1+\frac{1}{2}\lambda t) = mag(\cos\phi - \cos\beta)\dots(2)$.

From (1) we have, putting $a \phi$ for s,

$$ad\phi = (1 + \lambda t) ds',$$

and λ being by the hypothesis a small quantity,

$$ds' = ad\phi (1 - \lambda t) = ad\phi - a\lambda t d\phi \dots (3).$$

Now from (2) we get approximately

$$t = mag(\cos \phi - \cos \beta),$$

and therefore, substituting this value of t in the small term of the equation (3),

$$ds' = ad\phi - ma^2\lambda g (\cos \phi - \cos \beta) d\phi$$
;

integrating we get

$$s' + C = a\phi - ma^2\lambda g (\sin \phi - \phi \cos \beta);$$

but when $\phi = 0$, it is evident that s' = 0; hence C = 0, and we have

$$s' = a\phi - ma^2\lambda q (\sin \phi - \phi \cos \beta);$$

let $a\beta'$ be equal to the natural length of Aa; then evidently

$$a\beta' = a\beta - ma^2\lambda g (\sin \beta - \beta \cos \beta),$$

$$\beta' = \beta - ma\lambda g (\sin \beta - \beta \cos \beta);$$

but since $\beta = \beta'$ nearly, we may substitute β' for β in the coefficient of the small quantity λ ; thus we obtain

$$\beta = \beta' + ma\lambda g (\sin \beta' - \beta' \cos \beta'),$$

which determines the required length Aa.

(4) Two weights P, Q, (fig. 73), resting on two smooth inclined planes CA, CB, are connected by a given elastic string PQ; to find their position of equilibrium.

Let θ be the inclination of QP to the horizon; a, β , the inclinations to the horizon of the planes CA, CB; a = the natural length of the string PQ. Then the position of equilibrium will be defined by the two equations,

$$\tan \theta = \frac{P \cot \beta - Q \cot \alpha}{P + Q}, \quad PQ = a \left\{ 1 + \frac{\lambda P \sin \alpha}{\cos (\alpha - \theta)} \right\}.$$

(5) Two equal weights P, Q, (fig. 74), are connected by an elastic string PQ, of which the horizontal line BC is the natural length; to find the nature of the curves BP, CQ, on which they will always remain in equilibrium with the string parallel to the horizon; the plane of the curves being vertical.

Bisect BC in A, and draw AM vertical; let AB = a = AC, AM = x, MP = y = MQ; then the equation to each of the curves will be

$$(y-a)^2 = 2\lambda a P x,$$

or BP, CQ, are two semi-parabolas of which B, C, are the vertices.

(6) An elastic ring BC is placed round a vertical cone and descends by its own weight; to find the position of equilibrium.

Let O (fig. 75) be the centre of the ring in its position of equilibrium, $\angle OAB = \alpha$; $2\pi a =$ the natural length of the ring, and W = its weight; then

$$BO = a\left(1 + \frac{\lambda}{2\pi} W \cot a\right),\,$$

which determines the required position.

CHAPTER VI.

VIRTUAL VELOCITIES.

THE Principle of Virtual Velocities consists in the following general proposition:

"If any assignable system of bodies or points, solicited each of them by any forces whatever, be in equilibrium; and we conceive this system to experience consistently with its geometrical relations any small arbitrary displacement, by virtue of which each point describes an indefinitely small space; the sum of the forces multiplied each of them by the resolved part, parallel to its direction, of the space described by its point of application, will be always equal to zero; this resolved part being considered positive when it lies in the direction of its corresponding force, and negative when in an opposite direction."

The resolved parts of the spaces described by the points of application of the forces are called their Virtual Velocities. Let P, Q, R, \ldots denote any system of forces acting on a system of points consistently with equilibrium; and let a, β, γ, \ldots denote their virtual velocities; then, as far as the first powers of a, β, γ, \ldots are concerned,

$$Pa + Q\beta + R\gamma + S\delta + \ldots = 0 \ldots (A)$$

The Principle of Virtual Velocities was first detected by Guido Ubaldi as a property of the equilibrium of the lever and of moveable pullies. Its existence was afterwards recognized by Galileo in the inclined plane, and the machines depending upon it. The expression 'moment' of a force or weight acting on any

¹ Mechanicorum Liber; De Libra, De Cochlea.

² Della Scienza Mccanica, Opera, tom. 1. p. 265; Bologna, 1655.

machine, was used by Galileo to denote its energy or effort to set the machine in motion, who accordingly declared that for the equilibrium of a machine acted on by two forces, it is necessary that their moments should be equal, and should take place in opposite directions; he shewed moreover that the moment of a force is always proportional to the force multiplied by its virtual velocity. The word 'moment' was used in the same sense by Wallis', who adopted Galileo's principle of the equality of moments as the fundamental principle of Statics; and deduced from it the conditions for the equilibrium of the principal machines. Descartes has likewise reduced the whole science of Statics to a single principle, which virtually coincides with that of Galileo; it is presented however under a less general aspect. The principle is, that it requires precisely the same force to raise a weight P through an altitude a, as a weight Q through an altitude b, provided that P is to Q as b to a. From this it follows, that two weights attached to a machine will be in equilibrium when they are disposed in such a manner that the small vertical paths which they can simultaneously describe are reciprocally as the weights.

Torricelli³ is the author of another principle which may be immediately deduced from the principle of virtual velocities: the principle is, that when any two weights rigidly connected together are so placed that their centre of gravity is in the lowest position which it can assume consistently with the geometrical conditions to which they are subject, they will be in equilibrium. The principle of Torricelli has given birth to the following more general one, viz.—that any system whatever of heavy bodies will be in equilibrium when their centre of gravity is in its lowest or highest position.

John Bernoulli was the first to announce the principle of virtual velocities under its most general aspect in the form which we have given above, in a letter to Varignon⁴, dated Bâle, Jan. 26, 1717. The striking value of the principle,

¹ Mechanica, sive de Motu, Tractatus Geometricus.

^{*} Lettre 73, tom. 1. 1657; De Mechanica Tractatus, Opuscuia Posthuma.

^{*} De Motu gravium naturaliter descendentium, 1644.

¹ Nouvelle Mécanique, tom. 11. sect 9.

as an instrument of analytical generalization, has been splendidly exhibited by Lagrange in his *Mécanique Analytique*.

From the principle of virtual velocities may be immediately deduced the principle which was proposed by Maupertuis in the *Mémoires de l'Académie des Sciences de Paris* for the year 1740, under the name of the *Loi de Repos*; and which Euler has developed at large in the *Mémoires de l'Académie de Berlin* for the year 1751. Suppose that any number of forces P, Q, R, \ldots tending towards fixed centres and functional of their distances p, q, r, \ldots from the centres, to act on a system of points rigidly connected together. Then supposing the system of points to be slightly displaced, so that p, q, r, \ldots receive increments dp, dq, dr, \ldots we shall have, by the principle of virtual velocities,

$$Pdp + Qdq + Rdr + \ldots = 0.$$

Let $d\Pi$ denote the left-hand member of this equation; then

$$d\Pi = 0 \ldots (B)$$
.

From this it appears that if the system be so placed that Π may have a maximum or a minimum value, there will be equilibrium: this proposition constitutes Maupertuis' Principle of Rest. It does not however follow conversely that, whenever the system is at rest, Π shall have a maximum or minimum value, since by the principles of the differential calculus we know that the equation (B), although a necessary, is not the only condition for the existence of such a value. Lagrange¹ has shewn that if Π be a minimum the equilibrium will be stable, and if a maximum, unstable.

As an example of this theory, it is evident that, if any system be in equilibrium under the action of gravity, there will be stable or unstable equilibrium accordingly as the centre of gravity is in the lowest or highest position which is compatible with the geometrical relations to which the system is subject.

The principle of equilibrium developed by Courtivron² is likewise grounded upon the principle of virtual velocities; Courtivron's Principle asserts, that if a system of bodies be in motion under the action of any forces varying according to any assigned

^{&#}x27; Mécanique Analytique. Première Partie, sect. 5.

^{&#}x27; Mémoires de l'Académie des Sciences de Berlin, 1748, 1749.

laws, a position of the system corresponding to a maximum or minimum value of the vis viva will be a position of equilibrium; a maximum value of the vis viva corresponding to stable, and a minimum to unstable equilibrium.

Sect. 1. Equilibrium.

(1) A particle P (fig. 76) is attracted towards two centres of force A and B; to find the position of the particle that it may be in equilibrium.

Let A, B, denote the two forces; AP = r, BP = s, AB = a; draw PM at right angles to AB, and let AM = x, PM = y. Then, supposing P to receive some slight arbitrary displacement, the decrements dr, ds, of r, s, will be the virtual velocities of the forces A, B; hence, by the formula (A),

$$Adr + Bds = 0 \dots (1).$$
But
$$r = (x^2 + y^2)^{\frac{1}{2}}, s = \left\{ (a - x)^2 + y^2 \right\}^{\frac{1}{4}},$$

$$dr = \frac{xdx + ydy}{(x^2 + y^2)^{\frac{1}{4}}}, \quad ds = \frac{-(a - x)dx + ydy}{\left\{ (a - x)^2 + y^2 \right\}^{\frac{1}{4}}};$$

and therefore by (1),

$$A\frac{xdx + ydy}{(x^2 + y^2)^{\frac{1}{4}}} + B\frac{-(a-x)dx + ydy}{\{(a-x)^2 + y^2\}^{\frac{1}{4}}} = 0;$$

but since dx and dy are independent quantities, whatever be the small variation in the position of P, we have, equating their coefficients to zero,

$$\frac{Ax}{(x^{2}+y^{2})^{\frac{1}{4}}} - \frac{B(a-x)}{\{(a-x)^{2}+y^{2}\}^{\frac{1}{4}}} = 0 \dots (2),$$

$$\frac{Ay}{(x^{2}+y^{2})^{\frac{1}{4}}} + \frac{By}{\{(a-x)^{2}+y^{2}\}^{\frac{1}{4}}} = 0 \dots (3).$$

From (3) we have y = 0, and therefore from (2), we see that A = B; thus it appears that if any particle be acted on by two forces tending towards two fixed centres, the conditions for its

equilibrium are, first, that it shall lie in the straight line joining the two centres, and, secondly, that the two forces shall be equal.

Euler; Mémoires de l'Académie de Berlin, 1751, p. 184.

(2) A rigid rod AB (fig. 77) without weight, rests over a peg O, and against a smooth wall CD, and is acted on by a weight P suspended from the extremity A; to determine its position of equilibrium and the pressures on the wall and the peg.

Draw EOF horizontally; let AB = a, OB = x, OE = b, AF = y. Let R, S, denote the reactions of the wall and peg against the rod, of which the former will be horizontal, and the latter at right angles to AB. Conceive the rod AB to be slightly displaced from its position of rest by making its end B slide along CD, the peg O still touching the rod; then it is evident that the point B will have no motion parallel to B, and that the motion of the point B of the rod resolved parallel to B will be an infinitesimal of the second order. Hence of the three forces B, B, B alone will have a virtual velocity. We have then, by the principle of virtual velocities,

$$Pdy = 0$$
, or $dy = 0 \dots (1)$.

Now by similar triangles AFO, BEO, there is

$$AF = AO \cdot \frac{BE}{BO}$$

and therefore

$$y = \frac{a-x}{x} (x^2 - b^2)^{\frac{1}{4}};$$

differentiating this equation and performing obvious simplifications, we shall have

$$dy = \frac{ab^2 - x^3}{x^2 (x^2 - b^2)^{\frac{1}{4}}} dx;$$

and therefore, by (1),

$$ab^2-x^3=0, x=(ab^2)^{\frac{1}{3}}....(2),$$

which defines the position of equilibrium.

In order to determine S, conceive each point of the rod to receive the same vertical displacement β , the point B thus sliding along CD and the rod moving parallel to itself.

Then, putting $\angle BOE = \phi$, the virtual velocities of P, R, S, will be $-\beta$, 0, β cos ϕ , respectively, and therefore

$$S \cdot \beta \cos \phi = P \cdot \beta$$
, $S \cos \phi = P$;
 $\cos \phi = \frac{b}{x} = \left(\frac{b}{a}\right)^{\frac{1}{3}}$, by (2);
 $S = P\left(\frac{a}{b}\right)^{\frac{1}{3}}$.

hence

but

In order to find R, conceive the rod to be displaced along its length through a space β ; then the virtual velocities of P, R, S, being $-\beta \sin \phi$, $\beta \cos \phi$, 0, respectively, we have

$$R\beta\cos\phi=P\beta\sin\phi$$
, $R=P\tan\phi$;

but

$$\tan \phi = \frac{(x^2 - b^2)^{\frac{1}{4}}}{b} = \frac{(a^{\frac{3}{4}} - b^{\frac{3}{4}})^{\frac{1}{4}}}{b^{\frac{1}{4}}}, \text{ by (1)};$$

and therefore

$$R = P \frac{(a^{\frac{2}{3}} - b^{\frac{2}{3}})^{\frac{1}{4}}}{b^{\frac{1}{3}}}.$$

Euler; Mémoires de l'Académie de Berlin, 1751, p. 196.

(3) A particle is placed upon a smooth inclined plane AB, (fig. 78), at a point O, and acted on by a force P in a given direction; to determine the magnitude of P, that the particle may be at rest, and the pressure on the plane.

Let W be the weight of the particle, R the reaction of the plane, $\angle POB = \varepsilon$, $\alpha =$ the inclination of AB to the horizon AC.

Conceive the particle O to receive a displacement β along the plane AB; then, the virtual velocity of R being zero, the virtual velocities of P, W, will be β cos ϵ , $-\beta$ sin α , respectively. Hence, by the principle of virtual velocities,

$$P\beta \cos \varepsilon = W\beta \sin \alpha$$
, $P \cos \varepsilon = W \sin \alpha \dots (1)$,

which determines the value of P.

Next, displace the particle parallel to AC through a space β ; then, the virtual velocity of W being zero, the virtual velocities of P, R, will be respectively $\beta \cos (\alpha + \epsilon)$, $-\beta \sin \alpha$, and therefore $P\beta \cos (\alpha + \epsilon) = R\beta \sin \alpha,$

whence

$$R = \frac{P\cos(a+\epsilon)}{\sin a} = \frac{W\cos(a+\epsilon)}{\cos \epsilon}, \text{ by (1)}.$$

Euler; Mémoires de l'Académie de Berlin, 1751, p. 191.

(4) A rigid rod OA, (fig. 79), without weight, is acted on by a weight P hanging from its extremity A; the end O of the rod is fixed; also EF is a spring in the form of a circular are to a centre O, of which the force of contraction varies as the angle AOB, OB being a horizontal line; to find the position of the rod that it may be at rest.

Let OA = a, $\angle AOB = \phi$, OF = b; a the value of ϕ when the force of the spring's contraction is equal to E; then corresponding to the angle ϕ the force of contraction will be equal to

$$\frac{E_{\phi}}{a}$$
.

Let AO be displaced slightly through an angle $d\phi$ into the position Oa; draw ap at right angles to AP, and let f be the new position of F; then, by the principle of virtual velocities,

$$P \cdot Ap - \frac{E}{a} \phi \cdot Ff = 0 \cdot \dots (1);$$

but since Ap, Aa, are respectively at right angles to OB, OA, it is clear that $\angle aAp = \phi$, and therefore

$$Ap = Aa \cos \phi = ad\phi \cdot \cos \phi$$
;

also

$$\mathbf{F}\mathbf{f} = \mathbf{b}\mathbf{d}\mathbf{\phi}$$
:

hence from (1) we have

$$Pa \cos \phi - \frac{E}{a} \phi b = 0;$$

$$\frac{\cos \phi}{\phi} = \frac{Eb}{Paa},$$

or

the required condition of equilibrium, from which φ is to be determined.

Euler: Mémoires de l'Académie de Berlin, 1731, p. 196.

5 A smooth rod AB fig. 80 rests against two horizontal bars which pierce the vertical plane through the rod at right angles in the points A. A: the rod passes under the lower and over the higher bar, its lower extremity A being sustained upon

a smooth horizontal plane; to determine the pressures upon the two bars, and upon the horizontal plane.

The pressures upon the bars and upon the horizontal plane will be equal to their reactions upon the rod; the reactions of the bars upon the rod will be two forces R', R', at right angles to the rod; and the reaction of the horizontal plane will be a vertical force R. Let G be the centre of gravity of the rod, at which point we will suppose its whole weight to be collected. Thus we have four forces R, R', R'', W, acting respectively at the four points A, A', A'', G, rigidly connected together, so as to produce equilibrium.

Let AG = a, A'A'' = b, and a = the inclination of the rod to the horizon.

Conceive the rod to receive a small displacement of such a character that it still remains in contact with the two bars; then evidently the virtual velocities of W and R will be equal, the one being a positive and the other a negative velocity, and therefore, a denoting the magnitude of the virtual velocity of each, we have Ra - Wa = 0.

Next, conceive the rod to receive a slight displacement, as in (fig. 81), by revolving through a small angle ω about the point A' which is supposed to be kept stationary; the points a, a', g, being the new positions of A, A', G; from a draw am at right angles to the vertical line through A, and from g draw gn at right angles to the vertical Gn through G. Then, by the principle of virtual velocities,

and therefore, by (1),

$$W(Am - Gn) - R' \cdot A'a' = 0 \cdot \dots \cdot (2);$$
but
$$Am = Aa \cos a = AA'' \cdot \omega \cdot \cos a,$$
and
$$Gn = Gg \cos a = A''G \cdot \omega \cdot \cos a;$$
and therefore
$$Am - Gn = AG \cdot \omega \cos a = a\omega \cos a,$$
also
$$A'a' = A'A'' \cdot \omega = b\omega.$$
Hence from (2) we have substituting for $Am - Gn$ and $Am = Gn$ and

Hence from (2) we have, substituting for Am - Gn and A'a' their values, $Wa\omega \cos a - b\omega R' = 0$,

and therefore

$$R' = \frac{Wa \cos a}{h} \dots \dots (3).$$

Again, conceive the rod, as in (fig. 82), to be slightly displaced into the position Aa'ga''; draw gn horizontal and Gn vertical. Then, by the principle of virtual velocities,

$$R''$$
. $A''a'' - R'$. $A'a' - W$. $Gn = 0$;

but

$$Gn = Gg \cos a$$
,

hence

$$R'' \cdot A''a'' - R' \cdot A'a' - W \cdot Gg \cos a = 0$$
,

and therefore, observing that A''a', A'a', Gg, are evidently in the same proportion as AA'', AA', AG, we have

$$R'' \cdot AA'' - R' \cdot AA' - W \cdot AG \cdot \cos a = 0$$
;

and therefore by (3),

$$R'' \cdot AA'' = W\left(\frac{a \cos a}{b} AA' + AG \cos a\right)$$
$$= W\frac{a \cos a}{b} (AA' + b) = W\frac{a \cos a}{b} \cdot AA';$$

hence

$$R' = W \frac{a \cos a}{h} = R'$$
, by (1).

(6) A string of given length passes over a given pulley; it has attached to its two extremities two weights, one of which is capable of sliding freely on a given curve; to determine the curve on which the other ought to slide in order that in every position of the two weights they may be in equilibrium.

Let P, P', (fig. 83), denote the two weights in any position; A the pulley; and let AB be a vertical line through A; $AP = \rho$, $AP' = \rho'$; m, m', the masses of the bodies. Draw PM, P'M', at right angles to AB; let AM = x, AM' = x', $\angle PAB = \phi$, $\angle P'AB = \phi'$.

Then, supposing the two weights to receive displacements along two small arcs of their corresponding curves of constraint, we have, by the principle of virtual velocities,

$$mdx + m'dx' = 0;$$

and since this relation is true for all corresponding points of the two curves, we have, integrating,

$$mx + m'x' = c,$$

where c is some constant quantity; and therefore, in polar coordinates, $m\rho \cos \phi + m'\rho' \cos \phi' = c \dots (1)$.

Again, if I denote the length of the string,

$$\rho + \rho' = l \dots (2).$$

Supposing the curve on which P' moves to be the given one, we have $f(\rho', \phi') = 0 \dots (3)$,

where $f(\rho', \phi')$ denotes some known function of ρ', ϕ' .

Eliminate from the equations (1), (2), (3), the quantities ρ' , ϕ' , and we shall get for the equation of the required curve

$$\chi(\rho,\,\phi)=0,$$

 $\chi(\rho, \phi)$ denoting some function of ρ, ϕ .

John Bernoulli; Act. Erudit. 1695. Febr. p. 59. Leibnitz; Ib. April, p. 184. L'Hôpital; Act. Erudit. Suppl. tom. 11. sect. 6. p. 289. Fuss; Nova Acta Acad. Petrop. 1788, p. 197.

(7) Four uniform beams AB, BC, CD, DE, (fig. 84), connected together by smooth hinges, are placed in a position of equilibrium, the ends A and E being attached to two smooth hinges in the same horizontal line AE; the beam AB is equal to the beam ED, and the beam BC to the beam CD; to compare the angles BAE and CBD.

Let AB = DE = 2a, BC = CD = 2b, $\angle BAE = a$, $\angle CBD = \beta$, AE = c; h = the height of the centre of gravity of the beams above the line AE; m = the weight of each of the lower and n = that of each of the higher beams. Then

$$(2m + 2n)h = 2ma \sin a + 2n(2a \sin a + b \sin \beta)$$

 $(m + n)h = (m + 2n)a \sin a + nb \sin \beta;$

but for equilibrium h must be a maximum or a minimum; hence

 $0 = (m+2n) a \cos a da + nb \cos \beta d\beta \dots \dots (1).$

Again, it is evident by the geometry that

$$c = 4a \cos a + 4b \cos \beta$$

and therefore $0 = a \sin a \, da + b \sin \beta \, d\beta \dots (2)$.

Multiply (1) by $\sin a \sin \beta$, and then, by (2), we have

 $(m+2n)\cos a\sin \beta \cdot b\sin \beta d\beta = n\cos \beta\sin a \cdot b\sin \beta d\beta$, and therefore $(m+2n)\cos a\sin \beta = n\cos \beta\sin \alpha$,

$$\tan \alpha = \frac{m+2n}{n} \tan \beta.$$

If m = n, we have $\tan \alpha = 3 \tan \beta$.

(8) A beam AB, (fig. 85), rests with one end against a smooth vertical plane OK, and upon a smooth curve $a\beta$; the plane of the beam and the curve being at right angles to the plane OK; to determine the nature of the curve, that the beam may rest in any position.

From any point O in the section OK of the vertical plane draw OL horizontal; from the point of contact P corresponding to any position of the beam draw PM vertical; let G be the centre of gravity of the beam; draw GH vertical. Let OM = x, PM = y, AG = a. Then, from the geometry, we see that

$$GH = y + PG \cdot \frac{dy}{ds} = y + \left(a - x \frac{ds}{dx}\right) \frac{dy}{ds} = y - x \frac{dy}{dx} + a \frac{dy}{ds}.$$

Now since for the equilibrium of a material system acted on by gravity, it is necessary that its centre of gravity be in the highest or lowest possible position consistent with geometrical relations, it is clear that in the present problem, equilibrium being possible for every position of the beam, GH must be of

invariable magnitude. Hence, if
$$p = \frac{dy}{dx}$$
,

$$C = y - xp + \frac{ap}{\left(1 + p^2\right)^2};$$

differentiating with respect to x, and putting $\frac{dp}{dx} = q$,

$$0 = -xq + \frac{aq}{(1+p^2)^{\frac{3}{4}}};$$

hence

$$q = 0$$
, or $x = \frac{\alpha}{(1 + p^2)^{\frac{3}{4}}}$:

the former of these solutions gives a straight line for the locus of P; if we integrate the second equation, we shall get, for the equation to the locus of P,

$$C = (a^{\frac{q}{3}} - x^{\frac{q}{3}})^{\frac{3}{2}} + y.$$

Suppose that the origin of co-ordinates be so chosen that when x = a, y = 0, in which case O will be the intersection of the beam in its horizontal position with the line OK; then C = 0, and the equation will be $x^{\frac{3}{3}} + y^{\frac{3}{3}} = a^{\frac{3}{3}}.$

(9) A particle O is acted upon by three forces A, B, C, passing through three points A, B, C; to determine the conditions for the equilibrium of the particle by the principle of virtual velocities.

The three points A, B, C, must all lie in a single plane containing the particle; also the relative magnitude of the forces A, B, C, are given by any two of the three proportions,

 $A:B::\sin BOC:\sin AOC$, $A:C::\sin BOC:\sin AOB$, $B:C::\sin AOC:\sin AOB$.

Euler; Mémoires de l'Académie de Berlin, 1751, p. 185.

(10) A particle is acted on by any number of forces; to find the conditions to which their magnitudes and directions must be subject that the particle may be at rest.

From the particle draw straight lines representing the forces in magnitude and in direction; then, that the particle may be in equilibrium, its position must coincide with the centre of gravity of a number of equal particles placed at the extremities of the straight lines.

This celebrated theorem for the equilibrium of a particle is due to Leibnitz: Euler gave a demonstration by the aid of Maupertuis' Loi de Repos, and Lagrange by the Principle of Virtual Velocities. See also Poisson, Traité de Mécanique, tom. 1. No. 67. A more general theorem of forces, which comprehends this of Leibnitz as a particular case, has been given by Chasles: see Bulletins de l'Académie des Sciences et Belles-Lettres de Bruxelles, 1840, 2me partie, p. 261.

¹ Journal des Savans, 1693; Opera, tom. III. p. 283.

Mémoires de l'Académie de Berlin, 1751.

³ Mécanique Analytique, tom. 1. p. 106.

⁴ Correspondance Mathématique, tom. v. p. 106-108; 1829.

(11) A string of given length passes over a fixed point; it has attached to its two extremities two weights, one of which is capable of sliding freely along an inclined plane passing through the point; to determine the curve on which the other must be placed that in every position of the two weights they may be in equilibrium.

Let the angle which the inclined plane makes with the vertical be a; then, the notation remaining the same as in (6), the equation to the required curve will be

$$(m \cos \phi - m' \cos a) \rho = c - m'l \cos a,$$

which belongs to a conic section.

Fuss; Nova Acta Acad. Petrop. 1788.

(12) A beam PQ, (fig. 86), rests against a smooth vertical plane AB and a smooth curve AP; to find the nature of the curve that the beam may be at rest in all positions.

Let G be the centre of gravity of the beam; draw PM horizontal; let PQ = a, GP = c, AM = x, PM = y; then the equation to the curve will be

$$\frac{(c-x)^2}{c^2} + \frac{y^2}{a^2} = 1,$$

which is the equation to an ellipse, the centre of which coincides with Q when PQ is horizontal.

(13) A uniform beam AB, (fig. 87), rests upon a smooth horizontal plane Ca, and against a smooth vertical plane Cb; a string ACP is attached to the end A of the beam, and hangs through a small ring at C, with a weight P at its extremity; to find the position of the beam when at rest.

If W denote the weight of the beam, and θ the angle BAC, then

 $\tan \theta = \frac{W}{2P}.$

(14) A plane figure, bounded by a parabola, rests in a vertical plane, on two points in the same horizontal line, the centre of gravity of the figure being in the axis of the parabola at a given distance from the vertex; to find the position of equilibrium.

Let 2a be the distance between the two points, 4m the latus rectum, h the distance of the centre of gravity from the vertex,

and θ the inclination of the axis to the vertical in the position of equilibrium; then the equation

$$\sin \theta \left\{ 3a^2 \cos^4 \theta - 4m(h - m) \cos^2 \theta + 4m^2 \right\} = 0$$
 will give the positions of equilibrium.

(15) A particle is attracted towards each of two fixed centres of force varying inversely as the square of the distance; to find the equation to the surface on which it may remain at rest in every position.

If μ , μ' , be the absolute forces of attraction; r, r', simultaneous distances of the particle from the centres; and a, a', given values of r, r'; then the equation to the surface will be

$$\frac{\mu}{r} + \frac{\mu'}{r'} = \frac{\mu}{a} + \frac{\mu'}{a'}.$$

(16) To the extremity B of a rod AB, (fig. 88), which is able to revolve freely about A, is attached an indefinitely fine thread BCM, passing over a point C vertically above A, and sustaining a heavy particle at M on a smooth curve CMN in the vertical plane BAC; to determine the nature of the curve that for all positions of the rod and particle the system may be in equilibrium.

Let AB = 2a, AC = b, l = the length of the thread BCM, $\rho =$ the straight line CM, $\theta = \angle ACM$, m = the mass of the particle, m' = the mass of the rod. Then the equation to the curve will be

$$m'\rho^2 + 2(2mb\cos\theta - m'l)\rho = c,$$

where c is a constant quantity.

This problem was proposed by Sauveur to L'Hôpital, by whom a solution was published in the *Acta Eruditorum*, 1695, Febr. p. 56. The curve was shewn by John Bernoulli, *Ib.* p. 59, to be an Epitrochoid. See also James Bernoulli, *Ib.* p. 65.

SECT. 2. Stability and Instability of Equilibrium.

(1) AB (fig. 89) is a beam moveable about a hinge A; C is a small pulley in the vertical line through A, AC being equal to AG where G is the centre of gravity of AB; a fine string is attached to G which passes over C and has a weight P

suspended by it; to find the stable and unstable positions of equilibrium of the beam.

Let GA = CA = a, l = the length of the string GCP, W = the weight of the beam AB, $\angle GCA = \theta$; x = the vertical distance of the centre of gravity of P and the beam below the horizontal line through C.

Now, from the geometry,

$$CP = l - 2a \cos \theta;$$

and the distance of G below the horizontal line through C, is

$$a + a \cos 2\theta$$
;

hence, by the property of the centre of gravity of bodies,

$$(P+W) x = P(l-2a\cos\theta) + Wa(1+\cos 2\theta).$$

Now for equilibrium x must have a maximum or a minimum value; hence evidently

$$u = W \cos 2\theta - 2P \cos \theta$$

must have a maximum or minimum value; therefore

$$\frac{du}{d\theta} = -2 W \sin 2\theta + 2P \sin \theta = 0,$$

and therefore $\sin \theta (2 W \cos \theta - P) = 0$;

hence for equilibrium it is necessary that $\sin \theta = 0$, and there-

fore
$$\theta = 0$$
, or $\cos \theta = \frac{P}{2W}$.

Differentiating u a second time, we get

$$\frac{d^2u}{d\theta^2} = -4W\cos 2\theta + 2P\cos \theta;$$

if
$$\theta = 0$$
, we have $\frac{d^2u}{d\theta^2} = -4W + 2P$;

hence $\frac{d^2u}{d\theta}$ will be positive or negative, and therefore u a mini-

mum or maximum according as P is greater or less than 2W; hence if P be greater than 2W, $\theta = 0$ gives a position of unstable equilibrium, and if P be less than 2W, one of stability.

Again, if
$$\cos \theta = \frac{P}{2W}$$
, we shall have

$$\frac{d^2u}{d\theta^2} = -4W(2\cos^2\theta - 1) + 2P\cos\theta = \frac{4W^2 - P^2}{W};$$

if then 2 W be greater than P, $\frac{d^2u}{d\theta^2}$ is positive, and therefore the altitude of the centre of gravity of P and the beam is a maximum, and therefore the position will be one of unstable equilibrium; if 2W be less than P, $\cos \theta$ will be impossible, or the only position of equilibrium will be the unstable one given by $\theta = 0$.

(2) A uniform beam PQ (fig. 90) is placed upon two smooth inclined planes AB, AC; to find whether its position of equilibrium is one of stability or of instability.

Let G be the centre of gravity of the beam; from P and G draw PM, GH, at right angles to the horizontal plane bActhrough A. Let $\angle BAb = a$, $\angle CAc = \beta$, PG = QG = a, $\theta = \text{the}$ angle of inclination of PQ to the horizon, GH = z. Then, by the geometry,

$$z = a \sin \theta + PM = a \sin \theta + AP \sin \alpha$$

$$= a \sin \theta + \sin \alpha \cdot 2a \frac{\sin (\beta - \theta)}{\sin (\alpha + \beta)}$$

$$= \frac{a}{\sin (\alpha + \beta)} \{ \sin (\beta - \alpha) \sin \theta + 2 \sin \alpha \sin \beta \cos \theta \};$$

then, if z be a maximum or minimum,

$$u = \sin (\beta - a) \sin \theta + 2 \sin \beta \sin a \cos \theta$$

will be a maximum or minimum; hence

$$\frac{du}{d\theta} = \sin (\beta - a) \cos \theta - 2 \sin \beta \sin a \sin \theta = 0;$$

and therefore, for equilibrium,
$$\tan \theta = \frac{\sin (\beta - a)}{2 \sin \beta \sin a},$$

a positive quantity, if, as we will suppose, β be greater than a. Differentiating u a second time, we have

$$\frac{d^2u}{d\theta^2} = -\sin(\beta - a)\sin\theta - 2\sin\beta\sin\alpha\cos\theta;$$

from this it appears that, since θ is clearly less than $\frac{\pi}{2}$, $\frac{d^2u}{d\theta^2}$ will be negative, or that in the position of equilibrium the centre of gravity is at its maximum altitude; hence the equilibrium will be unstable.

(3) A uniform slender rod, acted on by gravity, is placed with its extremities against two planes (one horizontal and the other vertical), having at a point in their intersection an attractive force, varying inversely as the square of the distance, which at the centre of gravity of the rod is equal to half the force of gravity; to find the position of equilibrium of the rod, and to ascertain whether it is stable or unstable.

Let AB (fig 91) be the rod, G its centre of gravity, P any point in it; join OP, O being the centre of attraction; draw PM at right angles to the horizontal line OA. Let AG = a = BG, AP = s, OP = r, PM = y, $\angle OAB = \theta$; let the mass of a unit of the rod's length be taken as the unit of mass.

Then the attraction on an element δs of the rod at P will be equal to $g\delta s$ vertically downwards, and to $\frac{1}{2}g\frac{a^2}{r^2}\delta s$ towards the centre O. Hence, adopting the notation which was employed above in the enunciation of Maupertuis' Principle,

$$\Pi = \delta^{-1} d^{-1} \left\{ g \delta s \ dy + \frac{1}{2} g \frac{a^2}{r^2} \delta s \ dr \right\},\,$$

and therefore $d\Pi = g\delta^{-1}(\delta s \, dy) - \frac{1}{2} a^2 g \, d\delta^{-1} \, \frac{\delta s}{r}$.

Now, by the geometry,

$$y = s \sin \theta$$
, $dy = s \cos \theta d\theta$;

hence $\delta^{-1}(\delta s \, dy) = \delta^{-1}(s \delta s \cos \theta \, d\theta) = \delta^{-1}(s \delta s) \cos \theta \, d\theta$, and consequently, the limits of the integration being obviously 0, 2a, $\delta^{-1}(\delta s \, dy) = 2a^2 \cos \theta \, d\theta$.

Again, by the geometry, we see that

$$r^2 = s^2 - 4a \cos^2 \theta \ s + 4a^2 \cos^2 \theta;$$

hence we have

$$\delta^{-1} \frac{\delta s}{r} = \delta^{-1} \frac{\delta s}{\left(s^2 - 4as\cos^2\theta + 4a^2\cos^2\theta\right)^{\frac{1}{3}}}$$

= $C + \log \{s - 2a \cos^2 \theta + (s^2 - 4as \cos^2 \theta + 4a^2 \cos^2 \theta)^{\frac{1}{3}}\}$, and therefore, between the limits s = 0, s = 2a,

$$\delta^{-1} \frac{\delta s}{r} = \log \frac{\sin \frac{\theta}{r} (1 + \sin \frac{\theta}{r})}{\cos \frac{\theta}{r} (1 - \cos \frac{\theta}{r})} = \log \frac{\tan \frac{1}{4} (\pi + 2\theta)}{\tan \frac{1}{4} \theta}.$$

$$d\delta^{-1} \frac{\delta s}{r} = \frac{\frac{1}{2} \sec^2 \frac{1}{4} (\pi + 2\theta)}{\tan \frac{1}{4} (\pi + 2\theta)} - \frac{\frac{1}{2} \sec^2 \frac{1}{2} \theta}{\tan \frac{1}{2} \theta}$$
$$= \frac{1}{\cos \theta} - \frac{1}{\sin \theta}.$$

Putting for $\delta^{-1}(\delta s dy)$ and $d\delta^{-1} \frac{\delta s}{r}$ their values in the expression for $d\Pi$, we get

$$\frac{d\Pi}{d\theta} = \frac{1}{2} a^2 g \left(-\frac{1}{\cos \theta} + \frac{1}{\sin \theta} + 4 \cos \theta \right).$$

Now there will be equilibrium if II have a maximum or a minimum value, and therefore if

$$\sin \theta - \cos \theta - 4 \sin \theta \cos^2 \theta = 0$$
;

multiplying this equation by $\cos \theta + \sin \theta$, a quantity which cannot be equal to zero from $\theta = 0$ to $\theta = \frac{1}{2} \pi$, we get

$$-\cos 2\theta - (1 + \cos 2\theta)(1 + \sin 2\theta - \cos 2\theta) = 0,$$

$$\cos 2\theta + \sin 2\theta + \sin 2\theta (\cos 2\theta + \sin 2\theta) = 0,$$

$$(1 + \sin 2\theta)(\cos 2\theta + \sin 2\theta) = 0;$$

but it is evident that $1 + \sin 2\theta$ cannot become zero for any value of θ from 0 to $\frac{1}{2}\pi$; hence

$$\cos 2\theta + \sin 2\theta = 0$$
, $\tan 2\theta = -1$, $\theta = \frac{3}{8}\pi$,

which determines the position of equilibrium.

Again, differentiating the expression for $\frac{d\Pi}{d\theta}$, we have

$$\frac{d^2\Pi}{d\theta^2} = -\frac{1}{2}a^2g\left(\frac{\sin\theta}{\cos^2\theta} + \frac{\cos\theta}{\sin^2\theta} + 4\sin\theta\right),\,$$

which is evidently a negative quantity when $\theta = \frac{3}{8}\pi$; hence, for this value of θ , Π receives a maximum value, and therefore the equilibrium is one of instability.

(4) A square board hangs in a vertical plane by a string, which passing over a smooth nail has its ends fastened to two points symmetrically situated in one edge of the board. To investigate the positions and circumstances of equilibrium.

Let G (fig. 92) be the centre of gravity of the board, KCL the string passing over the nail C and attached to the board at

the points K, L; draw GH at right angles to KL; let fall KM, HQ, GR, LN, at right angles to the horizontal line passing through C in the plane of the board.

Let KL = a, c = the length of a side of the square, l = the length of the string, $CK = \lambda$, $CL = \lambda'$; ϕ , ϕ' , the inclinations of CK, CL, to the vertical, $\angle HGR = \theta$.

Then, from the geometry, putting $RG = \frac{1}{2}u$, we have

$$u = c \cos \theta + \lambda \cos \phi + \lambda' \cos \phi' \dots (1),$$

$$a \cos \theta = \lambda \sin \phi + \lambda' \sin \phi' \dots (2)$$

$$a \sin \theta = \lambda \cos \phi - \lambda' \cos \phi' \dots (3),$$

$$\lambda + \lambda' = l \dots (4).$$

Thus we have four equations connecting the six variables u, θ , λ , λ' , ϕ , ϕ' , and we may express u in terms of two independent variables λ , θ . We must proceed then to determine the maximum or minimum values of u, as a function of two independent variables.

Differentiating (1), (2), (3), with respect to λ , considering θ as constant, and observing that, from (4), $\frac{d\lambda'}{d\lambda} = -1$, we get, as conditions for a maximum or minimum value of u,

$$0 = \frac{du}{d\lambda} = \cos \phi - \cos \phi' - \lambda \sin \phi \frac{d\phi}{d\lambda} - \lambda' \sin \phi' \frac{d\phi'}{d\lambda} \dots (5),$$

$$0 = \sin \phi - \sin \phi' + \lambda \cos \phi \frac{d\phi}{d\lambda} + \lambda' \cos \phi' \frac{d\phi'}{d\lambda} \dots (6),$$

$$0 = \cos \phi + \cos \phi' - \lambda \sin \phi \frac{d\phi}{d\lambda} + \lambda' \sin \phi' \frac{d\phi'}{d\lambda} \dots (7).$$

From (5) and (7), by addition and subtraction,

$$\lambda \sin \phi \frac{d\phi}{d\lambda} = \cos \phi \dots (8),$$

$$\lambda' \sin \phi' \frac{d\phi'}{d\lambda} = -\cos \phi' \dots (9);$$

hence, from (6), it may easily be shewn that $\phi = \phi'$.

Again, differentiating (1), (2), (3), with respect to θ , λ being considered constant, and putting $\frac{du}{d\theta} = 0$, we get

$$0 = \frac{du}{d\theta} = -c \sin \theta - \lambda \sin \phi \frac{d\phi}{d\theta} - \lambda' \sin \phi' \frac{d\phi'}{d\theta} \dots (10),$$

$$-a \sin \theta = \lambda \cos \phi \frac{d\phi}{d\theta} + \lambda' \cos \phi' \frac{d\phi'}{d\theta} \dots (11),$$

$$a \cos \theta = -\lambda \sin \phi \frac{d\phi}{d\theta} + \lambda' \sin \phi' \frac{d\phi'}{d\theta} \dots (12).$$
From (10) and (12), since $\phi = \phi'$, we have
$$2\lambda \sin \phi \frac{d\phi}{d\theta} = -a \cos \theta - c \sin \theta \dots (13),$$
and
$$2\lambda' \sin \phi \frac{d\phi'}{d\theta} = a \cos \theta - c \sin \theta \dots (14);$$
hence, from (11), there is, effecting simplifications,
$$\sin \theta \left\{ a \tan \phi - c \right\} \dots (15);$$
also from (2), putting $\phi' = \phi$, $\lambda + \lambda' = l$, we have
$$a \cos \theta = l \sin \phi$$
, $(l^2 - a^2 \cos^2 \theta)^{\frac{1}{2}} \tan \phi = a \cos \theta$;
hence, from (15), for the positions of equilibrium,

 $\sin \theta \left\{ a^2 \cos \theta - c \left(l^2 - a^2 \cos^2 \theta \right)^{\frac{1}{4}} \right\} = 0.$

From this equation we get

$$\theta = 0$$
, $\theta = \pm \cos^{-1} \frac{cl}{u(a^2 + c^2)^{\frac{1}{2}}}$.

Thus, if l be less than $\frac{a}{c} (a^2 + c^2)^{\frac{1}{2}}$, there will be three positions of equilibrium; and, if it be greater, only one. In the former case it may be ascertained that, when $\theta = 0$, $\frac{d^2u}{d\lambda^2}$, $\frac{d^2u}{d\theta^2}$, are both positive, and in the latter case, both negative; and that, in both cases, $\frac{d^2u}{d\lambda^2} \cdot \frac{d^2u}{d\theta^2}$ is greater than $\left(\frac{d^2u}{d\lambda d\theta}\right)^2$. Hence, in the former case, $\theta = 0$ corresponds to unstable, and, in the latter, to stable equilibrium. Thus, since positions of stable and unstable equilibrium recur alternately, there will either be three positions of equilibrium of which two are stable and one unstable, or only a single position of stable equilibrium.

(5) Two heavy particles, connected together by a thread PAQ (fig. 93) passing over the convex side of a circle situated in a vertical plane, balance each other when placed at P and Q; to determine the position of P, Q, and to ascertain whether the equilibrium is stable or unstable, the weight of the thread being neglected.

Let O be the centre of the circle, OA a vertical radius; $\angle POQ = a$, $\angle POA = \theta$, $\angle QOA = \phi$; and let m, n, denote the masses of the particles. Then we shall have for the equilibrium, which will be unstable, the equation

$$\tan \frac{1}{2} (\phi - \theta) = \frac{m-n}{m+n} \tan \frac{1}{2} \alpha.$$

(6) A uniform rod passes through a hole in a spherical shell, and rests with one end against the internal surface, the length of the rod being equal to twice that of the diameter; having given the inclination of the rod to the vertical when it is in a position of stable equilibrium, to determine its inclinations to the vertical when in its positions of unstable equilibrium.

If a denote its inclination to the vertical when in its position of stable equilibrium, then its inclinations for its two positions of unstable equilibrium will be

$$\frac{1}{3}(\pi+a)$$
 and $\frac{1}{3}(\pi-a)$.

(7) A particle is placed in a position of equilibrium between two centres of attractive force, varying according to any power of the distance; to determine for what laws of force the equilibrium is stable and for what unstable.

The equilibrium will be stable or unstable according as the forces attract in direct or inverse powers respectively.

DYNAMICS.

CHAPTER I.

IMPACT AND COLLISION. SMOOTH SPHERICAL BODIES.

Conceive two spherical bodies, which are composed of the same material, to be moving in the same straight line, namely, in the line joining their centres, and at any time during their motion to impinge against each other. Let m, m', denote the masses of the two bodies; and let u, u', denote their velocities before and v, v', their velocities after collision; velocities being reckoned positive in one direction and negative in the other; then whatever be the absolute or relative magnitudes of the velocities u, u', or of the masses m, m',

$$u - u' : v' - v :: 1 : e,$$

or $v' - v = e(u - u') \dots (A),$

where e is a numerical quantity not greater than unity, which is invariable while the material of the bodies remains the same, but which changes generally with a change in their substance. The bodies are said to be inelastic if e be equal to zero; imperfectly elastic if it be equal to any fraction between zero and unity; and perfectly elastic if it be equal to unity.

The theory of collision furnishes us likewise with the following general relation,

$$m(u-v) = m'(v'-u') \dots (B).$$

The signification of the equation (A) is, that the relative velocity of the two bodies after collision bears a constant ratio to the relative velocity before collision, so long as the material of the bodies remains unchanged; and the equation (B) implies, that the momentum which one body gains by the collision in the positive direction of motion, is equal to that which the other loses. These are the two fundamental principles in the theory of collision.

Suppose that u' is equal to zero, and that m is inconsiderable in comparison with m'; then clearly, by (A) and (B),

$$v'-v=eu$$
 and $v'=u'=0$, and therefore $v=-eu$(C),

or the small body is reflected backwards with a velocity which is to the velocity of impact as e to 1; while the large body experiences no appreciable motion from the collision. This is evidently the case of bodies impinging and rebounding upon the surface of the earth, or upon other bodies firmly attached to it, the earth being regarded as stationary.

In the year 1639, J. Marc Marci de Crownland, a Hungarian physician, published at Prague a work entitled De Proportione Motus, seu Regula Sphymica, in which he has treated of the collision of perfectly elastic and perfectly inelastic bodies. He occupies himself principally with the consideration of perfectly elastic bodies, and lays down precisely the same rules for their collision which are now commonly adopted. This work, the earliest in which the theory of collision had been correctly propounded, having fallen into general oblivion in the scientific world, the subject was again correctly investigated by the independent efforts of Wallis, Wren, and Huvghens, who apparently had not the slightest knowledge even of the existence of the work by Marci. The laws of the collision of perfectly inelastic bodies were laid down by Wallis, Phil. Trans. 1668, p. 864, and of perfectly elastic bodies by Wren, Phil. Trans. 1668, p. 867, and Huyghens, Phil. Trans. 1669, p. 925, and Journal des Scarans of March 18, 1669. Wren and Lawrence Rook had, several years earlier than this, exhibited various experi-

¹ Montuela; Histoire des Mathématiques, tom. 11. p. 406.

ments before the Royal Society, in illustration of the principles of collision. The conclusions of Wallis, Wren, and Huyghens, which had been presented to the Royal Society in a very brief shape, were afterwards given more at large by Wallis, Mechanica, Pars Tertia, 1671; Keill, Introductio ad Veram Physicam, Lect. 12, 13, 14; and Mariotte, Traité de Percussion. There are some ingenious experiments by Smeaton on the theory of collision in the *Phil. Trans.*, April 18, 1782. The principles of the collision of imperfectly elastic bodies were first propounded by Newton, *Principia*, Lib. 1., Scholium to the Laws of Motion, who inferred experimentally the truth of the equation (A) for any value whatever of e between zero and unity; preceding philosophers having directed their attention to those cases alone in which e is supposed to be either zero or unity. The physical value of Newton's generalization is the more striking when it is considered that natural bodies are never actually endowed with perfect elasticity. For the mathematical formulæ in the theory of the collision of imperfectly elastic spheres, the reader is referred to Maclaurin, Choc des Corps, Prix de l'Academie, tom. 1. and to Bossut, Cours de Mathématique, tom. 111. The results of a series of experiments on the elasticity of bodies, by Mr. Hodgkinson, are to be found in vol. III. p. 534, of the Reports of the British Association for the Advancement of Science, where he has shewn that the quantity e, in the equation (A), is not, as we stated, and as we shall suppose for the sake of mathematical simplicity, entirely independent of the velocities of the impinging bodies, as Newton had concluded, but that it decreases as the relative velocity increases, assuming however a nearly constant value when the relative velocity of collision becomes considerable.

(1) Two inelastic bodies are moving in opposite directions with given velocities; to find their velocities after collision.

Let m, m', denote the masses of the bodies; a, a', their velocities before collision. Then, putting in the formulæ (A) and (B), the direction of the motion of m being taken as the positive one, u = a, u' = -a', e = 0,

we have
$$v' - r = 0$$
, $m(a - v) = m'(v' + a')$,

and therefore

$$m (a - v) = m' (a' + v),$$

$$v' = v = \frac{ma - m'a'}{m + m'}.$$

If then ma be greater than m'a', the bodies will, after collision, move along in the positive direction with a common velocity (ma - m'a') : (m + m'); and if ma be less than m'a', they will move in the negative direction with a common velocity (m'a' - ma) : (m' + m). If ma be equal to m'a', the collision will reduce both the bodies to rest.

Wallis; Mechan. Pars Tertia, de Percussione, Prop. IV.

(2) Two perfectly elastic bodies are moving in opposite directions with given velocities; to find their velocities after collision.

The notation being the same as in the preceding example, we have, e being in this case equal to unity,

$$v'-v=a+a', \quad m(a-v)=m'(a'+v').$$

Eliminating v' from these two equations,

$$v = \frac{ma - m'a}{m + m'} - \frac{2m'a'}{m + m'}.$$

Eliminating v, we have

$$v'=\frac{ma'-m'a'}{m+m'}+\frac{2ma}{m+m'}.$$

Wallis; Ib. de Elatere et Resilitione, Prop. x.

(3) Three bodies m, m', m'', are placed in a row. The body m receiving a given velocity towards m', to find the magnitude of m' that the velocity communicated to m'' by its intervention may be the greatest possible.

Let a be the velocity with which m is projected; a' the velocity which m' acquires on being struck by m, and a'' that which m'' receives on being struck by m'. Then

$$a' = \frac{2ma}{m+m'}, \quad a'' = \frac{2m'a'}{m'+m''},$$

and therefore

$$a'' = \frac{4mm'a}{(m+m')(m'+m'')}$$
.

Since a is to be a maximum, we must have

$$\left(\frac{m}{m'}+1\right)\left(m'+m'\right)$$
 a minimum;

hence, differentiating with respect to the variable m',

$$\frac{m}{m'} + 1 - \frac{m}{m'^2}(m' + m'') = 0$$

$$m'^2 - mm'' = 0, \quad m' = (mm'')^{\frac{1}{4}}.$$

Huyghens; Phil. Trans. 1669, p. 928. Wolff; Elementa Matheseos Universæ, tom. II. p. 158.

(4) A perfectly elastic sphere impinges with a given velocity and in a given direction against a smooth plane; to determine the velocity and direction of reflection.

Let u, v, denote the velocities of incidence and of reflection, and a, β , the angles which the directions of the motion before and after impact make with a normal to the plane.

The resolved parts of the velocities parallel to the plane are $u \sin a$ and $v \sin \beta$, and, at right angles to it, $u \cos a$ and $v \cos \beta$. But the plane being perfectly smooth will not affect the resolved part of the velocity parallel to itself, and therefore

$$v \sin \beta = u \sin \alpha$$
;

while the other resolved part of the incident velocity will be affected as if the impact had been direct, and therefore, by (C),

$$v \cos \beta = u \cos \alpha$$
.

From these two equations it is evident that

$$\tan \beta = \tan \alpha$$
, $\beta = \alpha$, and $v = u$,

or the angle of reflection is equal to that of incidence, and the velocity of reflection to the velocity of incidence.

Wallis; Mechan. Pars Tertia, De Elatere, &c. Prop. 11.

(5) Two smooth spheres moving with given velocities and in any given directions whatever, impinge against each other; the spheres being supposed perfectly elastic, to determine their velocities and the directions of their motions after collision.

Let AB, A'B', (fig. 94), be the directions of the motion of the two bodies before collision; and O, O, the positions of their

centres at the instant of contact; produce O' O indefinitely to a point C. Let a, a', denote their velocities before collision, and a, a', the angles AOC, A'O'C.

Then the resolved parts of the velocities of the spheres O, O', in the direction COO' will be $a \cos a$, $a' \cos a'$, and at right angles to OO' in the planes AOC, A'O'C, respectively, $a \sin a$, $a' \sin a'$. These latter resolved velocities will not be affected by the collision. The former will be affected exactly as if the sphere O moving along COO' with a velocity $a \cos a$ were to impinge directly upon the sphere O' moving with a smaller velocity $a' \cos a'$ estimated in the same direction. Hence if v, v', denote the resolved parts of the velocities after collision parallel to the line COO', we have, as may be readily ascertained by the principles of this chapter,

$$v = \frac{ma \cos a - m'a \cos a}{m + m'} + \frac{2m'a' \cos a'}{m + m'},$$

$$v' = \frac{m'a' \cos a' - ma' \cos a'}{m + m'} + \frac{2ma \cos a}{m + m'}.$$

Let V, V, denote the velocity of the spheres O, O', after collision, and ϕ , ϕ' , the angles which the directions of their motions make with OO'; then

$$V^2 = v^2 + a^2 \sin^2 a$$
, $V'^2 = v'^2 + a'^2 \sin^2 a'$,
 $\tan \phi = \frac{a \sin a}{v}$, $\tan \phi' = \frac{a' \sin a'}{v'}$,

their motions still taking place in the planes BOO', B'O'O.

Keill; Introductio ad Veram Physicam, Lect. 14.

(6) Two imperfectly elastic bodies are moving in the same direction along the same straight line with given velocities; the one overtakes the other and collision ensues; to find the velocities of the two bodies after collision.

If m, m', be the masses of the two bodies, e their common elasticity; a, a', their velocities before, and v, v', their velocities after collision;

$$v = \frac{ma + m'a'}{m + m'} - \frac{em'(a - a')}{m + m'},$$

$$v' = \frac{ma + m'a'}{m + m'} + \frac{em(a - a')}{m + m'}.$$

Maclaurin; Choc des Corps, p. 30, Prix de l'Académie, tom. 1.

(7) To find the sum of the vires vivæ of two perfectly elastic bodies after direct collision.

If a, a', be the velocities before, and v, v', after collision,

$$mv^2 + m'v'^2 = ma^2 + m'a'^2$$

Huyghens; De Motu Corporum ex Percuss. Prop. x1. John Bernoulli; Discours sur le Mouvement, chap. x.

(8) To find the sum of the vires vivæ of two imperfectly elastic bodies after direct collision.

The notation remaining the same as in the preceding example, and s denoting the elasticity,

$$mv^{2} + m'v'^{2} = ma^{2} + m'a'^{2} - \frac{(1 - e^{2})mm'(a - a')^{2}}{m + m'},$$

which shews that vis viva is lost by the collision.

(9) To find with what velocity a ball must impinge upon another equal ball moving with a given velocity, that the impinging ball may be reduced to rest by the collision, the common elasticity of the balls being known.

If e be the common elasticity, and a the velocity of the ball which is struck, the impinging ball must impinge with an opposite velocity equal to

$$\frac{1+e}{1-e}a.$$

(10) To determine the velocities of two bodies A and B of given elasticity and given masses moving in the same direction, that after collision A may remain at rest and B may move along with an assigned velocity.

If m, m', be the masses of A, B, e their elasticity, β the velocity which B is to have after collision, and a, b, the required velocities of A, B, before impact,

$$a = \frac{1+e}{e} \frac{m'\beta}{m+m'}, \quad b = \frac{(em'-m)\beta}{e(m+m')}.$$

Maclaurin; Choc des Corps, p. 52, Prix de l'Acad. tom. 1.

(11) A spherical body A impinges directly with a certain velocity upon a spherical body B at rest, the common elasticity of the two bodies being given; to find the mass of a third body which, moving with the velocity which A has before the collision, shall have the same momentum which B has after the collision.

If m, m', denote the masses of A, B; e the common elasticity of A, B; and m' the mass of the required body,

$$m'' = (1+e) \frac{mm'}{m+m'}.$$

(12) To find the elasticity of two spheres A and B, and the ratio between their masses, that when A impinges upon B at rest, A may be reduced to rest, and B move on with the nth part of A's velocity.

If m, m', denote the masses of A, B, and e their common elasticity, then

$$e=\frac{1}{n}, \quad \frac{m'}{m}=n.$$

(13) Two perfectly elastic spheres meet directly with equal velocities; to find the relation between their magnitudes, that after collision one of them may remain at rest.

If m, m', denote their masses, m' corresponding to the one which remains at rest,

(14) Any number of spheres of given elasticity being arranged in a straight line, and one of the extreme ones having a given velocity communicated to it so as to bring it into direct collision with the adjacent sphere of the series; to determine the velocity ultimately acquired by the last sphere.

If r be the number of the spheres, e their common elasticity, m_1 , m_2 , m_3 , m_r , their masses, and a the velocity with which the first is projected; then v being the velocity acquired at last by m_r ,

$$v = (1 + e)^r \frac{m_1}{m_1 + m_2} \cdot \frac{m_2}{m_2 + m_3} \cdot \frac{m_3}{m_3 + m_4} \cdot \cdot \cdot \cdot \cdot \cdot \frac{m_r}{m_r + m_{r+1}} \cdot a.$$

Maclaurin; Choc des Corps, p. 54, Prix de l'Académie, tom. 1.

(15) A number of equal spheres are placed on a smooth table in a straight line and close together; they are connected together by equal inelastic threads; a motion is given to the first in the direction of the line which they form so as to separate it from the second; to find the time which elapses before the last sphere is put in motion.

If n be the number of spheres, a the length of each of the connecting threads, and β the velocity with which the first sphere is projected, then

time required =
$$\frac{n(n-1)}{1 \cdot 2} \frac{a}{\beta}$$
.

(16) Between two spheres of given masses is placed a row of spheres; a velocity is communicated to one of the original spheres so as to bring it into direct collision with the nearest of the intermediate ones; to find the requisite magnitudes of the intermediate spheres in order that the velocity acquired by the last sphere may be the greatest possible, and to determine this velocity when their number becomes indefinitely great.

The intermediate spheres must be geometrical means between the two original ones; and if m, m', denote the masses of the original spheres, a the velocity communicated to m, and a' that acquired by m' when the number of the intermediate spheres becomes infinite,

$$a' = \left(\frac{m}{m'}\right)^{\frac{1}{4}} a.$$

(17) An imperfectly elastic sphere impinges upon a plane; to find the angles of incidence and of reflection, that the velocity before may be to the velocity after impact as $2^{\frac{1}{4}}$: 1, the elasticity being equal to $\frac{1}{2^{\frac{1}{4}}}$.

Angle of incidence = $\frac{1}{6}\pi$, angle of reflection = $\frac{1}{4}\pi$.

(18) An inelastic sphere A, moving with a given velocity, impinges upon an inelastic sphere B at rest, the line joining the centres of the two spheres at the instant of collision making a given angle with the direction of A's motion; to determine the velocity of A after collision.

If a be the given angle; m, m', the masses of the spheres A, B; and a, v, the velocities of A before and after collision,

$$v = a \left\{ \sin^2 a + \frac{m^2}{(m+m')^2} \cos^2 a \right\}^{\frac{1}{4}}.$$

(19) A sphere A in motion is struck by an equal one B moving with the same velocity, and in a direction making an angle a with that in which A is moving, in such a manner that the line joining their centres at the time of impact is in the direction of B's motion; to find the velocities of the spheres after impact, and to determine for what value of a that of A will be a maximum, the common elasticity of the spheres being supposed to be known.

If a denote the velocity of each of the spheres A, B, before, and u, v, their respective velocities after impact, then, e being their common elasticity,

$$u^{2} = \frac{1}{4} a^{2} \left\{ 1 + e + (1 - e) \cos a \right\}^{2} + a^{2} \sin^{2} a,$$

$$v^{2} = \frac{1}{4} a^{2} \left\{ 1 + e - (1 - e) \cos a \right\}^{2}.$$

When u is a maximum, $\cos a = \frac{1-e}{3-e}$.

(20) Three perfectly elastic spheres A, B, C, are placed at the three angles of a plane triangle of which the angles are known; to compare the magnitudes of the spheres, when A impinging obliquely upon B is reflected so as to strike C, and thence reflected to its first position; the lines joining the centres of the spheres A, B, and A, C, during collision, being respectively perpendicular to the opposite sides of the triangle, and their diameters being inconsiderable in comparison with the sides of the triangle.

If m, m', m'', denote the masses of the spheres A, B, C; and a, β , γ , the angles of the triangle at which A, B, C, are placed,

$$\frac{m'}{m} = \frac{\sin \beta}{\sin (\alpha - \gamma)}, \quad \frac{m''}{m} = \frac{\sin \gamma}{\sin (\beta - \alpha)}.$$

(21) A sphere A (fig. 95) moving in the direction EAF with an assigned velocity impinges upon a sphere B at rest, the two spheres having the same elasticity; supposing AF to be the

direction of A's motion after impact, and KABL to be a straight line passing through the centres of the two spheres at the time of collision, to find the value of the angles EAK and FAF when the latter angle has its greatest value.

If *n* be the ratio of the mass of *A* to that of *B*, *e* be the common elasticity of the spheres, $\angle EAK = \theta$, $\angle FAF' = \phi$; then

$$\tan \theta = \left(\frac{n-e}{n+1}\right)^{\frac{1}{4}}, \tan \phi = \frac{\frac{1}{2}(1+e)}{\left\{\left(n+1\right)\left(n-e\right)\right\}^{\frac{1}{4}}}.$$

CHAPTER II.

RECTILINEAR MOTION OF A PARTICLE.

THE determination of the circumstances of the motion of a material particle, which moves in a straight line under the action of a finite accelerating or retarding force, depends upon the two following differential equations, called the equations of motion of the particle

$$\frac{dx}{dt}=v,\,\frac{dv}{dt}=f,$$

where t denotes the time of the motion reckoned from an assigned epoch, x the distance of the particle at the end of this time from an assigned point in the line of its motion, v the velocity, and f the accelerating or retarding force.

From these two equations we readily deduce the two following,

$$\frac{d^2x}{dt^2} = f, \quad v \frac{dv}{dx} = f.$$

These equations, which constitute the complete expression of the circumstances of rectilinear motion in the language of the differential calculus for every condition of acceleration or retardation, are due to Varignon, and were published in the *Mém. de l'Acad. des Sciences de Paris*, 1700, p. 22. It may be observed however that, long before this, geometrical investigations of rectilinear motion for variable forces had been given by Newton.¹

From the formula vdv = fdx we see that dv^2 varies as fdx: an opinion however was expressed by Daniel Bernoulli, that there is no reason to consider this the only possible law of variation; for instance, that we might as well have $dv^n \propto fdx$, n being any quantity whatever. In opposition to Bernoulli's suggestion, Euler endeavoured to prove that the law of the square of the

¹ Principia, lib. 1. sect. 7; lib. 11. sect. 1.

^a Comment. Petrop. 1727, p. 136.

Mechanica, tom. 1. p. 62 et seq.

velocity is necessarily true; and D'Alembert shewed the truth of this law to depend simply upon the definition of the meaning of the symbol f.

The complete solution of a problem in rectilinear motion consists in the determination of relations between every two of the quantities x, v, f, t: now the general equations of rectilinear motion furnish us with only two independent relations between these four quantities; it is evident then that the data in every problem must consist in the expression of some particular equation, $\phi(x, v, f, t) = 0$ between x, v, f, t, so that we may have, in all, three equations connecting the four variables.

The function $\phi(x, v, f, t)$ may involve two, three, or all of the quantities x, v, f, t; and by the theory of combinations it is evident that there will be six varieties of the first, and four of the second class; hence the general problem of rectilinear motion resolves itself into eleven distinct classes of problems. We shall however confine ourselves to the consideration of those two classes in which the given function involves either x, f, alone; or x, f, v, alone: under the former head we shall exemplify the motion of a particle in vacuo; under the latter, in a resisting medium. The other classes are devoid of any physical interest.

SECT. 1. Motion in Vacuo.

(1) A particle is placed at a centre of repulsive force which varies as any power of the distance; to determine its velocity after receding to any distance from the centre, and the time of the motion.

Let μ represent the absolute force, x the distance of the particle from the centre of force after a time t, and v the velocity. Then, for the motion, we have

$$v \frac{dv}{dx} = \mu x^n.$$

Integrating with respect to x, and bearing in mind that v = 0 when x = 0, we have

$$\frac{1}{2} v^2 = \mu \int_0^a x^n \ dx = \frac{\mu}{n+1} x^{n+1},$$

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and therefore

$$v^2 = \frac{2\mu}{n+1} x^{n+1},$$

which gives the velocity for any value of x.

Again,
$$\frac{dx}{dt} = v = \left(\frac{2\mu}{1+n}\right)^{\frac{1}{2}} x^{\frac{1}{2}(n+1)}$$
:

hence, t being equal to zero when x = 0, there is

$$t = \left(\frac{1+n}{2\mu}\right)^{\frac{1}{2}} \int_0^{\pi} x^{-\frac{1}{2}(n+1)} dx$$
$$= \frac{2}{1-n} \left(\frac{1+n}{2\mu}\right)^{\frac{1}{2}} x^{\frac{1}{2}(1-n)}.$$

Euler; Mechanica, tom. 1. p. 123.

(2) A particle being attracted by a force varying inversely as the n^{th} power of the distance, to find the value of n when the velocity acquired from an infinite distance to a distance a from the centre is equal to the velocity which would be acquired from a to $\frac{1}{4}$ a.

Let μ denote the absolute force, x the distance of the particle from the centre of force after a time t, and v_1 , v_2 , the two velocities. Then, for the motion of the particle,

$$v \frac{dv}{dx} = -\frac{\mu}{x^n}.$$

Hence, for the former motion, v being equal to zero when $x = \infty$,

$$v_1^2 = -2\mu \int_{\infty}^a \frac{1}{x^n} dx = \frac{2\mu}{(n-1) a^{n-1}},$$

and, for the latter motion, since v = 0 when x = a,

$$v_2^2 = -2\mu \int_a^{\frac{1}{4}a} \frac{1}{x^n} dx = \frac{2\mu}{n-1} \left(\frac{4^{n-1}}{a^{n-1}} - \frac{1}{a^{n-1}} \right).$$

But by hypothesis v_1^2 is equal to v_2^2 ; hence

$$\frac{2\mu}{n-1} \frac{1}{a^{n-1}} = \frac{2\mu}{n-1} \frac{1}{a^{n-1}} (4^{n-1}-1),$$

and therefore

$$1 = 4^{n-1} - 1, 4^{n-1} = 2;$$

hence

$$n=\frac{3}{2}$$
.

(3) Four equal attractive forces are placed in the corners of a square, their intensity varying as any function of the distance; a particle is placed in one of the diagonals of the square very near to its centre; to find the time of an oscillation.

Let O be the centre of the square (fig. 96), E the position of the particle after any time t from the commencement of the motion; let OD = a, OE = x, AE = CE = r. Then, for the motion of the particle, taking the sum of the forces acting upon it in the line OD, we have

$$\frac{d^2x}{dt^2} = -2\phi(r)\frac{x}{r} + \phi(a-x) - \phi(a+x),$$

and therefore, neglecting powers of the small quantity x higher than the first, we get, by Taylor's theorem,

$$\frac{d^2x}{dt^2} = -2\left(\frac{\phi(a)}{a} + \phi'(a)\right)x,$$

or, putting the coefficient of x equal to -2k,

$$\frac{d^3x}{dx^2} = -2kx.$$

Multiply both sides of this equation by 2 $\frac{dx}{dt}$, and we get by integration

$$\frac{dx^2}{dt^2} = C - kx^2,$$

where C is an arbitrary constant.

Let β be the initial value of x; then, $\frac{dx}{dt}$ being initially equal

to zero, we have

$$0 = C - k\beta^2,$$

and therefore

$$\frac{dx^2}{dt^2} = k(\beta^2 - x^2)$$

$$dt = -\frac{1}{k^{\frac{1}{4}}} \frac{dx}{(\beta^2 - x^2)^{\frac{1}{4}}},$$

the negative sign being taken because x decreases as t increases.

Integrating, we get

$$t=\frac{1}{L^{\frac{1}{4}}}\cos^{-1}\frac{x}{\beta},$$

no constant being added because $x = \beta$ when t = 0; hence

$$x = \beta \cos(k^{\frac{1}{4}}t).$$

Now as soon as $k^{\frac{1}{4}}t$ becomes equal to π , x becomes equal to $-\beta$, its greatest negative value. Hence the time of a complete oscillation being T, we have

$$k^{\frac{1}{4}} T = \pi,$$

and therefore, substituting for k its value,

$$T = \pi \left\{ \frac{1}{a} \phi(a) + \phi'(a) \right\}^{-\frac{1}{4}}.$$

(4) A particle A attracts a particle B with a force always to that with which B attracts A in the ratio of μ' to μ ; the particles being originally at rest, to find their position as well as that of their centre of gravity after any time; the intensity of each force being directly as the distance between the particles.

Let O be a fixed point in the line of the motion of the particles, (fig. 97), and let OA = x, OB = x', at any time t.

Then, for the motion, we have

$$\frac{d^2x}{dt^2}=\mu(x'-x)\ldots(1),$$

$$\frac{d^2x'}{dt^2}=-\mu'(x'-x)\ldots(2).$$

Multiplying (1) and (2) by μ' and by μ respectively, and adding the results, we get

$$\mu' \frac{d^2x}{dt^2} + \mu \frac{d^2x'}{dt^2} = 0.$$

Integrating, and bearing in mind that $\frac{dx}{dt}$, $\frac{dx'}{dt}$, are both equal to

zero initially,

$$\mu' \frac{dx}{dt} + \mu \frac{dx'}{dt} = 0;$$

integrating again,

$$\mu'x + \mu x' = \mu'a + \mu a' \dots (3),$$

a, a', being the initial values of x, x'.

Again, subtracting (1) from (2),

$$\frac{d^2}{dt^2}(x'-x)+(\mu+\mu')(x'-x)=0,$$

and putting x' - x = z,

$$\frac{d^2z}{dt^2}+\left(\mu+\mu'\right)z=0;$$

multiplying by $2\frac{dz}{dt}$, and integrating,

$$\frac{dz^2}{dt^2} + (\mu + \mu')z^2 = C,$$

where C is an arbitrary constant.

But, initially, $\frac{dz}{dt} = 0$, and z = a' - a, and therefore

$$(\mu + \mu')(a' - a)^2 = C;$$
hence
$$\frac{dz^3}{dt^2} = (\mu + \mu')\left\{(a' - a)^2 - z^3\right\},$$

$$\frac{dt}{dz} = -\frac{1}{(\mu + \mu')^{\frac{1}{4}}} \cdot \frac{1}{\left\{(a' - a)^2 - z^2\right\}^{\frac{1}{4}}},$$

the negative sign being taken because z or x'-x decreases as t increases. Integrating, we have

$$t = \frac{1}{(\mu + \mu')^{\frac{1}{4}}} \cos^{-1} \frac{z}{a' - a} = \frac{1}{(\mu + \mu')^{\frac{1}{4}}} \cos^{-1} \frac{x' - x}{a' - a},$$

no constant being added because x' - x = a' - a when t = 0;

hence $x' - x = (a' - a) \cos \{(\mu + \mu')^{\frac{1}{4}} t\} \dots (4)$

From (3) and (4) we readily obtain

$$x = \frac{\mu' a + \mu a'}{\mu' + \mu} - \frac{\mu(a' - a)}{\mu' + \mu} \cos \left\{ (\mu' + \mu)^{\frac{1}{2}} t \right\},$$

$$x' = \frac{\mu' a + \mu a'}{\mu' + \mu} + \frac{\mu' (a' - a)}{\mu' + \mu} \cos \left\{ (\mu' + \mu)^{\frac{1}{2}} t \right\};$$

$$(m + m') \overline{x} = mx + m'x'$$

$$= \frac{m' + m}{\mu' + \mu} (\mu' a + \mu a') + \frac{\mu' m' - \mu m}{\mu' + \mu} (a' - a) \cos \left\{ (\mu' + \mu)^{\frac{1}{2}} t \right\},$$

where m, m', denote the masses of A, B, and \bar{x} the distance of their centre of gravity from O at any time t.

If μ' , μ , be proportional to m, m', respectively, then clearly from our general result

$$(m+m')\overline{x} = \frac{m'+m}{\mu'+\mu}(\mu'\alpha + \mu\alpha')$$
or $\overline{x} = \frac{\mu'\alpha + \mu\alpha'}{\mu'+\mu}$,

which shews that the centre of gravity remains stationary during the whole motion.

(5) A body not affected by gravity falls down the axis of a thin cylindrical tube infinite in length, the particles of which attract with a force which varies inversely as the square of the distance; to find the velocity acquired in falling through a given space.

Let k be the thickness of the tube, r the radius of its interior surface, x the distance of the particle P from the extremity of the tube after a time t; then the volume of a portion of the tube contained between slices at distances s and s + ds from P will be $2\pi rkds$, and therefore the attraction of this elemental portion on the particle along the axis of the tube will be, the unit of attraction being chosen to be the attraction of a unit of mass at a unit of distance,

 $2\pi \rho rkds \cdot \frac{1}{s^2+r^2} \cdot \frac{s}{(s^2+r^2)^{\frac{1}{4}}},$

where ρ denotes the density; and therefore for the motion of the particle we have

$$\frac{d^{2}x}{dt^{2}} = 2\pi\rho rk \int_{-\pi}^{+\infty} \frac{sds}{(s^{2} + r^{2})^{\frac{1}{4}}}$$

$$= 2\pi\rho rk \cdot \int_{-\pi}^{+\infty} \left\{ -\frac{1}{(s^{2} + r^{2})^{\frac{1}{4}}} \right\} = \frac{2\pi\rho rk}{(x^{2} + r^{2})^{\frac{1}{4}}}$$

multiplying by $2 \frac{dx}{dt}$ and integrating,

$$v^{2} = \frac{dx^{2}}{dt^{2}} = 4\pi\rho rk \log \left\{x + (x^{2} + r^{2})^{\frac{1}{2}}\right\} + C.$$

But v = 0 when x = 0; hence

$$0 = 4\pi \rho r k \log r + C,$$

and therefore

$$v^2 = 4\pi \rho r k \log \frac{x + (x^2 + r^2)^{\frac{1}{4}}}{r}.$$

(6) A particle is projected downwards with a certain volocity on a horizontal plane through a height of twenty feet; it rebounds ten feet and then falls again and rebounds four feet: to find the elasticity of the particle and the velocity of projection.

Let a denote the velocity of projection, and e the elasticity; then if v denote the velocity at the end of the first fall just before impact, we shall have

$$r^2 = a^2 + 40g.....(1)$$
.

The velocity at the beginning of the first rebound will be ev, and since this velocity is lost by an ascent through ten feet, there is

$$e^2 v^2 = 20g....(2)$$
.

Again, since $e^2 v$, the velocity at the beginning of the second rebound, is lost by an ascent through four feet, we get

$$e^4 v^2 = 8g.....(3).$$

From the equations (1), (2), (3), we readily see that

$$e = \left(\frac{2}{5}\right)^{\frac{1}{4}}, \ a = (10g)^{\frac{1}{4}}.$$

Taking for g its approximate value 32 $\frac{1}{2}$ feet, we obtain 18 feet nearly for the value of a.

(7) A particle falling in a straight line towards a centre of force, the intensity of which varies as the nth power of the distance, acquires a velocity β on arriving at a distance a from the centre; to find at what distance z from the centre of force it must have commenced its motion.

Let μ denote the absolute force; then z will be given by the equation

$$z^{n+1} - a^{n+1} = \frac{n+1}{2\mu} \beta^2.$$

Euler; Mechan. tom. 1. p. 109.

(8) A particle falls towards a centre of force, of which the intensity varies inversely as the cube of the distance; to find the whole time of descent.

Let μ denote the absolute force and a the initial distance; then

time of descent
$$=\frac{a^2}{\mu^4}$$
.

(9) A particle descends from an infinite distance towards a centre of force which varies inversely as the square of the distance; to find the velocity at a given distance from the centre of force.

Let μ be the absolute force and a the given distance; then

velocity =
$$\left(\frac{2 \mu}{a}\right)^{\frac{1}{4}}$$
.

(10) A particle is placed at an assigned point between two centres of force of equal intensity attracting directly as the distance; to determine the position of the particle at any time, and the period of its oscillations.

Let a denote the initial distance of the particle from the middle point of the line joining the two centres of force, x the distance after the expiration of a time t, and μ the absolute force of each centre. Then

$$x = a \cos \{(2\mu)^{\frac{1}{4}}t\}$$
, and period of an oscillation = $\frac{\pi}{(2\mu)^{\frac{1}{4}}}$.

(11) A particle acted upon by two central forces, each attracting with an intensity varying inversely as the square of the distance, is projected from an assigned point between them towards one of the centres; to find the velocity of projection that the particle may just arrive at the neutral point of attraction and remain at rest there.

Let μ , μ' , denote the absolute forces of the two centres; 2a, 2a', the initial distances of the particle from the two centres; and V the velocity of projection. Then

$$V^2 = \frac{\mu}{a} + \frac{\mu'}{a'} - \frac{(\mu^{\frac{1}{4}} + \mu'^{\frac{1}{4}})^2}{a+a'}.$$

(12) A centre of force C (fig. 98) moves along the straight line OA with a uniform velocity, attracting, with a force varying directly as the first power of the distance, a particle P which is moving in the same straight line; having given the initial position of C, and both the initial position and the initial velocity of P, to find the position of P at any time.

Let a, a', be the initial distances of C, P, from O; β the uniform velocity of C, and β' the initial velocity of P; x the distance of P from O after a time t; μ the absolute force of attraction.

Then
$$x = a + \beta t + \frac{\beta' - \beta}{\mu^{\frac{1}{4}}} \sin(\mu^{\frac{1}{4}}t) + (a' - a)\cos(\mu^{\frac{1}{4}}t)$$
.
Riccati; Bonon. Institut. tom. vi. p. 138; 1783.

(13) The circumstances remaining the same as in the preceding problem, except that the force is repulsive; to find the position of P at any time.

$$x = a + \beta t - \left\{ \mu^{\frac{1}{4}}(a - a') + (\beta - \beta') \right\} \frac{\epsilon^{\mu^{\frac{1}{2}}t}}{2\mu^{\frac{1}{4}}} - \left\{ \mu^{\frac{1}{4}}(a - a') - (\beta - \beta') \right\} \frac{\epsilon^{-\mu^{\frac{1}{4}}t}}{2\mu^{\frac{1}{4}}}.$$
Riccati; *Ib.* p. 151.

(14) Supposing the centre C to move along OA with a uniform acceleration, attracting directly as the distance; to determine the place of P at any time.

Let f represent the increment of C's velocity in each unit of time, and β its velocity at the commencement of the motion; then, the notation remaining the same as in the two preceding problems,

$$x = a - \frac{f}{\mu} + \beta t + \frac{1}{2} f t^2 + \frac{\beta' - \beta}{\mu^{\frac{1}{4}}} \sin(\mu^{\frac{1}{4}} t) + \left(a' - a + \frac{f}{\mu}\right) \cos(\mu^{\frac{1}{4}} t).$$
Riccati; *Ib.* p. 168.

(15) The circumstances and notation remaining the same as in the preceding example, except that the force is repulsive; to find the place of P at any time.

$$x = a + \frac{f}{\mu} + \beta t + \frac{1}{2}ft^{3}$$

$$-\left\{\mu^{\frac{1}{4}}\left(a - a' + \frac{f}{\mu}\right) + (\beta - \beta')\right\} \frac{\epsilon^{\mu^{\frac{1}{2}}t}}{2\mu^{\frac{1}{4}}}$$

$$-\left\{\mu^{\frac{1}{4}}\left(a - a' + \frac{f}{\mu}\right) - (\beta - \beta')\right\} \frac{\epsilon^{-\mu^{\frac{1}{2}}t}}{2\mu^{\frac{1}{4}}}.$$

Riccati; Ib. p. 182.

(16) A particle is placed at a given distance from a uniform thin plate of indefinite extent, every particle of which attracts with a force varying inversely as the square of the distance; to find the time in which the particle will arrive at the surface of the plate.

Let k denote the thickness of the plate, ρ its density, and a the initial distance of the particle from it. Then

time =
$$\left(\frac{a}{\pi \rho k}\right)^{\frac{1}{4}}$$
.

(17) A particle is placed at a small distance from the centre of a thin ring of uniform density and thickness, every molecule of which repels with a force varying inversely as the square of the distance; to determine the position of the particle at any time, and the period of its oscillations.

Let l be the initial distance of the particle from the centre of the ring, a the radius of the ring, k the area of a section, ρ the density, and x the distance of the particle from the centre at the end of a time t. Then, the repulsion of a unit of the ring's mass at a unit of distance being taken as the unit of repulsion,

$$x = l \cos \left\{ \frac{(\pi \rho k)^{\frac{1}{4}}}{a} t \right\}$$
 and period of an oscillation $= \left(\frac{\pi}{\rho k} \right)^{\frac{1}{4}} a$.

(18) A particle being placed at a given distance from a thin circular lamina of uniform density, in a line passing through its centre and perpendicular to its plane, to find the velocity which it will acquire by moving to the circle, the attractive force of each molecule of the circle varying inversely as the square of the distance.

Let a be the radius of the circle, k its thickness, ρ its density, b the given distance; then, the unit of attraction being the attraction of a unit of mass at a unit of distance, and V being the velocity required,

$$V^{2} = 4\pi\rho k \{a + b - (a^{2} + b^{2})^{\frac{1}{2}}\}.$$

(19) A body of known elasticity falls from a given altitude above a hard horizontal plane, and rebounds continually till its whole velocity is destroyed; to find the whole space described.

If a denote the first altitude, e the elasticity, and s the required space,

 $s=\frac{1+e^2}{1-e^2}a.$

(20) Two perfectly elastic balls beginning to descend from different points in the same vertical line impinge upon a perfectly hard plane inclined at an angle of 45°, and move along a horizontal plane with the velocities acquired; to find what distance they will move along the horizontal plane before collision takes place.

If a, a', denote the altitudes through which they fall, and s the distance required, $s = (aa')^{\frac{1}{4}}.$

(21) From a point A in a vertical line AB falls a particle from rest; at the same instant another particle is projected upwards from B with a given velocity; to find when and where the two particles will meet; the motion being supposed to take place in a vacuum, and gravity being the only force to which the particles are subject.

Let a be the length of the line AB, β the velocity of projection of the ascending particle, x the distance from A at which collision takes place, and t the time of this event from the commencement of the motion. Then

$$x = \frac{ga^2}{2\beta^2}, t = \frac{a}{\beta}.$$

Kurdwanowski; Mém. de l'Acad. des Sciences de Berlin, 1755, p. 394.

SECT. 2. Motion in Resisting Media.

The retardation experienced by a material particle in traversing a resisting medium of variable density, depends at any point of its path upon the density of the medium and the velocity of the particle, and will therefore be some function of these quantities. The nature of this function can be ascertained only by experiment. In mathematical investigations, for the sake of simplicity and as a probable approximation to the truth, the function is assumed to be of the form $k\rho\Omega$, where ρ denotes the density of the medium and Ω some function of the velocity of the particle; and where k is an invariable coefficient depending upon the nature of the particular medium in respect to the tenacity and the friction of its constituent molecules.

For the earliest mathematical development of the theory of the resistance of media to the motion of bodies, we are indebted to the labours of Newton and Wallis. The profound researches of Newton on this theory were published in the year 1687 in the second book of the Principia. In the same year, after the publication of Newton's investigations, Wallis, who had independently arrived at valuable conclusions on the subject, communicated his reflections to the Royal Society, which were inserted in the Philosophical Transactions for the year 1687. There is a paper by Leibnitz on the question of resisting media in the Acta Erudit. Lips. ann. 1689, in which he developes opinions which he declares to have been communicated by him twelve years before to the Royal Academy of Sciences. Huyghens also has discussed certain points of the theory at the end of his Discours de la Cause de la Pesanteur, published in the year 1690. Finally, all which these philosophers had communicated to the scientific world either with or without demonstration, was investigated analytically by Varignon in a series of papers in the Mémoires de l'Acad. des Sciences de Paris, for the years 1707, 1708, 1709, and 1710. There is an elaborate paper by Bouguer in the Mém. de l'Acad. des Sciences de Paris, 1731, p. 390, in which he investigates the motion of a particle in resisting media which are themselves in motion.

(1) A particle acted upon by no forces is projected with a given velocity in a resisting medium of uniform density, where the resistance varies directly as the velocity; to determine the velocity and the space described at the end of any time.

For the motion of the particle we have

$$v \frac{dv}{dx} = -\mu v,$$

where μ is some constant quantity; hence

$$\frac{dv}{dx}=-\mu,$$

$$v = C - \mu x$$
.

Let β denote the initial velocity when x is supposed to be zero; then $\beta = C$, and therefore

$$v=eta-\mu x$$
; whence, v being equal to $\frac{dx}{dt}$, we have
$$dt=\frac{dx}{eta-\mu x},$$

$$t = C - \frac{1}{\mu} \log (\beta - \mu x).$$

But t = 0 when x = 0; and therefore

$$0 = C - \frac{1}{\mu} \log \beta,$$
hence $t = \frac{1}{\mu} \log \frac{\beta}{\beta - \mu x},$

$$\varepsilon^{\mu t} = \frac{\beta}{\beta - \mu x},$$

$$x = \frac{\beta}{\mu} (1 - \varepsilon^{-\mu t}),$$

$$v = \beta \varepsilon^{-\mu t}.$$

and therefore

Newton; Principia, lib. 11. Prop. 1 and 2. Leibnitz; Acta Erudit. Lips. ann. 1689. Varignon; Mém. de l'Acad. des Sciences de Paris, ann. 1707, p. 391.

(2) A body falls towards a centre of force which varies as the inverse cube of the distance, in a medium of which the density varies also as the inverse cube, and of which the resistance varies as the square of the velocity; to find the velocity at any distance from the centre.

Let x represent the distance of the particle from the centre after a time t, and let a be the initial distance. Let k denote the force of resistance at a unit of distance for a unit of velocity, and μ the absolute force of attraction.

Then, for the motion of the particle,

$$\frac{d^3x}{dt^2} = -\frac{\mu}{x^3} + \frac{k}{x^3} \frac{dx^2}{dt^4},$$

$$2 \frac{dx}{dt} \frac{d^2x}{dt^2} = -\frac{2\mu}{x^3} \frac{dx}{dt} + \frac{2k}{x^3} \frac{dx^3}{dt^3},$$

$$\frac{d}{dt} \frac{dx^2}{dt^2} = \mu \frac{d}{dt} \frac{1}{x^2} - k \frac{dx^2}{dt^2} \frac{d}{dt} \frac{1}{x^2}.$$
Assume $\frac{dx^2}{dt^2} = w$ and $\frac{1}{x^2} = z$; then
$$\frac{dw}{dt} = \mu \frac{dz}{dt} - kw \frac{dz}{dt},$$

$$dw + kwdz = \mu dz,$$

$$d(\epsilon^{h}w) = \mu \epsilon^{h}dz,$$

$$\epsilon^{h}w = C + \frac{\mu}{k}\epsilon^{h}.$$

Hence, putting for w and z their values,

$$\varepsilon^{\frac{k}{\mu}} e^{2} = C + \frac{\mu}{k} \varepsilon^{\frac{k}{\mu}}.$$

But x = a when v = 0; hence

$$0 = C + \frac{\mu}{k} \epsilon^{\frac{k}{k^2}},$$

$$\epsilon^{\frac{k}{k^2}} \epsilon^{2} = \frac{\mu}{k} \left(\epsilon^{\frac{k}{k^2}} - \epsilon^{\frac{k}{k^2}} \right),$$

$$\epsilon^{2} = \frac{\mu}{k} \left\{ 1 - \epsilon^{-\frac{k}{k} \left(\frac{1}{k^2} - \frac{1}{k^2} \right)} \right\}.$$

When x becomes equal to infinity, suppose V to be the value of v; then $V^{z} = \frac{\mu}{L} \left(1 - e^{\frac{L}{2}}\right).$

V is called the terminal velocity of the particle, a technicality invented by Huyghens¹, to signify the ultimate velocity of a particle descending in a resisting medium to an indefinitely great depth.

(3) To determine the motion of a particle, not acted upon by any force, when the resistance varies as any power of the velocity.

For the determination of the relation between the velocity and the space,

$$v \frac{dv}{dx} = -kv^{n},$$

$$kdx = -\frac{dv}{v^{n-1}},$$

$$kx = C + \frac{1}{(n-2)v^{n-2}}.$$

Let x = 0 and $v = \beta$ initially; then

$$0 = C + \frac{1}{(n-2)\beta^{n-2}},$$

¹ Discours sur la Cause de la Pesanteur, p. 170.

$$(n-2) kx = \frac{1}{v^{n-2}} - \frac{1}{\beta^{n-2}} \dots (1).$$

Again, for the relation between the velocity and the time,

$$\frac{dv}{dt} = -kv^n, \quad kdt = -\frac{dv}{v^n},$$

$$kt = C + \frac{1}{(n-1)v^{n-1}}.$$

But $v = \beta$ when t = 0; hence

$$0 = C + \frac{1}{(n-1)\beta^{n-1}},$$

$$(n-1)kt = \frac{1}{v^{n-1}} - \frac{1}{\beta^{n-1}}.....(2).$$

If between (1) and (2) we eliminate v, we shall obtain a relation between s and t.

Varignon; Mém. de l'Acad. des Sciences de Paris, 1707, p. 404.

(4) A particle acted on by gravity falls from a given altitude in a medium of uniform density, where the resistance varies as the square of the velocity; on arriving at the lowest point of its descent it is reflected upwards with the velocity which it has acquired in its fall; after reaching its greatest altitude it again descends and is again reflected; and so on perpetually: to determine the altitude of ascent after any number of reflections.

Let the maximum altitudes of the particle be represented by $a_1, a_2, a_3, \ldots a_n$ being the altitude from which it originally falls. Let c denote the volume of the particle, and ρ, ρ' , the density of the particle and of the fluid.

For the descent down any of the altitudes there is

$$v \frac{dv}{dx} = \frac{cg\rho - cg\rho'}{c\rho} - kv^{2},$$

$$v \frac{dv}{dx} = g' - kv^{2}, \text{ where } g' = \left(1 - \frac{\rho'}{\rho}\right)g,$$

$$\frac{vdv}{g' - kv^{2}} = dx,$$

$$-\frac{1}{2k} \log (g' - kv^{2}) = x + C;$$

but, the origin of x being the highest point,

or

$$-\frac{1}{2k}\log g'=C,$$

hence

$$\frac{1}{2k}\log\frac{g'}{g'-kv^2}=x,$$

and therefore, if v_n denote the velocity acquired down the n^{th} altitude,

$$\log \frac{1}{1 - \frac{k}{g'} v_n^2} = 2ka_n,$$

$$\frac{k}{g'} v_n^2 = 1 - \varepsilon^{-2ka_n} \cdot \dots (1).$$

For the ascent up the (n + 1)th altitude, the origin of x being the lowest point,

$$v \frac{dv}{dx} = -g' - kv^{2}, \quad \frac{vdv}{g' + kv^{2}} = -dx,$$

$$\frac{1}{2k} \log (g' + kv^{2}) = C - x,$$

$$\frac{1}{2k} \log (g' + kv_{n}^{2}) = C,$$

$$\frac{1}{2k} \log g' = C - a_{n+1},$$

$$\frac{1}{2k} \log (1 + \frac{k}{g'} v_{n}^{2}) = a_{n+1},$$

$$\frac{k}{g'} v_{n}^{2} = \varepsilon^{2ka_{n+1}} - 1 \dots (2).$$

Hence, from (1) and (2),

$$\varepsilon^{2ka_{m+1}} + \varepsilon^{-2ka_m} = 2$$
:

assume $e^{2ka_n} = u_n$, and we have

$$u_{n+1} + \frac{1}{u_n} = 2,$$

$$u_n u_{n+1} + 1 = 2u_n.$$

Putting
$$u_n = v_n + 1$$
, we get
$$(v_n + 1)(v_{n+1} + 1) + 1 = 2 (v_n + 1),$$

$$v_n v_{n+1} + v_{n+1} - v_n = 0, \quad \frac{1}{v_{n+1}} - \frac{1}{v_n} = 1,$$

$$\Delta \frac{1}{v_n} = 1, \quad \frac{1}{v_n} = n + C,$$

$$\frac{1}{u - 1} = n + C, \quad u_n - 1 = \frac{1}{n + C}, \quad u_n = \frac{n + 1 + C}{n + C}.$$

But
$$\frac{1}{u_1 - 1} = 1 + C$$
; hence
$$u_n = \frac{n + \frac{1}{u_1 - 1}}{n - 1 + \frac{1}{u_1 - 1}} = \frac{nu_1 - n + 1}{(n - 1)u_1 - n + 2}.$$

Or, putting for u_1 , u_1 , their values,

$$\varepsilon^{2ka_n} = \frac{n\varepsilon^{2ka_1} - n + 1}{(n-1)\varepsilon^{2ka_1} - n + 2},$$

$$a_n = \frac{1}{2k}\log\frac{n\varepsilon^{2ka_1} - n + 1}{(n-1)\varepsilon^{2ka_1} - n + 2}.$$

If a, be equal to infinity,

$$a_n = \frac{1}{2k} \log \frac{n}{n-1}.$$

Euler; Mechan. tom. 1. p. 192.

(5) To determine the centripetal force that a particle may always descend to a given centre in the same time from whatever distance it commences its motion; the density of the medium in which the particle moves being known at every point in its path, and the resistance varying as the square of the velocity.

The equation of motion is

$$v\frac{dv}{dx}=-p+kv^2,$$

where p denotes the centripetal force, and k the density at any point. Multiplying by $2e^{-2/kdx} dx.$

the equation becomes

$$d\left(\varepsilon^{-2\int hdx}\ v^2\right) = -\ 2\ \varepsilon^{-2\int hdx}\ pdx.$$

Integrating,

$$\varepsilon^{-2\int hds} \ v^3 = C - 2\int \varepsilon^{-2\int hds} \ pdx.$$

Let a denote the initial distance of the particle from the centre of force; then, the velocity being initially equal to zero,

$$\varepsilon^{-2\int hdx} v^2 = A - X,$$
 where $A = 2 \int_0^{\pi} \varepsilon^{-2\int hdx} pdx$ and $X = 2 \int_0^{\pi} \varepsilon^{-2\int hdx} pdx$.

Therefore

$$v = (A - X)^{\frac{1}{6}} \epsilon^{\int h dx},$$

$$dt = -\frac{\epsilon^{-\int h dx} dx}{(A - X)^{\frac{1}{6}}}.$$

Now since X, k, are both functions of x, it is clear that we may assume $\epsilon^{-\int hdx} dx = \frac{dX}{R}$

where P is a function of X alone; hence

$$dt = -\frac{dX}{P(A-X)^{\frac{1}{4}}},$$

and the whole time of descent, since X = 0 when x = 0, will be equal to

$$-\int_{a}^{0} \frac{\varepsilon^{-f k d x}}{\left(A - X\right)^{\frac{1}{4}}} = -\int_{A}^{0} \frac{dX}{P\left(A - X\right)^{\frac{1}{4}}};$$

and since the value of this integral is to be the same for all values of a and therefore of A, the differential

$$\frac{dX}{P(A-X)^{\frac{1}{4}}}$$

must be of no dimensions in X, dX, and A. Hence we must have P, which clearly cannot involve a, equal to $\frac{X^{\frac{1}{4}}}{\beta}$, where β is some constant quantity; and therefore

$$\beta \frac{dX}{X^{\frac{1}{4}}} = \varepsilon^{-\int kdx} dx;$$

hence, X and x being simultaneously equal to zero,

$$2\beta X^{\frac{1}{4}} = \int_{0}^{x} \varepsilon^{-\int k dx} dx, \quad 4\beta^{2} X = \left\{ \int_{0}^{x} \varepsilon^{-\int k dx} dx \right\}^{2},$$

$$8\beta^{2} \int_{0}^{x} (\varepsilon^{-2\int k dx} p dx) = \left\{ \int_{0}^{x} \varepsilon^{-\int k dx} dx \right\}^{2},$$

$$4\beta^{2} \varepsilon^{-2\int k dx} p dx = \varepsilon^{-\int k dx} dx \int_{0}^{x} \varepsilon^{-\int k dx} dx,$$

$$p = \frac{1}{4\beta^{2}} \varepsilon^{\int k dx} \int_{0}^{x} \varepsilon^{-\int k dx} dx.$$

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If k be equal to zero, which corresponds to a perfect vacuum,

$$p=\frac{1}{4\beta^2}\int_0^x dx=\frac{x}{4\beta^2},$$

or the centripetal force varies as the distance.

If the medium be uniform, or k a constant quantity,

$$p = \frac{1}{4\beta^3} \epsilon^{ka} \int_0^a \epsilon^{-ka} dx$$
$$= \frac{1}{4\beta^3} \epsilon^{ka} \cdot \frac{1}{k} (1 - \epsilon^{-ka}) = \frac{1}{4\beta^3 k} (\epsilon^{ka} - 1).$$

Euler; Mechan. tom. 1. p. 220.

(6) A centre of force C, (fig. 99), moves along the straight line OA with a uniform velocity, repelling with a force varying directly as the first power of the distance; a particle P is moving along the same straight line OA in a medium resisting as the velocity; having given the initial position of C, and both the initial position and the initial velocity of P, to find the position of P at any time.

Let β be the uniform velocity of C, a its initial distance from O; x the distance of P from O at the end of the time t; μ the absolute force of repulsion; k the resistance of the medium for a unit of velocity. Then, the distance between C and P at the time t being $a + \beta t - x$, we have for the motion of P, so long as it continues to proceed in the direction OA,

$$\frac{dv}{dt} = -\mu (a + \beta t - x) - kv,$$

v being the velocity of P at the time t, estimated in the direction OA. Supposing the particle to be moving in the direction AO, then v being also estimated in this direction, we should have

$$\frac{do}{dt} = \mu (a + \beta t - x) - kv,$$

an equation deducible from the former one by the substitution of -v in place of v: hence the former equation applies to the motion of the particle under all circumstances, the quantity v being supposed to involve the direction-sign of the velocity implicitly.

But
$$v = \frac{dx}{dt}$$
, and therefore
$$\frac{d^2x}{dt^2} = -\mu (a + \beta t - x) - k \frac{dx}{dt},$$

$$\frac{d^2x}{dt^2} + k \frac{dx}{dt} = \mu (x - a - \beta t);$$

and therefore, putting $x - a - \frac{k\beta}{\mu} - \beta t = z$,

$$\frac{d^2z}{dt^2} + k\,\frac{dz}{dt} = \mu z.$$

Assume $z = A \epsilon^{\rho t}$, A and ρ being constants; then, substituting for z in the differential equation, we have

$$\rho^{2} + k\rho = \mu,$$

$$4\rho^{2} + 4k\rho + k^{2} = 4\mu + k^{2},$$

$$2\rho = -k \pm (4\mu + k^{2})^{\frac{1}{2}};$$

hence the complete integral is

$$z = C \varepsilon^{\gamma_t} + C' \varepsilon^{\gamma'_t},$$

where γ , γ' , are the two values of ρ obtained by the solution of the quadratic. Hence for the position of P at any time we have

$$x = a + \frac{k\beta}{\mu} + \beta t + C \varepsilon^{\gamma t} + C' \varepsilon^{\gamma' t}.$$

Suppose a', β' , to be the initial values of x, $\frac{dx}{dt}$; then clearly

$$a' = a + \frac{k\beta}{\mu} + C + C', \quad \beta' = \beta + \gamma C + \gamma' C';$$

from which two equations the values of the constants C and C' are immediately determined.

For further information on the subject of the rectilinear motion of a particle in a resisting medium under the action of a centre of force moving according to any assigned law, the reader is referred to Riccati; De motu rectilineo corporis attracti aut repulsi a centro mobili; Disquisitio quarta. Comment. Bonon. tom. vi. p. 212; 1783.

(7) A particle is projected with a given velocity, towards a centre of force attracting inversely as the cube of the distance,

in a medium of which the density varies inversely as the square of the distance from the centre of force; to determine the velocity of the particle at any distance from the centre, the resistance for a given density varying as the square of the velocity.

If β denote the velocity of projection, μ the absolute attracting force; k the retarding force of the medium, at a unit of distance from the centre of force, for a unit of velocity; a the initial distance of the particle, and x its distance corresponding to a velocity v,

$$\varepsilon^{-\frac{2k}{a}}v^2 - \varepsilon^{-\frac{2k}{a}}\beta^2 = \frac{\mu}{2k^2}\left(\frac{a-2k}{a}\varepsilon^{-\frac{2k}{a}} - \frac{x-2k}{x}\varepsilon^{-\frac{2k}{a}}\right).$$

(8) A particle is projected with a given velocity in a uniform medium, in which the resistance varies as the square root of the velocity; to find what time will elapse before the particle is reduced to rest.

If β be the velocity of projection, and k the resistance for a unit of velocity,

required time = $\frac{2\beta^{\frac{1}{4}}}{k}$.

(9) If t denote the time in which a particle falling from rest will acquire a certain velocity, and r the time in which, when projected vertically upwards, it will lose the same velocity; the motion in both cases taking place in a medium of uniform density, where the resistance varies as the square of the velocity; to investigate the relation between t and r.

If k denote the resistance for a unit of velocity,

$$\log \tan (\frac{1}{4}\pi + k^{\frac{1}{4}}g^{\frac{1}{4}}\tau) = 2g^{\frac{1}{4}}k^{\frac{1}{4}}t.$$

(10) A particle projected with a velocity of 1000 feet a second, loses half its velocity by passing through 3 inches of a resisting medium, in which the resistance is uniform; to find the time of passing through this space.

Required time = 3000th part of a second.

(11) A particle, attracted to a centre of constant attractive force, moves directly towards it from rest, through a medium of

which the resistance varies as the square of the velocity directly, and as the distance from the centre inversely; to find the velocity for any position of the particle during its approach towards the centre, and to ascertain its distance from the centre when its velocity is a maximum.

If f denote the constant central force, v the velocity for any distance x from the centre, a the initial value of x, k the resistance when x and v are each equal to unity, and x' the central distance when v is a maximum;

$$v^2 = \frac{2f}{1-2k} x^{2k} (a^{1-2k} - x^{1-2k}), \quad x' = (2k)^{\frac{1}{1-2k}} a.$$

(12) One particle begins to fall from the higher extremity of a vertical line, at the same instant in which another is projected upwards with a given velocity; the particles move in a uniform medium in which the resistance varies as the velocity; to find the time in which they will meet.

Let a denote the length of the vertical line, β the velocity with which the lower particle is projected upwards, k the resistance for a unit of velocity, and t the required time; then

$$t=\frac{1}{k}\log\frac{\beta}{\beta-ka}.$$

(13) A particle, of which the elasticity is e, falls from rest from an altitude a in a uniform medium, the resistance of which is kv^2 ; and impinging upon a perfectly hard horizontal plane, rises and falls alternately; to determine the whole space described before the motion ceases.

Required space =
$$a + \frac{1}{k} \log \frac{1 - e^2 e^{-2ka}}{1 - e^2}$$
.

Bordoni; Memorie della Societa Italiana, 1816, p. 162.

CHAPTER III.

FREE CURVILINEAR MOTION OF A PARTICLE.

SECT. 1. Forces acting in any directions in one Plane.

LET a particle moving in a plane curve under the action of any accelerating forces be referred to two fixed co-ordinate axes in the plane of its motion. Let x, y, be its co-ordinates at the end of a time t from an assigned epoch; and X, Y, the sum of the resolved parts of the accelerating forces parallel to the axes of x, y. Then the circumstances of the motion will be completely represented by the equations

$$\frac{d^2x}{dt^2}=X, \quad \frac{d^2y}{dt^2}=Y, \ldots \ldots (A).$$

The method of resolving parallel to fixed axes the accelerating forces which act upon a particle, and thus reducing the determination of the circumstances of its motion to the formulæ for rectilinear acceleration, was first given by Maclaurin, Treatise on Fluxions, Vol. 1. Art. 465, et sq., published in the year 1742. Before this time all problems in curvilinear motion were solved by the method of the tangential and normal resolutions, which, although more immediately suggested by the physical conception of the motion, is not generally so convenient in analysis as that of Maclaurin. The great work of Euler on Mechanics, which appeared in 1736, proceeds altogether by the ancient method of resolution. We shall devote the third section of this chapter to the illustration of the ancient equations of motion.

(1) A particle acted on by gravity is describing a path KABL, (fig. 100); having given the resolved part of the velocity at A at right angles to the chord AB, to find the resolved part at B taken in the same direction.

Let u be the given resolved part of the velocity at A; v the

velocity at right angles to AB at any point of the path corresponding to a time t from leaving A, and x the perpendicular distance of the particle from AB at this time; also let a be the inclination of AB to the horizon. Then, for the motion of the particle,

$$x = ut - \frac{1}{2}g \cos a \cdot t^2,$$

and

$$v = u - g \cos a \cdot t$$
.

Now at the point B, x = 0, and therefore

$$u - \frac{1}{2}g \cos a \cdot t = 0$$
, $g \cos a \cdot t = 2u$;

hence for the value of v at B we have

$$v = -u$$
.

Thus we see that the velocity of the particle at B, resolved at right angles to AB, is equal to the similarly resolved part at A, but of an opposite direction.

(2) A particle revolves in a parabola about a centre of force situated in the point of intersection of the directrix with the axis; to find the force at any point of the path of the particle.

Take the centre of force as the origin of co-ordinates, the axis of the parabola as the axis of x, and the directrix as the axis of y; let 4 m be the principal parameter, r the distance of the particle at any time from the origin, and F the force estimated repulsively.

The equations of motion will be

$$\frac{d^2x}{dt^2} = \frac{Fx}{r}, \quad \frac{d^2y}{dt^2} = \frac{Fy}{r}, \quad \dots (1).$$

The equation to the parabola will be

$$y^2 = 4m(x-m)....(2);$$

hence

$$y\frac{dy}{dt}=2m\frac{dx}{dt}....(3),$$

$$y \frac{d^2y}{dt^2} + \frac{dy^2}{dt^2} = 2m \frac{d^2x}{dt^2};$$

and therefore, from (1)

$$F(2mx-y^2)=r\frac{dy^2}{dt^2}....(4).$$

Again, eliminating F between the equations (1), we have

$$x\frac{d^2y}{dt^2}-y\frac{d^2x}{dt^2}=0;$$

integrating, and adding a constant c which will represent twice the area described in a unit of time about the centre of force, we obtain

 $x\frac{dy}{dt}-y\frac{dx}{dt}=c;$

and therefore, by (3),

$$(2mx - y^2)\frac{dy}{dt} = 2mc:$$

hence, from (4),

$$F = \frac{4m^2c^2r}{(2mx - y^2)^3} = \frac{c^2r}{2m(2m - x)^3}$$
, by (2).

(3) A particle P, moving with a uniform angular velocity round a fixed point O, is acted on by a force always perpendicular to the straight line OP; to find the polar equation to the orbit, O being the pole, and the particle having initially no velocity in the direction of the radius vector.

Let the axis of x coincide with the prime radius vector; $\angle POx = \theta$, OP = r, F = the force, $\omega =$ the invariable angular velocity. Then for the motion of the particle there is

$$\frac{d^2x}{dt^2}=-\frac{Fy}{r}, \quad \frac{d^2y}{dt^2}=\frac{Fx}{r};$$

hence we have

$$x\frac{d^2x}{dt^2}+y\frac{d^2y}{dt^2}=0.....(1).$$

But, $\frac{d\theta}{dt}$ being equal to the constant quantity ω ,

$$x = r \cos \theta$$
, $\frac{dx}{dt} = \cos \theta \frac{dr}{dt} - \omega r \sin \theta$,

$$\frac{d^2x}{dt^2} = \cos\theta \frac{d^2r}{dt^2} - 2\omega \sin\theta \frac{dr}{dt} - \omega^2r \cos\theta \dots (2);$$

and

$$y = r \sin \theta$$
, $\frac{dy}{dt} = \sin \theta \frac{dr}{dt} + \omega r \cos \theta$,

$$\frac{d^2y}{dt^2} = \sin\theta \frac{d^2r}{dt^2} + 2\omega\cos\theta \frac{dr}{dt} - \omega^2r\sin\theta \dots (3).$$

From (1), (2), (3), we get

$$r \frac{d^2r}{dt^2} - \omega^2 r^2 = 0, \frac{d^2r}{dt^2} - \omega^2 r = 0;$$

multiplying by $2 \frac{dr}{dt}$ and integrating,

$$\frac{dr^2}{dt^2}-\omega^2r^2=C;$$

let a be the initial value of r; then, since r = a, $\frac{dr}{dt} = 0$, simultaneously, $C = -\omega^2 a^2$; hence

$$\frac{dr^2}{dt^2}=\omega^2\left(r^2-a^2\right);$$

or, since $\frac{d\theta}{dt} = \omega$,

$$\frac{dr^2}{d\theta^2} = r^2 - a^2, \ d\theta = \frac{dr}{(r^2 - a^2)^{\frac{1}{2}}}.$$

Integrating, we have

$$\theta + C = \log \{r + (r^2 - a^2)^{\frac{1}{4}}\};$$

but $\theta = 0$, r = a, simultaneously, and therefore $C = \log a$; hence

$$\theta = \log \left\{ \frac{r}{a} + \left(\frac{r^2}{a^2} - 1 \right)^{\frac{1}{2}} \right\},$$

$$r = \frac{1}{3} a \left(\epsilon^{\theta} + \epsilon^{-\theta} \right).$$

(4) A particle moves with a uniform velocity under the action of two centres of force, each varying inversely as the distance, and of equal intensity at the same distance; to find the path of the particle, its initial position coinciding with the middle point of the straight line which joins the two centres of force.

Let the initial position O (fig. 101) of the particle be taken as the origin of co-ordinates; let C, C', be the two centres of force; OCx the axis of x, Oy at right angles to Ox the axis of y; OC = OC' = a; OM = x, PM = y, OP = s, CP = r, C'P = r'; β = the constant velocity of the particle.

Then, by hypothesis,

$$\frac{ds^2}{dt^2} \text{ or } \frac{dx^2}{dt^2} + \frac{dy^2}{dt^2} = \beta^2.$$

Differentiating, we have

$$\frac{dx}{dt}\frac{d^2x}{dt^2}+\frac{dy}{dt}\frac{d^2y}{dt^2}=0\ldots\ldots(1).$$

Again, μ denoting the absolute force of attraction of each centre of force, we have, by Maclaurin's Equations (A),

$$\frac{d^2x}{dt^2} = \frac{\mu}{r} \frac{a-x}{r} - \frac{\mu}{r'} \frac{a+x}{r'},$$

$$\frac{d^2y}{dt^2} = -\frac{\mu}{r} \frac{y}{r} - \frac{\mu}{r'} \frac{y}{r'},$$

and therefore, from (1),

$$\frac{dx}{dt}\left(\frac{a+x}{r'^2} - \frac{a-x}{r^2}\right) + \frac{dy}{dt}\left(\frac{y}{r^2} + \frac{y}{r'^2}\right) = 0,$$

$$\frac{(x+a) dx + ydy}{(x+a)^2 + y^2} + \frac{(x-a) dx + ydy}{(x-a)^2 + y^2} = 0.$$

Integrating, we have

$$\log \{(x+a)^2+y^2\} + \log \{(x-a)^2+y^2\} = \log C,$$
$$\{(x+a)^2+y^2\} \{(x-a)^2+y^2\} = C.$$

But x = 0, y = 0, simultaneously; hence $C = a^4$, and therefore

$$\{(x+a)^2+y^2\} \{(x-a)^2+y^2\} = a^4,$$

$$(a^2+x^2+y^2)^2-4a^2x^2=a^4,$$

$$2a^2(x^2+y^2)+(x^2+y^2)^2-4a^2x^2=0,$$

$$(x^2+y^2)^2=2a^2(x^2-y^2),$$

which is the equation to the Lemniscata of James Bernoulli, $2^{\frac{1}{4}}a$ being the semiaxis of the corresponding equilateral hyperbola.

(5) A particle attracting with a force varying directly as the distance moves uniformly in a straight line; to determine the motion of another particle situated in the same plane and subject to its influence, the initial circumstances of the latter particle being given.

Let the initial position of the attracting particle be taken as the origin of co-ordinates, and let the axes of x and y make angles of 45° on each side of the line of its motion.

Let x', y', be the co-ordinates of the attracting, and x, y, of the attracted particle at any time t, then the equations of motion will be d^2x d^2y

$$\frac{d^2x}{dt^2} = \mu^2(x'-x), \quad \frac{d^2y}{dt^2} = \mu^2(y'-y),$$

 μ^2 denoting the absolute force of attraction.

But if a, β , denote the resolved parts of the velocity of the attracting particle parallel to the axes of x, y, which are by hypothesis invariable,

and therefore

$$x' = at, \quad y' = \beta t,$$

$$\frac{d^2x}{dt^2} = \mu^2 (at - x) \dots (1),$$

$$\frac{d^2y}{dt^2}=\mu^2(\beta t-y)\ldots (2).$$

From the equation (1)

$$\frac{d^2}{dt^2}(x-at)+\mu^2(x-at)=0.$$

The integral of this equation is

$$x = A \cos(\mu t) + B \sin(\mu t) + at \dots (3);$$

where A, B, are arbitrary constants. Differentiating we have

$$\frac{dx}{dt} = -A\mu \sin (\mu t) + B\mu \cos (\mu t) + a \dots (4).$$

Let a, m, be the initial values of x, $\frac{dx}{dt}$; then, from (3) and (4),

$$a=A, m=B\mu+a,$$

and therefore, from (3),

$$x = a \cos(\mu t) + \frac{m-a}{\mu} \sin(\mu t) + at.$$

In precisely the same way, b and n being the initial values of

$$y \text{ and } \frac{dy}{dt}$$
, $y = b \cos(\mu t) + \frac{n-\beta}{\mu} \sin(\mu t) + \beta t$.

(6) A particle urged towards a plane by a force varying as the perpendicular distance from it, is projected at right angles to the plane from a given point in it with a given velocity; to determine the force which must act at the same time on the particle parallel to the plane, that it may move in a given parabola having its axis in the plane, and to find the co-ordinates of the particle at any epoch of the motion in terms of the time.

Let the initial place of the particle be taken as the origin of co-ordinates, the axis of the parabola as the axis of x, and a straight line at right angles to the plane through the origin as

the axis of y. Then, since the required force must evidently act parallel to the axis of x, we have, by Maclaurin's Equations (A), $d^{2}x$

$$\frac{d^3x}{dt^2}=X,\ldots (1),$$

$$\frac{d^2y}{dt^2} = -\mu y \dots (2),$$

where X is the required force and μ a constant quantity. Also the equation to the parabola will be

$$y^2 = 4mx \dots (3).$$

Differentiating this equation,

$$y\frac{dy}{dt}=2m\frac{dx}{dt},$$

$$\frac{dy^2}{dt^2} + y \frac{d^3y}{dt^2} = 2m \frac{d^3x}{dt^2} = 2mX$$
, by (1),(4).

Multiplying (2) by $2 \frac{dy}{dt}$, and integrating,

$$\frac{dy^2}{dt^2} = C - \mu y^2.$$

Let V be the velocity of projection; then, y being initially equal to zero, $V^* = C$, and therefore

$$\frac{dy^2}{dt^2} = V^2 - \mu y^2 \dots (5).$$

Hence, from (2), (4), (5),

$$2mX = V^2 - 2\mu y^2,$$

$$X = \frac{V^2}{2m} - \frac{\mu}{m} y^2 = \frac{V^2}{2m} - 4\mu x$$
, by (3)....(6)

which determines the required force.

Again, from (5),

$$\frac{dt}{dy} = \frac{1}{(V^2 - \mu y^2)^{\frac{1}{4}}} = \frac{1}{\mu^{\frac{1}{4}}} \frac{1}{(V^2 - y^2)^{\frac{1}{4}}}.$$

Integrating, we have

$$t = C + \frac{1}{\mu^{\frac{1}{4}}} \sin^{-1} \frac{\mu^{\frac{1}{4}} y}{V}:$$

but y = 0 when t = 0, and therefore C = 0; hence

$$t = \frac{1}{\mu^{\frac{1}{4}}} \sin^{-1} \frac{\mu^{\frac{1}{4}} y}{V}, \quad y = \frac{V}{\mu^{\frac{1}{4}}} \sin (\mu^{\frac{1}{4}} t), \quad \dots (7).$$

Also, from (1) and (6)

$$2\frac{dx}{dt}\frac{d^2x}{dt^2} = \frac{V^2}{m}\frac{dx}{dt} - 8\mu x\frac{dx}{dt}.$$

Integrating,

$$\frac{dx^2}{dt^2} = \frac{V^2}{m} x - 4\mu x^2,$$

$$\frac{dt}{dx} = \frac{1}{\left(\frac{V^2}{m} x - 4\mu x^2\right)^{\frac{1}{2}}} = \frac{1}{2\mu^{\frac{1}{2}}} \left(\frac{1}{4\mu m} x - x^2\right)^{\frac{1}{2}}.$$

Integrating,

$$t = \frac{1}{2\mu^{\frac{1}{4}}} \text{ vers}^{-1} \frac{8\mu mx}{V^2},$$

no constant being added because x = 0 when t = 0;

$$x = \frac{V^2}{8\mu m} \operatorname{vers} (2\mu^{\frac{1}{2}}t) \dots (8).$$

From the equations (7) and (8) we see that when

$$t = 0, \quad \text{then } x = 0, \qquad y = 0;$$

$$t = \frac{\pi}{2\mu^{\frac{1}{2}}}, \quad x = \frac{V^2}{4\mu m}, \quad y = \frac{V}{\mu^{\frac{1}{2}}};$$

$$t = \frac{\pi}{\mu^{\frac{1}{2}}}, \quad x = 0, \quad y = 0;$$

$$t = \frac{3\pi}{2\mu^{\frac{1}{2}}}, \quad x = \frac{V^2}{4\mu m}, \quad y = -\frac{V}{\mu^{\frac{1}{2}}};$$

$$\vdots \quad \vdots \quad \vdots \quad \vdots \quad \vdots$$

and, generally,

when
$$t = \frac{n\pi}{\mu^{\frac{1}{2}}}$$
, then $x = 0$, $y = 0$, $t = \frac{(2n+1)\pi}{2\mu^{\frac{1}{2}}}$, $x = \frac{V^2}{4\mu m}$, $y = (-)^n \frac{V}{\mu^{\frac{1}{2}}}$.

From these relations it appears that the particle oscillates continually in a portion of the parabola cut off by a double ordinate at a distance $\frac{V^2}{4\mu m}$ from the vertex; and that the period of a complete oscillation is $\frac{\pi}{u^{\frac{1}{4}}}$.

(7) An imperfectly elastic particle, subject to the action of gravity, is projected from an assigned point in a horizontal plane with a given velocity and in a given direction; to find the velocity of incidence and of reflection, and also the total range with the corresponding time of flight, after the particle has described by rebounding any number of parabolic arcs.

Let e be the elasticity of the particle; u_s the velocity at each end of the x^{th} parabolic arc, and a_s the inclination of the curve at these points to the horizon; t_s the time which elapses before the x^{th} impact; s_s the distance of the point of x^{th} impact from the initial position of the particle.

By the theory of impact we have

$$u_{z+1}\cos a_{z+1}=u_z\cos a_z:\ldots\ldots(1),$$

$$u_{x+1} \sin a_{x+1} = e u_x \sin a_x \dots (2);$$

and, by the properties of the motion of projectiles,

$$\Delta t_x = \frac{2}{g} u_{x,1} \sin a_{x,1} \dots (3),$$

$$\Delta s_x = u_{x+1} \cos a_{x+1} \Delta t_x \ldots (4).$$

From (1) it is evident that

$$u_z \cos a_z = u_1 \cos a_1 \ldots (5),$$

where u_1 is the given velocity and a_1 the given angle of projection.

Again, from (2), putting $u_x \sin a_x = v_x$, we have

$$\Delta v_{s} = (e-1) v_{s};$$

and therefore, integrating,

$$v_z = C e^x$$
,

where C is an arbitrary constant. But x = 1, $v_x = v_1$, simultaneously; hence $Ce = v_1$, and therefore

$$v_x = v_1 e^{x-1}, \quad u_x \sin a_x = u_1 \sin a_1 e^{x-1}.......(6).$$

From (5) and (6) we get

$$\tan a_x = \tan a_1 \cdot e^{x-1}, \quad u_x = u_1 \cos a_1 \cdot \left\{1 + \tan^2 a_1 \cdot e^{2(x-1)}\right\}^{\frac{1}{2}},$$

by which the circumstances of the projection are determined for each of the parabolic paths.

Again, from (3) and (6),

$$\Delta t_x = \frac{2}{g} u_1 \sin a_1 e^x \dots (7);$$

integrating and adding a constant,

$$t_{z} = \frac{2u_{1} \sin a_{1}}{q} \frac{e^{z}}{e-1} + C;$$

but $t_0 = 0$, hence

$$0 = \frac{2u_1 \sin a_1}{g} \frac{1}{e-1} + C,$$

and therefore

$$t_x = \frac{2u_1 \sin a_1}{g} \frac{e^x - 1}{e - 1}$$
.

Again, from (4) and (7)

$$\Delta s_x = \frac{2}{\sigma} u_1 \sin a_1 \cdot e^x \cdot u_{x+1} \cos a_{x+1},$$

and therefore, by (5),
$$\Delta s_x = \frac{u_1^2 \sin 2a_1}{g} e^x,$$

whence, integrating and observing that $s_0 = 0$, we shall have

$$s_{x} = \frac{u_{1}^{2} \sin 2a_{1}}{q} \frac{1 - e^{x}}{1 - e}.$$

Bordoni; Memorie della Societa Italiana, tom. XVII. P. 1. p. 191; 1816.

(8) A particle is projected obliquely from a point A at an angle a with the horizon, so as to hit a point B, AB being inclined at an angle β to the horizon; and the velocity of projection is such that with it the particle would describe the straight line AB uniformly in n seconds; to find the time of flight.

Required time = $n \frac{\cos \beta}{\cos \alpha}$.

(9) If two particles be projected from the same point, at the same instant, with velocities v and v', and in directions a, a', to find the time which elapses between their transits through the other point which is common to both their paths.

Time =
$$\frac{2}{g} \frac{vv' \sin (a \sim a')}{v \cos a + v' \cos a'}$$
.

(10) If a be the angle between the two tangents at the extremities of any arc of the parabolic path of a particle acted on by gravity; v, v', the velocities at these two points, and v_1 the velocity at the vertex; to find the time through the arc.

The required time =
$$\frac{vv' \sin a}{gv_1}$$
.

(11) A particle describes an ellipse under the action of a force at right angles to the axis major; to find the force at any point of the path.

Let a, b, be the semiaxes major and minor, y the distance of the particle at any point of its path from the axis major, β the velocity of the particle parallel to the axis major which will remain invariable during the whole motion. Then

force required =
$$\frac{b^4 \beta^2}{a^2 y^3}$$
.

If b = a, or the ellipse become a circle,

force =
$$\frac{a^2\beta^2}{y^3}$$
.

Riccati; Comment. Bonon. tom IV. p. 149; 1757. Newton; Principia, lib. 1. sect. 2, prop. 8.

(12) A particle describes the arc of a cycloid under the action of a force parallel to its base; to find the law of the force.

If the equations to the cycloid be

$$x = a \text{ vers } \theta, y = a (\theta + \sin \theta),$$

and F, β , denote the force required and the velocity parallel to the axis of the cycloid,

$$\frac{1}{F} = \frac{a}{2\beta^2} (2 \sin \theta - \sin 2\theta).$$

(13) A particle is projected with a given velocity parallel to a given straight line towards which it is always attracted with a force proportional to its perpendicular distance from it; to determine the position of the particle at any time and the equation to its path.

Let A (fig. 102) be the initial position of the particle; Ox the given straight line, yAO at right angles to Ox; Ox, Oy, the

axes of x, y; P the position of the particle after a time t; OM = x, PM = y; AO = b, β = the velocity of projection; μ^2 the absolute force of attraction.. Then

$$x = \beta t$$
, $b \cos \frac{\mu x}{\beta} = y = b \cos (\mu t)$.

Riccati; Comment. Bonon. tom. IV. p. 155; 1757.

(14) A particle is projected from a point x = 0, y = b, with a velocity β parallel to the axis of x, and is subject to the action of a force tending towards the axis of x parallel to the axis of y, and varying inversely as the square of the distance; to find the equation to the path of the particle.

Let μ denote the attracting force at a unit of distance; then the equation to the path will be

$$\left(\frac{2\mu}{b}\right)^{\frac{1}{4}}\frac{x}{\beta} = \frac{1}{2}b\left\{\pi - \operatorname{vers}^{-1}\frac{2y}{b}\right\} + (by - y^{2})^{\frac{1}{4}}.$$
Riccati; *Ib.* p. 159.

(15) A particle (fig. 103) is projected with a given velocity in the direction Oy, and is acted on by a centre of force, which attracts directly as the distance and moves uniformly with a given velocity along Ox at right angles to Oy; to determine the position of the particle when its motion first becomes parallel to Ox.

Let μ^2 denote the absolute force; a the initial distance of the centre of force from O; β the velocity with which the particle is projected, β' the uniform velocity of the centre of force along Ox, and x', y', the co-ordinates of the required position of the particle; then

$$x' = a + \frac{\beta'}{2\mu}(\pi - 2), y' = \frac{\beta}{\mu}.$$

(16) A particle which is placed at rest initially in a given position, is acted on by two forces, one repulsive and varying as the distance from a given point, and the other constant and acting in parallel lines; to determine the position of the particle at any time and the equation to its path.

Let the centre of the central force be taken as the origin of co-ordinates, and let the directions of the axes be so chosen that the direction of the constant force makes an angle of 45° with each of them. Then if a, b, be the co-ordinates of the initial position of the particle, μ^2 the absolute force of repulsion, and f the constant force, we shall have, putting $f = 2^{\frac{1}{4}}\mu^2 m$,

$$\frac{x+m}{a+m}=\frac{1}{2}\left(\dot{\epsilon^{\mu t}}+\dot{\epsilon^{-\mu t}}\right)=\frac{y+m}{b+m}.$$

(17) Four equal particles attracting directly as the distance are fixed in the corners of a square; to find the path of a particle projected from any point and in any direction in the plane of the square.

Let the centre of the square be taken as the origin of co-ordinates, and let the axes be at right angles to the two pairs of opposite sides of the square. Then if a, b, be the co-ordinates of the initial position of the particle, 2m, 2n, the resolved parts of its velocity of projection parallel to the axes of x, y, respectively, and μ^2 the absolute force of attraction of each of the fixed particles, the equation to the path of the free particle will be

$$\sin^{-1}\frac{\mu y}{(n^2+\mu^2b^2)^{\frac{1}{4}}}-\sin^{-1}\frac{\mu b}{(n^2+\mu^2b^2)^{\frac{1}{4}}}=\sin^{-1}\frac{\mu x}{(m^2+\mu^2a^2)^{\frac{1}{4}}}-\sin^{-1}\frac{\mu a}{(m^2+\mu^2a^2)^{\frac{1}{4}}}.$$

(18) A spherical particle, of which e is the elasticity, is projected with a velocity v at an angle of elevation a, and at the instant of attaining its greatest altitude strikes a similar and equal particle falling downwards with a velocity $\frac{1}{2}v$ at the point of collision; to find the distance of the particles from each other at the end of t seconds after the impact.

Distance required =
$$\frac{1}{2} vt \left(1 + 4e^2 \cos^2 a\right)^{\frac{1}{4}}$$
.

(19) A heavy particle, having been projected at a given angle to the inclined plane AB, (fig. 104), proceeds to ascend this plane by bounding in a series of parabolic arcs; to determine the angles of incidence and reflection after any number of impacts.

Let ι be the inclination of the plane AB to the horizon; a_x the angle of reflection in the x^{th} arc, β_{x-1} the angle of incidence in the $(x-1)^{th}$; and e the elasticity of the particle. Then

$$\tan \, a_s = \frac{(1-e) \, e^{s-1} \, \tan \, a_1}{1-e-2 \, (1-e^{s-1}) \, \tan \, a_1} = e \, \tan \, \beta_{s-1}.$$

Bordoni; Memorie della Societa Italiana, tom. xvII. P. 1, p. 191; 1816.

(20) A ball, of which the elasticity is e, is projected with a velocity V in a direction making an angle $a + \iota$ with the horizon, and rebounds from a plane inclined to the horizon at an angle ι and passing through the point of projection. To determine the relation between R_s , R_{s-1} , R_{s-2} , three consecutive ranges upon the inclined plane after x, x+1, x+2, rebounds respectively, and to find the sum of all the ranges on the inclined plane before the ball begins to slide down the plane.

If $\cot \beta = (1 - e) \cot \alpha$, and S denote the sum of the ranges;

$$R_{s+2} - (e + e^2) R_{s+1} + e^3 R_s = 0,$$

$$S = \frac{2 V^2 \sin \beta \sin \alpha \cos (\alpha + \beta)}{g \sin \alpha \cdot \cos^2 \beta}.$$

SECT. 2. Central Forces.

Let the force which acts on a particle tend always towards a fixed centre, which we will take as the origin of co-ordinates. Let F denote the force at any distance from the centre; x, y, the co-ordinates of the particle at the end of a time t reckoned from an assigned epoch, r its distance from the centre of force, and θ the inclination of this distance to any fixed line in the plane of x, y. Then, by Maclaurin's Equations, the plane of co-ordinates being identical with the plane of the motion,

$$\frac{d^2x}{dt^2} = -\frac{Fx}{r}, \quad \frac{d^2y}{dt^2} = -\frac{Fy}{r}.$$

From these equations may be obtained the following formulæ:

$$r^2d\theta = hdt$$
 (I),
 $v = \frac{h}{p}$ (II),

$$v^{2} = h^{2} \left\{ \left(\frac{d}{d\theta} \frac{1}{r} \right)^{2} + \frac{1}{r^{3}} \right\} \dots (III),$$

$$v^{2} = v^{2} - 2 \int_{r'}^{r} F dr \dots (IV),$$

$$F = \frac{h^{2}}{r^{3}} \left\{ \frac{d^{2}}{d\theta^{2}} \left(\frac{1}{r} \right) + \frac{1}{r} \right\} \dots (V),$$

$$F = \frac{h^{2}}{p^{3}} \frac{dp}{dr} \dots (VI),$$

$$F = r \frac{d\theta^{2}}{dt^{2}} - \frac{d^{2}r}{dt^{2}} \dots (VII).$$

In these formulæ h represents twice the area swept out by the radius vector about the centre of force in a unit of time, p the perpendicular from the centre upon the tangent at any point of the orbit, v the velocity of the particle, and v', r', any simultaneous values of v, r. If the central force, instead of being attractive as we have been supposing, be repulsive, we must replace F in these formulæ by -F.

The equation (I) shews that the area swept out by the radius vector varies as the time, and either of the equations (II) and (III) that the velocity at any point of the orbit varies inversely as the perpendicular from the centre of force upon the tangent to the orbit at that point: these two propositions were first established by Newton¹. The equation (IV) shews that the velocity of the particle at any point of its path depends only upon the distance of the point from the centre, the velocity of projection and the prime radius vector, whatever be the course which it may have pursued; the discovery of this proposition is likewise due to Newton². The formula (V), by which the path of the particle may be determined when we know the law of the central force and conversely, Ampére' ascribes to Binet. The formula (VI) was communicated without demonstration to John Bernoulli by De Moivre in the year 1705; a proof of the formula was returned to him by Bernoulli in a letter dated

¹ Principia, lib. 1. Prop. 1.

¹ Ib. lib. t. Prop. 40.

² Annales de Gergonne, tom. xx. p. 53.

Basle, Feb. 1706. The formula (VII) was given much about the same time by Clairaut¹ and by Euler², and signifies that the acceleration of the radius vector is equal to the excess of the centrifugal above the attractive force.

(1) To find the law of the force by which a particle may be made to describe the Lemniscata of James Bernoulli, the centre of force coinciding with the node, and to investigate the time of describing one of the ovals.

The polar equation to the Lemniscata is

hence
$$\frac{1}{r} = \frac{1}{a(\cos 2\theta)^{\frac{1}{4}}}, \quad \frac{d}{d\theta} \left(\frac{1}{r}\right) = \frac{\sin 2\theta}{a(\cos 2\theta)^{\frac{3}{4}}},$$
$$\frac{d^2}{d\theta^2} \left(\frac{1}{r}\right) = \frac{2}{a(\cos 2\theta)^{\frac{1}{4}}} + \frac{3(\sin 2\theta)^2}{a(\cos 2\theta)^{\frac{3}{4}}} = \frac{3}{a(\cos 2\theta)^{\frac{3}{4}}} - \frac{1}{a(\cos 2\theta)^{\frac{3}{4}}},$$

and therefore

$$\frac{d^2}{d\theta^2}\left(\frac{1}{r}\right) + \frac{1}{r} = \frac{3}{a\left(\cos \frac{2\theta}{r}\right)^{\frac{4}{3}}} = \frac{3a^4}{r^5}.$$

Hence, by the formula (V),

$$F = \frac{3a^4h^2}{r^7}.$$

Again, by the formula (I) and the equation (1), we have $hdt = r^2d\theta = a^2\cos 2\theta \ d\theta,$

and therefore, if P denote the required periodic time,

$$P = \frac{a^2}{h} \int_{-1\pi}^{+\frac{1}{4}\pi} \cos 2\theta \ d\theta = \frac{a^2}{h}.$$

Let μ denote the value of F when r = 1; then we have

$$\mu = 3a^4h^2$$
, $h = \frac{\mu^{\frac{1}{4}}}{3^{\frac{1}{4}}a^2}$, $P = \frac{3^{\frac{1}{4}}a^4}{\mu^{\frac{1}{4}}}$.

(2) A particle moves in an equiangular spiral under the action of a force tending towards the pole; to find the law of force and the velocity at any point of the orbit.

¹ Théorie de la Lune, p. 2; the first edition of which appeared in 1752, from a MS. sent to St. Petersburg in 1750.

^{*} Nov. Comment. Petrop. 1752, 1753, p. 164.

If β be the invariable angle, r the radius vector, and p the perpendicular from the pole upon the tangent,

$$p = r \sin \beta \ldots \ldots (1).$$

Differentiating with respect to r, we have $\frac{dp}{dr} = \sin \beta$, and therefore, from (VI),

$$F = \frac{h^2}{p^3} \sin \beta = \frac{h^2}{r^3 \sin^2 \beta}$$
, by (1), (2).

Let c be the velocity corresponding to a given radius vector r'; then, by (II) and (1),

$$h = cr' \sin \beta$$
.

Hence, from (2),

$$\cdot F = \frac{c^2 r'^2}{r^3} ,$$

and, from (II) and (1),

$$v = \frac{cr' \sin \beta}{r \sin \beta} = \frac{cr'}{r}.$$

(3) A particle describes an equilateral hyperbola round a centre of force situated in the centre; to find the law of the force and the angle which the particle will describe about the centre from the apse in a given time.

The equation to the hyperbola being

$$\left(\frac{1}{r}\right)^2 = \frac{\cos 2\theta}{a^2} \dots \dots (1),$$

we have

$$\frac{1}{r}\frac{d}{d\theta}\left(\frac{1}{r}\right) = -\frac{\sin 2\theta}{a^2} \dots (2),$$

$$\frac{1}{r}\frac{d^2}{d\theta^2}\left(\frac{1}{r}\right) + \left\{\frac{d}{d\theta}\left(\frac{1}{r}\right)\right\}^2 = -\frac{2\cos 2\theta}{a^2};$$

and therefore, from (2),

$$\frac{1}{r}\frac{d^2}{d\theta^2}\left(\frac{1}{r}\right)+\frac{r^2}{a^4}\sin^2 2\theta=-\frac{2\cos 2\theta}{a^2},$$

and thence, by (1),

$$\frac{1}{r}\frac{d^2}{d\theta^2}\left(\frac{1}{r}\right) + \frac{r^2}{a^4} - \frac{1}{r^2} = -\frac{2}{r^2},$$

$$\frac{d^2}{d\theta^2}\binom{1}{r}+\frac{1}{r}=-\frac{r^3}{a^4}.$$

Hence, by the formula (V), $F = -\frac{h^2 r}{a^4}.$

$$F = -\frac{h^2r}{a^4}.$$

The negative sign shews that the force must be repulsive; let $-\mu$ be the absolute force, that is, the value of F at a unit of distance. Then

 $\mu = \frac{h^2}{a^4}, \quad F = -\mu r.$

Putting for h its value $a^2\mu^{\frac{1}{4}}$ in the formula (I), and $\frac{a^2}{\cos 2\theta}$ for r^2 , we get

$$\frac{d\theta}{\cos 2\theta} = \mu^{\frac{1}{4}}dt, \quad \frac{\cos 2\theta \ d\theta}{1 - \sin^2 2\theta} = \mu^{\frac{1}{4}}t, \quad \frac{d \sin 2\theta}{1 - \sin^2 2\theta} = 2\mu^{\frac{1}{4}}dt;$$

integrating, and supposing the time to be reckoned from apsidal passage, we have

$$\frac{1}{2}\log\frac{1+\sin 2\theta}{1-\sin 2\theta}=2\mu^{\frac{1}{4}}t,$$

whence, writing μ' in place of $4\mu^{\frac{1}{2}}$, we obtain

$$\sin 2\theta = \frac{\epsilon^{\mu''}-1}{\epsilon^{\mu''}+1}.$$

(4) A particle is revolving in a parabola about a centre of force in the focus, and when it arrives at a given distance from the focus, the absolute force is suddenly doubled; to determine the nature of the subsequent path of the particle.

Let 4m be the latus rectum of the parabola, r the radius vector at any point, and p the perpendicular from the focus upon the tangent. Then, by the nature of the parabola,

$$\frac{1}{p^2} = \frac{1}{mr}$$
, $\frac{2}{p^3} \frac{dp}{dr} = \frac{1}{mr^2}$,

and therefore, by (VI),

$$F = \frac{h^2}{2mr^2}.$$

But, after the absolute force has been doubled, we shall have for the motion

$$F = \frac{h^2}{mr^2},$$

and therefore, by (VI),

$$\frac{h^2}{mr^2} = \frac{h^2}{p^3} \frac{dp}{dr}, \quad \frac{1}{mr^2} = \frac{1}{p^3} \frac{dp}{dr}.$$

Integrating, we have

$$\frac{1}{mr}=C+\frac{1}{2p^2}.$$

Let c be the value of r at the instant when the absolute force is doubled; then, p being then common both to the parabola and to the new path, we have

$$\frac{1}{mc}=C+\frac{1}{2mc}, \quad C=\frac{1}{2mc},$$

and therefore, for the equation to the new path, there is

$$\frac{1}{mr} = \frac{1}{2mc} + \frac{1}{2p^2}, \quad \frac{mc}{p^2} = \frac{2c}{r} - 1,$$

which is the equation to an ellipse, 2c being the major and $(mc)^{\frac{1}{c}}$ the minor axis.

Since the ellipse touches the parabola when r = c, the semi-axis major, it follows from the nature of the ellipse that the point of contact is an extremity of the semi-axis minor, and therefore that the axis major of the ellipse is parallel to the tangent at the point r = c of the parabola. But the sine of the angle of inclination of the tangent of the parabola at this point to its axis is

$$\frac{p}{r} = \frac{m^{\frac{1}{4}}}{r^{\frac{1}{4}}}$$
, when $r = c$, that is, $= \frac{m^{\frac{1}{4}}}{c^{\frac{1}{4}}}$, and therefore the inclination

of the major axis of the ellipse to the axis of the parabola is

$$\sin^{-1}\left\{\frac{m^{\frac{1}{4}}}{c^{\frac{1}{4}}}\right\}.$$

(5) A particle is describing a curve about a certain centre of force, the velocity of the particle varying inversely as the n^{th} power of its distance from the centre of force; to find the law of the force and the equation to the path.

We shall have, μ denoting some constant quantity,

$$v = \frac{\mu}{r^n}$$
.

Hence, from (IV), there is

$$\frac{\mu^2}{r^{2n}} = v'^2 - 2 \int_{r'}^{r} F dr,$$

and therefore, differentiating,

$$F=\frac{n\mu^2}{r^{2n+1}},$$

which determines the law of the force.

Again, from (III) there is

$$\frac{\mu^{2}}{r^{2n}} = h^{2} \left\{ \left(\frac{d}{d\theta} \frac{1}{r} \right)^{2} + \frac{1}{r^{2}} \right\},$$

$$\left(\frac{\mu^{2} - h^{2} r^{2n-2}}{r^{2n}} \right)^{\frac{1}{4}} = h \frac{d}{d\theta} \frac{1}{r},$$

$$(n-1) d\theta = -\frac{dr^{n-1}}{\left(\frac{\mu^{2}}{h^{2}} - r^{2n-2} \right)^{\frac{1}{4}}},$$

$$(n-1) \theta = C + \cos^{-1} \frac{hr^{n-1}}{\mu}.$$

Suppose $\theta = 0$, when r = a; then, k denoting a constant quantity, $(n-1)\theta = \cos^{-1}(kr^{n-1}) - \cos^{-1}(ka^{n-1}).$

Riccati; Comment. Bonon. tom. IV. p. 184.

(6) If the force vary as the n^{th} power of the distance, and a particle be projected from an apsidal distance, with a velocity of which the square is equal to $1 - \varepsilon$ times the square of the velocity in a circle about the same centre of force with a radius equal to the apsidal distance; to find the equation to the orbit, ε being a very small quantity.

Let a be the apsidal distance; then r = a - x, where x is a small quantity, because the path of the particle, as is evident from the initial circumstances, will be nearly circular. Then, approximately,

$$\frac{1}{r} = \frac{1}{a-x} = \frac{1}{a} \left(1 + \frac{x}{a} \right), \quad \frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) = \frac{1}{a^2} \frac{d^2x}{d\theta^2}.$$

Also, μ denoting the absolute force of attraction,

$$\frac{Fr^2}{h^2} = \frac{\mu r^{n+2}}{h^2} = \frac{\mu}{h^2} (a-x)^{n+2} = \frac{\mu a^{n+2}}{h^2} \left\{ 1 - \frac{(n+2)x}{a} \right\}.$$

Hence, by the formula (V),

$$\frac{1}{a^2}\frac{d^2x}{d\theta^2} + \frac{x}{a^2} + \frac{1}{a} - \frac{\mu a^{n+2}}{h^2} \left\{ 1 - \frac{(n+2)x}{a} \right\} = 0,$$

$$\frac{d^2x}{d\theta^2} + \left\{ 1 + (n+2)\frac{\mu a^{n+3}}{h^2} \right\} x + a - \frac{\mu a^{n+4}}{h^2} = 0 \dots (1).$$

Let V be the velocity of projection, and v the velocity in a circle about the same centre of force with a radius a; then

$$V^2 = (1 - \varepsilon) v^2 = (1 - \varepsilon) \mu a^{n+1} \dots (2).$$

But, by the formula (II),

$$h^2=a^2V^2,$$

because the motion is initially at right angles to the radius vector, and a, V, are the initial values of the radius vector and of the velocity. Hence, from (2),

$$h^{2} = \mu (1 - \varepsilon) a^{n+3}, \ \frac{\mu}{h^{2}} = \frac{1}{(1 - \varepsilon) a^{n+3}},$$

and therefore, from (1), the product of ε and x being neglected,

$$\frac{d^3x}{d\theta^2}+(n+3)x-\epsilon a=0.$$

Multiplying by 2 $\frac{dx}{d\theta}$ and integrating,

$$\frac{dx^2}{d\theta^2} + (n+3)x^2 - 2\varepsilon ax = 0,$$

no constant being added because $\frac{dx}{d\theta}$ is by hypothesis equal to zero when x = 0; hence

$$d\theta = \frac{dx}{(n+3)^{\frac{1}{2}} \left(\frac{2\epsilon a}{n+3} x - x^2\right)^{\frac{1}{2}}}.$$

Integrating and considering zero to be the initial value of θ ,

$$(n+3)^{\frac{1}{4}}\theta = \operatorname{vers}^{-1}\frac{(n+3)x}{4},$$

whence for the polar equation to the curve we have

$$r = a - \frac{\epsilon_a}{n+3} \text{ vers } \{(n+3)^{\frac{1}{4}}\theta\} \dots (3).$$

The general condition for an apsidal distance is evidently that $\frac{dr}{d\theta}$ shall be equal to zero: differentiating the equation (3), we get for the determination of apsides,

$$\frac{dr}{d\theta} = -\frac{\epsilon a}{(n+3)^{\frac{1}{4}}} \sin \left\{ (n+3)^{\frac{1}{4}} \theta \right\} = 0.$$

Hence

$$(n+3)^{\frac{1}{4}}\theta=\lambda\pi,$$

where λ is any integer: let θ' , θ'' , be the values of θ for two consecutive apsidal distances; then

$$(n+3)^{\frac{1}{3}}\theta'=\lambda\pi,\ (n+3)^{\frac{1}{2}}\theta''=(\lambda+1)\pi,$$

and therefore, ϕ denoting the angle between any two consecutive apsides,

 $\phi = \theta' - \theta' = \frac{\pi}{(n+3)^{\frac{1}{4}}}.$

(7) A particle, attached to one end of a slightly extensible string which is extended to its natural length and has its other end fixed, is projected at right angles to it; to determine the extension of the string at any time and the path of the particle, the string resting on a smooth horizontal plane and being indefinitely fine.

Let a, c, denote the initial length of the string and the initial velocity of the particle. Let r = a + x represent the length of the string at the end of any time t, the corresponding angular co-ordinate being θ , and the initial position of the string the prime radius vector. Let m denote the mass of the particle, and p the tension necessary to stretch the string to a length $a + \beta$. Then, the extension being, by Hooke's law, proportional to the tension,

we have for the tension at the time t the expression $\frac{px}{B}$. Hence,

by the formula (VII), we have

$$\frac{d^2r}{dt^2}=r\,\frac{d\theta^2}{dt^2}-\frac{px}{m\beta}\,,$$

and therefore, by (I),
$$\frac{d^2r}{dt^2} = r\frac{d\theta}{dt^2} - \frac{px}{m\beta}$$

$$\frac{d^2r}{dt^2} = \frac{h^2}{r^3} - \frac{px}{m\beta}.$$
But from the circumstances of the projection

But from the circumstances of the projection of the particle it is clear by the formula (II) that h = ac; hence

$$\frac{d^2r}{dt^2} = \frac{a^2c^2}{r^3} - \frac{px}{m\beta},$$

$$\frac{d^2x}{dt^2} = \frac{a^2c^2}{(a+x)^3} - \frac{px}{m\beta}.$$

Multiplying by $2 \frac{dx}{dt}$ and integrating,

$$\frac{dx^2}{dt^2} = C - \frac{a^2c^2}{(a+x)^2} - \frac{px^3}{m\beta}:$$

but, initially, x=0, $\frac{dx}{dt}=0$, hence $C=c^2$, and therefore

$$\frac{dx^{3}}{d\ell} = c^{3} - \frac{a^{2}c^{3}}{(a+x)^{3}} - \frac{px^{2}}{m\beta}$$

$$= c^{3} - c^{3}\left(1 - \frac{2x}{a}\right) - \frac{px^{3}}{m\beta}$$

$$= \frac{2c^{3}x}{a} - \frac{px^{3}}{m\beta} \cdot \dots (1),$$

where small quantities of the first order only are retained, $\frac{x}{\beta}$. x being of the first order because x and β are of the same order. Extracting the root, we have

$$\frac{dt}{dx} = \left(\frac{m\beta}{p}\right)^{1}_{i} \left(2 \frac{m\beta c^{2}}{ap} x - x^{2}\right)^{-\frac{1}{4}};$$

whence, integrating and bearing in mind that x = 0 when t = 0,

$$t = \left(\frac{m\beta}{p}\right)^{\frac{1}{4}} \text{ vers}^{-1} \left(\frac{pax}{mc^{2}\beta}\right),$$
$$x = \frac{mc^{2}\beta}{pa} \text{ vers} \left\{\left(\frac{p}{m\beta}\right)^{\frac{1}{4}} t\right\},$$

or

which gives the extension of the string at any time during the motion.

We proceed now to the determination of the equation to the path of the particle. From (I) and (1) there is

$$\frac{dx^2}{d\theta^2} = \frac{r^4}{h^2} \left(\frac{2c^2x}{a} - \frac{px^2}{m\beta} \right)$$

$$= \frac{(a+x)^4}{a^2c^2} \left(\frac{2c^2x}{a} - \frac{px^2}{m\beta}\right)$$

$$= 2ax - \frac{pa^2x^2}{mc^2\beta}, \text{ approximately,}$$

$$= \frac{pa^2}{mc^2\beta} \left(2\frac{mc^2\beta}{pa}x - x^2\right),$$

$$\frac{d\theta}{dx} = \left(\frac{mc^2\beta}{pa^2}\right)^{\frac{1}{4}} \left(2\frac{mc^2\beta}{pa}x - x^2\right)^{-\frac{1}{4}}.$$

Integrating, and remembering that $\theta=0$, x=0, simultaneously,

$$\theta = \left(\frac{mc^2\beta}{pa^2}\right)^{\frac{1}{2}} \text{ vers}^{-1} \frac{pax}{mc^2\beta},$$

$$r = a + x = a + \frac{mc^2\beta}{pa} \text{ vers} \left\{ \left(\frac{pa^2}{mc^2\beta}\right)^{\frac{1}{2}} \theta \right\} \dots (2),$$

which is the polar equation to the orbit.

From this equation we have, for the determination of apsides,

$$\frac{dx}{d\theta} \propto \sin \left\{ \left(\frac{pa^2}{mc^2\beta} \right)^{\frac{1}{4}} \theta \right\} = 0:$$

hence the angle between consecutive apsides is

$$\frac{\pi c}{a} \left(\frac{m\beta}{p} \right)^{\frac{1}{4}}$$
.

In the solution of this problem James Bernoulli takes as the approximate equation of motion

$$\frac{d^2r}{dt^2}=\frac{c^2}{\rho}-\frac{px}{m\beta},$$

where ρ denotes the radius of curvature at any point of the orbit. The difference, however, between $\frac{c^2}{\rho}$ and $\frac{\hbar^2}{r^3}$ is of the first order of small quantities, and therefore his approximate equation is essentially erroneous. Instead of the equation (2) he gets $r = a + \frac{mc^2\beta}{na} \text{ vers } \left\{ \left(\frac{pa^2}{2mc^2\beta} \right)^{\frac{1}{2}} \theta \right\},$

which makes the angle between the apsides greater than it should be in the ratio of $2^{\frac{1}{4}}$ to 1.

James Bernoulli; Nov. Act. Acad. Petrop. 1783, p. 213.

(8) A particle moves in a spiral, of which the equation is $r = a \left(\sec \frac{\theta}{n} \right)^n$, about a centre of force in the pole; to find the law of the force.

If μ denote the absolute force, then

$$F=\frac{\mu}{r^{\frac{3n-2}{n}}}.$$

(9) To find the law of force by which the Cissoid of Diocles may be described about a centre of force in the cusp.

If θ be the angle which the radius vector r makes with the axis, then $\frac{1}{r} \cos^2 \theta$

 $F \propto \frac{\csc^2 \theta}{r^3}$.

(10) A particle revolves in a circle about a centre of force situated at a point in its circumference; to determine the force and the velocity at any point of the path.

If μ denote the absolute force,

$$F=\frac{\mu}{r^5},\ v^2=\frac{\mu}{2r^4}.$$

Newton; Princip. lib. 1. Prop. 7. Riccati; Comment. Bonon. tom. 1v. p. 175.

(11) A particle is moving about a centre of force, its velocity at each point of its path varying inversely as its distance from the centre of force; to determine the path of the particle.

The path will be a logarithmic spiral.

Riccati; Ib. p. 184.

(12) A particle, attracted towards a point by a force equal to $\frac{r}{m^2} + \frac{h^2}{r^3}$, is projected from an apse at the distance $m^{\frac{1}{4}} h^{\frac{1}{4}}$, h being twice the area described in a unit of time; to find the polar equation to the orbit, and the time of describing any angle about the centre of force.

If θ be the angle described about the centre of force in any time t after the projection,

$$r^2 = \frac{mh}{1+\theta^2}, \ t = m \tan^{-1} \theta.$$

(13) At a distance a from a centre of force, a particle is projected at an angle of 45° to the distance, and with a velocity which is to that in a circle at the same distance as $2^{\frac{1}{2}}$ to $3^{\frac{1}{2}}$; the central force at any distance r is equal to $\frac{2a^2n}{r^5} + \frac{n}{r^3}$; to find the equation to the orbit.

If the angular co-ordinate be estimated from the initial radius vector, the equation to the orbit will be

$$r=a(1-\theta).$$

(14) A particle is revolving in a circle about a centre of force in its geometrical centre, the force varying inversely as the square of the distance; the absolute force is suddenly increased in the ratio of m to 1 when the particle is at any assigned point of its path, and when the particle arrives again at the same point, the absolute force is again increased in the same ratio: to find the path which the particle will describe after the second change in the absolute force.

The path will be an ellipse of which the eccentricity is $\frac{m^2-1}{m^2}$.

(15) A particle is revolving in an ellipse about a centre of force in the focus; supposing that, every time the particle arrives at the lower apse, the absolute force is diminished in the ratio of 1 to 1-n, to find the eccentricity of the elliptic orbit after p revolutions, the original eccentricity being e.

The required eccentricity is equal to

$$\frac{1}{(1-n)^p}-1.$$

(16) An indefinitely fine clastic string, extended to its natural length, is fastened at one end and has a particle of matter attached to the other; the particle is projected at right angles to the string with an angular velocity such that, if it were revolving in a circle with this angular velocity, the length of the string must have been stretched to twice its natural length; to find the path which the particle will ultimately describe after an indefinite number of revolutions.

If a be the natural length of the string, the particle will ultimately move in a circle the value of the radius r of which is a root of the equation $a^{2} \qquad r = a$

 $\frac{a^3}{r^2}=2\frac{r-a}{r+a}.$

(17) If r be the radius vector at any point of a curve described round a centre of force which varies inversely as the square of the distance; and c, c', be the external and internal spaces due to the velocity at that point respectively; to find the chord of curvature at that point drawn through the centre of force. c - c'

Chord of curvature = $4r \frac{c-c'}{c+c'}$.

(18) A particle moves in an equilateral hyperbola about a centre of force in the centre; to find the locus of the points to which the particle must move from the curve under the action of the force to acquire the velocity in the curve.

If a be the semiaxis of the equilateral hyperbola in which the particle is moving, the required locus will also be an equilateral hyperbola having a semiaxis equal to $2^{\frac{1}{4}}$. a, the centres of the two hyperbolas being coincident and their axes in the same straight lines.

(19) A particle, acted on by a central force varying as any function of the distance, is projected at an apse with a velocity nearly equal to that requisite for a circular orbit; to find the distance, from the centre of force, of the apse at which the particle next arrives.

Let the force at any distance r be equal to $\frac{1}{r^2}\phi\left(\frac{1}{r}\right)$, where $\phi\left(\frac{1}{r}\right)$ is any function of $\frac{1}{r}$; let a be the initial distance of the particle from the centre of force, a' the distance of the apse at which it next arrives, and let the velocity of projection be to the velocity in a circle about the same centre as 1 to 1 + m; then

$$a' = a \left\{ 1 - \frac{4m\phi\left(\frac{1}{a}\right)}{\phi\left(\frac{1}{a}\right) - \frac{1}{a}\phi'\left(\frac{1}{a}\right)} \right\}.$$

SECT. 3. Tangential and Normal Resolutions.

The method of the solution of the general problem of the curvilinear motion of a particle in one plane, by the principle of the tangential and normal resolutions, is expressed by the equations

 $v\frac{dv}{ds}=T\ldots\ldots(A)_{s}$

$$\frac{v^2}{\rho}=N.....(B),$$

where v denotes the velocity at any point of the path, ds an element of the curve, ρ the radius of curvature, T the sum of the resolved parts of the forces which act on the particle estimated in the direction of its motion, and N the sum of the resolved parts along the normal on the concave side of the curve in the neighbourhood of the particle.

(1) A particle is projected with a given velocity and in a given direction, and is acted upon by a constant force in parallel lines; to determine the path of the particle.

Let the axis of x be taken so as to pass through the initial place of the particle, and let the axis of y be taken parallel to the constant force which acts towards the axis of x. Let f denote the constant force. Then the tangential resolved part

being $-f\frac{dy}{ds}$, and the normal one being $f\frac{dx}{ds}$, we have for the

motion of the particle,

$$v\frac{dv}{ds}=-f\frac{dy}{ds}\ldots\ldots(1),$$

$$\frac{v^2}{\rho} = f \frac{dx}{ds} \dots (2).$$

Integrating (1),

$$v^2 = C - 2fy.$$

Let V be the initial velocity; then, y being zero initially, $V^2 = C$, and therefore $v^2 = V^2 - 2fy$.

Hence, substituting this expression for v^2 in (2),

$$\frac{1}{\rho}(V^2-2fy)=f\frac{dx}{ds};$$

but
$$\rho = -\frac{\frac{ds^3}{dx^3}}{\frac{d'y}{dx^2}},$$
 hence
$$-\frac{d^2y}{dx^3} (V^2 - 2fy) = f \frac{ds^2}{dx^2} = f \left(1 + \frac{dy^2}{dx^2} \right),$$

$$(V^3 - 2fy) \frac{d}{dx} \left(1 + \frac{dy^2}{dx^2} \right) - \left(1 + \frac{dy^2}{dx^2} \right) \frac{d}{dx} (V^2 - 2fy) = 0.$$
 Integrating, we have
$$C \left(1 + \frac{dy^2}{dx^2} \right) = V^2 - 2fy,$$

where C is an arbitrary constant. Let β be the angle which the direction of projection makes with the axis of x; then

$$C(1 + \tan^2 \beta) = V^2:$$
hence
$$V^2 \left(1 + \frac{dy^2}{dx^2}\right) = \sec^2 \beta \left(V^2 - 2fy\right),$$

$$V^2 \frac{dy^2}{dx^2} = V^2 \tan^2 \beta - 2f \sec^2 \beta y,$$

$$V dy = \left(V^2 \tan^2 \beta - 2f \sec^2 \beta y\right)^{\frac{1}{2}} dx;$$

whence, by integration,

$$C - V(V^2 \tan^2 \beta - 2f \sec^2 \beta y)^{\frac{1}{4}} = f \sec^2 \beta x.$$

But x = 0, y = 0, simultaneously; hence

$$C-V^2\tan\beta=0,$$

and therefore

$$V^2 \tan \beta - V(V^2 \tan^2 \beta - 2f \sec^2 \beta y)^{\frac{1}{2}} = f \sec^2 \beta x.$$

Clearing the equation of radicals, and simplifying,

$$y = \tan \beta \cdot x - \frac{f \sec^2 \beta}{2 V^2} x^2.$$

Euler; Mechan. tom. 1. p. 232.

(2) A particle, always acted on by a force in parallel lines, describes a given curve; to determine the nature of the force, the velocity and direction of projection being given.

Let the force be represented by Y, which we will suppose to

act towards the axis of x parallel to the axis of y. The equations of motion will be

$$v \frac{dv}{ds} = -Y \frac{dy}{ds} \dots (1),$$

$$\frac{v^2}{\rho} = Y \frac{dx}{ds} \dots (2).$$

Eliminating Y, we have

$$\frac{1}{v}\frac{dv}{ds} = -\frac{1}{\rho}\frac{dy}{dx} = \frac{\frac{dy}{dx}\frac{d^3y}{dx^2}}{\frac{ds^3}{dx^3}},$$

$$\frac{dy}{dx}\frac{d^3y}{dx}\frac{dy}{dx}\frac{d^2y}{dx}$$

$$\frac{1}{v}\frac{dv}{dx} = \frac{\frac{dy}{dx}\frac{d^3y}{dx^2}}{\frac{ds^2}{dx^3}} = \frac{\frac{dy}{dx}\frac{d^3y}{dx^3}}{1 + \frac{dy}{dx^3}}.$$

Integrating, we get

$$\log v = C + \frac{1}{2} \log \left(1 + \frac{dy^2}{dx^2} \right) = C + \log \frac{ds}{dx}.$$

Let V denote the initial velocity, and β the angle which the direction of projection makes with the axis of x; then

$$\log V = C + \log \sec \beta,$$

$$ds$$

and therefore

$$\log \frac{v}{V} = \log \frac{\frac{ds}{dx}}{\sec \beta},$$

$$v = V \cos \beta \frac{ds}{dx}.$$

Substituting this value of v in (2), we have

$$Y = \frac{V^2}{\rho} \frac{\cos^2 \beta}{\rho} \frac{ds^3}{dx^3}.$$

Euler; Mechan. tom. 1. p. 240.

(3) A particle describes a given curve about a centre of force; to determine the motion of the particle and the law of the force.

Let APB (fig. 105) be the path of the particle, S the centre of force; P the position of the particle at any time; PT a tangent at the point P, and SY perpendicular to PT. Let F

denote the force along PS, and ϕ the angle SPT; then, the direction of the motion at P being towards B,

$$v \frac{dv}{ds} = -F \cos \phi \dots \dots (1),$$

$$\frac{v^2}{a} = F \sin \phi \dots \dots (2).$$

Now, since

$$ds \cos \phi = dr$$
,

and

$$\rho \sin \phi = \frac{rdr}{dp} \sin \phi = p \frac{dr}{dp},$$

where p denotes SY, we have, by (1) and (2),

$$vdv = - Fdr \dots (3),$$

$$v^2 = Fp \frac{dr}{dp} \dots (4).$$

Eliminating F between (3) and (4),

$$\frac{dv}{v} = -\frac{dp}{p}, \log v = C - \log p.$$

Let V, P, be the initial values of v, p; then

$$\log V = C - \log P,$$

and therefore

$$\log \frac{v}{V} = \log \frac{P}{p}, \quad v = \frac{V \cdot P}{p} \cdot \dots \cdot (5).$$

Again, if t denote the time of the motion,

$$\frac{ds}{dt} = v = \frac{V \cdot P}{p}, \text{ by (5)},$$

$$pds = VP dt;$$

but pds is equal to dh', where h' represents twice the area swept out by the radius vector in its motion from some assigned position; hence $dh' = VP dt, \quad h' = VPt \dots (6),$

the area being supposed to commence with the time.

Again, by (2), we have

$$F = \frac{v^2}{\rho \sin \phi} = \frac{v^2}{p} \frac{dp}{dr} = \frac{V^2 P^2}{p^3} \frac{dp}{dr}, \text{ by (5), } \dots (7).$$

Suppose now that h represents twice the area swept out in a

unit of time; then, since by [6] h is equal to ΓP , we have by (6), (5), (7), $h = ht \dots (8),$ $\tau = \frac{h}{p} \dots (9),$

$$F = \frac{h^2}{p^3} \frac{dp}{dr} = \frac{h^2r}{\rho p^3} \dots (10).$$

The formula '8, and '9, were given by Newton.' Formula (10) was discovered by De Moivre in the year 1705, by whom it was communicated without demonstration to John Bernoulli. A proof of the formula was obtained by Bernoulli and forwarded to De Moivre in a letter dated Basle, Feb. 16, 1706. Demonstrations were afterwards given by Keill,' and by Hermann.' See De Moivre's Miscell. Analyt. lib. viii., and John Bernoulli, Opera, tom. 1. p. 477.

Integrating the equation (3) we get another expression for the velocity

 $v^2 = V^2 - 2 \int_R^r F dr,$

where R denotes the prime radius vector. This result shews that the velocity of the particle depends only upon its distance from the centre of force, and not upon the path described; a theorem proved by Newton.

Euler; Mechan. tom. 1. p. 240.

SECT. 4. Motion in Resisting Media.

(1) A particle acted on by gravity is projected in a uniform medium, of which the resistance varies as the velocity, with a given velocity and at a given angle of inclination to the horizon; to find after what interval of time the particle will arrive at its greatest altitude.

Let k be the resistance for a unit of velocity, u the velocity and a the angle of projection, and let y be the height through

¹ Principia, lib. 1. Prop. 1.

² Phil. Trans. Num. 317; 1708.

³ Phoronomia, p. 70.

⁴ Princip. lib. 1. Prop. 40.

which the particle has ascended at the end of the time t. Then

$$\frac{d^2y}{dt^2} = -g - k \frac{ds}{dt} \frac{dy}{ds} = -g - k \frac{dy}{dt},$$

$$\log\left(g+k\,\frac{dy}{dt}\right)=C-kt\;;$$

but $\frac{dy}{dt} = u \sin a$ when t = 0; hence

 $\log\left(g+ku\sin\,a\right)=\,C,$

and therefore

$$\log \frac{g + ku \sin a}{g + k \frac{dy}{dt}} = kt.$$

When y is a maximum, $\frac{dy}{dt} = 0$, and therefore the required value of t will be equal to

$$\frac{1}{k}\log\left(1+\frac{k}{q}u\sin a\right).$$

(2) A particle moving in a resisting medium is acted on by a given force in parallel lines; to find the resistance that any proposed curve may be described, and conversely.

Let the positions of the particle be referred to two rectangular axes Ox, Oy, (fig. 106), OM = x, PM = y, AP = s; P being the position of the particle at any time, and APB the curve of its motion; also let Y denote the accelerating force at P parallel to Oy, v the velocity of the particle, and R the resistance of the medium.

Then, by the equations of tangential and normal resolution given in the preceding section, we have

$$v \frac{dv}{ds} = -R + Y \frac{dy}{ds} \dots (1),$$
$$\frac{v^2}{\rho} = Y \frac{dx}{ds} \dots (2);$$

where ρ denotes the radius of curvature at P. But

$$\rho = \frac{\frac{ds^3}{dx^3}}{\frac{d^3y}{dx^2}};$$

hence, from (2),
$$v^2 = \frac{\frac{ds^2}{dx^2}}{\frac{d^2y}{dx^2}} Y;$$

differentiating, we have, since
$$\frac{ds^2}{dx^2} = 1 + \frac{dy^2}{dx^2}$$
,

$$v \frac{dv}{dx} = Y \frac{dy}{dx} + \frac{1}{2} \frac{ds^2}{dx^3} \frac{d}{dx} \left\{ \frac{Y}{d^2y} \right\},\,$$

$$v \frac{dv}{ds} = Y \frac{dy}{ds} + \frac{1}{2} \frac{ds}{dx} \frac{d}{dx} \left\{ \frac{Y}{d^2y} \right\};$$

hence, from (1),
$$R = -\frac{1}{2} \frac{ds}{dx} \frac{d}{dx} \left\{ \frac{Y}{d^2y} \right\}$$
,

which gives the resistance if the curve be given, and conversely.

The solution of this problem, which Newton had given in the first edition of the *Principia*, involved certain errors, which at the suggestion of John Bernoulli were afterwards corrected.

Cor. If the resistance vary as the square of the velocity for a uniform density; then, Q denoting the density generally, we have

we have
$$R = Qv^2 = Q\rho \frac{dx}{ds} Y, \text{ by } (2),$$
 and therefore
$$Q = \frac{R}{Y} \frac{\frac{d^2y}{dx^2}}{\frac{ds^2}{dx^2}}$$

$$= -\frac{1}{2} \frac{dx}{ds} \frac{\frac{d^2y}{dx}}{\frac{dx}{ds}} \frac{d}{dx} \left\{ \frac{Y}{\frac{d^2y}{dx^2}} \right\}$$

$$= -\frac{1}{2} \frac{dx}{ds} \frac{d}{dx} \log \left\{ \frac{Y}{\frac{d^2y}{dx^2}} \right\},$$

which gives the density at any point within the medium; or, if the density be given, determines the curve. Cor. It is evident that $\rho \frac{dx}{ds}$ is equal to half the chord of curvature at P parallel to Oy, or in the direction of the force Y; let q denote this chord of curvature. Then

$$v^2 = 2 Y \cdot \frac{1}{4} q$$
;

and therefore the space due to the velocity, supposing the force to continue constant, is equal to one-fourth of the chord of curvature.

Newton; Princip. lib. 11. prop. 10. John Bernoulli; Act. Erudit. Lips. 1713; Opera, tom. 1. p. 514.

(3) A particle moves in a resisting medium under the action of a given force always tending towards a fixed centre; to determine the law of resistance when the path of the particle is given, and conversely.

Let APB (fig. 107) be the path of the particle, S the centre of force, AP = s, SP = r, p = the perpendicular from S upon the tangent at P, v = the velocity at P, ρ the radius of curvature, P the central force, and R the resistance of the medium.

Then, by the equations of tangential and normal resolution,

we have

$$v \frac{dv}{ds} = -R - P \frac{dr}{ds} \dots (1),$$

$$\frac{v^{2}}{\rho} = \frac{p}{r} P \dots (2).$$

$$r \frac{dr}{ds} \text{ we have from } (2)$$

Since ρ is equal to $r \frac{dr}{dp}$, we have, from (2),

$$v^2 = p \frac{dr}{dn} P \dots (3);$$

and therefore, differentiating with respect to s,

$$v \frac{dv}{ds} = \frac{1}{2} \frac{d}{ds} \left(p \frac{dr}{dp} P \right)$$

$$= \frac{1}{2} \frac{d}{ds} \left\{ \left(p^3 \frac{dr}{dp} P \right) \frac{1}{p^3} \right\}$$

$$= -\frac{1}{p^3} \frac{dp}{ds} \left(p^3 \frac{dr}{dp} P \right) + \frac{1}{2p^2} \frac{d}{ds} \left(p^3 \frac{dr}{dp} P \right)$$

$$= -\frac{dr}{ds} P + \frac{1}{2p^2} \frac{d}{ds} \left(p^3 \frac{dr}{dp} P \right).$$

Hence, by substituting this value of $v \frac{dv}{ds}$ in (1), we have

$$R = -\frac{1}{2p^2} \frac{d}{ds} \left(p^3 \frac{dr}{dp} P \right) \dots (4);$$

which determines the law of resistance when the curve is known, and conversely.

Cor. Supposing the resistance to vary as the density of the medium multiplied by the square of the velocity of the particle, we have, Q denoting the density,

$$R = Qv^2 = Qp \frac{dr}{dp} P$$
, by (3),

and therefore, by (4),

$$Q = -\frac{1}{2} \frac{\frac{d}{ds} \left(p^3 \frac{dr}{dp} P \right)}{p^3 \frac{dr}{dp} P} = -\frac{1}{2} \frac{d}{ds} \log \left(p^3 \frac{dr}{dp} P \right) \dots (4),$$

which determines the law of the density when the curve is given, and conversely.

Cor. From the equation (2) we have

$$v^2 = \frac{p}{r} \rho P = \frac{1}{2} q P = 2 (\frac{1}{4} q) P$$
,

where q denotes the chord of curvature through S. This result shews that the velocity at any point of the curve is that due to falling in vacuo towards the centre of force, continued constant, through a quarter of the chord of curvature.

Cor. From (4) we have

$$p^{3}\frac{dr}{dp}P=h^{2}\epsilon^{-2\int Qds}, \quad P=\frac{h^{2}\frac{dp}{dr}}{p^{3}}\epsilon^{-2\int Qds},$$

where h is some constant quantity. This formula gives the central force when the law of the density and the orbit are given. It is easily shewn that, if $ASP = \theta$,

$$\frac{h^2}{p^3}\frac{dp}{dr} = \frac{h^2}{r^2}\left\{\frac{1}{r} + \frac{d^2}{d\theta^2}\left(\frac{1}{r}\right)\right\},\,$$

and therefore
$$P = \frac{h^2}{r^2} \left\{ \frac{1}{r} + \frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) \right\} \epsilon^{-2f \, Qds}.$$
If $Q = 0$, we get
$$P = \frac{h^2}{r^2} \left\{ \frac{1}{r} + \frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) \right\},$$

which is Binet's formula for the central force in vacuo.

Newton; *Principia*, lib. 11. Prop. 17, 18; John Bernoulli; *Opera*, tom. 1v. p. 347. Euler; *Mechan.* tom. 1. p. 428 et sq., p. 451 et sq.

(4) A particle is projected with a given velocity in a uniform medium, in which the resistance varies as the square of the velocity; the particle is acted on by gravity, and the direction of its projection makes a very small angle with the horizon; to determine approximately the equation to the portion of the path which lies above the horizontal plane passing through the point of projection.

Let Ox and Oy (fig. 108) be the horizontal and vertical axes of x and y, O being the point of projection; P the position of the particle at any time; OM = x, PM = y, v = the velocity at P, s = the arc OP, k = the resistance for a unit of velocity; then, by the tangential and normal resolutions,

$$v \frac{dv}{ds} = -kv^2 - g \frac{dy}{ds} \dots (1),$$

$$v^2 = g \frac{dx}{ds} \rho = -g \frac{1+p^2}{q} \dots (2),$$

where $p = \frac{dy}{dx}$ and $q = \frac{dp}{dx}$. Hence, eliminating r between these two equations, we have

$$\frac{d}{ds} \frac{1+p^{2}}{q} + 2k \frac{1+p^{2}}{q} - 2 \frac{dy}{ds} = 0,$$

$$\frac{d}{dx} \frac{1+p^{2}}{q} + 2k \frac{1+p^{2}}{q} \frac{ds}{dx} - 2p = 0,$$

$$\frac{d}{dx} \log \frac{1+p^{2}}{q} + 2k \frac{ds}{dx} - \frac{2pq}{1+p^{2}} = 0.$$

Integrating, we get

$$\log \frac{1+p^2}{q} + 2ks + \log C - \log (1+p^2) = 0,$$

C being some constant quantity;

$$\log \frac{C}{q} + 2ks = 0, \quad q = Ce^{2ks} \cdot \ldots (3).$$

Let u be the velocity and a the angle of projection; then, initially, by (3) and (2),

$$q = C$$
, $u^2 = -g \frac{1 + \tan^2 a}{q}$,

and therefore

$$C=-\frac{g}{u^2\cos^2a};$$

hence by (3),

$$q=-\frac{g}{u^2\cos^2 a}e^{2ka}:$$

but, the angle of projection being small, we may neglect all powers of p beyond the first, and therefore

$$s = \int_0^x (1 + p^2)^{\frac{1}{4}} dx = \int_0^x dx = x \text{ nearly };$$

hence

$$q = -\frac{g}{u^2 \cos^2 a} e^{2kx} \text{ nearly.}$$

Multiplying by dx, and integrating,

$$p = C - \frac{g}{2ku^2\cos^2 a} e^{2kr};$$

but $p = \tan a$ when x = 0; hence

$$\tan a = C - \frac{g}{2ku^2\cos^2 a},$$

and therefore $p = \tan a + \frac{g}{2ku^2 \cos^2 a} (1 - e^{2kx})$.

Integrating again,

$$y = C + x \tan \alpha + \frac{g}{2ku^2} \frac{g}{\cos^2 \alpha} \left(x - \frac{1}{2k} e^{2kx}\right);$$

but x = 0, y = 0, simultaneously; hence

$$0 = C - \frac{g}{4k^2u^2\cos^2 a},$$

and therefore $y = x \tan a + \frac{g}{4k^2u^2\cos^2 a}(1 + 2kx - e^{2kx})$.

Moreau; Journal de l'Ecole Polytech. Cahier xi. p. 215.

The general problem of the path of a projectile in a uniform resisting medium where the resistance varies as the square of the velocity, was proposed by Keill as a trial of skill to John Bernoulli, by whom the challenge was received in February 1718. Keill, trusting to the complexity of the analysis, which had probably deterred Newton from attempting any regular solution of the problem in the second book of the Principia, was in hopes that the exertions of Bernoulli would prove unsuccessful. Bernoulli however, having expeditiously effected a solution, not only of Keill's problem, but likewise of the more general one where the resistance varies as the n^{μ} power of the velocity, expressed a determination not to publish his investigation until he had received intimation that his antagonist had himself been able to solve his own problem. He gave Keill till the following September to exercise his talents, declaring that if he received by that time no satisfactory communication, he should feel himself entitled to question the ability of his adversary. At the request of a mutual friend, Bernoulli consented to extend the interval to the first of November. It turned out, however, that Keill was unable to obtain a solution. At length Nicholas Bernoulli, Professor of Mathematics at Padua, communicated to John Bernoulli a solution of Keill's problem, which the author afterwards extended to the more general one. Finally, on the 17th of November, information was received by John Bernoulli, from Brook Taylor, to the effect that he had obtained a solution. John Bernoulli published his own analysis, together with that of his nephew Nicholas, in the Acta Erudit. Lips. 1719 mai. p. 216; see also his Opera, tom. 11. p. 393. For further information on this celebrated problem, the reader may avail himself of the labours of Euler, Borda, Legendre, Templehoff, and Moreau.

(5) A particle moves in a semicircle acted on by gravity, in a medium where the resistance varies as the density and the square of the velocity; to find the law of the resistance and density.

Let OA, OB, (fig. 109), be horizontal and vertical radii of the

¹ Mém. de l'Acad. de Berlin, 1753.
² Ib. 1769.
³ Ib. 1782.
⁴ Ib. 1788—89.

Journal de l' Ecole Polytech. Cahier xi. p. 204.

semicircle, OAx and OBy being the axes of x and y; a the radius of the circle, OM = x, and gravity = g. Then

$$R = \frac{3}{2} \frac{gx}{a}, \quad v^2 = g \left(a^2 - x^2\right)^{\frac{1}{2}}, \quad Q = \frac{3x}{2a} \left(a^2 - x^3\right)^{-\frac{1}{4}}.$$

Newton; Princip. lib. 11. Prop. 10. Ex. 1. John Bernoulli; Act. Erudit. Lips. 1713. Euler; Mechan. tom. 11. p. 392.

(6) A particle acted on by gravity moves in a parabola of any order; to find the law of resistance.

Let the equation to the parabola ABC (fig. 109) be $y = b - cx^{n}$, Oy being vertical; then

$$R = \frac{(n-2)(1+n^2c^2x^{2n-2})^{\frac{1}{2}}g}{2n(n-1)cx^{n-1}}.$$

(7) A particle moves in an hyperbola of any order, having one of its asymptotes vertical; to find the law of the density, the resistance varying as the density into the square of the velocity.

Let APB (fig. 110) denote the path of the particle, Oy the vertical asymptote being taken as the axis of y, and Ox, which is horizontal, as the axis of x; then, if the equation to the hyperbola be β^{n+1}

 $y = ax + \frac{\beta^{n+1}}{x^n},$

we shall have

$$Q = \frac{\frac{1}{2}(n+2)x^{n}}{\left\{x^{2n+2} + (ax^{n+1} - n\beta^{n+1})^{2}\right\}^{\frac{1}{2}}}.$$

Euler; Mechan. tom. 11. p. 400.

(8) A particle moves in a circle about a centre of force in the circumference, the force being attractive and varying as any power of the distance; to determine the resistance of the medium and the law of the density, supposing the resistance to be equal to the product of the density and the square of the velocity.

Let P (fig 111) be the position of the particle at any time, its motion taking place towards S the centre of force; let C be the centre of the circle; SP = r, $\angle PSA = \theta$, central force = μr^n ; then

$$R = \frac{1}{4} \mu (5 + n) r^{n} \sin \theta, \ Q = \frac{1}{2} (n + 5) \frac{\sin \theta}{r}.$$

(9) A particle moves in an equiangular spiral about a centre of force in the pole which varies as any power of the distance from the pole; to find the law of the resistance and of the density of the medium, the resistance being considered equal to the product of the density and the square of the velocity.

If a be the constant angle, r the radius vector at any point, μr^n the attractive central force, and the particle be so moving as to approach the centre of force;

$$R = \frac{1}{2} \mu(n+3) r^n \cos \alpha, \ Q = \frac{1}{2} (n+3) \frac{\cos \alpha}{r}.$$

Newton; *Princip*. lib. 11. Prop. 15, 16. John Bernoulli; *Opera*, tom. 1v. p. 350.

(10) A particle moves in the circumference of a circle about a centre of force in the centre; the resistance of the medium in which the motion takes place is constant; to determine the law of the force, the velocity at any time of the motion, and the time which elapses, as well as the space which is described, before the particle is reduced to rest.

Let β be the initial velocity of the particle, a the radius of the circle, c the constant value of the resistance, s the arc described from the beginning of the motion, P the central force; then

$$v^2 = \beta^2 - 2cs$$
, $\dot{P} = \frac{1}{a}(\beta^2 - 2cs)$;

when the particle is reduced to rest,

$$s=\frac{\beta^2}{2c}, \quad t=\frac{\beta}{c}.$$

(11) A particle is moving along the curve of an equiangular spiral so as to be approaching the pole; the motion takes place in a medium where the resistance varies as any power of the distance from the pole; to find the law of the central attractive force in the pole.

Let a be the constant angle, β the initial velocity, a the initial distance, kr^a the resistance at a distance r, P the required force; then $(n+3)\cos r r^2(2^2+2^2/r^{n+3}-r^{n+3})$

$$P = \frac{(n+3)\cos a \ a^{2}\beta^{2} + 2k(r^{n+3} - a^{n+3})}{(n+3)\cos a \ r^{3}}.$$

Euler; *Mechan*. tom. 1. p. 442.

(12) A particle is projected with a velocity u, and at an inclination a to the horizon, in a uniform medium where the resistance varies as the velocity; to determine the time which elapses before the direction of the motion is inclined to the horizon at an angle β .

If k represent the resistance for a unit of velocity, t will be found from the equation

$$g\cos\beta(\epsilon^{kt}-1)=ku\sin(a-\beta).$$

(13) Two particles, subject to the action of gravity, are simultaneously projected at equal angles of inclination to the horizon, and with equal velocities, the one in vacuo and the other in a medium where the resistance varies as the velocity; to determine a relation between the times in which the particles describe two arcs so related to each other that the tangents at their extremities shall make equal angles with the horizon.

If k denote the resistance of the medium for a unit of velocity, and t_1 , t_2 , denote corresponding times in vacuo and in the medium; then $e^{kt_0} = 1 + kt_1$.

(14) Having given the co-ordinates of the highest point of the path described by a particle acted on by gravity and projected in vacuo at a known angle of inclination to the horizon; to find the decrements of these co-ordinates when the particle is projected in a rare medium in which the resistance varies as the velocity.

Let a, b, be the given co-ordinates, k the resistance for a unit of velocity, and β the angle of projection; then

$$\delta a = -k \left(\frac{a^3 \tan \beta}{g} \right)^{\frac{1}{2}}, \quad \delta b = -k \left(\frac{8b^3}{9g} \right)^{\frac{1}{2}}.$$

SECT. 5. Impossible Mechanics.

The determination of the motion of a material particle in a plane XOY, (fig. 112), under the action of assigned forces in that plane, depends upon the integration of two simultaneous differential equations

$$\frac{d^2x}{dt^2}=\phi(x,y), \quad \frac{d^2y}{dt^2}=\chi(x,y) \ldots (1),$$

where x, y, are the co-ordinates of the position of the particle, at the end of any time t, referred, we will suppose, to rectangular axes OX, OY, and $\phi(x, y)$, $\chi(x, y)$, are certain functions of x, y, depending upon the nature of the forces. Suppose that, for certain particular forms of the functions, having effected the integrations and determined the arbitrary constants from the initial circumstances of the motion, we are enabled to obtain the relations $x = \phi_1(t), \quad y = \chi_1(t), \dots (2)$,

where $\phi_1(t)$, $\chi_1(t)$, represent certain functions of t. In the determination of these relations is generally supposed to consist the complete solution of the problem of the motion of the particle. There are particular cases, however, in which they cease, after the lapse of a certain time, to correspond to the physical motion of the particle. It occasionally happens that as t increases indefinitely from zero by continuous gradation, one or sometimes both of the functions $\phi_1(t)$, $\chi_1(t)$, after remaining possible up to a certain magnitude of t, at length assumes impossible values.

Before proceeding to the consideration of the mechanical circumstances connected with this anomaly, it will be necessary to make a few observations on the geometrical signification of impossible values of x and y.

By the theory of impossible quantities we know that for all values of x and y, possible or impossible, we may assume

$$x = (\cos \theta + -\frac{1}{3} \sin \theta) \alpha,$$

$$y = (\cos \phi + -\frac{1}{3} \sin \phi) \beta,$$

where a, β , θ , ϕ , are all possible and positive quantities. Suppose now that $\theta = 2\lambda \pi$, where λ is an integer; then we have x = a; also, if $\theta = (2\lambda + 1) \pi$, where λ is still supposed to be integral, then x = -a. It is evident then that whatever geometrical interpretation may be given to the general form of the value of x, it must be such that, when θ is any even multiple of π , x may denote a distance a estimated along the axis OX in the positive direction; and that, when θ is any odd multiple of π , x may denote a distance a estimated negatively from O, that is, in the direction XO. Similar remarks, mutatis mutandis, are evidently applicable to the interpretation of the values of y. These

are the only conditions to which we are subject in our selection of a geometrical interpretation of impossible values of the variables. From what has been said then, it appears that the conventional significations of the signs + and -, leave that of the more general sign $\cos \theta + -\frac{1}{4} \sin \theta$ in a great measure arbitrary. In fact $x = (\cos \theta + -\frac{1}{4} \sin \theta) a$

may be taken to represent any line OA of a length a drawn from O at an inclination θ to the direction OX; thus the locus of the extremity A will be a circle described with a radius $a \sin \theta$ about a centre in OX, at a distance $a \cos \theta$ from O, the plane of the circle being at right angles to OX. In like manner

$$y = (\cos \phi + -\frac{1}{4} \sin \phi) \beta$$

may be taken to represent a straight line $OB = \beta$, inclined at an angle ϕ to OY. Complete the parallelogram OAPB; then P will be the geometrical point corresponding to the values of x and y. From this construction it is clear that P is not limited to a particular point, but that it may assume an infinite number of different positions.

In order that, for every pair of values of x and y, the point P may receive a definite position, it will be necessary to adopt, for the construction of the conjugate axes OA, OB, some general law which shall be consistent with the circular loci of the points A, B. For instance, we might assume A and B always to coincide with the intersections of their circular loci with the plane XOY; or, drawing OZ at right angles to the plane XOY, with the intersections of these loci with the planes XOZ, YOZ, or, in fact, with any assigned planes whatever passing through OX, OY. It is evident that the locus described by the point P in accordance with the equations (2), will vary, as far as impossible values of the variables are concerned, with variations in the law of constructing the conjugate axes.

Now, reverting to the consideration of the mechanical value of the equations (2), it is evident that, so long as they perfectly represent the physical motion of the particle, there can be no indeterminateness in their geometrical interpretation, it being impossible for a particle to describe any but one path in accordance with given forces and a given projection. We will proceed to ascertain whether it be possible to adopt such a system of construction for the conjugate axes, as to render the geometrical locus of the equations (2) coincident with the physical path of the particle under the circumstances of the problem. Let μ , ν , denote the inclinations of the planes AOX, BOY, to the plane XOY. Then, since the particle can never deviate from the plane XOY under the circumstances of the problem, we shall have, by restricting the construction of the axes OA, OB, to the fulfilment of this condition, a relation

$$F(\alpha, \beta, \theta, \phi, \mu, \nu) = 0 \ldots (3).$$

Also, the effective accelerating forces on the point P parallel to OX, OY, must (when x and y become impossible) continue to coincide with the physical law of the forces. Hence we shall have two more requisite relations,

$$G(\alpha, \beta, \theta, \phi, \mu, \nu) = 0, \quad H(\alpha, \beta, \theta, \phi, \mu, \nu) = 0 \dots (4).$$

Now a, β , θ , ϕ , are each known functions of t; hence, between the relation (3) and the two relations (4), we may eliminate μ , ν , and thus we shall get an equation involving t and known quantities. Since, then, t is restricted to particular values, it is evident that the equations (2) cannot, for any law of constructing the axes, be enabled to represent the physical motion of the particle for impossible values of the variables.

Whenever, then, x or y assumes an impossible value for any value of t, we must conclude that the physical motion of the particle cannot be represented by any single pair of equations (2). It will be necessary, on arriving at such a critical value of t, again to revert to the equations (1), to integrate them anew, and to determine the values of the arbitrary constants in accordance with the values which x, y, $\frac{dx}{dt}$, $\frac{dy}{dt}$, have acquired at the conclusion of the preceding stage of the motion. It will be necessary to repeat this process whenever t makes x or y impossible. Corresponding to each integration there will be a new equation in x and y, resulting from the elimination of t between the corresponding pair of equations (2). A portion only of each of the curves belonging to the equations in x and y

will be described by the particle, which will therefore pursue a path consisting of fragments of a series of distinct curves.

(1) A particle placed in the plane XOY is attracted towards the axes OX and OY by forces varying inversely as the cubes of the distances; to determine the motion of the particle.

This problem gives rise to the differential equations

$$\frac{d^2x}{dt^2} = -\frac{m^4}{x^3}$$
, $\frac{d^2y}{dt^2} = -\frac{n^4}{y^3}$,....(1).

Integrating these equations, we have

$$\frac{dx^2}{dt^2} = A + \frac{m^4}{x^2}, \quad \frac{dy^2}{dt^2} = B + \frac{n^4}{y^2},$$

A, B, being arbitrary constants. Now, initially, $\frac{dx}{dt}$ and $\frac{dy}{dt}$ are both equal to zero, and therefore, if a, b, denote the initial values of x, y, we have

$$0 = A + \frac{m^4}{a^2}, \quad 0 = B + \frac{n^4}{b^2}.$$

Hence

$$\frac{dx^2}{dt^2} = m^4 \left(\frac{1}{x^2} - \frac{1}{a^2}\right), \quad \frac{dy^2}{dt^2} = n^4 \left(\frac{1}{y^2} - \frac{1}{b^2}\right), \quad \dots (2).$$

Now, since the particle will move towards OY, $\frac{dx}{dt}$ will be negative, and thus

$$\frac{dx}{dt} = -m^2 \frac{(a^2 - x^2)^{\frac{1}{4}}}{ax}; \frac{axdx}{(a^2 - x^2)^{\frac{1}{4}}} = -m^2 dt;$$

and therefore, A

$$A' - a(a^2 - x^2)^{\frac{1}{4}} = -m^2t$$

where A' is an arbitrary constant; but x = a when t = 0, and therefore A' = 0: hence

$$a(a^{2}-x^{2})^{\frac{1}{2}}=m^{2}t$$

$$x^{2}=a^{2}-\frac{m^{4}t^{2}}{a^{2}}:$$

$$y^{2}=b^{2}-\frac{n^{4}t^{2}}{k^{2}}.$$

$$(3).$$

similarly

As soon as t becomes greater than $\frac{a^2}{m^2}$, it is evident, from the former of the equations (3), that x becomes impossible; and,

when t becomes greater than $\frac{b^2}{n^2}$, it appears in like manner from the latter that y becomes impossible. We will commence with the consideration of the motion parallel to the axis of x. When x becomes impossible, the former of the equations (3) can no longer represent the physical motion of the particle parallel to OX, and we must revert to the equation

$$\frac{d^2x}{dt^2}=-\frac{m^4}{x^3},$$

where x is now to be estimated in the negative direction.

Integrating, and adding an arbitrary constant A', we have

$$\frac{dx^2}{dt^2} = A' + \frac{m^4}{x^2} \dots (4).$$

But in the beginning of the second stage of the motion, and at the end of the first, the velocity is the same. Hence, by (2) and (4),

$$m^4 \left\{ \frac{1}{(x=0)^3} - \frac{1}{a^3} \right\} = A' + \frac{m^4}{(x=0)^3}, \quad A' = -\frac{m^4}{a^2};$$

and therefore, from (4),

$$\frac{dx^2}{dt^2}=m^4\left(\frac{1}{x^2}-\frac{1}{a^2}\right).$$

Extracting the square root, we obtain

$$m^{2}dt = + \frac{axdx}{(a^{2} - x^{2})^{\frac{1}{4}}}$$
$$m^{2}dt = - \frac{axdx}{(a^{2} - x^{2})^{\frac{1}{4}}}.$$

or

Integrating these two equations,

$$m^{2}t + C_{1} = -a(a^{2} - x^{2})^{\frac{1}{4}},$$

 $m^{2}t + C_{2} = +a(a^{2} - x^{2})^{\frac{1}{4}}.$

The former of these equations corresponds to the motion from the axis of y, and the latter to the motion which afterwards takes place towards it. To determine C_1 , we have x = 0 when $t = \frac{a^2}{m^2}$, which gives $C_1 = -\frac{a^2}{2}$; and thus, for the first part of

the motion,
$$m^2t - 2a^2 = -a(a^2 - x^2)^{\frac{1}{4}}$$
.

When $t = \frac{2a^2}{m^2}$, then x = a, and the motion of return begins which is defined by the second of the equations; and $C_2 = -2a^2$. Thus both integrals are comprehended by the equation

$$a^{2}(a^{2}-x^{2})=(m^{2}t-2a^{2})^{2}.$$

After a time $\frac{3a^2}{m^2}$ it is evident from this equation that x becomes impossible, and then the third period of the motion of the particle takes place. It is evident that the particle will again cross the axis of y, and perform on the other side of it, parallel to the axis of x, a motion exactly similar to that of the second stage of the motion. Similarly for the fourth, fifth, &c. stages of the motion parallel to the axis of x.

From the preceding conclusions it is manifest that, dividing the time from its zero value into a series of intervals of which the first is equal to $\frac{a^3}{m^2}$ and all the following ones to $\frac{2a^3}{m^2}$, the motion of the particle at right angles to OY will be represented for each interval respectively by the following equations,

the general formula for the pth interval being

$$a^{2} \{a^{2} - (-p^{-1} x)^{2}\} = \{m^{2}t - 2(p-1)a^{2}\}^{2} \dots (5).$$

In like manner, for the motion of the particle parallel to the axis of y, if we divide the time into intervals of which the first is equal to $\frac{b^2}{n^2}$ and the rest to $\frac{2b^2}{n^2}$, we shall have for the motion in the q^{th} interval,

$$b^{2} \{b^{2} - (-q^{-1} y)^{2}\} = \{n^{2}t - 2(q-1) b^{2}\}^{2} \dots (6).$$

Eliminating t between (5) and (6), we obtain as a general

formula for the equations to the series of curves of which portions are successively traversed by the particle,

$$an^{2}\left[\left\{a^{2}-\left(-\frac{p-1}{2}x\right)^{2}\right\}^{\frac{1}{4}}+2\left(p-1\right)a\right]=bm^{2}\left[\left\{b^{2}-\left(-\frac{q-1}{2}y\right)^{2}\right\}^{\frac{1}{4}}+2\left(q-1\right)b\right].$$

As t keeps continuously increasing from zero, we must keep putting for p and q integral values next greater respectively than

$$\frac{t+\frac{a^2}{m^2}}{\frac{2a^2}{m^2}}=\frac{m^2t+a^2}{2a^2},$$

and

$$\frac{t + \frac{b^2}{n^2}}{\frac{2b^2}{n^2}} = \frac{n^2t + b^2}{2b^2}.$$

Thus the formula will give us the series of equations representing the successive curves. The intersections of the successive curves, if we pay attention to the signs $-^{p-1}$ and $-^{q-1}$, will constitute the limits of the portions of each which the particle actually describes.

(2) To investigate the path of a particle the law of the motion of which is expressed by the integrals of the equations

$$\frac{d^2x}{dt^2} = -\frac{m^4}{x^3}, \quad \frac{d^2y}{dt^2} = -\frac{n^4}{y^3}.....(1),$$

corrected in accordance with the conditions that initially x = a, y = b, $\frac{dx}{dt} = 0$, $\frac{dy}{dt} = 0$; the conjugate axes OA, OB, being supposed to be always at right angles to OY, OX, respectively.

This problem will, as we know, correspond to a path very different, after the lapse of a certain time, from the course pursued by the particle in the preceding problem. Integrating the equations (1), as in the first stage of the motion in the preceding problem, we have

$$x^2 = a^2 - \frac{m^4 t^2}{a_{\infty}^2}, \ y^2 = b^2 - \frac{n^4 t^2}{b^2} \dots (2).$$

Assume, as the most general forms of x and y,

$$x = (\cos \theta + -\frac{1}{3}\sin \theta) a, \quad y = (\cos \phi + -\frac{1}{3}\sin \phi) \beta;$$

then, from the former of the equations (2),

$$(\cos 2\theta + -\frac{1}{3}\sin 2\theta) a^3 = a^2 - \frac{m^4t^3}{a^2};$$

equating the coefficients of like affections, we get

$$a^2 \cos 2\theta = a^2 - \frac{m^2 t^2}{a^2} \dots (3),$$

$$\sin 2\theta = 0 \dots (4).$$

From (4) we see that

$$2\theta = \lambda \pi$$
,

where λ is some integer, and therefore, by (3),

$$a^2\cos\lambda\pi=a^2-\frac{m^4t^2}{a^2}.$$

If t be less than $\frac{a^2}{m^2}$, $a^2 - \frac{m^4 t^2}{a^2}$ is positive, and therefore $\cos \lambda \pi$ must be positive; and if t be greater than $\frac{a^2}{m^2}$, $a^2 - \frac{m^4 t^2}{a^2}$ is negative, and therefore $\cos \lambda \pi$ must be negative. Hence, if t be less than $\frac{a^2}{m^2}$,

$$a^2=a^2-\frac{m^4t^2}{a^2}, \quad \theta=\mu\pi ;$$

and, if t be greater than $\frac{a^2}{m^2}$,

$$a^2 = \frac{m^4t^2}{a^2} - a^2, \quad \theta = \frac{1}{2}(2\mu + 1)\pi;$$

μ being an integral number.

Similarly, if t be less than $\frac{b^2}{n^2}$,

$$\beta^2 = b^2 - \frac{n^4 t^2}{b^2}, \quad \phi = \mu' \pi,$$

and, if t be greater than $\frac{b^2}{n^2}$,

$$\beta^2 = \frac{n^4t^2}{b^2} - b^2$$
, $\phi = \frac{1}{2}(2\mu' + 1)\pi$.

Suppose now that $\frac{a^2}{m^2}$ is less than $\frac{b^2}{n^2}$; then, so long as t is less than $\frac{a^2}{m^2}$,

$$\left\{ a^{2} = a^{2} - \frac{m^{4}t^{2}}{a^{2}}, \quad \beta^{2} = b^{2} - \frac{n^{4}t^{2}}{b^{2}}, \\
 \theta = \mu\pi, \qquad \qquad \phi = \mu'\pi,
 \right\}...(5);$$

while t is greater than $\frac{a^2}{m^2}$, but less than $\frac{b^2}{n^2}$,

$$a^{2} = \frac{m^{4}t^{2}}{a^{3}} - a^{3}, \qquad \beta^{2} = b^{3} - \frac{n^{4}t^{2}}{b^{2}}, \\ \theta = \frac{1}{2}(2\mu + 1) \pi, \qquad \phi = \mu'\pi, \end{cases}$$
 (6);

and, when t is greater than $\frac{b^2}{n^2}$,

$$a^{2} = \frac{m^{4}t^{2}}{a^{2}} - a^{2}, \quad \beta^{2} = \frac{n^{4}t^{2}}{b^{2}} - b^{2}, \\ \theta = \frac{1}{2}(2\mu + 1), \quad \phi = \frac{1}{2}(2\mu' + 1), \end{cases} \dots (7).$$

Suppose now that x', y', z', denote rectangular co-ordinates of the particle; x', y', being estimated along OX, OY, and z' along OZ at right angles to the plane XOY. Then it will easily be seen that

$$x' = a \cos \theta$$
, $y' = \beta \cos \phi$, $z' = a \sin \theta + \beta \sin \phi$.

In the case of the equations (5),

$$x' = a \cos \mu \pi$$
, $y' = \beta \cos \mu' \pi$, $z' = 0$,

and therefore

$$x'^{2} = a^{2} - \frac{m^{4}t^{2}}{a^{2}}, \quad y'^{2} = b^{2} - \frac{n^{4}t^{2}}{b^{2}}.$$

$$\frac{a^{3}x'^{2}}{m^{4}} - \frac{b^{2}y'^{2}}{n^{4}} = -\left(\frac{b^{4}}{n^{4}} - \frac{a^{4}}{m^{4}}\right), \quad z' = 0,$$

$$(8).$$

In the case of the equations (6),

$$x'=0$$
, $y'=\beta\cos\mu'\pi$, $z'=a\cos\mu\pi$,

and therefore

$$y'^{2} = b^{2} - \frac{n^{4}t^{2}}{b^{2}}, \quad z'^{2} = a^{3} - \frac{m^{4}t^{2}}{a^{2}}, \\ \frac{b^{2}y'^{2}}{n^{4}} + \frac{a^{2}z'^{2}}{m^{4}} = \frac{b^{4}}{n^{4}} - \frac{a^{4}}{m^{4}}, \quad x' = 0, \end{cases}$$
....(9).

In the case of the equations (7),

$$x' = 0$$
, $y' = 0$, $z' = \left(\frac{m^4t^2}{a^3} - a^2\right)^{\frac{1}{4}} + \left(\frac{n^4t^2}{b^2} - b^2\right)^{\frac{1}{4}} \dots (10)$.

The equations (8) shew that the particle moves from its initial position to the axis of y' in the arc of an hyperbola, the time of the motion being $\frac{a^2}{m^2}$. The equations (9) shew that, on arriving at the axis of y', the particle subsequently moves in the arc of an ellipse to the axis of z', in a time equal to $\frac{b^2}{n^2} - \frac{a^2}{m^2}$. And the equations (10) shew that, after arriving at the axis of z', it perpetually ascends this axis, with which its path for the future entirely coincides. The position of the particle at any assigned time in each of the three discontinuous portions of its path, is given by the corresponding pair of relations between the variables and the time.

CHAPTER IV.

CONSTRAINED MOTION OF A PARTICLE.

SECT. 1. Motion of a Particle within smooth immoveable Tubes.

LET a particle, under the action of any forces in one plane, move within an indefinitely thin curvilinear tube APB, (fig. 113). Let x, y, be the co-ordinates of the place P of the particle, after a time t from the commencement of the motion; and let AP = s, where A is some assigned point in the tube. Let X, Y, represent the resolved parts of the accelerating force acting on the particle parallel to the axes Ox, Oy, and S the resolved part along the tangent to the curve APB at P. Then the equation for the motion of the particle will be

$$\frac{d^3s}{dt^2} = X \frac{dx}{ds} + Y \frac{dy}{ds} = S \dots (A);$$

or, by integration, v denoting the velocity of the particle at the point P,

$$v^{3}$$
 or $\frac{ds^{3}}{dt^{3}} = 2 \int (Xdx + Ydy) + C = 2 \int Sds + C \dots (B),$

where C is an arbitrary constant, introduced by the integration, which may be determined if we know the initial velocity and the initial position of the particle.

If the force acting on the particle be a central force; then, P representing its intensity at a distance r, we have, taking the centre of force as the origin of x, y,

$$Xdx + Ydy = Pdr,$$

and the formulæ (A), (B), become

$$\frac{d^2s}{dt^2} = P \frac{dr}{ds} \dots (C),$$

$$v^2$$
 or $\frac{ds^2}{dt^2} = 2 \int P dr + C \dots (D)$

(1) A particle, acted on by gravity, descends from rest down a given circular arc, the tangent to which at the lowest point is horizontal; to compare the initial accelerating force estimated along the curve with that which would correspond to motion down the chord, when the arc is indefinitely diminished.

Let A (fig. 114) be the lowest point of the arc, C the centre of the circle, P the position of the particle at any time, T the intersection of the tangent PT with the vertical line CA produced. Draw PM horizontally, and join AP, CP. Let F = the accelerating force at P down the arc PA, f = that down the chord PA, $\angle CTP = a$, $\angle CAP = a'$, CA = a, AM = x, PM = y; g = the force of gravity.

Then $F = g \cos a$, $f = g \cos a'$, and therefore $\frac{F}{f} = \frac{\cos a}{\cos a'}$.

But, by the nature of the circle,

$$\cos a = \cos \angle CPM = \frac{y}{a}.$$

Also, observing that $y^2 = 2ax - x^2$, we get

$$\cos a' = \frac{AM}{AP} = \frac{x}{(x^2 + y^2)^{\frac{1}{4}}} = \left(\frac{x}{2a}\right)^{\frac{1}{4}}.$$

Hence $\frac{F}{f} = \frac{y}{a} \left(\frac{2a}{x}\right)^{\frac{1}{4}} = \frac{(2ax - x^2)^{\frac{1}{4}}}{a} \left(\frac{2a}{x}\right)^{\frac{1}{4}} = \left\{\frac{2}{a}(2a - x)\right\}^{\frac{1}{4}};$

and therefore, when x = 0, we have, for the required ratio,

$$F = 2f$$
.

Saurin; Mémoires de l'Académie des Sciences de Paris, 1724, p. 70. Louville; Ib. p. 128. Courtivron, Ib. 1744, p. 384.

(2) The highest point of the circumference of a circle in a vertical plane is connected by means of a chord with some other point in the curve; to determine the time in which a particle, under the action of gravity, will fall down this chord.

Let AB (fig. 115) be the vertical diameter through the highest point A of the circle; AC the chord down which the particle descends. Join BC, and let P be the position of the

particle after a time t from the commencement of its motion. Let AP=s, AB=2a, $\angle BAC=a$, AC=l. Then, the resolved part of g, the force of gravity, along AC, being $g \cos a$, we have, for the motion of the particle,

$$\frac{d^2s}{dt^2} = g\cos a.$$

Integrating, we get

$$\frac{ds}{dt} = gt \cos \alpha + C;$$

but $\frac{ds}{dt} = 0$, t = 0, simultaneously; hence C = 0; and therefore

$$\frac{ds}{dt} = gt \cos a.$$

Integrating again, and observing that s = 0 when t = 0, we have $s = \frac{1}{2}gt^2 \cos a$.

Let T denote the whole time of descent down AC; then

$$l = \frac{1}{2}gT^2 \cos a;$$

but, from the geometry, $l = 2a \cos a$;

hence

$$2a\cos a = \frac{1}{2}gT^2\cos a,$$

and therefore

$$T=2\left(\frac{a}{y}\right)^{\frac{1}{4}}.$$

This result, being independent of the inclination of the chord to the vertical, shews that the descents down all such chords are performed in the same time; a proposition established by Galileo.

Wolff; Elementa Matheseos Universæ, tom. 11. p. 58.

(3) From one extremity of the horizontal diameter of a circle in a vertical plane, two chords are drawn subtending angles a, 2a, at the centre; given that the time down the latter chord is n times as great as that down the former, to find the value of a.

Let β , β' , be the inclinations of the two chords to the horizon; l, l', their lengths, and t, t', their times of description. Then, as in the preceding problem, it will be found that

and therefore
$$l = \frac{1}{2}gt^{2}\sin \beta, \quad l' = \frac{1}{2}gt^{2}\sin \beta',$$

$$\frac{l'}{l} = \frac{t'^{2}}{t^{2}}\frac{\sin \beta'}{\sin \beta} = n^{2}\frac{\sin \beta'}{\sin \beta}.$$

But from the geometry it is evident that

$$l = 2a \cos \beta$$
, $l' = 2a \cos \beta'$;

hence

$$\frac{\cos \beta'}{\cos \beta} = n^2 \frac{\sin \beta'}{\sin \beta}, \quad \tan \beta = n^2 \tan \beta';$$

but it is clear that

hence

$$2\beta = \pi - a, \quad 2\beta' = \pi - 2a,$$

$$\cot \frac{1}{2} a = n^{2} \cot a,$$

$$\frac{2 \tan \frac{1}{2} a}{1 - \tan^{2} \frac{1}{2} a} = n^{2} \tan \frac{1}{2} a,$$

$$\frac{1 + \cos a}{\cos a} = n^{2},$$

and therefore $\cos a = \frac{1}{n^2 - 1}$, $a = \cos^{-1} \frac{1}{n^2 - 1}$.

(4) A particle is placed anywhere within a thin rectilinear tube, and is acted on by a force tending always towards a fixed centre, and varying directly as the distance; to find the time of an oscillation.

Let x be the distance of the particle, at any time t, from its position of equilibrium, r its corresponding distance from the centre of force, μ^2 the absolute force of attraction. Then, $\mu^2 r$ being the central force at the time t,

$$\frac{d^{2}x}{dt^{2}} = -\mu^{2}r \frac{x}{r} = -\mu^{2}x,$$

$$2 \frac{dx}{dt} \frac{d^{2}x}{dt^{2}} = -2\mu^{2}x \frac{dx}{dt},$$

$$\frac{dx^{2}}{dt^{2}} = C - \mu^{2}x^{2}.$$

Let a be the initial value of x; then, $\frac{dx}{dt}$ being initially equal to

zero, $0 = C - \mu^2 a^2;$

hence

$$\frac{dx^2}{dt^2}=\mu^2(a^2-x^2).$$

Taking the square root, we have, for the motion of the particle from its initial position to its position of rest,

$$-\frac{dx}{(a^2-x^2)^{\frac{1}{2}}}=\mu dt;$$

integrating, we get $\cos^{-1}\frac{x}{a}=C+\mu t$;

but x = a, t = 0, simultaneously; hence C = 0, and therefore

$$\cos^{-1}\frac{x}{a}=\mu t,\ x=a\,\cos\,\mu t.$$

Now, when x acquires its greatest negative value, $\mu t = \pi$; hence, T denoting the period of a complete oscillation, we have

$$T=\frac{\pi}{\mu}$$
.

Euler; Mechan. tom. II. p. 91. Cor. 4.

(5) A particle is constrained to move in a straight line, and is attached to one end of an indefinitely fine elastic string, the other end of which is fixed at a distance from the straight line equal to the unstretched length of the string; to find the time of a small oscillation.

Let a = the natural length of the string, m = the mass of the particle; s = its distance at any time t from its position of equilibrium, T = the tension of the string, and l = its length at the same time. Then, for the motion of the particle,

$$m \frac{d^2s}{dt^2} = -\frac{Ts}{l} \dots \dots \dots (1).$$

Again, by Hooke's law of extension,

$$l = a(1 + \epsilon T) \dots (2),$$

where ϵ is a constant quantity depending upon the extensibility of the string.

But, s being a small quantity,

$$l = (a^2 + s^2)^{\frac{1}{4}} = a\left(1 + \frac{1}{2}\frac{s^2}{a^2}\right)$$
, nearly;

hence, from (2),

$$a\left(1+rac{1}{2}\,rac{s^2}{a^2}
ight)=a\left(1+\epsilon T
ight), \quad T=rac{s^2}{2\epsilon a^2};$$

and therefore, by (1),

$$m \frac{d^2s}{dt^2} = -\frac{s^3}{2sa^2l} = -\frac{s^2}{2sa^3}$$
, nearly;

multiplying by $2\frac{ds}{dt}$, and integrating, we get

$$m \frac{ds^2}{dt^2} = C - \frac{s^4}{4\epsilon a^3};$$

let c be the value of s when $\frac{ds}{dt} = 0$; then

$$0=C-\frac{c^4}{4\epsilon a^3},$$

and therefore

$$m \frac{ds^2}{dt^2} = \frac{1}{4\epsilon a^3} (c^4 - s^4);$$

hence, supposing s to be diminishing as t increases,

$$2a\left(\varepsilon am\right)^{\frac{1}{4}}\frac{ds}{\left(c^4-s^4\right)^{\frac{1}{4}}}=-dt:$$

put $s = c \cos \phi$, and our equation becomes

$$\frac{2a}{c}\left(\varepsilon am\right)^{\frac{1}{4}}\frac{d\phi}{\left(1+\cos^2\phi\right)^{\frac{1}{4}}}=dt,$$

and therefore, 0, π , being the values of ϕ corresponding to the values c, -c, of s, the time of a complete oscillation will be equal to

$$\frac{2^{\frac{1}{4}}a}{c}\left(\epsilon am\right)^{\frac{1}{4}}\int_{0}^{\pi}\frac{d\phi}{\left(1-\frac{1}{6}\sin^{2}\phi\right)^{\frac{1}{4}}},$$

an elliptic function of which the modulus is $\sin \frac{1}{4} \pi$.

(6) A particle moves from rest from a distance a along a thin spiral tube towards a centre of force in the pole attracting inversely as the square of the distance; to find the whole time which will elapse before the particle will arrive at the centre of force, the equation to the spiral being

$$\log \frac{c}{r} = \frac{\theta}{a}.$$

Let μ denote the absolute force; then, by (D), we have, for the velocity at any time,

$$r^2 = C + 2 \int_{\overline{r^2}}^{\mu} dr = C + \frac{2\mu}{r}$$
,

but v = 0 when r = a; hence

$$v^{2} = 2\mu \left(\frac{1}{r} - \frac{1}{a}\right),$$

$$\frac{dr^{2}}{dt^{2}} + r^{2}\frac{d\theta^{2}}{dt^{2}} = 2\mu \left(\frac{1}{r} - \frac{1}{a}\right).$$

or

But, from the equation to the spiral,

$$d\theta = -a \frac{dr}{r};$$

hence

$$\frac{dr^3}{d\ell^2}\left(1+\alpha^2\right)=2\mu\left(\frac{1}{r}-\frac{1}{a}\right);$$

extracting the square root, and taking the negative sign because r decreases with the increase of t, we get

$$(1+a^2)^{\frac{1}{4}}\frac{dr}{\left(\frac{1}{r}-\frac{1}{a}\right)^{\frac{1}{4}}}=-2^{\frac{1}{4}}\mu^{\frac{1}{4}}dt,$$

whence

$$(1+a^2)^{\frac{1}{4}}\int \frac{dr}{\left(\frac{1}{r}-\frac{1}{a}\right)^{\frac{1}{4}}}=C-2^{\frac{1}{4}}\mu^{\frac{1}{4}}t.$$

But

$$\int \frac{dr}{\left(\frac{1}{r} - \frac{1}{a}\right)^{\frac{1}{4}}} = a^{\frac{1}{4}} \int \frac{rdr}{(ar - r^{2})^{\frac{1}{4}}}$$

$$= -a^{\frac{1}{4}} \int \frac{(\frac{1}{2}a - r)dr}{(ar - r^{2})^{\frac{1}{4}}} + \frac{1}{2}a^{\frac{3}{4}} \int \frac{dr}{(ar - r^{2})^{\frac{1}{4}}}$$

$$= -a^{\frac{1}{4}} (ar - r^{2})^{\frac{1}{4}} + \frac{1}{2}a^{\frac{3}{4}} \text{ vers}^{-1} \frac{2r}{a}.$$

Hence we have

$$a^{\frac{1}{4}}(1+a^2)^{\frac{1}{4}}\left\{\frac{1}{2}a \text{ vers}^{-1}\frac{2r}{a}-(ar-r^2)^{\frac{1}{4}}\right\}=C-2^{\frac{1}{4}}\mu^{\frac{1}{4}}t.$$

Now t = 0, r = a, simultaneously; hence

$$\frac{1}{2}\pi a^{\frac{3}{4}}(1+a^2)^{\frac{1}{4}}=C;$$

also, T denoting the whole time of the approach to the pole,

$$0=C-2^{\frac{1}{4}}\mu^{\frac{1}{4}}T;$$

hence we have

$$\frac{1}{2}\pi a^{\frac{1}{4}} (1 + a^2)^{\frac{1}{4}} = 2^{\frac{1}{4}} \mu^{\frac{1}{4}} T,$$

$$T = \frac{\pi a^{\frac{1}{4}} (1 + a^2)^{\frac{1}{4}}}{2^{\frac{1}{4}} \mu^{\frac{1}{4}}}.$$

(7) A particle descends by the action of gravity down a tube AO (fig. 116) in the form of a semi-cubical parabola of which the axis Ox is vertical, and the cusp the lowest point; to investigate the time of falling from a given point A to the cusp O.

Let
$$OM = x$$
, $PM = y$; then by (B), since $X = -g$, $Y = 0$,
$$\frac{ds^2}{dt^2} = -2gx + C.$$

Let h be the initial value of OM, then, $\frac{ds}{dt}$ being initially zero,

we have

$$0 = -2gh + C,$$

and therefore

$$\frac{ds^2}{dt^2} = 2g(h-x)....(1).$$

Now the equation to the curve is $ay^2 = x^3$,

and therefore

$$a^{\frac{1}{4}}y = x^{\frac{3}{4}}, \ a^{\frac{1}{4}}dy = \frac{3}{2}x^{\frac{3}{4}}dx,$$

 $ady^2 = \frac{9}{4}xdx^2, \ ads^2 = \frac{1}{4}(9x + 4a) \ dx^2.$

Hence, by (1), there is

$$\frac{9x+4a}{4a}\frac{dx^2}{dt^2}=2g(h-x),$$

and therefore

$$dt = -\frac{1}{2(2aq)^{\frac{1}{4}}} \left(\frac{9x+4a}{h-x}\right)^{\frac{1}{4}} dx,$$

the negative sign being taken because x decreases as t increases.

Assume $z^2 = 9x + 4a$, and our equation becomes

$$dt = -\frac{1}{2(2ag)^{\frac{1}{4}}} \frac{z}{\left(h + \frac{4a}{9} - \frac{z^{2}}{9}\right)^{\frac{1}{4}}} \cdot \frac{2}{9} zdz$$

$$= -\frac{1}{3(2ag)^{\frac{1}{4}}} \frac{z^{2}dz}{(4a + 9h - z^{2})^{\frac{1}{4}}}$$

$$= -\frac{1}{3(2ag)^{\frac{1}{4}}} \frac{z^{2}dz}{(\beta^{2} - z^{2})^{\frac{1}{4}}}, \text{ where } \beta^{2} = 4a + 9h,$$

$$t = C - \frac{1}{3(2ag)^{\frac{1}{4}}} \int \frac{z^2 dz}{(\beta^2 - z^2)^{\frac{1}{4}}}.$$
But
$$\int \frac{z^3 dz}{(\beta^3 - z^2)^{\frac{1}{4}}} = -z (\beta^3 - z^2)^{\frac{1}{4}} + \int (\beta^2 - z^2)^{\frac{1}{4}} dz$$

$$= -z (\beta^3 - z^2)^{\frac{1}{4}} + \beta^3 \int \frac{dz}{(\beta^2 - z^2)^{\frac{1}{4}}} - \int \frac{z^2 dz}{(\beta^2 - z^2)^{\frac{1}{4}}}$$

$$= -\frac{1}{2} z (\beta^2 - z^2)^{\frac{1}{4}} + \frac{1}{2} \beta^2 \sin^{-1} \frac{z}{\beta};$$
hence
$$t = C - \frac{1}{3(2ag)^{\frac{1}{4}}} \left\{ -\frac{1}{2} z (\beta^2 - z^2)^{\frac{1}{4}} + \frac{1}{2} \beta^2 \sin^{-1} \frac{z}{\beta} \right\};$$

but, initially, x = h, $z^2 = \beta^2$, and therefore

$$0 = C - \frac{1}{3(2ag)^{\frac{1}{2}}} \cdot \frac{1}{2} \beta^2 \frac{\pi}{2} :$$

hence, eliminating C, we have

$$t = \frac{1}{3(2ag)^{\frac{1}{4}}} \left\{ \frac{1}{2} z (\beta^2 - z^2)^{\frac{1}{4}} + \frac{1}{2} \beta^2 \cos^{-1} \frac{z}{\beta} \right\},\,$$

and, substituting for z and β their values,

$$t = \frac{1}{3(2aq)^{\frac{1}{4}}} \left\{ \frac{3}{2} \left(9x + 4a \right)^{\frac{1}{4}} \left(h - x \right)^{\frac{1}{4}} + \frac{1}{2} \left(9h + 4a \right) \cos^{-1} \left(\frac{9x + 4a}{9h + 4a} \right)^{\frac{1}{4}} \right\}.$$

When the particle arrives at the cusp, x = 0, and therefore the whole time of descent is equal to

$$\frac{1}{3(2ag)^{\frac{1}{4}}}\left\{8\ a^{\frac{1}{4}}h^{\frac{1}{4}}+\frac{1}{2}\left(9h+4a\right)\cos^{-1}\frac{2a^{\frac{1}{4}}}{\left(9h+4a\right)^{\frac{1}{4}}}\right\}.$$

(8) Two particles P, P, (fig. 117), of which the masses are m, m', are connected by a straight rigid rod without weight, and being constrained to move in two straight grooves Aa, Aa', which are inclined to the horizon at given angles and are in the same vertical plane, make small oscillations; to find the length of the isochronous pendulum.

Let AP, AP', make angles a, a', with the vertical line through A, and let $\angle PAP'$ be equal to a. Let T denote the

mutual action of the two particles communicated along the rod. AP = x, AP' = x', PP' = c, $\angle PP' = a$, $\angle PP' = a$.

Then, for the motion of the two particles, we have, resolving forces along AP and AP,

$$m \frac{d^3x}{dt^2} = mg \cos a - T \cos \phi,$$

$$m' \frac{d^3x'}{dt^2} = m'g \cos a' - T \cos \phi'.$$

Eliminating T, we get

$$m \cos \phi' \frac{d^3x}{dt^2} - m' \cos \phi \frac{d^3x'}{dt^2} = mg \cos a \cos \phi' - m'g \cos a' \cos \phi.$$

But from the geometry we evidently have

$$c \cos \phi = x' \cos \iota - x$$
, $c \cos \phi' = x \cos \iota - x'$;

hence
$$-m(x'-x\cos\iota)\frac{d^3x}{dt^2}+m'(x-x'\cos\iota)\frac{d^3x'}{dt^2}$$

$$= -mg\cos\alpha(x'-x\cos\iota) + m'g\cos\alpha'(x-x'\cos\iota)....(1).$$

Let a, a', be the values of x, x', when there is equilibrium; then, $\frac{d^2x}{dt^2}$ and $\frac{d^2x'}{dt^2}$ being both equal to zero, we have

$$0 = -m \cos a (a' - a \cos \iota) + m' \cos a' (a - a' \cos \iota) \dots (2).$$

Assume x = a + v, x' = a' + v', v and v' being by the hypothesis small quantities. Then, from the equation (1), as far as small quantities of the first order, we have, by the aid of (2),

$$-m(a'-a\cos\iota)\frac{d^2v}{dt^2}+m'(a-a'\cos\iota)\frac{d^2v'}{dt^2}$$

$$= -mg\cos a (v' - v\cos \iota) + m'g\cos a' (v - v'\cos \iota) \dots (3).$$

Now, by the geometry,

$$c^2 = x^2 + x'^2 - 2xx'\cos \alpha$$

hence
$$0 = x \delta x + x' \delta x' - (x \delta x' + x' \delta x) \cos \iota.$$

But $\delta x = v$, $\delta x' = v'$, x = a + v, x' = a' + v'; hence, neglecting small quantities higher than of the first order,

$$0 = (a - a' \cos \iota) v + (a' - a \cos \iota) v' \dots (4).$$

Let r represent the length of the isochronous pendulum; then

$$v = k \sin \left\{ \left(\frac{g}{r} \right)^{\frac{1}{4}} t + \epsilon \right\}, \quad v' = k' \sin \left\{ \left(\frac{g}{r} \right)^{\frac{1}{4}} t + \epsilon \right\},$$

where k, k', ε , are constants: substituting these values of v, v', in the equations (3) and (4), we have

$$m(a'-a\cos\iota)\frac{k}{r}-m'(a-a'\cos\iota)\frac{k'}{r}$$

 $= k (m \cos a \cos \iota + m' \cos a') - k' (m' \cos a' \cos \iota + m \cos a),$ and $(a - a' \cos \iota) k + (a' - a \cos \iota) k' = 0.$

Eliminating k and k' between these two equations,

$$\frac{m}{r}(a'-a\cos\iota)^3+\frac{m'}{r}(a-a'\cos\iota)^3=ma\cos\alpha\sin^2\iota+m'a'\cos\alpha'\sin^2\iota;$$

and therefore

$$r = \frac{m(a' - a\cos \iota)^2 + m'(a - a'\cos \iota)^2}{(ma\cos a + m'a'\cos a')\sin^2 \iota} \dots (5).$$

From P draw PO at right angles to AP, meeting P'O drawn from P' at right angles to AP'. Then the projection of OP upon the line AP' is equal to $OP \sin \iota$, and the projection of PP' on the line AP' is equal to $a' - a \cos \iota$; but these two projections are evidently coincident; hence

$$OP^a \sin^2 \iota = (a' - a \cos \iota)^2;$$

similarly

$$OP^2 \sin^2 \iota = (a - a' \cos \iota)^2.$$

Again, let G be the centre of gravity of m and m' in their position of equilibrium, and H the point in which a vertical line through G will cut a horizontal line through A; then we have

$$(m + m') GH = ma \cos a + m'a' \cos a'.$$

Hence, from (5), we obtain

$$r=\frac{m\cdot OP^2+m'\cdot OP'^2}{(m+m')\ GH}.$$

(9) A particle is placed within a thin rectilinear tube, and is attracted by a force always tending towards a fixed point without the tube, and varying as some function of the distance; to find the time of a small oscillation of the particle.

If f(r) denote the intensity of the force at a distance r, and a be the perpendicular distance of the centre of force from the tube, the time of a small oscillation will be equal to

$$\frac{\pi a^{\frac{1}{4}}}{\left\{f(a)\right\}^{\frac{1}{4}}}.$$

(10) Two equal particles, attracting each other with forces varying inversely as the square of the distance, are constrained to move in two straight lines at right angles to each other; supposing their motions to commence from rest, to find the time in which each of them will arrive at the intersection of the two straight lines.

If a denote the initial distance between the particles, and μ the absolute attracting force of each, they will arrive simultaneously at the intersection of the straight lines in a time equal to

$$\frac{\pi a^{\frac{3}{4}}}{2\left(2\mu\right)^{\frac{1}{4}}}.$$

The particles would arrive simultaneously at the intersection of their paths for any other law of mutual attraction.

(11) A particle acted on by gravity descends from any point in the arc of an inverted cycloid, of which the axis is vertical, to the lowest point of the curve; to find the whole time of descent.

If a be the radius of the generating circle, the required time will be equal to $\pi\left(\frac{a}{g}\right).$

This result, being independent of the initial position of the particle, shews that the time of descent will be the same from whatever point in the curve the motion commences. This elegant mechanical property of the Cycloid, from which it has received the name of a Tautochronous Curve, was first discovered by Huyghens, *Horolog. Oscill.* Pars 11.

(12) To compare the times in which a particle acted on by gravity will descend down the chord and the arc AB, (fig. 118), from the base of an inverted cycloid of which the axis is vertical.

From A let fall a perpendicular upon the normal BC at B; then

Time down chord AB: time down arc AB:: chord AB: AC.

Sault; Phil. Trans. 1698, Num. 246, p. 425.

(13) A particle under the action of gravity falls down an arc OB (fig. 119) of one of the loops of a Lemniscata, of which the axis OA is inclined at an angle of 45° to the horizon; to determine the time of the descent.

Let a denote the semi-axis of the corresponding equilateral hyperbola, θ the angle between the chord OB and the axis OA of the loop, and T the required time. Then

$$T = \left(\frac{8^{\frac{1}{2}}a}{g}\right)^{\frac{1}{4}} \left\{ \tan \left(\frac{1}{4}\pi - \theta\right) \right\}^{\frac{1}{4}}.$$

This expression for the time is the same as that for the descent of a particle down the chord OB; a mechanical property of the Lemniscata which was discovered by Saladini, *Memorie dell'* Istituto Nazionale Italiano, tom. 1. parte 2.

(14) A particle slides down a plane of given length, inclined at an angle θ to the horizon, and is reflected by the horizontal plane; to determine the value of θ that the range on the horizontal plane may be the greatest possible, the particle being perfectly elastic. $\theta = \sin^{-1} \left\{ \left(\frac{a}{3} \right)^{\frac{1}{2}} \right\}.$

(15) Two equal spherical particles of given elasticity are placed at two points in the circumference of a vertical circle, the radii of these two points making angles of 60° on each side of the radius which tends vertically downwards; to determine the sum of the chords of the arcs described by each particle before it ceases to move.

If a = the radius of the circle, and e = the common elasticity of the particles, the required space will be equal to

$$\frac{a}{1-e}$$

(16) A spherical particle A impinges with a velocity u in a horizontal direction upon a spherical particle B, which is resting

at the lowest point of an inverted cycloid with its axis vertical; to determine the velocities of A and B after any number of impacts, the volumes of the particles being equal, while their masses differ in any proposed degree.

The velocities of A, B, after x impacts, will be respectively, e denoting their common elasticity,

$$\frac{A-A(-e)^{*}}{A+B}u, \frac{A+B(-e)^{*}}{A+B}u.$$

Sect. 2. Pressure of a moving Particle on immoveable plane Curves.

The general value of the reaction of a curve against a particle which is moving along the curve, is given by the formula

$$R = Y \frac{dx}{ds} - X \frac{dy}{ds} \pm \frac{1}{\rho} \frac{ds^2}{dt^2} = N \pm \frac{v^2}{\rho} \dots (A),$$

where N represents the resolved part of the whole accelerating force on the particle estimated along the normal in an opposite direction to that in which the reaction R exerts itself, and ρ denotes the radius of curvature of the curve. In this formula the positive or the negative sign is to be taken accordingly as the particle is moving on the concave or on the convex side of the curve.

This formula was first given by L'Hôpital¹ in the discussion of John Bernoulli's problem of the Curve of Equal Pressure.

When the expression for R becomes equal to zero, the particle will either leave the curve or will move along it freely without experiencing any reaction; and the analytical condition

$$\frac{v^2}{\rho} = \mp N$$

shews that, on the commencement of free motion, the normal accelerating force and the centrifugal force of the particle must be equal and opposite.

(1) A particle, starting with a given velocity from the vertex of a parabola, of which the axis is vertical, descends down the convex side of the curve by the action of gravity; to find the reaction of the curve at any point of the descent.

¹ Mém. de l'Acad. des Sciences de Paris, 1700, p. 9.

The resolved part of the force of gravity along the normal in a direction opposite to the reaction is $g \frac{dy}{ds}$, and therefore by (A), the particle moving on the convex side of the curve,

$$R = g \frac{dy}{ds} - \frac{1}{\rho} \frac{ds^2}{dt^2}.$$

Now the equation to the parabola being $y^2 = 4mx$,

$$\frac{dy}{ds} = \frac{m^{\frac{1}{4}}}{(m+x)^{\frac{1}{4}}} \text{ and } \rho = \frac{2}{m^{\frac{1}{4}}} (m+x)^{\frac{3}{4}}.$$

Also, if h be the altitude due to the initial velocity of the particle, we have $\frac{ds^2}{dt^2} = 2g(x+h).$

Hence $R = \frac{gm^{\frac{1}{4}}}{(m+x)^{\frac{1}{4}}} - \frac{gm^{\frac{1}{4}}}{(m+x)^{\frac{3}{4}}}(x+h)$ $= m^{\frac{1}{4}}g\frac{m-h}{(m+x)^{\frac{3}{4}}}.$

If h = m, then the pressure during the whole motion will be equal to zero; and the particle will describe the parabola freely. If h were greater than m, since, from the nature of the case, R cannot have any negative value, the particle would from the first proceed in a path different from the parabola in question. If, instead of supposing the particle to move on a mere curve, we were to conceive it to be moving within an indefinitely thin parabolic tube, R might be negative; and in fact always would be negative, supposing h to be greater than m, when the motion would be the same as if the particle were moving along the concave side of the parabolic curve.

Euler; Mechan. tom. 11. p. 64.

(2) A particle attracted towards two centres of force, varying inversely as the square of the distance, moves in an hyperbolic groove, of which the foci are the centres of force; to find the pressure on the groove at any point, the particle being supposed to move on the concave side.

Let P (fig. 120) be the position of the particle at any time t; S, H, the foci of the hyperbola, SP = r, HP = r'; a the semi-transverse axis; PT a tangent at P; $\angle SPT = \phi = \angle HPT$; μ , μ' , the absolute forces towards S, H.

Resolving forces at right angles to the tangent at P, we have, by the equation (A),

$$R = \left(\frac{\mu'}{r'^2} - \frac{\mu}{r^2}\right) \sin \phi + \frac{v^2}{\rho} \dots (1);$$

also, for the value of v at any time, there is

$$v^{2} = 2 \int \left(-\frac{\mu'}{r'^{2}} dr' - \frac{\mu}{r^{2}} dr \right) + C$$
$$= \frac{2\mu'}{r'} + \frac{2\mu}{r} + C.$$

Let f, f', be the initial values of r, r', and β the initial value of v; then $2u' \quad 2u$

$$\beta^2 = \frac{2\mu'}{f} + \frac{2\mu}{f} + C,$$

and therefore

$$v^2 = \frac{2\mu'}{r'} + \frac{2\mu}{r} - \frac{2\mu'}{f'} - \frac{2\mu}{f} + \beta^2.$$

Hence, from (1), we have

$$R\rho = \left(\frac{\mu'}{r'^2} - \frac{\mu}{r^2}\right)\rho \sin \phi + \frac{2\mu'}{r'} + \frac{2\mu}{r} - \frac{2\mu'}{f'} - \frac{2\mu}{f} + \beta^2.$$

But $2\rho \sin \phi$ is equal to the chord of curvature through S, which, by the nature of the hyperbola, is equal to $\frac{2rr'}{a}$; hence

$$R\rho = \left(\frac{\mu'}{r'^2} - \frac{\mu}{r^2}\right) \frac{rr'}{a} + \frac{2\mu'}{r'} + \frac{2\mu}{r} - \frac{2\mu'}{f'} - \frac{2\mu}{f} + \beta^2$$

$$= \frac{\mu'}{ar'} \frac{(r+2a)}{ar} - \frac{\mu}{ar} \frac{(r'-2a)}{ar} - \frac{2\mu'}{f'} - \frac{2\mu}{f} + \beta^2$$

$$= \frac{\mu'}{a} - \frac{\mu}{a} - \frac{2\mu'}{f'} - \frac{2\mu}{f} + \beta^2$$

$$= \frac{\mu'}{af'} \frac{(f'-2a)}{af} - \frac{\mu(f+2a)}{af} + \beta^2$$

$$= \frac{\mu'f}{af'} - \frac{\mu f'}{af} + \beta^2.$$

If the initial velocity be zero, and the particle be attracted at the commencement of its motion with equal intensity by the two centres of force; then $\beta = 0$, $\frac{\mu}{f^2} = \frac{\mu'}{f'^2}$, and therefore R = 0 during the whole motion. Hence the particle would under these circumstances describe the hyperbola freely.

(3) A particle is moving along the convex side of an equiangular spiral, towards the pole of which it is attracted by a force varying as any power of the distance; to determine the reaction of the curve at any time during the motion.

Let r be the distance of the particle from the pole at any time, μr^n the attractive force, a the constant angle between the curve and the radius vector, β the initial velocity, and a the initial value of r. Then, by the formula (A), N being equal to $\mu r^n \sin a$, we have

 $R = \mu x^n \sin a - \frac{x^2}{\rho} \dots \dots \dots \dots (1).$

Again, estimating the velocity v of the particle in a direction corresponding to an increase of r, and denoting by ds an element of its path, we have

Now, by the nature of the curve, $\cos a \, ds = dr$; hence, from (2),

$$v\frac{dv}{dr}=-\mu r^{\alpha};$$

integrating and observing that β , a, are the initial values of v, r, we get 2μ

 $v^2 - \beta^2 = -\frac{2\mu}{n+1} (r^{n+1} - a^{n+1}) \dots (3).$

Again, p denoting the perpendicular from the pole upon the tangent to the curve, we have, since $p = r \sin a$,

$$\rho = r \frac{dr}{dp} = \frac{r}{\sin a} \dots (4).$$

From (1), (3), (4), we obtain

$$R = \mu r^{n} \sin a - \frac{\sin a}{r} \left\{ \beta^{2} - \frac{2\mu}{n+1} \left(r^{n+1} - a^{n+1} \right) \right\}$$
$$= \mu \frac{n+3}{n+1} r^{n} \sin a - \frac{\beta^{2} \sin a}{r} - \frac{2\mu \sin a}{(n+1)r} a^{n+1}.$$

Euler; Mechan. tom. II. p. 86.

(4) A particle, starting from rest, descends down the convex side of a circle from a given point in its circumference; to find where it will leave the curve.

Let O (fig. 121) be the centre of the circle, AO being a vertical radius. Let P be the initial position of the particle, Q its point of departure; PM, QN, horizontal lines. Join OQ, and let $\angle AOQ = \phi$; a =the radius of the circle.

Then, the centrifugal force at Q being equal to the normal component of gravity, we have

$$\frac{v^2}{a}=g\,\cos\phi\,;$$

but, denoting MN by x,

$$v^2=2gx;$$

hence, putting MO = c,

$$2x = a \cos \phi = c - x$$
, $x = \frac{1}{3}c$.

Fontana; Memorie della Societa Italiana, 1782, p. 175.

(5) A particle acted on by gravity oscillates in a circular arc; to find the reaction of the curve at any point.

Let O (fig. 122) be the centre of the circle; P the position of the particle at any time; A the lowest point of the circle; $\triangle AOP = \theta$. Then, θ being initially equal to α and the velocity zero, we have

$$\mathbf{R}=g\left(3\cos\,\theta-2\,\cos\,a\right).$$

(6) A particle descends down the convex side of a logarithmic curve placed with its asymptote parallel to the horizon; to find where it leaves the curve.

Let P (fig. 123) be the point in which the particle is placed, and Q the point of its departure; OM = h, ON = x. Then, the equation to the curve being $y = \log_a x$, we have, putting $A = \log_a a$,

$$x = \frac{1}{A} \left\{ Ah + (1 + A^2h^2)^{\frac{1}{2}} \right\}.$$

Fontana; *Ib.* p. 182.

(7) A particle descends from rest down the convex side of an ellipse with its major axis vertical, from a given point in the curve; to determine where it will leave the ellipse.

Let the highest point of the ellipse be taken as the origin of co-ordinates, the axis of x being vertical, and that of y horizontal.

Let a, b, denote the semi-axes major and minor; h the initial distance of the particle from the axis of y, and x the distance of the point at which it leaves the curve. Then the value of x will be a root of the cubic equation

$$(a^2 - b^2)(x^3 - 3ax^2) - 3a^2b^2x + a^3(b^2 + 2ah) = 0.$$

Fontana; *Ib.* p. 175.

(8) A particle descends from rest down the convex side of the Cissoid of Diocles, which is so placed as to have its asymptote vertical; the initial place of the particle being known, to find the point in which it will leave the curve.

Let P (fig. 124) be the initial position of the particle, and Q its place on leaving the curve. PS, QN, at right angles to Ox, Oy; PS = h, QN = x. Then, a being the radius of the generating circle, the value of x will be a root of the cubic equation

$$x^{3} - \frac{16a}{9} x^{2} + \frac{64a^{2} + 36h^{2}}{81} x - \frac{8ah^{2}}{9} = 0.$$

If the motion commence at the cusp O, h = 0, and therefore

$$x = \frac{8}{9} a$$
. Fontana; *Ib.* p. 181.

(9) A particle is projected with a given velocity along the convex side of a parabola from a given point in the curve; to determine the reaction of the curve at any time of the motion, the particle being always attracted to the focus by a force varying inversely as the square of the distance.

Let S (fig. 125) be the focus of the parabola; B the point from which the particle is initially projected in the direction of the tangent BT; P the position of the particle after any time, SP = r, SB = a, SA = m, β = the velocity of projection, μ = the absolute force towards S. Then, at P

$$R = \left(\frac{m}{r^3}\right)^{\frac{1}{4}} \left(\frac{\mu}{a} - \frac{\beta^2}{2}\right).$$

(10) A particle is projected with a given velocity at the highest point of a circle in a vertical plane along the concave side of the curve; to determine the pressure on the curve at any point in its path.

Let AOB (fig. 126) be the vertical diameter, O being the centre of the circle; P the position of the particle at any time, OP = a, $\angle AOP = \theta$; β the velocity of projection at A; then, for the pressure at P,

$$R = \frac{\beta^2}{a} + g (2 - 3 \cos \theta).$$

Suppose that R = 0 initially; then $\frac{\beta^2}{a} = g$, and

$$R = 3g \text{ vers } \theta$$

= $6g$, when $\theta = \pi$;

which shews that when the particle arrives at the lowest point the reaction is six times the force of gravity.

Euler; Mechan. tom. 11. p. 65, Cor. 7.

(11) A particle moves along the convex side of an ellipse under the action of two forces tending to the foci and varying inversely as the square of the distance, and a third force tending to the centre and varying as the distance; to find the reaction of the curve at any point.

Let R denote the reaction of the curve on the particle at any point, ρ the radius of curvature; f, f' the initial focal distances, and μ , μ' the corresponding absolute forces; μ'' the absolute force to the centre, 2a the axis major of the ellipse, and β the initial velocity. Then

$$R\rho = \frac{\mu f'}{af} + \frac{\mu' f}{af'} + \mu'' f f' - \beta^2.$$

If β' , β'' , denote the velocities which the particle ought to have initially to revolve freely round the three centres of force taken separately,

$$\beta'^2 = \frac{\mu f'}{af}, \quad \beta''^2 = \frac{\mu' f}{af'}, \quad \beta''^2 = \mu' f f';$$

and therefore, when the forces are taken conjointly, it will revolve about them freely when

$$\beta^2 = \beta'^2 + \beta'^2 + \beta''^2$$
.

(12) An ellipse is placed with its major axis in a vertical position; to find the velocity with which a particle must be projected vertically upwards from the extremity of the minor axis

along the interior of the elliptic arc, so that after quitting the curve it may pass through the centre.

If a, b, denote the semi-axes major and minor, the required velocity will be equal to

$$\left\{ \frac{(8a^2+b^2)g}{3a_13^{\frac{1}{4}}} \right\}^{\frac{1}{4}}.$$

SECT. 3. Inverse Problems on the Motion of a Particle along immoveable plane Curves.

(1) To find a curve EPF (fig. 127) such that, A and B being two given points in the same horizontal line, the sum of the times in which a particle will descend by the action of gravity down the straight lines AP, BP, may be the same whatever point in the curve P may be.

Bisect AB in O; let Ox, a vertical line, be the axis of x, and OAy, which is horizontal, the axis of y; let AB = 2a. Then, x, y, being the co-ordinates of P, the times down AP, BP, will be respectively equal to

$$\left\{\frac{x^2+(a-y)^2}{\frac{1}{2}gx}\right\}^{\frac{1}{4}}, \quad \left\{\frac{x^2+(a+y)^2}{\frac{1}{2}gx}\right\}^{\frac{1}{4}}.$$

Hence, k denoting the sum of the times,

$$\left(\frac{1}{3}k^3qx\right)^{\frac{1}{4}}=\left\{x^2+(a-y)^2\right\}^{\frac{1}{4}}+\left\{x^2+(a+y)^2\right\}^{\frac{1}{4}}\ldots\ldots(1).$$

Putting $\frac{1}{2}k^2g = 4c$, and squaring both sides of the equation, we have

$$2cx = a^{2} + x^{2} + y^{2} + \left\{x^{2} + (a - y)^{2}\right\}^{\frac{1}{4}} \left\{x^{2} + (a + y)^{2}\right\}^{\frac{1}{4}},$$

$$(2cx - a^{2} - x^{2} - y^{2})^{2} = \left\{x^{2} + (a - y)^{2}\right\} \left\{x^{2} + (a + y)^{2}\right\}.$$

Developing both sides of the equation, and simplifying, we shall readily find that

$$c^2x^2 - a^2cx - cx^3 - cxy^2 = -a^2y^2$$
, and, therefore, $y^2 = cx \frac{x^2 - cx + a^2}{a^2 - cx} \cdot \dots (2)$,

which is the equation to the required curve.

If we trace this curve, we shall find it to consist of a branch VOV' having an asymptote parallel to the axis of y, and of an

oval EFP. The oval is the portion of the curve which corresponds to the problem which we are considering. The infinite branch VOV' would correspond to the condition that the times down AP', BP', shall have a constant difference: in which case we should have had, instead of the equation (1),

$$\left(\frac{1}{3}k^2gx\right)^{\frac{1}{4}}=\left\{x^2+(a+y)^2\right\}^{\frac{1}{4}}-\left\{x^2-(a-y)^2\right\}^{\frac{1}{4}};$$

whence, by the involution, we should have obtained the same equation (2). The curve has pretty much the shape of the Conchoid, although its equation is essentially different.

Fuss; Mémoires de l' Acad. de St. Pétersb. 1819.

(2) A particle, not acted on by any forces, is constrained to move within a thin tube of such a form that the acceleration of the particle parallel to a given straight line is invariable; to determine the equation to the path of the particle.

Let the axis of x be taken parallel to the given line; let c be the constant acceleration of the particle parallel to the axis of x, and β its velocity within the tube, which will be invariable. Then

$$\frac{dx^3}{dt^3} + \frac{dy^2}{dt^3} = \beta^3. \dots (1);$$

but, by the condition of the problem, $\frac{d^2x}{dt^2} = c$, and therefore, the

axis of y being so chosen that $\frac{dx}{dt} = 0$ when x = 0,

$$2\frac{dx}{dt}\frac{d^2x}{dt^2}=2c\frac{dx}{dt}, \frac{dx^2}{dt^2}=2cx \dots (2).$$

Eliminating dt between the equations (1) and (2), we get

$$2cx\left(1+\frac{dy^2}{dx^3}\right)=\beta^2,$$

or, putting $\frac{\beta^2}{4c} = a$,

$$\frac{dy}{dx} = \left(\frac{2a-x}{x}\right)^{1};$$

whence, by integration, the position of the axis of x being supposed such that x = 0 when y = 0,

$$y = (2ax - x^2)^{\frac{1}{4}} + a \text{ vers}^{-1} \frac{x}{a};$$

which is the equation to a cycloid of which the axis is parallel to the given line.

There is an elaborate investigation by Euler, in the *Mémoires* de l'Académie de St. Pétersb., tom x. p. 7, on the nature of the curve of constraint when the particle is subject to the action of gravity, and the direction of uniform acceleration is horizontal. A notice of this problem may be seen in the Bulletin des Sciences de Bruxelles, tom. IX.

(3) To determine the curve down which a particle may descend by the action of gravity, so as to describe equal vertical spaces in equal times, the tangent to the curve at the point where the motion commences being vertical.

Let O (fig. 128) be the point where the motion commences, Ox the axis of x touching the required curve OA at O, Oy the axis of y at right angles to Ox; OM = x, PM = y, β = the invariable velocity of the particle parallel to Ox.

Then
$$\frac{ds^2}{dt^2} = C + 2gx$$
, C being a constant quantity,
$$\frac{ds^2}{dx^2} \frac{dx^2}{dt^2} = C + 2gx.$$

But $\frac{dx}{dt} = \beta$; hence

$$\beta^2 + \beta^2 \frac{dy^2}{dx^2} = C + 2gx$$

But when x = 0, $\frac{dy}{dx} = 0$; and therefore

$$\beta^2 = C, \quad \beta^2 \frac{dy^2}{dx^2} = 2gx,$$

$$\frac{dy}{dx} = \frac{(2g)^{\frac{1}{4}}}{\beta}x^{\frac{1}{4}}, \quad y = \frac{2(2g)^{\frac{1}{4}}}{3\beta}x^{\frac{3}{4}},$$

no constant being added because x = 0, y = 0, simultaneously. The required curve OA is therefore the semi-cubical parabola, O being the cusp, and Ox the axis.

This curve is called the Isochrone. It was proposed by Leibnitz,* as a challenge to the disciples of Des Cartes, who, from an

^{*} Nouvelles de la Republique des Lettres, Septembre 1687.

excessive attachment to the geometry of their master, affected to despise the methods of the Differential Calculus. No solution was communicated by any of the Cartesians. Huyghens alone successfully accepted the challenge, by whom a geometrical solution was given in the Nouvelles de la Republique des Lettres, Octobre 1687. The solution by Leibnitz appeared for the first time in the Acta Erudit., Lips. 1689, p. 196, et sq. The solutions both of Huyghens and of Leibnitz were synthetical. An analytical solution was given afterwards for the first time by James Bernoulli.*

(4) A particle is projected with a given velocity from a point Δ (fig. 129) along a horizontal line ΔO towards a point O; to find the curve along which it must be constrained to move that it may approach the point O uniformly; the particle being acted on by gravity, and ΔO being a tangent to the required curve.

Let P be the position of the particle at any time; AO = a, OP = r, $\angle AOP = \theta$; $\beta =$ the velocity of the particle at its initial position A. For the motion of the particle at any point in its descent there is

$$\frac{dr^3}{dt^2} + r^3 \frac{d\theta^3}{dt^2} = \beta^2 + 2gr \sin \theta,$$

or

$$\frac{dr^3}{dt^2}\left(1+r^3\frac{d\theta^3}{dr^3}\right)=\beta^3+2gr\sin\theta.$$

But, by the condition of the problem, $\frac{dr}{dt} = C$, a constant quantity: hence

$$C^{2}\left(1+r^{2}\frac{d\theta^{2}}{dr^{2}}\right)=\beta^{2}+2gr\sin \theta.$$

But, initially, $\theta = 0$, $r \frac{d\theta}{dr} = 0$; hence $C^* = \beta^2$, and therefore

$$\beta^2 r^3 \frac{d\theta^3}{dr^2} = 2gr \sin \theta,$$

$$\frac{dr}{r^{\frac{1}{4}}} = \frac{\beta}{(2g)^{\frac{1}{4}}} \frac{d\theta}{(\sin\theta)^{\frac{1}{4}}}:$$

^{*} Act Erudit., Lips. 1690, p. 217.

integrating, and observing that $\theta = 0$, r = a, initially; we have

$$r^{\frac{1}{4}} - a^{\frac{1}{4}} = \frac{\beta}{2(2g)^{\frac{1}{4}}} \int_{0}^{\theta} \frac{d\theta}{(\sin \theta)^{\frac{1}{4}}},$$

which is an equation for the construction of the path of the particle.

The particle will move from A along ABCO to the point O with a uniform velocity of approach; it will afterwards move from O along OcBa with a uniform velocity of recession. When it has arrived at a it will proceed uniformly along Oa produced.

This curve has been called the Paracentric Isochrone by Leibnitz, by whom the problem was originally proposed as a challenge to the mathematicians of the day, in the Acta Erudit., Lips. 1689, p. 198. Several years elapsed before the problem received a solution. At length James Bernoulli succeeded in obtaining one, which appeared in the Acta Erudit., Lips. 1694, p. 277. Solutions were shortly afterwards published by Leibnitz and John Bernoulli, in the Acta Erudit., Lips. 1694, p. 371, 394. The problem was afterwards generalized by Varignon in the Mémoires de l'Académie des Sciences de Paris, 1699, p. 9, et sq.

(5) To find the nature of the curve OPA (fig. 130) such that a particle acted on by gravity will descend down any arc OP in the same time as down its chord.

Let Ox be vertical, Oy horizontal, PM parallel to yO. Let OP = r, $\angle xOP = \theta$, arc OP = s, $OM = r \cos \theta$. Then, since the velocity acquired down the arc OP is the same as that which is due to falling freely down OM,

$$\frac{ds^2}{dt^2} = 2gr\cos\theta, dt = \frac{1}{(2g)^{\frac{1}{4}}} \frac{ds}{(r\cos\theta)^{\frac{1}{4}}};$$

and therefore the whole time of descent down OP is equal to

$$\frac{1}{(2g)^{\frac{1}{4}}} \int_{a}^{s} \frac{ds}{(r \cos \theta)^{\frac{1}{4}}}.$$

But the time of descent down the chord OP is equal to

$$\left(\frac{2}{g}\right)^{\frac{1}{q}} \left(\frac{r}{\cos \theta}\right)^{\frac{1}{q}};$$

and, therefore, by hypothesis,

$$\left(\frac{2}{g}\right)^{\frac{1}{2}} \left(\frac{r}{\cos\theta}\right)^{\frac{1}{2}} = \frac{1}{(2g)^{\frac{1}{2}}} \int_{0}^{s} \frac{ds}{(r\cos\theta)^{\frac{1}{2}}},$$

$$2\left(\frac{r}{\cos\theta}\right)^{\frac{1}{2}} = \int_{0}^{s} \frac{ds}{(r\cos\theta)^{\frac{1}{2}}}.$$

Differentiating both sides of the equation, we have

$$\left(\frac{\cos \theta}{r}\right)^{\frac{1}{2}} \frac{\cos \theta \ dr + r \sin \theta d\theta}{(\cos \theta)^{2}} = \frac{ds}{(r \cos \theta)^{\frac{1}{2}}},$$

 $\cos\theta dr + r \sin\theta d\theta = \cos\theta (dr^3 + r^3d\theta^3)^{\frac{1}{4}}.$

Squaring both sides and simplifying,

$$2 \sin \theta \cos \theta \, r dr d\theta = r^2 \cos 2\theta \, d\theta^2,$$

$$\frac{dr}{r} = \frac{\cos 2\theta}{\sin 2\theta} d\theta.$$

Integrating,

$$\log r^2 = \log a^2 + \log \sin 2\theta, r^2 = a^2 \sin 2\theta,$$

where a^2 is some constant quantity.

From O draw OE, bisecting the angle xOy, and let $\angle POE = \phi$; then, since $\theta = \frac{1}{4}\pi - \phi$, we have

$$r^2=a^2\cos 2\phi;$$

which is the equation to the Lemniscata of James Bernoulli, O being the centre and A the vertex of the equilateral hyperbola. This very beautiful problem is due to Saladini.

Saladini; Memorie dell' Istituto Nazionale Italiano, tom. 1. parte 2. Fuss; Mémoires de l'Acad. de St. Pétersb. 1819.

(6) To find the equation to the tautochrone when a particle is acted on by any forces whatever in one plane.

A tautochrone is a curve along which a particle acted on by any assigned forces will arrive in the same time at a given point from whatever point in the curve its motion commences. Let A (fig. 131) be any assigned point, and E any point whatever in the curve AEB; then the time from E to A is to be independent of the position of E.

Let P be any point in AE; AP = s, AE = a, S= the sum of the resolved parts of the accelerating forces on the particle along the tangent PT at the point P. Then, for the motion of the particle, $\frac{d^3s}{dA} = -S, \quad \frac{ds^2}{dA} = C - 2\int S ds.$

•

But, the particle being supposed to have no initial velocity,

$$0 = C - 2 \int^a S ds;$$

and therefore

$$\frac{ds^2}{dt^2} = 2 \int_s^a S ds \dots (1),$$

$$dt = -\frac{1}{2^{\frac{1}{4}}} \frac{ds}{\left(\int_{s}^{a} S ds\right)^{\frac{1}{4}}}.$$

The time from E to A is equal to

$$\frac{1}{2^{\frac{1}{4}}} \int_{0}^{a} \frac{ds}{\left(\int_{s}^{a} Sds\right)^{\frac{1}{4}}};$$

and this formula must be independent of a. Hence we must

$$\int \frac{ds}{\left(\int_{s}^{a} S ds\right)^{\frac{1}{4}}} = \phi\left(\frac{s}{a}\right),$$

where $\phi\left(\frac{s}{a}\right)$ denotes some function of $\frac{s}{a}$. Hence

$$\frac{1}{\left(\int_{s}^{a} S ds\right)^{\frac{1}{2}}} = \frac{1}{a} \phi'\left(\frac{s}{a}\right), \quad \int_{s}^{a} S ds = \frac{a^{2}}{\left\{\phi'\left(\frac{s}{a}\right)\right\}^{2}};$$

and therefore, differentiating with respect to s,

$$- Sds = d \frac{a^3}{\left\{\phi'\left(\frac{s}{a}\right)\right\}^2}.$$

But the tautochrone AB being an invariable curve, whatever be the value of a, it is manifest that a must not appear in this equation; hence

$$\left\{\phi'\left(\frac{s}{a}\right)\right\}^2 = -\frac{a^2}{As^2}$$
, where A is a constant quantity,

and therefore

$$S = ks \dots (2),$$

k being some constant quantity.

Hence, by (1) and (2),

$$\frac{ds^2}{dt^2}=k\left(a^2-s^2\right),$$

and therefore, if τ denote the time the motion from E to A,

$$\tau = -\frac{1}{k^{\frac{1}{4}}} \int_{a}^{0} \frac{ds}{(a^{2} - s^{2})^{\frac{1}{4}}} = \frac{\pi}{2k^{\frac{1}{4}}}, \ k = \frac{1}{4} \pi^{2} \tau^{2}.$$

Hence we have, from (2),

$$S = \frac{1}{4} \pi^2 \tau^2 s,$$

which is a differential equation to the tautochrone.

The direct problem of Tautochronism in the case when gravity is the accelerating force, was first considered by Huyghens, in his *Horolog. Oscill.*, where he proves the inverted cycloid with its axis vertical to be tautochronous. The inverse problem was first considered by Newton, *Princip.* lib. 1. sect. 10. See also Euler, *Comment. Petrop.* 1729, and *Mechan.* tom. 11. p. 211.

(7) A particle is acted on by an attractive force tending towards a fixed centre, and varying as the distance; to find the tautochrone.

Let μ denote the absolute force of attraction, r the radius vector at any point of the curve, p the perpendicular from the pole upon the tangent at the point, ϕ the inclination of the tangent to the radius vector.

Then, by the formula of the preceding general problem, we have, putting $\mu r \cos \phi$ for S,

$$\mu r \cos \phi = \frac{1}{4} \pi^2 \tau^2 s$$
.

whence

$$\mu d(r\cos\phi) = \frac{1}{4}\pi^2\tau^2 ds;$$

but $ds \cos \phi = dr$; hence we have

$$\mu r \cos \phi \ d (r \cos \phi) = \frac{1}{4} \pi^2 \tau^2 r dr;$$

integrating, we get

$$\mu r^2 \cos^2 \phi + C = \frac{1}{4} \pi^2 \tau^2 r^2$$

or

$$\mu (r^2 - p^2) + C = \frac{1}{4} \pi^2 \tau^2 r^2$$

Let c be the value of r when s = 0, and therefore when $\phi = \frac{1}{2}\pi$; then also p = c, and consequently

 $C = \frac{1}{4} \pi^2 c^3 \tau^3.$ $\mu (r^3 - p^3) + \frac{1}{4} \pi^2 c^3 \tau^2 = \frac{1}{4} \pi^3 \tau^2 r^3,$ $\mu p^2 = (\mu - \frac{1}{4} \pi^3 \tau^3) r^3 + \frac{1}{4} \pi^3 c^2 \tau^3,$

Hence

which is the differential equation to the curve.

Euler; Mechan. tom. II. p. 208.

(8) An infinite number of similar curves originate at a given point; to determine the corresponding synchronous curve, or the curve which shall cut them in such a manner that a particle acted on by gravity may describe the intercepted arcs in equal times.

Let O (fig. 132) be the given point, and CPD the synchronous curve intercepting the arc OP of the curve OPQ, which is one of the similar curves. Let Ox, a vertical line, be taken as the axis of x, and Oy, at right angles to it, as the axis of y. Let OM = x, PM = y, OP = s. Then, if k denote the time down OP, which by hypothesis is constant for every point P in the curve CPD, we have

$$k = \int_0^s \frac{ds}{(2gx)^{\frac{1}{2}}} = \int_0^s \frac{(1+p^2)^{\frac{1}{2}}}{(2gx)^{\frac{1}{2}}} dx \dots (1),$$

where p is equal to $\frac{dy}{dx}$.

Now, by the nature of similar curves, the equation to the curve OPQ is of the form

$$F\left(\frac{x}{a}, \frac{y}{a}\right) = 0$$
, or $y = af\left(\frac{x}{a}\right) \dots (2)$,

where F, f, denote certain functions of the quantities to which they are prefixed, a being the value of the general parameter of the class of similar curves for the individual curve OPQ. Hence, assuming

 $x = a\tau$, and therefore $y = af(\tau)$, by (2), (3), we have, from (1),

$$k = a^{\frac{1}{4}} \int_{0}^{\tau} \frac{(1+T^{2})^{\frac{1}{4}}}{(2a\tau)^{\frac{1}{4}}} d\tau = a^{\frac{1}{4}} \phi(\tau) \dots (4),$$

where $T = \frac{dy}{dx} = \frac{d}{d\tau} f(\tau)$, and $\phi(\tau)$ is some function of τ . Hence, from (3) and (4), there is

Eliminating τ between these two last equations, we shall obtain an equation in x, y, the required equation to the sychronous curve.

If the integration indicated in the equation (1) can be effected, then it is needless to have recourse to the subsidiary symbol τ . We have merely in this case to eliminate, after the performance of the integration, the parameter a, by the aid of the equation (2). It rarely happens, however, that we can execute the operation of integration, and under these circumstances the equations (5) will enable us to construct the synchronous curve by the method of quadratures; a pair of values of x, y, and therefore a point in the synchronous curve, being ascertained approximately for every numerical value which we may assign to τ .

The problem of Synchronous Curves was first discussed by John Bernoulli, in the *Act. Erudit. Lips.* 1697, Mai. p. 206. The subject was afterwards investigated by Saurin, and by Euler.¹

(9) An assemblage of circles in the plane xOy, (fig. 132), all touch Ox in the point O; to determine the sychronous curve, Ox being vertical, and gravity the accelerating force; the descent being supposed to commence from O.

The equation to any one of the circles, its radius being a, will be $x^2 = 2ay - y^2$,

or
$$y = a \left\{ 1 - \left(1 - \frac{x^2}{a^2} \right)^{\frac{1}{4}} \right\}.$$

Adopting the notation of the preceding general problem, we have

$$f(\tau) = 1 - (1 - \tau^2)^{\frac{1}{4}},$$

$$T = \frac{df(\tau)}{d\tau} = \frac{\tau}{(1 - \tau^2)^{\frac{1}{4}}},$$

¹ Mechan. tom. 11. p. 47; Mém. de l'Acad. de St. Pétersb. 1819-1820, p. 20, 35.

$$\phi(\tau) = \int_0^{\tau} \frac{(1+T^2)^{\frac{1}{4}}}{(2g\tau)^{\frac{1}{4}}} d\tau = \frac{1}{(2g)^{\frac{1}{4}}} \int_0^{\tau} \frac{d\tau}{(\tau-\tau^3)^{\frac{1}{4}}}.$$

$$x = \frac{2gk^3\tau}{\left\{\int_0^{\tau} \frac{d\tau}{(\tau-\tau^3)^{\frac{1}{4}}}\right\}^2}, \quad y = \frac{2gk^3\left\{1-(1-\tau^2)^{\frac{1}{4}}\right\}}{\left\{\int_0^{\tau} \frac{d\tau}{(\tau-\tau^3)^{\frac{1}{4}}}\right\}^2},$$

Hence

whence the required curve may be constructed by the method of quadratures.

Euler; Mechan. tom. 11. p. 52.

(10) A particle acted on by any assigned accelerating forces in one plane moves along a curve from one given point to another; to determine the form of the curve that the whole time of the motion between the two points may be the least possible.

Let P (fig. 133) be any point in the required curve; OM = x, PM = y; A and B the two given points, AP = s; also let a, β , be the values of x at the points A, B. Then, v being the velocity of the particle at P,

$$dt = \frac{ds}{v} = \frac{(1+p^s)^{\frac{1}{v}}}{v} dx$$
, where $p = \frac{dy}{dx}$,

and the whole time from A to B will be equal to

Assume $\frac{(1+p^2)^{\frac{1}{4}}}{v} = V$; then, that the expression (1) may be a minimum, we have, by the Calculus of Variations, since V involves only p and v, of which the latter is a function of only x and y, $N - \frac{dP}{dx} = 0 \dots (2),$

where N, P, denote respectively the partial differential coefficients of V with regard to y, p; $\frac{dP}{dx}$ representing the total differential coefficient of P with respect to x.

Now
$$N = -\frac{1}{v^2} (1 + p^2)^{\frac{1}{4}} \frac{dv}{dy} \dots (3),$$

where $\frac{dv}{dy}$ signifies the partial differential coefficient of v with

regard to
$$y$$
; but $vdv = Xdx + Ydy \dots (3)$,

where X, Y, represent the resolved parts of the whole accelerating force on the particle parallel to Ox, Oy; and therefore, in

(3),
$$\frac{dv}{dy} = \frac{Y}{v}$$
. Hence

$$N = -\frac{Y}{v^3}(1+p^3)^{\frac{1}{4}} = -\frac{Y}{v^3}\frac{ds}{dx}.$$

Again,

$$P=\frac{p}{v(1+p^2)^{\frac{1}{2}}}=\frac{1}{v}\frac{dy}{ds}.$$

Hence, substituting for N and P in (2), we have

$$\frac{Y}{v^3}\frac{ds}{dx} + \frac{d}{dx}\left(\frac{1}{v}\frac{dy}{ds}\right) = 0,$$

$$\frac{Y}{v^3}\frac{ds}{dx} - \frac{1}{v^2}\frac{dv}{dx}\frac{dy}{ds} + \frac{1}{v}\frac{d}{dx}\frac{dy}{ds} = 0.$$

But, from (3),
$$\frac{dv}{dx} = \frac{1}{v} \left(X + Y \frac{dy}{dx} \right).$$

Hence
$$\frac{Y}{v^3}\frac{ds}{dx} - \frac{1}{v^3}\left(X + Y\frac{dy}{dx}\right)\frac{dy}{ds} + \frac{1}{v}\frac{d}{dx}\frac{dy}{ds} = 0;$$

and therefore, after a few obvious simplifications,

$$v^2 \frac{d}{dx} \frac{dy}{ds} = X \frac{dy}{ds} - Y \frac{dx}{ds} \dots \dots \dots (4).$$

If from this equation we eliminate v by the aid of (3), we shall obtain a differential equation of the second order, which is the equation to the required curve. The two arbitrary constants introduced by the integration are to be determined from the conditions that the curve shall pass through the two given points A and B.

The equation (4) is equivalent to

$$\frac{v^2}{\rho} = Y \frac{dx}{ds} - X \frac{dy}{ds},$$

where ρ denotes the radius of curvature at the point P; a result which shows that the pressure on the curve due to the centrifugal

force, is equal to that which arises from the accelerating forces which act upon the particle.

The curve in question belongs to a class of mechanical curves called Brachystochrones, which are characterized by the general property that a particle under the action of assigned accelerating forces, shall move along them between given limits in the least time possible.

The problem of the Brachystochrone between two given points, when gravity is the accelerating force, was proposed by John Bernoulli', as a challenge to the mathematicians of the day. Six months was the time allotted for its solution. Leibnitz was immediately successful, and communicated his good fortune by letter to Bernoulli. No other solution however having made its appearance within the prescribed time, Bernoulli, in conformity with the desire of Leibnitz, consented to prorogue the term of the challenge to the following Easter, the results obtained by himself and Leibnitz being suppressed for that interval. programme was accordingly published at Groningen, in January 1697, again announcing the problem and repeating the challenge. In consequence of this delay solutions were obtained by three other mathematicians: by Newton3, anonymously; by James Bernoulli; and by L'Hôpital. The solution of Leibnitz was announced in the Acta Erudit. Lips. Mai. 1697, p. 203. The conclusions of Newton, Leibnitz, and L'Hôpital, were given without the analysis. John Bernoulli gave two different solutions, one direct and the other indirect. The latter was published in the Act. Erudit. Lips. Mai. 1697, p. 207; the former was not made public till the year 1718, in a Memoir on Isoperimetrical problems, in the Mémoires de l'Académie des Sciences de Paris, p. 136; see also his works, tom. II. p. 266. A solution of the problem was afterwards given by Craig, who had merely seen Newton's result without consulting the analysis which had been given by John and by James Bernoulli.

¹ Act. Erudit, Lips. 1696, Jun. p. 269.

² Commerc. Epistol. Leibnitii et Bernoullii, Epist. 28.

³ Phil. Trans. 1697, Num. 224, p. 389.

⁴ Act. Erudit. Lips. Mai. 1697, p. 212.

³ Act. Erudit. Lips. ib. 217.

^a Phil. Trans. 1701, vol. xx11. p. 746.

(11) To find the brachystochrone when a particle, acted on by a central force attracting with an intensity which varies inversely as the square of the distance, moves along a curve from one given point to another.

Let A (fig. 134) be the point where the motion commences, and B the point at which the particle is to arrive in the shortest time possible. Let P be any point in the brachystochrone, S the centre of force; SP = r, p = the perpendicular from S upon the tangent at P, SA = a, $\phi =$ the angle between SP and the tangent at P, $\mu =$ the absolute force of attraction, v = the velocity, and $\rho =$ the radius of curvature at P. Then, since the pressure on the curve due to the centrifugal force must be equal to that due to the attraction, we have

$$\frac{v^2}{\rho} = \frac{\mu}{r^2} \sin \phi \dots \dots \dots (1).$$

But

$$v^2 = C - 2 \int \frac{\mu}{r^3} dr = C + \frac{2\mu}{r}$$
,

or, since v = 0 when r = a,

$$\mathbf{v}^2 = 2\mu \left(\frac{1}{r} - \frac{1}{a}\right) \dots \dots \dots (2).$$

Also, the curve being convex towards S,

$$\rho = -r \frac{dr}{dp} \dots (3).$$

From (1), (2), (3), we have

$$\frac{2\mu\left(\frac{1}{a}-\frac{1}{r}\right)}{r\frac{dr}{dp}} = \frac{\mu}{r^2}\frac{p}{r},$$

$$2\left(\frac{1}{a} - \frac{1}{r}\right) = \frac{p}{r^2} \frac{dr}{dp}, \ 2\frac{dp}{p} = \frac{adr}{r(r-a)};$$

integrating, we get

$$2\log p = \log C + \log \frac{r-a}{r},$$

$$\log p^2 = \log \left(C \frac{r-a}{r} \right);$$

but C must be a negative quantity, because, as will appear from the equation (2), a is greater than r; hence, putting -A for C,

we have, for the differential equation to the brachystochrone,

$$p^2=A\;\frac{a-r}{r}\;.$$

If from this equation we were to obtain, by integration, a relation between r and an angular co-ordinate θ , we should introduce another constant into the equation in addition to A. Both these constants would have to be determined by the conditions that the curve must pass through both A and B.

Euler; *Mechan.* tom. 11. p. 191.

(12) To find the inclination of a thin tube to the horizon, so that a descending particle may describe the greatest horizontal space in a given time.

Required angle of inclination = 45°.

(13) A particle having been placed at the point A, (fig. 135), moves along a thin tube APS towards a centre of attractive force in S which varies as any function of the distance; to find the nature of the curve of the tube that the time through any arc AP may be n times as great as through a portion Ap of the prime radius vector SA, Sp being equal to SP.

Let SP = r, SA = a, $\angle ASP = \theta$; then the equation to the curve will be $r = a \cdot \epsilon^{-\frac{\theta}{(a^2-1)^{\frac{1}{\theta}}}}.$

(14) A particle is projected with a given velocity from a point A (fig. 136) along a curve APO in which it is constrained to move, and is acted upon by a force always tending to O, and varying directly as the distance; to find the nature of this curve in order that the angular velocity of the radius vector OP may be invariable.

Let AO = a, OP = r, $\angle AOP = \theta$, μ^2 = the absolute force of attraction, ω = the angular velocity of OP, β = the initial velocity of the particle; then the equation to the curve will be

$$\left(\frac{\mu^2+\omega^2}{\omega^2}\right)^{\frac{1}{4}}\theta=\cos^{-1}\left\{\left(\frac{\mu^2+\omega^2}{\mu^2a^2+\beta^2}\right)^{\frac{1}{4}}r\right\}-\cos^{-1}\left\{\left(\frac{\mu^2+\omega^2}{\mu^2a^2+\beta^2}\right)^{\frac{1}{4}}a\right\}.$$

Euler; Mechan. tom. 11. p. 138.

(15) A particle is acted on by an attractive force tending to a

centre, and varying inversely as the square of the distance; to find the tautochrone.

If τ denote the time of the motion, and the notation remain the same as in problem (7), the differential equation to the tautochrone will be

$$p^2 = r^2 - \frac{\pi^2 \tau^2}{2\mu c} (r - c) r^{\delta}.$$

Euler; Mechan. tom. 11. p. 209.

(16) To find the tautochrone when the central attractive force is constant.

If f denote the constant central force, the equation to the tautochrone will be $p^2 = \left(1 + \frac{\pi^2 \tau^2 c}{2f}\right) r^3 - \frac{\pi^2 \tau^2}{2f} r^3.$

Euler; Mechan. tom. 11. p. 210.

(17) An infinite number of straight lines originate at a single point and lie in one plane; to determine the synchronous curve, gravity being the accelerating force.

The given point being taken as the origin of co-ordinates, the axis of x extending vertically downwards, and that of y being horizontal; the synchronous curve will be a circle of which the equation is $x^2 + y^2 = \frac{1}{2} gk^2x,$

where k denotes the common time of descent.

Euler; Mém. de l'Acad. de St. Pétersb. 1819, 1820, p. 22.

(18) There is an infinite number of cycloids, of which the bases all commence at the origin of co-ordinates, and coincide with the axis of y, which is horizontal; to find the synchronous curve, gravity being the accelerating force, and the motion commencing from the origin.

Let k denote the constant time of descent; then, the axis of x being vertical, the equation to the required curve depends upon the elimination of a between the two equations

$$x = a \text{ vers}^{-1} \left\{ k \left(\frac{g}{a} \right)^{\frac{1}{4}} \right\}, \ \ y = a \text{ vers}^{-1} \frac{x}{a} - (2ax - x^2)^{\frac{1}{4}},$$

and will cut all the cycloids at right angles.

John Bernoulli; Act. Erudit., Lips. 1697, Mai. p. 206.

(19) A particle, acted on by a central force attracting directly as the distance, moves along a curve from one given point to

another; to find the nature of the curve when it is brachystochronous.

Let A (fig. 134) be the point where the motion commences, and B the point where the particle is to arrive in the shortest time possible. Let P be any point in the brachystochrone; SP = r, p = the perpendicular from S, the centre of force, upon the tangent at P, a = SA. Then the equation to the curve between p and r will be $p^2 = A(r^2 - a^2)$,

where A is a constant quantity, which is the equation to the hypocycloid.

If from this equation we were to obtain by integration a relation between r and an angular co-ordinate θ , we should have another constant in the equation in addition to A. Both these constants would have to be determined by the conditions that the curve must pass through both A and B.

Euler; Mechan. tom. II. p. 191.

Sect. 4. Inverse Problems on the Pressure of a Particle on Smooth Fixed Curves.

(1) A particle descends down a curve line in a vertical plane by the action of gravity; to find the nature of the curve that the pressure may be invariable.

Let OA (fig 137) be the required curve; Ox, vertical, the axis of x, Oy, horizontal, the axis of y; P any point in the curve, OM = x, PM = y, OP = s; k the constant pressure; β the initial velocity of the particle, O being its initial position. Then, by formula (A) of sect. (II.), we have

$$k = g \frac{dy}{ds} + \frac{1}{\rho} \frac{ds^2}{dt^2} \dots (1),$$

where ρ denotes the magnitude of the radius of curvature at P.

But
$$\frac{ds^2}{dt^2} = \beta^2 + 2gx;$$

also, s being taken as the independent variable,

$$\frac{1}{\rho} = \frac{\frac{d^2y}{ds^2}}{\frac{dz}{ds}}.$$

Hence, from (1), we have

$$k=g\,rac{dy}{ds}+(2gx+eta^2)rac{\dfrac{d^2y}{ds^2}}{\dfrac{dx}{ds}},$$
 $\dfrac{k}{(2gx+eta^2)^{rac{1}{4}}}rac{d^2y}{ds^2}+\dfrac{g}{(2gx+eta^2)^{rac{1}{4}}}rac{dy}{ds}rac{dx}{ds}.$

Integrating, we have

$$\frac{k}{g}(2gx + \beta^{2})^{\frac{1}{4}} = (2gx + \beta^{2})^{\frac{1}{4}}\frac{dy}{ds} + C,$$

$$\frac{dy}{ds} = \frac{k}{g} - \frac{C}{(2gx + \beta^{2})^{\frac{1}{4}}},$$

where C is an arbitrary constant. Assume a to be the inclination of the curve to the vertical at the origin; then

and therefore,
$$\frac{dy}{ds} = \frac{k}{g} - \frac{\beta}{g} \frac{k - g \sin \alpha}{(2ax + \beta^2)^{\frac{1}{4}}} \dots (2).$$

The relation between x and y may be obtained by a second integration, but the result is of little value in consequence of its complexity. For the investigation of the form of the curve which corresponds to the differential equation (2), the reader is referred to Whewell's *Dynamics*, part 11. p. 95; or, Earnshaw's *Dynamics*, p. 129.

The problem of the Curve of Equal Pressure, in the case of gravity, was first proposed by John Bernoulli¹, and solved by L'Hôpital². Various problems of a similar character were afterwards discussed by Varignon³.

Commerc. Epistolic. Leibnitii et Bernoullii, Epist. vII.

(2) A particle, acted on by gravity, descends from a point O (fig. 137) down a curve OA, which it presses at each point of

¹ Act. Erudit. Suppl. tom. ii. Sect. 6. p. 291.

⁸ Mém. de l'Acad. des Sciences de Paris, 1700, p. 9.

³ Mém. de l'Acad. des Sciences de Paris, 1710, p. 196.

its descent with a force varying as the square of its distance below the horizontal line through O; to find the nature of the curve OA, the initial velocity of the particle being zero.

Let the axes Ox, Oy, be taken vertical and horizontal; let kbe the pressure on the curve when x is equal to unity. Then, by the formula (A) of Sect. (II.),

> $kx^2 = g\frac{dy}{ds} + \frac{2gx}{s};$ $\rho = -\frac{\left(1 + \frac{dx^2}{dy^2}\right)^{\frac{3}{4}}}{d^2x};$

but

hence

hence
$$kx^{3} = \frac{g}{\left(1 + \frac{dx^{2}}{dy^{3}}\right)^{\frac{1}{2}}} - \frac{2gx}{\left(1 + \frac{dx^{2}}{dy^{3}}\right)^{\frac{1}{2}}},$$

$$kx^{\frac{1}{2}} \frac{dx}{dy} = \frac{gx^{-\frac{1}{2}} \frac{dx}{dy}}{\left(1 + \frac{dx^{2}}{dy^{3}}\right)^{\frac{1}{2}}} - \frac{2gx^{\frac{1}{2}} \frac{dx}{dy} \frac{d^{2}x}{dy^{2}}}{\left(1 + \frac{dx^{2}}{dy^{3}}\right)^{\frac{1}{2}}};$$
integrating, we have
$$\frac{2}{3} kx^{\frac{1}{2}} = \frac{2gx^{\frac{1}{2}}}{\left(1 + \frac{dx^{2}}{dy^{2}}\right)^{\frac{1}{2}}},$$

no constant being added because the curve passes through the

origin. Putting $\frac{g}{k} = a$, we get

$$x^{3} \left(1 + \frac{dx^{2}}{dy^{3}}\right)^{\frac{1}{2}} = 5a,$$

$$(25a^{2} - x^{4})^{\frac{1}{2}} dy = x^{2}dx,$$

which is the equation to the Elastic Curve of James Bernoulli'.

Varignon; Mémoires de l'Académie des Sciences de

Paris, 1710, p. 151.

(3) A particle, acted upon by a force parallel to the axis of x, is constrained to move along a given curve OPA (fig. 137); to find the law of the force that the curve may experience an invariable pressure.

1 Act. Erudit., Lips. 1694, p. 272; 1695, p. 538.

Let k denote the constant pressure, β the velocity of the particle at O, which we will take as the origin of co-ordinates, and X the force, at any point P of the curve, parallel to Ox. Then, by formula (A) of section (II.) and formula (D) of section (I.), we have

$$k = X \frac{dy}{ds} + \frac{1}{\rho} \left\{ \beta^2 + 2 \int_0^x X dx \right\}.$$

Taking s as the independent variable, we have

$$\frac{1}{\rho} = \frac{\frac{d^3y}{ds^3}}{\frac{ds}{ds}};$$

and the equation becomes

$$k\frac{dx}{ds} = \beta^2 \frac{d^3y}{ds^2} + X \frac{dx}{ds} \frac{dy}{ds} + 2 \frac{d^3y}{ds^2} \int_0^x X dx.$$

Multiplying by $\frac{dy}{ds}$ ds, and integrating

$$k\int \frac{dx}{ds} \frac{dy}{ds} ds = \frac{1}{2} \beta^2 \frac{dy^2}{ds^2} + \frac{dy^2}{ds^2} \int_0^x X dx,$$

$$\int_0^x X dx = -\frac{1}{2} \beta^2 + \frac{k}{dy^2} \int \frac{dx}{ds} \frac{dy}{ds} ds;$$

and therefore, putting $\frac{dy}{dx} = p$,

$$\int_0^x X dx = -\frac{1}{2}\beta^2 + k\left(\frac{1}{p^2} + 1\right) \left\{ \int \frac{p dx}{(1+p^2)^{\frac{1}{4}}} + C \right\} \dots (1).$$

Differentiating with respect to x, we obtain the required expression for the force

$$X = \frac{k}{p} (1 + p^2)^{\frac{1}{3}} - \frac{2k}{p^3} \frac{dp}{dx} \left\{ \int \frac{p \, dx}{(1 + p^2)^{\frac{1}{3}}} + C \right\} \dots (2).$$

If we put x = 0, we have, from (1),

$$2k\left(\frac{1}{p^{2}}+1\right)\left\{\int_{\left(1+p^{2}\right)^{\frac{1}{4}}}^{\frac{p}{4}}+C\right\}=\beta^{2},$$

a condition which will determine the value of the arbitrary constant C.

Euler; Mechan. tom. 11. p. 101.

(4) A particle moves along a parabola OA, of which the axis is Oy, under the action of a force always parallel to Ox which is at right angles to Oy; to determine the law of the force that the particle may exert the same pressure on the curve during the whole of its motion.

Let Ox, Oy, be the co-ordinate axes, k the constant pressure, and $x^2 = ay$ the equation to the parabola. Then, by the formula for X given in the preceding problem, since $p = \frac{2x}{a}$, we have

$$X = \frac{ka}{2x} \left(1 + \frac{4x^3}{a^3} \right)^{\frac{1}{4}} - \frac{ka^3}{2x^3} \left\{ \int \frac{2x \, dx}{\left(a^3 + 4x^2 \right)^{\frac{1}{4}}} + C \right\}$$
$$= \frac{k}{2x} \left(a^3 + 4x^3 \right)^{\frac{1}{4}} - \frac{ka^3}{2x^3} \left\{ \frac{1}{2} \left(a^3 + 4x^2 \right)^{\frac{1}{4}} + C \right\},$$

where C is a constant quantity,

$$=-\frac{ka^{2}}{2x^{3}}C-\frac{k}{4x^{3}}(a^{2}-2x^{2})(a^{3}+4x^{2})^{\frac{1}{4}}....(1).$$

Again, by the formula (1) in the preceding problem,

$$\int_0^x X dx = -\frac{1}{2}\beta^2 + k\left(\frac{a^2}{4x^2} + 1\right) \left\{\frac{1}{2}(a^2 + 4x^2)^{\frac{1}{2}} + C\right\},\,$$

$$4x^2\int_0^{\pi} X \ dx = -2\beta^2x^2 + k(a^2 + 4x^2)\left\{\frac{1}{2}(a^2 + 4x^2)^{\frac{1}{2}} + C\right\};$$

hence, putting x = 0, we get

$$0=\tfrac{1}{2}a+C,$$

and therefore, by (1),

$$X = \frac{ka^3}{4x^3} - \frac{k}{4x^3} (a^2 - 2x^3)(a^3 + 4x^3)^{\frac{1}{3}}.$$

Euler; Mechan. tom. 11. p. 103.

(5) A particle, under the action of gravity, descends from a point O down a curve OA, (fig. 188), which it presses, at each point of its descent, with a force varying as its perpendicular distance from the horizontal line through O; to find the nature of the curve OA, the initial velocity of the particle being zero.

Take Ox, Oy, the axes of co-ordinates, vertical and horizontal; let k be the pressure on the curve when x is equal to unity; then putting $a = \frac{g}{L}$, the equation to the curve will be

$$x^2 = 6ay - y^2,$$

the equation to a circle of which OE the diameter is equal to 6a. Varignon; Mém. de l'Acad. des Sciences de Paris, 1710. p. 151.

(6) To find the curve when the pressure varies as the square root of the distance.

The equation to the curve is

$$y = 2a \text{ vers}^{-1} \frac{x}{2a} - (4ax - x^2)^{\frac{1}{4}},$$

which belongs to a cycloid OBA, (fig. 139), the radius of the generating circle being 2a.

Varignon; Ib. p. 152.

(7) A particle acted on by gravity descends from rest down a curve; to find the nature of the curve that the pressure at any point due to the centrifugal force may vary as any power of the distance of the particle below the horizontal line passing through its initial position.

Let k denote the pressure due to centrifugal force when x is equal to unity, the axis of x being vertical, as in fig. 137; then the differential equation to the curve will be, putting $\frac{g}{h} = a$,

$$(4n^2a^2 - x^{2n})^{\frac{1}{4}} dy = x^n dx.$$

Varignon; *Ib.* p. 156.

(8) To find the curve when the part of the pressure which is due to gravity varies as the nth power of the depth of the descent.

The notation being the same as in the preceding problem, except that k denotes the pressure due to gravity alone when x = 1, the differential equation to the curve will be

$$(a^2 - x^{2n})^{\frac{1}{2}} dy = x^n dx.$$
 Varignon; *Ib.* p. 160.

(9) A particle descends from rest by the action of gravity down a curve line; to determine the nature of the curve when the part

of the pressure due to centrifugal force bears a constant ratio to that due to gravity.

Let Oy (fig. 137) be horizontal and Ox vertical; then, if $\frac{m}{n}$ denote the constant ratio, the differential equation to the curve will be

$$(a^{\frac{m}{n}}-x^{\frac{m}{n}})^{\frac{1}{4}}dy=x^{\frac{m}{2n}}dx,$$

where a is a constant quantity. If m = n, the curve will be an inverted cycloid with its base horizontal.

Varignon; Ib. p. 161.

(10) A particle acted on by a force parallel to Ox, (fig. 139), moves along the arc OB of a cycloid; to determine this force that the curve may always experience the same pressure.

If k denote the constant pressure, a the radius of the generating circle, X the required force, and OM = x; then

$$X = \frac{k}{3} \left(\frac{2a}{x}\right)^{\frac{1}{4}}.$$
Euler; *Mechan.* tom. II. p. 104.

Sect. 5. Motion of Particles acted on by smooth constraining lines moveable according to assigned geometrical conditions.

Let Ox (fig. 140) be the axis of x, and Oy, at right angles to it, that of y. Let P be the position of the particle in the plane xOy at any time t from the commencement of the motion; OM = x, PM = y. Let X, Y, denote the resolved parts of the accelerating forces on the particle parallel to the axes of co-ordinates, and X', Y', the resolved parts of the force of constraint at any point of its path. Then the circumstances of its motion will depend upon the differential equations

$$\frac{d^2x}{dt^2}=X+X', \quad \frac{d^2y}{dt^2}=Y+Y'....(A).$$

The complete solution of the general problem of the motion of the particle consists in the determination of x and y in terms of t. But the equations (A) involve, in addition to x and y, the two unknown quantities X' and Y'. Hence it appears that the general consideration of the motion affords us only two equations

involving four unknown quantities. From this it is clear that the analytical expression of the conditions to which the motion of the constraining line in any particular problem is subject, must be virtually equivalent to two more equations involving only x, y, X', Y'.

Let r be the distance PO of the particle at the time t from the origin of co-ordinates, $\angle POx = \theta$; then, in case the action of the constraining line always take place in a direction at right angles to OP, and F denote the sum of the resolved parts of the accelerating forces on the particle along OP, we may obtain from the equations (A) the formula

$$\frac{d^2r}{dt^2}-r\frac{d\theta^2}{dt^2}=F.....(B).$$

The formula (B) was given by Ampère, Annales de Gergonne, tom. xx, p. 37, et. sq.

(1) To find the path of a particle upon a smooth horizontal plane, fastened by a thread to a point of which the motion is uniform and rectilinear in that plane.

Let Q (fig. 141) be constrained to move uniformly along the line Ox, and let P be the position of the particle in the plane xOy at any time t; PQ being the thread by which P is attached to Q. Let OM = x, PM = y, OQ = x', PQ = h, $\angle PQO = \theta$. Then, R denoting the tension of the thread, we have, by the equations

(A), since $X' = \frac{R}{m'} \cos \theta$, $Y' = -\frac{R}{m'} \sin \theta$, where m' denotes the mass of the particle,

$$\frac{d^2x}{dt^2} = \frac{R}{m'}\cos\theta, \quad \frac{d^2y}{dt^2} = -\frac{R}{m'}\sin\theta.$$

Hence, eliminating R.

$$\sin \theta \frac{d^2x}{dt^2} + \cos \theta \frac{d^2y}{dt^2} = 0 \dots (1);$$

again, it is clear that

$$x' = x + h \cos \theta = mt + n \dots (2),$$

where m denotes the uniform velocity of Q, and n its initial distance from O; and therefore

$$\frac{d^2x}{dt^2} + h\frac{d^2}{dt^2}\cos\theta = 0;$$

also, by the geometry,

$$y = h \sin \theta \dots (3),$$

and therefore

$$\frac{d^2y}{dt^2}=h\;\frac{d^2}{dt^2}\sin\;\theta.$$

Hence, from (1),

$$\cos \theta \frac{d^2}{dt^2} \sin \theta - \sin \theta \frac{d^2}{dt^2} \cos \theta = 0;$$

integrating, we obtain

$$\cos \theta \frac{d}{dt} \sin \theta - \sin \theta \frac{d}{dt} \cos \theta$$
, or $\frac{d\theta}{dt} = \omega$,

where ω is a constant quantity, which shews that the angular velocity of the thread PQ about Q is invariable.

Integrating again, $\theta = a + \omega t \dots (4)$, a being the initial value of θ .

Hence we have, from (2) and (3),

$$x = mt + n - h \cos(a + \omega t),$$

$$y = h \sin(a + \omega t) \dots (5),$$

which give the values of x and y at any time during the motion. Eliminating t, we get, as the equation to the path of the particle in rectangular co-ordinates,

$$x = \frac{m}{\omega} \sin^{-1} \frac{y}{h} - (h^2 - y^2)^{\frac{1}{4}} + n - \frac{ma}{\omega}$$

The equations (2) and (4) however furnish us with the most convenient conception of the motion of the particle. In fact they shew that P's motion may be perfectly represented by supposing it to move with a uniform velocity ωh in the circumference of a circle of which the radius is h, and of which the centre moves along Ox with a uniform velocity m. The path of P is therefore a trochoid.

From (5) we have, by differentiation,

$$\frac{d^2y}{dt^2}=-\hbar\omega^2\sin{(\alpha+\omega t)}.$$

Hence, from (4), and the original equations of motion,

$$-h\omega^2\sin(a+\omega t)=-\frac{R}{m'}\sin(a+\omega t),$$

and therefore,

$$\frac{R}{m'}=h\omega^2, R=m'h\omega^2,$$

which shews that the tension of the string is invariable.

This is an example of a class of curves called Tractories, which are traced by a material particle attached to one extremity of a string while the other is constrained to move along some assigned curve with a given velocity. The curve in which the end of the string is constrained to move is called the Directrix.

This problem formed the subject of a controversy between Fontaine and Clairaut; the solution given by Fontaine depended upon the assumption that the string would be always a tangent to the path of the particle, an hypothesis which Clairaut declared to be erroneous, and which, in fact, virtually involves a neglect of centrifugal force. Fontaine's assumption would be admissible for the motion of a particle on a perfectly rough plane, where its motion would be destroyed the moment it was generated.

Clairaut; Mémoires de l'Académie des Sciences de Paris, 1736, p. 4. Euler; Nova Acta Acad. Petrop. 1784.

(2) A thin rectilinear tube is constrained to move in a horizontal plane round a vertical axis passing through one extremity with a uniform angular velocity: to find the motion of a particle sliding freely within the tube.

By the formula (B), since no accelerating forces act on the particle,

$$\frac{d^2r}{dt^2}=r\frac{d\theta^2}{dt^2}=\omega^2r,$$

where ω denotes the invariable angular velocity.

Multiplying by $2 \frac{dr}{dt}$, and integrating,

$$\frac{dr^2}{dt^2} = C + \omega^2 r^2.$$

Let a, $\dot{\beta}$, be the initial values of r, $\frac{dr}{dt}$; then

$$\beta^2 = C + \omega^2 a^2;$$

and therefore, eliminating C,

$$\frac{dr^{2}}{dt^{2}} = \omega^{2}r^{2} + \beta^{2} - \omega^{2}a^{2},$$

$$dt = \frac{dr}{(\omega^{2}r^{2} + \beta^{2} - \omega^{2}a^{2})^{\frac{1}{2}}}:$$

integrating, $t + C = \frac{1}{\omega} \log \left\{ \omega r + (\omega^2 r^2 + \beta^2 - \omega^2 a^2)^{\frac{1}{4}} \right\}$.

But, since r = a when t = 0,

$$C = \frac{1}{\omega} \log \{ \omega a + \beta \}.$$

Hence

$$t = \frac{1}{\omega} \log \frac{\omega r + (\omega^2 r^2 + \beta^2 - \omega^2 a^2)^{\frac{1}{2}}}{\beta + \omega a},$$

$$(\boldsymbol{\beta} + \boldsymbol{\omega}\boldsymbol{a}) \, \boldsymbol{\varepsilon}^{\boldsymbol{\omega} t} = \boldsymbol{\omega} \boldsymbol{r} + (\boldsymbol{\omega}^2 \boldsymbol{r}^2 + \boldsymbol{\beta}^2 - \boldsymbol{\omega}^2 \boldsymbol{a}^2)^{\frac{1}{2}} \dots \dots (1).$$

Taking the reciprocals of both sides of the equation, we have

$$\frac{\varepsilon^{-\omega t}}{\beta + \omega a} = \frac{\left(\omega^2 r^2 + \beta^2 - \omega^2 a^2\right)^{\frac{1}{4}} - \omega r}{\beta^2 - \omega^2 a^2},$$

$$(\beta - \omega a) \ \varepsilon^{-\omega t} = \left(\omega^2 r^2 + \beta^2 - \omega^2 a^2\right)^{\frac{1}{4}} - \omega r. \ldots (2).$$

Hence, by (1) and (2), we have

$$2\omega r = (\beta + \omega a) \, \varepsilon^{\omega t} - (\beta - \omega a) \, \varepsilon^{-\omega t},$$

which gives the value of r at any time during the motion.

The equation to the path of the particle is, putting $\theta = \omega t$,

$$2\omega r = (\beta + \omega a) \, \varepsilon^{\theta} - (\beta - \omega a) \, \varepsilon^{\theta}.$$

John Bernoulli; Opera, tom. IV. p. 248. Clairaut; Mém.

Acad. Paris, 1742, p. 10.

(3) A material particle is placed within a thin circular tube which is constrained to revolve with a uniform angular velocity in a horizontal plane about a point in its circumference; to investigate the motion of the particle.

Let O (fig. 142) be the point about which the circle APO is constrained to revolve; C its centre at any time t, and P the position of the particle; R the action of the circle on the particle, which will take place in the direction PC. Let Ox, Oy, be the axes of co-ordinates, x, y, being the co-ordinates of P.

$$\angle POx = \theta$$
, $\angle OPC = \angle COP = \phi$, $OP = r$, $OC = a$.

Then, since no accelerating forces act on the particle, we have, by the formulæ (A),

$$\frac{d^2x}{dt^2} = -R\cos(\theta - \phi), \quad \frac{d^2y}{dt^2} = -R\sin(\theta - \phi)...(1);$$

multiplying these equations by $\sin (\theta - \phi)$, $\cos (\theta - \phi)$, and subtracting, we have

$$\sin (\theta - \phi) \frac{d^2x}{dt^2} - \cos (\theta - \phi) \frac{d^2y}{dt^2} = 0,$$

or, since $x = r \cos \theta$, $y = r \sin \theta$,

$$\sin (\theta - \phi) \frac{d^2}{dt^2} (r \cos \theta) - \cos (\theta - \phi) \frac{d^2}{dt^2} (r \sin \theta) = 0.$$

But, from the geometry, it is evident that

$$r=2a\cos\phi\ldots\ldots(2)$$
;

hence

$$\sin (\theta - \phi) \frac{d^2}{dt^2} (\cos \theta \cos \phi) - \cos (\theta - \phi) \frac{d^2}{dt^2} (\sin \theta \cos \phi) = 0,$$

$$\sin\left(\theta-\phi\right)\frac{d^2}{dt^2}\left\{\cos(\theta+\phi)+\cos(\theta-\phi)\right\}-\cos(\theta-\phi)\frac{d^2}{dt^2}\left\{\sin(\theta+\phi)+\sin(\theta-\phi)\right\}=0$$

But, supposing ω to be the angular velocity of the diameter OCA of the circle about O, and $\angle AOx$ to be initially zero, it is clear that

Hence, putting ωt for $\theta + \phi$,

$$\sin (\theta - \phi) \frac{d^2}{dt^2} \cos (\theta - \phi) - \cos (\theta - \phi) \frac{d^2}{dt^2} \sin (\theta - \phi)$$

$$+ \sin (\theta - \phi) \frac{d^2}{dt^2} \cos \omega t - \cos (\theta - \phi) \frac{d^2}{dt^2} \sin \omega t = 0,$$

$$\frac{d}{dt} \left\{ \sin (\theta - \phi) \frac{d}{dt} \cos (\theta - \phi) - \cos (\theta - \phi) \frac{d}{dt} \sin (\theta - \phi) \right\}$$

$$- \omega^2 \sin (\theta - \phi) \cos \omega t + \omega^2 \cos (\theta - \phi) \sin \omega t = 0,$$

$$- \frac{d^2}{dt^2} (\theta - \phi) - \omega^2 \sin (\theta - \phi - \omega t) = 0.$$

But, by (3), we have $\theta = \omega t - \phi$; hence

$$2\frac{d^2\phi}{dt^2}+\omega^2\sin 2\phi=0.$$

Multiplying by $2 \frac{d\phi}{dt}$, and integrating,

$$2\frac{d\phi^2}{dt^2}-\omega^2\cos 2\phi=C\ldots\ldots(4).$$

For the sake of simplicity we will suppose that initially P coincides with A, and that its velocity is zero; hence when t=0 we have $\phi=0$, and since from (3)

$$\frac{d\theta}{dt} + \frac{d\phi}{dt} = \omega,$$

we have also initially $\frac{d\phi}{dt} = \omega$. Hence, from (4),

$$2\omega^2-\omega^2=C,$$

and therefore

Integrating,

$$\log\frac{1+\sin\phi}{1-\sin\phi}=2\omega t,$$

no constant being added because $\phi = 0$ when t = 0.

From this equation we have

$$\frac{1+\sin\phi}{1-\sin\phi}=\epsilon^{2\omega t}, \quad \sin\phi=\frac{\epsilon^{\omega t}-\epsilon^{-\omega t}}{\epsilon^{\omega t}+\epsilon^{-\omega t}}=\frac{\epsilon^{\psi}-\epsilon^{-\psi}}{\epsilon^{\psi}+\epsilon^{-\psi}}$$

where ψ is equal to $\angle AOx$.

When $t = \alpha$, $\sin \phi = 1$, and therefore $\phi = \frac{\pi}{2}$, a value towards

which ϕ indefinitely tends as its limit. Thus it appears that after an infinite time the particle will arrive at the point O.

Again, since $r = 2a \cos \phi$, we may readily get

$$r = \frac{4a}{\varepsilon^{\omega_t} + \varepsilon^{-\omega_t}} = \frac{4a}{\varepsilon^{\psi} + \varepsilon^{-\psi}},$$

which gives the distance of the particle from O at any time during the motion.

From the above equations we may obtain for the pressure on the circle corresponding to any position of the particle,

$$R = 2\omega^2 a \cos \phi (3 \cos \phi - 2).$$

(4) P and Q (fig. 143) are two particles connected together by an inflexible rod PQ without weight; P is capable of moving along a smooth horizontal groove Ox, and Q may move any where upon a smooth horizontal plane passing through the groove; having given the initial circumstances of the particles, to determine their motions at any time after the commencement of the motion.

Let T be the tension of the rod at any time t; θ the inclination of the rod to the line xO; ON = x', QN = y', where O is an assigned point in Ox, and QN at right angles to ON; OP = x; ω = the initial angular velocity of Q about P, β the initial velocity of P, α the initial value of θ ; m, m', the masses of P, Q; α the length of the rod.

For the motion of P there is

$$m\frac{d^3x}{dt^2}=-T\cos\theta.....(1);$$

and for the motion of Q,

$$m'\frac{d^2x'}{dt^2}=T\cos\theta\ldots(2),$$

$$m'\frac{d^2y'}{dt^2} = -T\sin\theta\dots(3).$$

Adding together the equations (1) and (2),

$$m\frac{d^2x}{dt^2}+m'\frac{d^2x'}{dt^2}=0\ldots\ldots(4).$$

Multiplying (1) by $\sin \theta$, (3) by $\cos \theta$, and subtracting the latter of the resulting equations from the former,

$$m \sin \theta \frac{d^2x}{dt^2} - m' \cos \theta \frac{d^2y'}{dt^2} = 0 \dots (5).$$

Again, it is evident that

$$x' = x - a \cos \theta \dots (6),$$

$$y' = a \sin \theta \dots (7)$$

From (4) and (6) we have

$$(m+m')\frac{d^2x}{dt^2}-m'a\frac{d^2}{dt^2}\cos\theta=0,$$

and from (5) and (7),

$$m \sin \theta \frac{d^3x}{dt^2} - m'a \cos \theta \frac{d^3x}{dt^2} \sin \theta = 0$$
:

eliminating $\frac{d^2x}{dt^2}$ between the two last equations,

$$(m + m')\cos\theta \frac{d^2}{dt^2}\sin\theta - m\sin\theta \frac{d^2}{dt^2}\cos\theta = 0$$
;

multiplying by 2 $\frac{d\theta}{dt}$, and integrating,

$$(m + m')^{\epsilon} \left(\frac{d}{dt} \sin \theta\right)^{2} + m \left(\frac{d}{dt} \cos \theta\right)^{2} = C,$$

$$(m+m'\cos^2\theta)\frac{d\theta^2}{dt^2}=C;$$

but, initially, $\theta = a$, $\frac{d\theta}{dt} = \omega$; hence

$$(m + m' \cos^2 a) \omega^2 = C,$$

and therefore

$$\frac{d\theta^2}{dt^2} = \omega^2 \frac{m + m' \cos^2 \alpha}{m + m' \cos^2 \theta} \dots (8).$$

Again, integrating (4), we get

$$m\frac{dx}{dt}+m'\frac{dx'}{dt}=C,$$

and therefore, by (6),

$$(m+m')\frac{dx}{dt}+m'a\sin\theta\frac{d\theta}{dt}=C;$$

but, at the commencement of the motion, $\frac{dx}{dt} = \beta$, $\theta = a$, $\frac{d\theta}{dt} = \omega$;

hence
$$(m + m') \beta + m'a\omega \sin a = C_3$$

and therefore
$$\frac{dx}{dt} = \beta + \frac{m'a\omega \sin a}{m+m'} - \frac{m'a\sin \theta}{m+m'} \frac{d\theta}{dt}$$
;

whence, by (8),

$$\frac{dx}{dt} = \beta + \frac{m'a\omega \sin a}{m+m'} - \frac{m'a\omega \sin \theta}{m+m'} \left(\frac{m+m'\cos^2 a}{m+m'\cos^2 \theta}\right)^{\frac{1}{4}} \cdot \cdots \cdot (9).$$

The equations (8) and (9) give us the velocity of P along Ox, and the angular velocity of Q about P, for any assignable inclination of the rod to the line Ox. If between these two equations we eliminate dt we shall obtain a differential equation to the path of Q in x and θ .

From (8) we have

we have
$$t = \frac{\omega}{(m+m'\cos^2 a)^{\frac{1}{3}}} \int_a^{\theta} (m+m'\cos^2 \theta)^{\frac{1}{3}} d\theta,$$

an elliptic transcendent for the determination of t for any value of θ .

Clairaut; Mém. de l'Acad. des Sciences de Paris, 1736. p. 10.

(5) A string is completely coiled round the circumference of a circular lamina, and has a particle attached to one extremity which is free, the other extremity being fixed to the lamina: every particle of the lamina repels the free particle with a force varying inversely as the distance; to find the velocity of the particle at any time after its departure from the circumference of the lamina.

Let a denote the radius of the lamina, r the distance of the particle from its centre at any time, and f the initial repulsive force experienced by the particle. Then, as may be ascertained by the performance of the appropriate integrations, the repulsive force on the particle at any time from the centre of the lamina will be $\frac{af}{r}$. Hence the particle may be considered as moving along a curve which is the locus of the free extremity of the string acted on by a central repulsive force $\frac{af}{r}$; and therefore, by the formula (D) of section (I),

$$v^{2} = C + 2 \int \frac{af}{r} dr$$

$$= C + af \log r^{2} = C + af \log (\rho^{2} + a^{2}),$$

if ρ = the length of the string set free.

But, initially,
$$v = 0$$
, $\rho = 0$; hence
$$0 = C + af \log a^2$$

and therefore

$$v^2 = af \log \frac{\rho^2 + a^2}{a^2}.$$

Let θ denote the angle subtending the arc of the circumference of the lamina from which the string has been unfolded; then $\rho = a\theta$, and we have

$$v^2 = af \log (1 + \theta^2).$$

(6) Two particles connected together by a rigid rod without weight are projected along a smooth horizontal plane; to determine their motion.

Let the plane of co-ordinates coincide with the plane of the motion. Let m, n, be the resolved parts of the initial velocity of the centre of gravity of the two particles parallel to the axes of x, y, and let a, b, be its initial co-ordinates. Let ω be the initial angular velocity of the rod, θ its inclination to the axis of x at the end of the time t, and ε at the beginning of the motion. Then the position of the centre of gravity is given at any time t by the equations

$$x = mt + a$$
, $y = nt + b$;

and the inclination of the rod to the axis of x, by the equation

$$\theta = \omega t + \varepsilon$$
.

Clairaut; Mémoires de l'Académie des Sciences de Paris, 1736, p. 7. Euler; Act. Acad. Petrop. 1780, P. 1.; Opuscula, De motu corporum flexibilium, tom. 111. p. 91.

(7) A spherical particle moves within a smooth tube which revolves about one extremity with a uniform angular velocity in a vertical plane, the capacity of the tube being just sufficiently great for the reception of the particle; to determine the motion of the particle.

Let Ox, (fig. 140), which is horizontal, be the initial position of the tube, and P the position of the particle in the tube after a time t. Let ω denote the angular velocity of the tube about O, θ the inclination of OP to Ox, and OP = r. Then, supposing the initial velocity of the particle to be zero, and that r = a initially, the value of r at any time t is given by the equation

$$r = \frac{g}{2\omega^2} \sin \omega t + \frac{2a\omega^2 - g}{4\omega^2} \epsilon^{\omega t} + \frac{2a\omega^2 + g}{4\omega^2} \epsilon^{-\omega t},$$

and the polar equation to the path of the particle will result from the substitution of θ for ωt in this equation. When t becomes very great, the polar equation becomes

$$r = \frac{2a\omega^2 - g}{4\omega^2} \, \epsilon^{\theta},$$

which is the equation to an equiangular spiral.

The solution of this problem was attempted by M. Le Barbier, in the *Annales de Gergonne*, tom. xix. p. 285, who omitted to take into consideration the centrifugal force, an oversight which entirely vitiated his results. The correct solution was given in tom. xx. by Ampère.

(8) A material particle P (fig. 144) is fixed to one end of a rigid rod PQ without weight lying upon a smooth horizontal plane. The end Q is constrained to move with a uniform velocity in the circumference of a circle ABQ; to find the velocity of the increase of the angle PQR, O being the centre of the circle, and OQR a straight line.

If PQ = h, OQ = a, $\angle PQR = \psi$ at any time t, a = the initial value of ψ , $\omega =$ the angular velocity of OQ, $\beta =$ the initial value of $\frac{d\psi}{dt}$, then

$$h\left(\frac{d\psi^2}{dt^2}-\beta^2\right)=2a\omega^2(\cos\psi-\cos\alpha).$$

Clairaut; Mém. de l'Acad. des Sciences de Paris, 1736, p. 14.

(9) QBA (fig. 145) is a circle on a horizontal plane, and QP a string touching it at the point Q; P is a particle attached to the end of the string. Supposing the particle P to be projected at right angles to QP with a given velocity so as to cause QP to be gradually wrapped about the circumference QBA; to find the velocity of the particle at any time during the motion, and the time which will elapse before the particle reaches the circumference.

Let β be the velocity of projection, v the velocity at any time during the motion, b the length of the string PQ, a the radius of the circle, T the time required. Then

$$v = \beta$$
, $T = \frac{b^2}{2a\beta}$.

(10) A circular horizontal lamina of matter ABC, (fig. 146), every particle of which attracts with a force varying inversely as the distance, is made to revolve with a uniform angular velocity round an axis through its centre O at right angles to its plane, the motion taking place in the direction of the arrows; to find the equation to the groove Aa which must be carved in the circular lamina that it may be described freely by a particle subject to the attraction of the lamina; the initial position of the particle being a point A in the circumference of the circle, and its initial velocity being zero.

Let P be any point in the groove, OP = r, OA = a, $\angle POA = \theta$, $\omega =$ the angular velocity of the lamina about O, and f = the attraction of the lamina on a particle in its circumference. Then the equation to the groove Aa will be

$$r = a \cos \left\{ \left(\frac{f}{\omega^2 a} \right)^{\frac{1}{4}} \theta \right\}.$$

(11) Two small equal bodies A, B, connected together by a rigid line, are placed in a narrow rectilinear tube, in which they can move without friction; the tube is then made to revolve with a uniform angular velocity round a vertical axis which passes through a point C of the tube, this point C lying initially between A and B at a distance a from A and b from B; to find the time of A's arriving at C, and the tension of the rigid line at any time, a being considered less than b.

If ω denote the angular velocity of the tube, m the mass of each particle, t the required time, and T the tension; then

$$t = \frac{1}{\omega} \log \frac{b^{\frac{1}{4}} + a^{\frac{1}{4}}}{b^{\frac{1}{4}} - a^{\frac{1}{4}}}, \quad T = \frac{1}{2} m\omega^{2} (a + b).$$

(12) A particle is drawn up an indefinitely thin cycloidal tube, the axis of the cycloid being vertical, by means of an equal particle, to which the former particle is attached by a thread passing over a pulley at the highest point of the arc; to find the time of ascending to the highest point.

If T represent the required time, and t the time of a semi-oscillation in the cycloid, $T = 2^{\frac{1}{4}}t.$

but, by the nature of the cycloid, if $\frac{1}{4}a$ be the radius of the generating circle, $dx = \frac{s}{a}ds$; hence

$$d \cdot v^2 - \frac{2ds}{k} v^2 = -\frac{2g}{a} sds + 2hds.$$

Multiplying both sides of the equation by $\epsilon^{-\frac{2s}{k}}$, we have

$$d\left(v^{2} \varepsilon^{\frac{2s}{k}}\right) = 2h\varepsilon^{\frac{2s}{k}} ds - \frac{2g}{a} \varepsilon^{\frac{2s}{k}} s ds.$$

Integrating, we have

$$v^2 \, \epsilon^{-\frac{2s}{h}} = C - hk \, \epsilon^{-\frac{2s}{h}} - \frac{2g}{a} \int \epsilon^{-\frac{2s}{h}} s ds;$$

but

$$\int e^{-\frac{2s}{k}} s \, ds = -\frac{1}{2} \, k \, e^{-\frac{2s}{k}} \, s + \frac{1}{2} \, k \int e^{-\frac{2s}{k}} \, ds$$

$$= -\frac{1}{2}k \, \epsilon^{-\frac{2s}{k}} \, s - \frac{1}{4}k^2 \, \epsilon^{-\frac{2s}{k}};$$

hence

$$v^{2} \varepsilon^{-\frac{2s}{k}} = C - hk \varepsilon^{-\frac{2s}{k}} + \frac{1}{a} (gks + \frac{1}{2}gk^{2}) \varepsilon^{-\frac{2s}{k}},$$

$$= C + \frac{1}{a} (gks + \frac{1}{2}gk^{2} - ahk) \varepsilon^{-\frac{2s}{k}}.$$

But, initially, v = 0, s = c; hence

$$0 = C + \frac{1}{a} \left(g k c + \frac{1}{2} g k^2 - ahk \right) \epsilon^{\frac{2c}{k}}.$$

Let v_1 be the value of v when s = 0; then

$$v_1^2 = C + \frac{1}{a} \left({}_{2}^{1} g k^2 - a h k \right);$$

and therefore $v_1^2 = \frac{k}{a}(\frac{1}{2}gk - ah) - \frac{k}{a}(gc + \frac{1}{2}gk - ah) e^{-\frac{2c}{k}}$.

Again,
$$v^2 e^{-\frac{2s}{k}} = \frac{k}{a} (gs + \frac{1}{2}gk - ah) e^{-\frac{2s}{k}} - \frac{k}{a} (gc + \frac{1}{2}gk - ah) e^{-\frac{2c}{k}}$$
,

$$v^{2} = \frac{k}{a} (gs + \frac{1}{2}gk - ah) - \frac{k}{a} (gc + \frac{1}{2}gk - ah) \varepsilon^{\frac{2}{k}(s-c)}.$$

When v is a maximum,

$$0 = \frac{gk}{a} - \frac{2}{a} \left(gc + \frac{1}{2} gk - ah \right) \epsilon^{\frac{2}{k}(s-c)},$$

$$\begin{split} \frac{e^{\frac{2}{k}(s-c)}}{e^{\frac{1}{k}}gk-ah}, \\ s &= c + \frac{k}{2}\log\frac{\frac{1}{2}gk}{gc + \frac{1}{2}gk-ah} = c - \frac{k}{2}\log\frac{gc + \frac{1}{2}gk-ah}{\frac{1}{2}gk}, \end{split}$$

which gives the position of the particle for a maximum velocity.

Euler; Mechan. tom. 11. p. 292.

(3) From a given point O, (fig. 149), an infinite number of straight lines OP are drawn in a vertical plane; to determine the nature of the curve APD, such that a particle descending down any line OP may always acquire the same velocity on arriving at P, the medium in which the motion takes place being uniform, and its resistance varying as any power of the velocity.

Let β be the velocity at P, v at any point p in OP; OP = r, Op = z; $\angle POx = \theta$, Ox being vertical; draw PM horizontally, and let OM = x; then, k being the resistance for a unit of velocity, and m the index of its power,

$$v \frac{dv}{dz} = g \cos \theta - kv^{m},$$

$$dz = \frac{vdv}{g \cos \theta - kv^{m}}.$$

Integrating, we have

$$r = \int_{0}^{\beta} \frac{v dv}{g \cos \theta - k v^{m}} = \int_{0}^{\beta} \frac{\beta d\beta}{g \cos \theta - k \beta^{m}}.$$

But $x = r \cos \theta$; hence

But β being a constant quantity while x and θ vary, we have, from (1),

$$dx = kd (\sec \theta) \cdot \int_{0}^{\beta} \frac{\beta^{m-1}d\beta}{(g - k\beta^{m} \sec \theta)^{2}};$$
and therefore, by (2),
$$\beta^{2} \qquad m \sec \theta dx$$

$$2x = \frac{\beta^2}{g - k \beta^m \sec \theta} - \frac{m \sec \theta \, dx}{d \sec \theta};$$

or, putting
$$\frac{r}{x}$$
 for $\sec \theta$,
$$m \frac{r}{x} dx + 2x d\left(\frac{r}{x}\right) = \frac{\beta^3 d\left(\frac{r}{x}\right)}{g - k \frac{r}{x} \beta^m},$$

and therefore $(m-2) r dx + 2x dr = \beta^2 \frac{xdr - rdx}{gx - k\beta^m r}$;

which is the differential equation to the curve in x and r.

Euler; Mechan. tom. 11. p. 246.

(4) To find the tautochrone in a medium the resistance of which varies as the square of the velocity, the particle being acted on by gravity.

Let O (fig. 150) be the point to which the particle is always to descend in the same time, AO being the tautochrone. Take Oy horizontal as the axis of y, Ox vertical as the axis of x. Let OM = x, OP = s; v = the velocity of the particle at P, and k = the resistance of the medium for a unit of a velocity.

The equation for the motion along the curve will be

$$vdv = -gdx + kv^2ds:$$

multiplying by $2\epsilon^{-2ks}$, we have

$$d\left(v^{2}\varepsilon^{-2ks}\right)=-2g\varepsilon^{-2ks}\ dx.$$

Integrating, we obtain

$$v^2 \varepsilon^{-2ks} = C - 2g \int \varepsilon^{-2ks} dx.$$

Suppose the velocity of the particle on its arrival at O to be that due to an altitude h in vacuo; then

$$2gh = C - 2g \int_0^0 \varepsilon^{-2ks} dx;$$

hence

$$v^2 \varepsilon^{-2ks} = 2g \left\{ h - \int_0^x \varepsilon^{-2ks} \ dx \right\} \dots (1),$$

and therefore, v being equal to $-\frac{ds}{dt}$ at any time t,

$$dt = -\frac{1}{(2g)^{\frac{1}{4}}} \frac{\varepsilon^{-ks} ds}{(h-u)^{\frac{1}{4}}},$$

$$u = \int_{0}^{x} \varepsilon^{-2ks} dx \dots (2).$$

where

Now, s being some function of x and therefore of u, we may assume $\varepsilon^{-t} ds = \phi(u) du$, and thus

$$dt = -\frac{1}{(2g)^{\frac{1}{4}}} \frac{\phi(u) \ du}{(h-u)^{\frac{1}{4}}}, \quad t = -\frac{1}{(2g)^{\frac{1}{4}}} \int \frac{\phi(u) \ du}{(h-u)^{\frac{1}{4}}}.$$

But t=0 when v=0, and therefore, by (1) and (2), when u=h; and it will denote the whole time of descent to O from the beginning of the motion when x=0, and therefore, by (2), when u=0; hence the whole time of the descent is equal to

$$\frac{1}{(2g)^{\frac{1}{4}}}\int_0^h \frac{\phi(u)\,du}{(h-u)^{\frac{1}{4}}}\dots\dots(3),$$

a result which, for the condition of tautochronism, must evidently be independent of h. From this it is plain that

$$\int \frac{\phi(u)\ du}{(h-u)^{\frac{1}{4}}}$$

must be of no dimensions in u and h together, and that consequently its differential $\frac{\phi(u)}{(h-u)^{\frac{1}{4}}}$ must be of no dimensions in u, h, du;

hence, since $\phi(u)$ evidently does not involve h, we must have

$$\phi(u)=\frac{a}{u^{\frac{1}{4}}},$$

where a is some constant quantity. Hence, putting for ϕ (u) its value, du

 $\epsilon^{-ks} ds = a \frac{du}{u^{\frac{1}{2}}}.$

Integrating, we have

$$C-\frac{1}{k}\ \epsilon^{-ks}=2au^{\frac{1}{2}}.$$

But, by (2), when s and therefore x is equal to zero, u = 0;

hence
$$C - \frac{1}{k} = 0$$
;
hence $\frac{1}{k} (1 - \epsilon^{-ks}) = 2au^{\frac{1}{4}}, \quad \frac{1}{k^2} (1 - \epsilon^{-ks})^2 = 4a^2u,$
 $\frac{1}{k} (1 - \epsilon^{-ks}) \epsilon^{-ks} = 2a^2 \frac{du}{ds} = 2a^2 \epsilon^{-2ks} \frac{dx}{ds}, \text{ by (2)};$

and therefore the equation to the tautochrone will be

$$2a^2k \frac{dx}{ds} = \epsilon^{ks} - 1.$$

Since $\phi(u) = \frac{a}{u^3}$, we have, from (3), if τ denote the whole time of descent,

 $\tau = \frac{\pi a}{(2g)^{\frac{1}{4}}}, \quad a^2 = \frac{2g\tau^2}{\pi^2};$

and therefore the equation to the tautochrone for the time τ will be

 $4gk\tau^2\frac{dx}{ds}=\pi^2\left(\varepsilon^{ks}-1\right)\ldots\ldots(4).$

Euler; Comment. Petrop. 1729; Mechan. tom. II. p. 392.
John Bernoulli; Mém. de l'Acad. des Sciences de Paris, 1730; Opera, tom. III. p. 173. See also the Cambridge Mathematical Journal, vol. II. p. 153, where Mr. Leslie Ellis has reduced the solution of this problem to that of the Tautochrone in vacuo.

(5) A particle oscillates in an inverted cycloid, of which the axis is vertical, in a uniform medium where the resistance varies as the velocity; having given the first arc of descent, to find the whole space described by the particle before the motion ceases.

Let c denote the first arc of descent, k the resistance when the velocity is unity, a the radius of the generating circle; then the whole space will be equal to

$$c \frac{\epsilon^{\phi} + 1}{\epsilon^{\phi} - 1}$$
, where $\phi = \frac{k\pi}{\left(\frac{g}{a} - k^2\right)^{\frac{1}{2}}}$.

(6) An inelastic particle descends down the sides of a plane equilateral and equiangular polygon in a vertical plane, the medium in which the motion takes place being uniform and its resistance varying as the square of the velocity; to determine the velocity of the particle when it has arrived at the end of any side of the polygon, the side down which the particle first descends being vertical.

Let π -a be the magnitude of each of the angles of the polygon, l the length of each of its sides, kv^2 the resistance for a velocity v; v the velocity at the end of the x^{th} side. Then

$$v_{e} = (\cos a)^{a} \epsilon^{-kla} \left\{ A^{2} + \frac{g(\epsilon^{2kl} - 1)}{k(\cos a)^{2}} \cdot \frac{M\cos xa + N\sin xa}{M^{2} + N^{2}} \cdot \frac{\epsilon^{2klx}}{(\cos a)^{2x}} \right\},$$
where
$$M = \frac{\epsilon^{2kl} - \cos a}{\cos a}, \quad N = \frac{\sin a}{(\cos a)^{2}} \epsilon^{2kl};$$

and A is an arbitrary constant which will easily be determined if we know the value of v_a for any value of x.

Bordoni; Memorie della Societa Italiana, 1816, p. 173.

(7) From a given point O (fig. 149) an infinite number of straight lines OP are drawn in a vertical plane; to determine the nature of the curve APD, so that a particle descending down any line OP may always acquire the same velocity on arriving at P; the motion taking place in a medium of uniform density where the resistance varies as the square of the velocity.

Let OA be vertical and be represented by a; OP = r, OM = x, k =the resistance for a unit of velocity; then the equation to the curve will be $x = r \frac{\epsilon^{2kr}}{\epsilon^{2ka}} \frac{\epsilon^{2ka} - 1}{\epsilon^{2kr} - 1}.$

Euler; *Mechan.* tom. 11. p. 251.

(8) From a given point O (fig. 151) an infinite number of straight lines OP are drawn in a vertical plane; to determine the nature of the curve OPA so that a particle may descend through all its chords OP in the same time; the motion taking place in a medium of uniform density where the resistance varies as the square of the velocity.

Let OA be vertical and be equal to a; OP = r, $\angle AOP = \theta$; then the polar equation to the curve OPA will be

$$(\cos \theta)^{\frac{1}{4}} = \frac{\log \left\{ \varepsilon^{hr} + \left(\varepsilon^{2hr} - 1 \right)^{\frac{1}{4}} \right\}}{\log \left\{ \varepsilon^{hn} + \left(\varepsilon^{2hn} - 1 \right)^{\frac{1}{4}} \right\}}.$$
Euler; *Mechan*. tom. II. p. 256.

(9) A particle, acted on by gravity, is ascending the curve MNA, (fig. 152), in a medium where the resistance varies as the

square of the velocity; to find the nature of the curve that the velocity of the particle at any point N may be the same as that which it would acquire by falling in the same medium down a vertical line IN, the length of which is equal to the arc AN of the curve measured from a fixed point A.

Draw through A a vertical line AB, and let fall NQ at right angles to AB; let AQ = x, AN = s, and k = the resistance of the medium for a unit of velocity. Then the differential equation to the curve will be

$$k(x+s)=1-\epsilon^{-2ks}.$$

This problem was proposed to Clairaut on his journey to Lapland, by Klingstierna, Professor of Mathematics at Upsal, at which place he called on his way; Klingstierna's construction, together with his own solution, was published by Clairaut in the Mémoires de l'Académie des Sciences de Paris, 1740, p. 254.

CHAPTER V.

MOMENT OF INERTIA.

THE Moment of Inertia of a body with regard to any axis, is the sum of all the products resulting from the multiplication of each element of the mass by the square of its distance from the axis. If M denote the whole mass of the body, the Moment of Inertia may be represented by the expression Mk^2 , where k is a line called the Radius of Gyration. The term Moment of Inertia was first made use of by Euler. "Ratio hujus denominationis ex similitudine motus progressivi est desumpta: quemadmodum enim in motu progressivo, si a vi secundum suam directionem sollicitante acceleretur, est incrementum celeritatis ut vis sollicitans divisa per massam seu inertiam; ita in motu gyratorio, quoniam loco ipsius vis sollicitantis ejus momentum considerari oportet, eam expressionem fradM, quæ loco inertiæ in calculum ingreditur, momentum inertiæ appellemus, ut incrementum celeritatis angularis simili modo proportionale fiat momento vis sollicitantis diviso per momentum inertiæ."1

Sect. 1. A Material Curve revolving about an Axis within its own Plane.

 To find the moment of inertia and radius of gyration of a circular arc about a radius through its vertex.

Let HAK (fig. 153) be the circular arc, A its vertex, C the centre of the circle. Take any point P in the arc; draw PM at right angles to the radius CA; join HK intersecting CA in E. Let CM = x, PM = y, CA = a, arc AP = s, HE = c = KE. Then, the density of the arc and the indefinitely small area of the section of it made by a plane through C, at right angles to its own plane, being represented respectively by ρ and a, we shall have $Mk^2 = a\rho \int y^2 ds;$

1 Euler; Theoria Motus Corporum Solidorum, p. 167.

but, from the equation to the circle,

$$ds^2 = dx^2 + dy^2 = \frac{y^2 dy^2}{a^2 - y^2} + dy^2 = \frac{a^2 dy^2}{a^2 - y^2}$$
:

hence, introducing the appropriate limits,

$$Mk^2 = a\rho a \int_{-c}^{+c} \frac{y^2 dy}{(a^2 - y^2)^{\frac{1}{2}}}.$$

Now, integrating by parts,

$$\int \frac{y^2 dy}{(a^3 - y^2)^{\frac{1}{4}}} = -(a^2 - y^2)^{\frac{1}{4}} y + \int (a^2 - y^2)^{\frac{1}{4}} dy$$

$$= -(a^2 - y^2)^{\frac{1}{4}} y + \int \frac{a^2 dy}{(a^2 - y^2)^{\frac{1}{4}}} - \int \frac{y^2 dy}{(a^2 - y^2)^{\frac{1}{4}}}$$

$$= C - \frac{1}{2} y (a^2 - y^2)^{\frac{1}{4}} + \frac{1}{2} a^2 \sin^{-1} \frac{y}{a} :$$

$$\frac{Mk^2}{aa} = a^3 \sin^{-1} \frac{c}{a} - ac (a^2 - c^2)^{\frac{1}{4}};$$

hence

$$\frac{d}{d\rho} = a^{\alpha} \sin^{-\alpha} - ac (a^{\alpha} - c^{\alpha})^{-\alpha}$$

and therefore, since $\frac{M}{a\rho} = 2a \sin^{-1} \frac{c}{a}$, we have

$$k^{2} = \frac{1}{2}a^{2} - \frac{c(a^{2} - c^{2})^{\frac{1}{4}}}{2\sin^{-1}\frac{c}{a}}.$$

If the arc be a semicircle, c = a, and if a circle, c = 0; in both cases $k^2 = \frac{1}{2}a^2$.

(2) To find the radius of gyration of a material straight line OB, (fig. 154), about an axis OA, to which it is inclined at a given angle, the density at any point of OB varying as some power of its distance from O.

Take any point P in OB; draw PM at right angles to OA; let PM=y, OP=s, OB=l, $\angle AOB=a$, ρ = the density at P: then $\rho = \mu s^n$, where μ is a constant quantity. Hence

$$Mk^{2} = a \int \rho y^{2} ds = a\mu \sin^{2} a \int_{0}^{t} s^{n+2} ds = \frac{a\mu \sin^{2} a}{n+3} \frac{l^{n+3}}{n}.$$
Also,
$$M = a \int \rho ds = a\mu \int_{0}^{t} s^{n} ds = \frac{a\mu l^{n+1}}{n+1}.$$

Hence we get

$$k^2=\frac{n+1}{n+3}\,l^2\,\sin^2\,a.$$

If the density be invariable, n = 0, and $k^2 = \frac{1}{3} l^2 \sin^2 a$.

SECT. 2. Material Line revolving about an Axis at Right Angles to its own Plane.

(1) To find the radius of gyration of a straight line AB, (fig. 155) about an axis through D at right angles to the plane ADB.

Let C be the middle point of the line; join CD. Let AC = a = BC, CD = b; k = the radius of gyration about the axis through D, and k' = that about an axis parallel to this through C. Then

$$k^2=k'^2+b^2.$$

But, $2a\rho a$ being the mass of AB,

$$2a\rho a k^{2} = \rho a \int_{-a}^{+a} s^{3} ds = \frac{2\rho a}{3} a^{3},$$

$$k^{2} = \frac{1}{3} a^{2}.$$

$$k^{3} = \frac{1}{4} a^{3} + b^{3}.$$

Hence

(2) To find the radius of gyration of a circular arc about an axis perpendicular to its plane through its centre of gravity.

Let k be the radius of gyration about the required axis, k' about an axis parallel to this through the centre of the circle, and h the distance between the centre of gravity of the arc and the centre of the circle. Then

$$k^2=k^2+h^2.$$

But, r denoting the radius of the circle, c the chord, and a the length of the arc,

$$k^2 = r^2$$
, $h^2 = \frac{c^3 r^2}{a^2}$.

Hence, for the required radius of gyration,

$$k^2=\frac{r^3}{a^2}(a^2-c^3).$$

(3) To find the radius of gyration of a circular arc about an axis perpendicular to its plane through its vertex.

If r = the radius of the circle, a = the length and c = the chord of the arc,

 $k^2 = 2r^2\left(1 - \frac{c}{a}\right).$

(4) If the density of a straight rod AB vary as the nth power of the distance from one end A, and k, k', be the radii of gyration of the rod round axes at right angles to its length through A and B respectively; to compare the values of k and k', and to ascertain the value of n so that k may be equal to 6k'.

$$\frac{k^3}{k'^2} = \frac{(n+1)(n+2)}{2}$$
; $n = 7$, or $= -10$.

SECT. 3. A Plane Area revolving about an Axis within or parallel to the Plane.

(1) To find the radius of gyration of an elliptic area, of uniform thickness and density, about its principal axes.

Let ρ represent the uniform density of the area, and τ its indefinitely small thickness; then, x, y, denoting the co-ordinates of any point of the curve referred to a, b, as axes of co-ordinates, we have for the moment of inertia, about the axis a, of a quadrant of the ellipse,

$$\begin{aligned} Mk^2 &= \rho \tau \int_0^b y^2 \cdot x dy, \\ &= \frac{\rho \tau a}{b} \int_0^b y^2 \cdot (b^2 - y^2)^{\frac{1}{4}} dy : \\ &\int_0^b y^2 \cdot (b^2 - y^2)^{\frac{1}{4}} dy = \frac{1}{3} \int_0^b (b^2 - y^2)^{\frac{3}{4}} dy \\ &= \frac{1}{4} b^2 \int_0^b (b^2 - y^2)^{\frac{1}{4}} dy \\ &= \frac{1}{16} \pi b^4. \end{aligned}$$

but

Hence the moment of inertia of the whole ellipse will be equal to

$$4Mk^{2} = \frac{1}{4}\pi\rho\tau ab^{3};$$

 $4M = \pi\rho\tau ab;$
 $k^{2} = \frac{1}{4}b^{2}.$

but hence

If k' denote the radius of gyration about the axis b, we shall have, by similar reasoning,

$$k'^2 = \frac{1}{4} a^2.$$

(2) To find the radius of gyration of a circular area revolving about a straight line parallel to its plane, at a distance c from its centre.

If a be the radius of the circle, and k the required radius of gyration, $k^2 = \frac{1}{2}a^2 + c^2$.

(3) To find the radius of gyration of an isosceles triangle about a perpendicular let fall from its vertex upon its base.

If 2b = the length of the base,

$$k^3 = \frac{1}{6}b^2$$
.

Sect. 4. Plane Area about a Perpendicular Axis.

(1) To find the radius of gyration of a triangular lamina ABC, (fig. 156), about an axis through A at right angles to its plane.

Take two points P, p, indefinitely near to each other in the side AB, and draw PM, pm, parallel to BC. Take P', p', in PM, pm, and construct the indefinitely small parallelogram P'p', two of the sides of which are parallel to AC. Let AM = x, PM = y, PM = y', Am = x + dx, p'm = y' + dy', $\angle ACB = C$; a, b, c, the three sides of the triangle.

Then, Mk^2 denoting the moment of inertia about A, we have, τ denoting the indefinitely small thickness, and ρ the density of the lamina,

$$Mk^{2} = \int_{0}^{b} \int_{0}^{y} (x^{2} + y'^{2} - 2xy' \cos C) \rho \tau \sin C \, dx \, dy'$$

$$= \rho \tau \sin C \int_{0}^{b} (x^{2}y + \frac{1}{3}y^{3} - xy^{2} \cos C) \, dx$$

$$= \rho \tau \sin C \int_{0}^{b} \left(\frac{a}{b}x^{3} + \frac{1}{3}\frac{a^{3}}{b^{3}}x^{2} - \frac{a}{b^{2}}x^{3} \cos C\right) dx$$

$$= \frac{1}{12} \cdot \frac{1}{2}\rho \tau \, ab \sin C \left(6b^{2} + 2a^{2} - 6ab \cos C\right)$$

$$= \frac{1}{12} M \left\{6b^{2} + 2a^{2} - 3\left(a^{2} + b^{2} - c^{2}\right)\right\},$$

and therefore

$$k^2 = \frac{1}{12} (3b^2 + 3c^2 - a^2).$$

(2) To find the radius of gyration of a triangular lamina ABC about a perpendicular through its centre of gravity G.

Let AG, BG, CG, be represented by a, β , γ ; and BC, AC, AB, by a, b, c. Then, M denoting the mass of the whole

triangle ABC, $\frac{1}{3}$ M will be the mass of each of the triangles BGC, AGC, AGB. Hence, by the preceding problem, the moment of inertia of these three triangles respectively about the axis through G will be

$$\frac{1}{36} M(3\beta^2 + 3\gamma^2 - a^2),$$

$$\frac{1}{36} M(3a^2 + 3\gamma^2 - b^2),$$

$$\frac{1}{36} M(3a^2 + 3\beta^2 - c^2);$$

and therefore the moment of inertia of the whole triangle about G will be equal to

$$\frac{1}{36} M \left\{ 6 \left(a^2 + \beta^2 + \gamma^2 \right) - \left(a^2 + b^2 + c^2 \right) \right\};$$

or, by a property of the centre of gravity of a triangle, to

$$\frac{1}{36} M \left\{ 2 (a^2 + b^2 + c^2) - (a^2 + b^2 + c^2) \right\}$$

$$= \frac{1}{36} M (a^2 + b^2 + c^2).$$

Hence the radius of gyration will be equal to

$$\frac{1}{36}(a^2+b^2+c^3).$$

Euler; Theoria Motus Corporum Solidorum, cap. vi. Prob. 32. Cor. 1.

(3) To find the radius of gyration of an elliptic area about a perpendicular axis through its centre.

If M be the mass of the area, the moment of inertia about the two axes of the ellipse will be

$$\frac{1}{4} Mb^2$$
, $\frac{1}{4} Ma^2$.

But the moment of inertia of a plane area, about any perpendicular axis, is equal to the sum of the moments of inertia about any two lines, at right angles to each other in the plane area, passing through the point in which the axis meets the area. Hence, in the present problem, the moment of inertia about the proposed axis is equal to

$$\frac{1}{4}M(a^2+b^2),$$

and the radius of gyration = $\frac{1}{4}(a^2 + b^2)$.

(4) To find the radius of gyration of an annulus about a perpendicular axis through the centre.

Let r be the distance of any point of the annular area from the centre of the circle, θ the angular co-ordinate, ρ the density, and

 τ the indefinitely small thickness of the area; then, a, b, being the radii of the two concentric circles,

$$Mk^{2} = \int_{0}^{2\pi} \int_{a}^{b} r^{2} \cdot \rho \tau r d\theta dr$$

$$= \frac{1}{4} \rho \tau \int_{0}^{2\pi} (b^{4} - a^{4}) d\theta = \frac{1}{2} \pi \rho \tau (b^{4} - a^{4}).$$
But
$$M = \int_{0}^{2\pi} \int_{a}^{b} \rho \tau r d\theta dr = \frac{1}{2} \rho \tau \int_{0}^{2\pi} (b^{3} - a^{2}) d\theta$$

$$= \pi \rho \tau (b^{2} - a^{2});$$
hence
$$k^{2} = \frac{1}{2} (a^{2} + b^{2}).$$

(5) To find the radius of gyration of a parallelogram about an axis perpendicular to it through its centre of gravity.

If 2a, 2b, be the lengths of two adjoining sides of the parallelogram, then, whatever be the angle of their inclination,

$$k^2 = \frac{1}{3}(a^2 + b^2).$$

Euler; Theoria Motus Corp. Solid. cap. vi. Prob. 35.

(6) To find the radius of gyration of a regular polygon about an axis perpendicular to it through the centre. If n be the number of sides, and c the length of each,

$$k^{3} = \frac{1}{12} c^{3} \frac{2 + \cos \frac{2\pi}{n}}{1 - \cos \frac{2\pi}{n}}.$$

(7) To find the radius of gyration of a portion of a parabola bounded by a double ordinate to the axis about a perpendicular line through its vertex.

If x, y, represent the extreme co-ordinates of the portion,

$$k^3 = \frac{3}{7} x^2 + \frac{1}{5} y^2.$$

SECT. 5. Symmetrical Solid about its Axis.

(1) To find the radius of gyration of a homogeneous sphere about a diameter.

Let x, x + dx, be the distances of the circular faces of a thin circular slice of the sphere, at right angles to the diameter, from

the centre, and let y be the radius of the section. Then, ρ denoting the density of the sphere, the moment of inertia of this slice about the diameter will be equal to

$$\frac{1}{2}\pi\rho y^4dx$$
;

and therefore the moment of inertia of the whole sphere, a being its radius, will be equal to

$$\frac{1}{2} \pi \rho \int_{-a}^{+a} y^4 dx = \frac{1}{2} \pi \rho \int_{-a}^{+a} (a^3 - x^3)^2 dx = \frac{8}{15} \pi \rho a^5.$$

But the mass of the sphere is equal to $\frac{4}{3}\pi\rho a^3$; hence

$$k^2 = \frac{2}{5} a^3$$
.

Euler; Theoria Motus Corporum Solidorum, p. 198.

(2) To find the radius of gyration of a right cone about its axis.

If a denote the radius of the base of the cone,

$$k^2 = \frac{3}{10} a^2$$
.

Euler; Ib. p. 197.

(3) To find the radius of gyration of a hollow sphere about a diameter.

If a, b, be the external and internal radii,

$$k^2 = \frac{2}{5} \frac{a^5 - b^5}{a^3 - b^3}.$$

Euler; Ib. p. 203.

(4) To find the radius of gyration of a solid cylinder about its axis.

If a denote the radius of the cylinder,

$$k^2 = \frac{1}{3} a^2.$$

Euler; Ib. p. 200.

Sect. 6. Moment of Inertia of a Solid not Symmetrical with respect to the Axis of Gyration.

(1) To find the radius of gyration of a solid cylinder about an axis perpendicular to its own through its middle point.

Let x be the distance of any thin circular slice of the cylinder from the middle point of its axis; dx the thickness of the slice;

 ρ the density of the cylinder, b its radius, and 2a its length. Then, the moment of inertia of the slice about any diameter being equal to $\frac{1}{4}\pi\rho b^{4}dx$, its moment of inertia about the axis of gyration of the present

problem will be equal to

$$\pi \rho b^2 dx \cdot (x^2 + \frac{1}{4} b^2).$$

Hence, Mk² denoting the moment of inertia of the whole cylinder about the proposed axis, we have

$$Mk^{2} = \pi \rho b^{2} \int_{-a}^{+a} (x^{2} + \frac{1}{4} b^{2}) dx$$
$$= \pi \rho b^{2} (\frac{2}{3} a^{3} + \frac{1}{2} ab^{2})$$
$$= 2\pi \rho a b^{2} (\frac{1}{3} a^{2} + \frac{1}{4} b^{2});$$

and therefore, M being equal to $2\pi\rho ab^2$, we have

$$k^2 = \frac{1}{2} a^2 + \frac{1}{4} b^2.$$

Euler; Theoria Motus Corporum Solidorum, p. 196.

(2) To find the radius of gyration of a right cone about an axis at right angles to the axis of the cone and passing through its centre of gravity.

If a be the altitude of the cone, and c the radius of its base; then $k^2 = \frac{3}{80} (a^2 + 4c^2)$. Euler; Ib. p. 197.

(3) To find the radius of gyration of a right cone about an axis through its vertex at right angles to its geometrical axis.

If a = the altitude of the cone, and c = the radius of the base, $k^3 = \frac{3}{20} (4a^2 + c^3)$.

(4) To find the radius of gyration of a double convex lens about its axis, and about a diameter to the circle in which its two spherical surfaces intersect; the two surfaces having equal radii.

If a = the semi-axis of the lens, b = the radius of the circular intersection of the two surfaces; k = the radius of gyration of the lens about its axis, and k' about a diameter of the circle; we shall have

$$k^{2} = \frac{1}{10} \frac{a^{4} + 5a^{2}b^{2} + 10b^{4}}{a^{2} + 3b^{2}}, \quad k'^{2} = \frac{1}{20} \frac{7a^{4} + 15a^{2}b^{2} + 10b^{4}}{a^{2} + 3b^{2}}.$$
Euler; *Ib.* p. 201.

CHAPTER VI.

D'ALEMBERT'S PRINCIPLE.

A GENERAL method for the determination of the motion of a material system, acted on by any forces, was laid down by D'Alembert in his *Traité de Dynamique*, published in the year 1743¹, from which we have extracted the following passage in exposition of the Principle.²

" Problème Général.

"Soit donné un système de corps disposés les uns par rapport aux autres d'une manière quelconque; et supposons qu'on imprime à chacun de ces corps un mouvement particulier, qu'il ne puisse suivre à cause de l'action des autres corps; trouver le mouvement que chaque corps doit prendre.

" Solution.

¹ Sec also his Recherches sur la Précession des Equinoxes, p. 35, published in 1749.

D'Alembert's Principle was first enunciated by him in memoir which he read before the Academy of Sciences at the end of the year 42.

c' est-à-dire que, si les corps n' avoient reçu que les mouvemens α , β , κ , etc. ces mouvemens auroient dû se détruire mutuellement, et le système demeurer en repos.

"De là résulte le principe suivant, pour trouver le mouvement de plusieurs corps qui agissent les uns sur les autres. Décomposez les mouvemens a, b, c, etc. imprimés à chaque corps, chacun en deux autres a, a; b, β ; c, κ ; etc. qui soient tels, que si l on l eut imprimé aux corps que les mouvemens a, b, c, etc. ils eussent pu conserver ces mouvemens sans se nuire réciproquement; et que si on ne leur eût imprimé que les mouvemens a, β , κ , etc. le système fût demeuré en repos; il est clair que a, b, c, seront les mouvemens que ces corps prendront en vertu de leur action. Ce qu' il falloit trouver."

The idea of the general method developed by D'Alembert for the determination of the motion of material systems, had occurred somewhat earlier to Fontaine, as we are informed in the Table des Mémoires, prefixed to his Traité de Calcul Différentiel et Integral, having been communicated by him to the Academy of Sciences in the year 1739, and subsequently to several mathematicians. His views, however, on this subject were not made public till long after the appearance of the Traité de Dynamique; and in all probability D'Alembert, who did not become a member of the Academy before the year 1741, was not aware of Fontaine's generalization. D'Alembert, however, was the first to shew the wonderful fertility of the Principle by applying it to the solution of a great variety of difficult problems, among which may be mentioned that of the Precession of the Equinoxes, which had been inadequately attempted by Newton, and of which D'Alembert was the first to obtain a complete solution.

The earliest step towards the discovery of D'Alembert's Principle is to be met with in a memoir by James Bernoulli in the Acta Eruditorum, 1686, Jul. p. 356, entitled "Narratio Controversiæ inter Dn. Hugenium et Abbatem Catelanum agitatæ de Centro Oscillationis quæ loco animadversionis esse poterit in Responsionem Dn. Catelani, num. 27. Ephem. Gallic. anni 1684, insertam." Let m, m', denote two equal bodies attached to an

¹ Mémoires de l'Académie des Sciences de Paris, 1770.

inflexible straight line which is capable of motion in a vertical plane about one extremity which is fixed; let r, r', denote the distances of m, m', respectively, from the fixed extremity; v, v', their velocities for any position of the inflexible line in its descent from an assigned position, u, u'; the velocities which they would have acquired by descending down the same arcs unconnectedly. Then, in consequence of the connection of the bodies, a velocity u - v will be lost by m and a velocity v' - u' gained by m' in their descent. Bernoulli proposes it to the consideration of mathematicians whether, according to the statical relation of two forces in equilibrium on a lever, the proportion u - v : v' - u' :: r' : r be an accurate expression of the circumstances of the motion. This idea of Bernoulli's, although not free from error, contains however the first germ of the Principle of reducing the determination of the motions of material systems to the solution of statical problems. L'Hôpital, in a letter addressed to Huyghens, correctly observed that instead of considering the velocities acquired in a finite time, he should have considered the infinitesimal velocities acquired in an instant of time, and have compared them with those which gravity tends to impress upon the bodies during the same instant. He takes a complex pendulum, consisting of any two bodies attached to an inflexible straight line, and considers equilibrium to subsist between the quantities of motion lost and gained by these bodies in any instant of time, that is, between the differences of the quantities of motion which the bodies really acquire in this instant, and those which gravity tends to impress on them. He applies this Principle, which agrees with the general Principle of D'Alembert, to the determination of the Centre of Oscillation of a pendulum consisting of two bodies attached to an inflexible straight line oscillating about one extre-He then extends his theory to a greater number of bodies in a straight line, and determines their Centre of Oscillation on the supposition, the truth of which is not however sufficiently obvious without demonstration, that any two of them may be collected at their particular Centre of Oscillation. On the publication of L'Hôpital's letter, James Bernoulli2 reverted

¹ Histoire des Ouvrages des Scavans, 1690, Juin. p. 440.

² Acta Erudit. Lips. 1691. Jul. p. 317. Opera, tom. 1. p. 460.

to the subject of the Centre of Oscillation, and at length succeeded in obtaining a direct and rigorous solution of the problem in the case where all the bodies are in one line, by the application of the principle laid down by L'Hôpital. Bernoulli¹ afterwards extended his method to the general case of the oscillations of bodies of any figure.

An ingenious investigation of the Centre of Oscillation, a problem from the beginning intimately connected with the development of D'Alembert's Principle, was shortly afterwards given by Brook Taylor³ and John Bernoulli, 3 between whom arose an angry controversy respecting priority of discovery; the method given by these mathematicians, although depending upon the statical principles of the lever, did not however involve, in an explicit form, L'Hôpital's Principle of Equilibrium. Finally, Hermann determined the Centre of Oscillation by the principle of the statical equivalence of the solicitations of gravity, and the vicarious solicitations applied in opposite directions, or, as it is expressed by modern mathematicians, by the equilibrium subsisting between the impressed forces of gravity and the effective forces applied in opposite directions; a method of investigation virtually coincident with that given by James Bernoulli. The idea of L'Hôpital became still more general in the hands of Euler,6 in a memoir on the determination of the oscillations of flexible strings printed in the year 1740. From the above historical sketch it will be easily seen that in the enunciation of a general Principle of Motion, Fontaine and D'Alembert had little more to do than to express in general language what had been distinctly conceived in the prosecution of particular researches by L'Hópital, James and John Bernoulli, Brook Taylor, Hermann, and Euler. For additional information on the historical development of D'Alembert's Principle, the reader is referred

¹ Mémoires de l'Académie des Sciences de Paris, 1703, 1704.

^{*} Philosophical Transactions, 1714, May. Methodus Incrementorum.

^{*} Acta Erudit. Lips. 1714. Jun. p. 257; Mém. Acad. Par. 1714, p. 208. Opera, tom. 11. p. 168.

⁴ Act. Erudit, Lips. 1716, 1718, 1719, 1721, 1722.

^{*} Phoronomia; lib. t. cap. 5.

^{*} Comment. Petrop. tom. VII.

to Lagrange's Mécanique Analytique, Seconde Partie, Section 1; Montucla's Histoire des Mathématiques, part. v. liv. 3, part. 1v. liv. 7; and Whewell's History of the Inductive Sciences, vol. 11.

In modern treatises on Mechanics, D'Alembert's Principle is expressed under one or other of the following forms:

- (1) When any material system is in motion under the action of any forces, the moving forces lost by the different molecules of the system must be in equilibrium.
- (2) If the effective moving forces of the several particles of a system be applied to them in directions opposite to those in which they act; they will, conjointly with the impressed moving forces, constitute a system of forces statically disposed.

The former of these enunciations it will be seen is substantially the same as that given by D'Alembert, while the latter is a generalization of the idea developed by Hermann in his investigations on the particular problem of the Centre of Oscillation.

SECT. 1. Motion of a single Particle.

The object of this section is to apply D'Alembert's Principle to the exemplification of a general method for the determination of the motion of a particle within tubes and between contiguous surfaces, of which either the position, or the form, or both, are made to vary according to any assigned law whatever, the particle being acted on by given forces. Several of the problems of this section have been solved by particular methods in Chapter IV.

I. We will commence with the consideration of the motion of a particle along a tube, and, for the sake of perfect generality, we will suppose the tube to be one of double curvature. The tube is considered in all cases to be indefinitely narrow and perfectly smooth, and every section at right angles to its axis to be circular.

Let the particle be referred to three fixed rectangular axes, and let x, y, z, be its co-ordinates at any time t; let x, y, z, become $x + \delta x$, $y + \delta y$, $z + \delta z$, when t becomes $t + \delta t$; δt , and

¹ The substance of this Section was published in the Cambridge Mathematical Journal, vol. 111, p. 49.

consequently δx , δy , δz , being considered to be indefinitely small. Then the effective accelerating forces on the particle parallel to the three fixed axes will be, at the time t,

$$\frac{\delta^2 x}{\delta t^2}$$
, $\frac{\delta^2 y}{\delta t^2}$, $\frac{\delta^2 z}{\delta t^2}$.

Also, let X, Y, Z, represent the impressed accelerating forces on the particle resolved parallel to the axes of x, y, z; and let x + dx, y + dy, z + dz, be the co-ordinates of a point in the tube very near to the point x, y, z, which the particle occupies at the time t. Then, observing that the action of the tube on the particle is always at right angles to its axis at every point and therefore, at the time t, to the line joining the two points x, y, z, and x + dx, y + dy, z + dz, we have, by D'Alembert's Principle, combined with the Principle of Virtual Velocities,

$$\left(\frac{\delta^2 x}{\delta t^2} - X\right) dx + \left(\frac{\delta^2 y}{\delta t^2} - Y\right) dy + \left(\frac{\delta^2 z}{\delta t^2} - Z\right) dz = 0 \dots (A).$$

Again, since the form and position of the tube are supposed to vary according to some assigned law, it is clear that when t is known the equations to the tube must be known; hence it is evident that, in addition to the equation (A), we shall have, from the particular conditions of each individual problem, a number of equations equivalent to two of the form

$$\phi(x, y, z, t) = 0, \quad \chi(x, y, z, t) = 0 \dots (B),$$

where ϕ and χ are symbols of functionality depending upon the law of the variations of the form and position of the tube.

The three equations (A) and (B) involve the four quantities x, y, z, t, and therefore, in any particular case, if the difficulty of the analytical processes be not insuperable, we may ascertain x, y, z, each of them in terms of t; in which consists the complete solution of the problem.

If the tube remain during the whole of the motion within one plane, then, the plane of x, y, being so chosen as to coincide with this plane, the three equations (A) and (B) will evidently be reduced to the two

$$\left(\frac{\delta^{3}x}{\delta t^{3}}-X\right)dx+\left(\frac{\delta^{2}y}{\delta t^{3}}-Y\right)dy=0,\ldots, (C),$$

$$\phi\left(x,y,t\right)=0,\ldots, (D).$$

We proceed to illustrate the general formulæ of the motion by the discussion of a few problems.

(1) A rectilinear tube revolves with a uniform angular velocity about one extremity in a horizontal plane; to find the motion of a particle within the tube.

Let ω be the constant angular velocity; r the distance of the particle at any time t from the fixed extremity of the tube; then the plane of x, y, being taken horizontal, and the origin of co-ordinates at the fixed extremity of the tube, we shall have, supposing the tube initially to coincide with the axis of x,

$$x = r \cos \omega t \dots (1),$$

 $y = r \sin \omega t \dots (2).$

From (1) we have

$$dx = dr \cos \omega t,$$

and from (2),

$$dy = dr \sin \omega t$$
.

Again, from (1) we have

$$\frac{\delta x}{\delta t} = \frac{\delta r}{\delta t} \cos \omega t - \omega r \sin \omega t,$$

$$\frac{\delta^2 x}{\delta t^2} = \frac{\delta^2 r}{\delta t^2} \cos \omega t - 2\omega \frac{\delta r}{\delta t} \sin \omega t - \omega^2 r \cos \omega t;$$

and from (2).

$$\frac{\delta y}{\delta t} = \frac{\delta r}{\delta t} \sin \omega t + \omega r \cos \omega t,$$

$$\frac{\delta^2 y}{\delta t^2} = \frac{\delta^2 r}{\delta t^2} \sin \omega t + 2\omega \frac{\delta r}{\delta t} \cos \omega t - \omega^2 r \sin \omega t.$$

Substituting in the general formula (C) the values which we have obtained for dx, dy, $\frac{\delta^2 x}{\delta \ell^2}$, $\frac{\delta^2 y}{\delta \ell^2}$, we have, since X = 0, Y = 0,

$$\cos \omega t \left(\frac{\delta^2 r}{\delta t^2} \cos \omega t - 2\omega \frac{\delta r}{\delta t} \sin \omega t - \omega^2 r \cos \omega t \right)$$

$$+ \sin \omega t \left(\frac{\delta^2 r}{\delta t^2} \sin \omega t + 2\omega \frac{\delta r}{\delta t} \cos \omega t - \omega^2 r \sin \omega t \right) = 0 ;$$
and therefore
$$\frac{\delta^2 r}{\delta t^2} - \omega^2 r = 0 :$$

the integral of this equation is

$$r = C \epsilon^{\omega t} + C' \epsilon^{-\omega t}$$

Let r = a when t = 0; then

$$a = C + C';$$

also let $\frac{\delta r}{\delta t} = \beta$ when t = 0; then

$$\beta = C\omega - C'\omega$$
;

from the two equations for the determination of C and C', we have

 $C=\frac{a\omega+\beta}{2\omega}$, $C'=\frac{a\omega-\beta}{2\omega}$;

hence for the motion of the particle along the tube

$$2\omega r = (a\omega + \beta) \, \varepsilon^{\omega t} + (a\omega - \beta) \, \varepsilon^{-\omega t}.$$

This problem, which is the earliest problem of the motion of a particle subject to the constraint of a curve moving according to a prescribed law, is due to John Bernoulli¹. A solution of this problem is given also by Clairaut², to whom it had probably been proposed by Bernoulli.

(2) Supposing the tube to revolve in a vertical instead of a horizontal plane, we shall have, by the same process, the axis of y being now taken vertical, observing that X = 0, $Y = -g \sin \omega t$, if the time be reckoned from the moment of coincidence of the tube with the axis of x which is horizontal,

$$\frac{\delta^2 r}{\delta t^2} - \omega^2 r = -g \sin \omega t.$$

The integral of this equation is

$$r = C \varepsilon^{\omega t} + C' \varepsilon^{-\omega t} + \frac{g}{2\omega^3} \sin \omega t;$$

and if we determine the constants from the conditions that r, $\frac{\delta r}{\delta t}$, shall have initially values a, β , we shall get for the motion along the tube,

$$2\omega r = \left(a\omega + \beta - \frac{g}{2\omega}\right)\varepsilon^{\omega t} + \left(a\omega - \beta + \frac{g}{2\omega}\right)\varepsilon^{-\omega t} + \frac{g}{\omega}\sin \omega t.$$

1 Opera, tom. IV. p. 248.

^a Mémoires de l'Académie des Sciences de Paris, 1742, p. 10.

This problem, which had been erroneously attempted by Barbier in the Annales de Gergonne, tom. XIX., was correctly solved, in the following volume, by Ampère. In the Cambridge Mathematical Journal, vol. III. p. 42, a solution is given by Professor Booth, who has discussed at length the more interesting cases of the motion.

(3) Suppose the tube to revolve in a horizontal plane about a fixed extremity with such an angular velocity, that the tangent of its angle of inclination to the axis of x is proportional to the time.

The equation to the tube at any time t will be

$$y = mtx \dots (1),$$

where m is some constant quantity; hence

$$dy = mt dx$$

and therefore from (C), since X = 0 and Y = 0,

$$\frac{\delta^2 x}{\delta t^2} + mt \frac{\delta^2 y}{\delta t^2} = 0 \dots (2).$$

But from (1) we have

$$\frac{\delta y}{\delta t} = mt \frac{\delta x}{\delta t} + mx,$$

$$\frac{\delta^2 y}{\delta t^2} = mt \frac{\delta^2 x}{\delta t^2} + 2m \frac{\delta x}{\delta t};$$

hence, from (2),

$$(1 + m^2t^2)\frac{\delta^2x}{\delta t^2} + 2m^2t\frac{\delta x}{\delta t} = 0,$$

$$\frac{\delta^2x}{\frac{\delta t^2}{\delta x}} + \frac{2m^2t}{1 + m^2t^2} = 0.$$

Integrating, we have

$$\log \frac{\delta x}{\delta t} + \log (1 + m^2 t^2) = \log C,$$

$$\frac{\delta x}{\delta t} (1 + m^2 t^2) = C.$$

Let β be the initial value of $\frac{\delta x}{\delta t}$, which will be the velocity of

projection along the tube; then $C = \beta$, and therefore

$$\frac{\delta x}{\delta t} (1 + m^2 t^2) = \beta, \quad \delta x = \frac{\beta \delta t}{1 + m^2 t^2};$$

integrating, we get

$$x + C = \frac{\beta}{m} \tan^{-1} (m\ell).$$

Let x = a when t = 0; then a + C = 0, and therefore

$$x = a + \frac{\beta}{m} \tan^{-1} (m\ell),$$

and consequently, from (1),

$$y = amt + \beta t \tan^{-1}(mt).$$

If θ be the inclination of the tube to the axis of x at any time, and r be the distance of the particle from the fixed extremity,

$$r=\frac{am+\beta\theta}{m\cos\theta}.$$

(4) A circular tube is constrained to move in a horizontal plane with a uniform angular velocity about a fixed point in its circumference; to determine the motion of a particle within the tube, which is placed initially in the extremity of the diameter passing through the fixed point.

Let the fixed point be taken as the origin of co-ordinates, and let the axis of x coincide with the initial position of the diameter through this point; let ω be the angular velocity of the revolution of the circle, a the radius; also let θ be the angle at any time t between the distance of the particle and of the extremity of the diameter through the origin from the centre of the circle.

Then it will be easily seen that

$$x = a \cos \omega t + a \cos (\omega t - \theta) \dots (1),$$

$$y = a \sin \omega t + a \sin (\omega t - \theta) \dots (2)$$

From (1) we have

$$dx = ad\theta \sin(\omega t - \theta),$$

and, from (2), $dy = -ad\theta \cos (\omega t - \theta)$

Hence, from (C), observing that X = 0 and Y = 0,

$$\sin (\omega t - \theta) \frac{\delta^2 x}{\delta t^2} - \cos (\omega t - \theta) \frac{\delta^2 y}{\delta t^2} = 0 \dots (3).$$

$$\frac{\delta x}{\delta t} = -a\omega \sin \omega t + a\left(\frac{\delta \theta}{\delta t} - \omega\right) \sin (\omega t - \theta),$$

$$\frac{\delta^2 x}{\delta t^2} = -a\omega^2 \cos \omega t - a\left(\frac{\delta \theta}{\delta t} - \omega\right)^2 \cos (\omega t - \theta) + a\frac{\delta^2 \theta}{\delta t^2} \sin (\omega t - \theta);$$
and, from (2),

$$\frac{\delta y}{\delta t} = a\omega \cos \omega t - a\left(\frac{\delta \theta}{\delta t} - \omega\right) \cos (\omega t - \theta),$$

$$\frac{\delta^2 y}{\delta t^2} = -a\omega^2 \sin \omega t - a\left(\frac{\delta \theta}{\delta t} - \omega\right)^2 \sin (\omega t - \theta) - a\frac{\delta^2 \theta}{\delta t^2} \cos (\omega t - \theta);$$

and therefore, by (3),

$$a\omega^{2}\left\{\sin \omega t \cos(\omega t-\theta)-\cos \omega t \sin(\omega t-\theta)\right\}+a\frac{\delta^{2}\theta}{\delta t^{2}}=0,$$

$$\omega^2 \sin \theta + \frac{\delta^2 \theta}{\delta t^2} = 0:$$

multiplying by $2\frac{\delta\theta}{\delta t}$, and integrating,

$$\frac{\delta\theta^2}{\delta t^2} = C + 2\omega^2 \cos \theta.$$

But, the absolute velocity of the particle being initially zero, it is clear that 2ω will be the initial value of $\frac{\delta\theta}{\delta t}$; and therefore, θ being initially zero, we have

$$4\omega^2=C+2\omega^2,\quad C=2\omega^2,$$

and therefore

$$\frac{\delta\theta}{\delta t^2} = 2\omega^2 \left(1 + \cos\theta\right) = 4\omega^2 \cos^2\frac{\theta}{2}, \quad \frac{\delta\theta}{\delta t} = 2\omega \cos\frac{\theta}{2},$$

$$\frac{\cos\frac{\theta}{2}\delta\theta}{\cos^2\frac{\theta}{2}} = 2\omega\delta t, \quad \frac{\delta\sin\frac{\theta}{2}}{1 - \sin^2\frac{\theta}{2}} = \omega\delta t.$$

Integrating, we have

$$\log \frac{1+\sin\frac{\theta}{2}}{1-\sin\frac{\theta}{2}} = 2\omega t + C;$$

but $\theta = 0$ when t = 0; hence C = 0, and we have

$$\frac{1+\sin\frac{\theta}{2}}{1-\sin\frac{\theta}{2}}=\epsilon^{2\omega t},$$

and therefore

$$\sin\frac{\theta}{2} = \frac{\varepsilon^{\omega^t} - \varepsilon^{-\omega^t}}{\varepsilon^{\omega^t} + \varepsilon^{-\omega t}},$$

which determines the position of the particle within the tube at any time. When $t=\infty$, we have $\sin\frac{\theta}{2}=1$, and therefore $\theta=\pi$, which shews that, after the lapse of an infinite time, the particle will arrive at the point of rotation.

(5) If we pursue the same course as in the solution of the problems (1), (2), (4), we may obtain a convenient formula for the following more general problem: a plane curvilinear tube of any invariable form whatever revolves in its own plane about a fixed point with a uniform angular velocity; to determine the motion of a particle, acted on by any forces, within the tube.

Let ω be the constant angular velocity of the tube about the fixed point; r the distance of the particle at any time from this point; ϕ the angle between the simultaneous directions of r and of a line joining an assigned point of the tube with the fixed point of rotation; ds an element of the length of the tube at the place of the particle, and S the accelerating force on the particle resolved along the element ds; then the equation for the motion of the particle will be

$$r^2 \frac{\delta \phi^2}{\delta t^2} + \frac{\delta r^2}{\delta t^2} - \omega^2 r^2 = 2 \int S \frac{ds}{d\phi} \, \delta \phi :$$

but since, the form of the tube being invariable, $\delta \phi$, δr , may evidently be replaced by $d\phi$, dr, we have, putting, for the sake of uniformity of notation, dt in place of δt ,

$$r^2\frac{d\phi^2}{dt^2}+\frac{dr^2}{dt^2}-\omega^2r^2=2\int Sds.$$

If w be zero, the formula will become

$$r^2\frac{d\phi^2}{dt^2}+\frac{dr^2}{dt^2}=2\int Sds,$$

the well-known formula for the motion of a particle under the action of any forces within an immoveable plane tube.

(6) In the foregoing examples the position of the tube varies with the time, the form however remains invariable. We will now give an example in which the form changes with the time.

A particle is projected with a given velocity within a circular tube, the radius of which increases in proportion to the time while the centre remains stationary; to determine the motion of the particle, the tube being supposed to lie always in a horizontal plane.

The equation to the circle will be

$$x^2 + y^2 = a^2(1 + at)^2 \dots (1),$$

where a and a are some constant quantities; hence

$$x\,dx+y\,dy=0,$$

and therefore, by the general formula (C),

$$y\frac{\delta^2x}{\delta t^2}-x\frac{\delta^2y}{\delta t^2}=0;$$

integrating, we have

$$y \frac{\delta x}{\delta t} - x \frac{\delta y}{\delta t} = C.$$

Let the axis of x be so chosen as to coincide with the initial distance of the particle from the centre, and let β be the initial velocity of the particle along the tube; then $C = -a\beta$, and therefore

$$x \frac{\delta y}{\delta t} - y \frac{\delta x}{\delta t} = a\beta....(2);$$

again, from (1), we have

$$x\frac{\delta x}{\delta t}+y\frac{\delta y}{\delta t}=a^2a(1+at)...(3);$$

multiplying (2) by y and (3) by x, and subtracting the former result from the latter, we have

$$(x^2+y^2)\frac{\delta x}{\delta t}=a^2a(1+at)x-a\beta y,$$

and therefore, by (1),

$$a (1 + at)^{2} \frac{\delta x}{\delta t} = aa (1 + at) x - \beta \left\{ a^{2} (1 + at)^{2} - x^{2} \right\}^{\frac{1}{4}}$$

Put $1 + at = \tau$; then

$$aa\tau^{2}\frac{\delta x}{\delta \tau} = aa\tau x - \beta (a^{2}\tau^{2} - x^{2})^{\frac{1}{2}};$$

again, put $x = m\tau$, and there is

$$aa\tau^{3}\left(m+\tau\frac{\delta m}{\delta \tau}\right) = aam\tau^{3} - \beta\tau(a^{3}-m^{2})^{\frac{1}{4}},$$

$$aa\tau^{3}\frac{\delta m}{\delta\tau} = -\beta\tau(a^{3}-m^{2})^{\frac{1}{4}},$$

$$-aa\frac{\delta m}{(a^{2}-m^{2})^{\frac{1}{4}}} = \beta\frac{\delta\tau}{\tau^{2}};$$

integrating,

$$C + aa \cos^{-1}\frac{m}{a} = -\frac{\beta}{\tau},$$

or, putting for m its value,

$$C + aa \cos^{-1} \frac{x}{a\tau} = -\frac{\beta}{\tau},$$

and putting for τ its value 1 + at,

$$C + aa \cos^{-1} \frac{x}{a(1+at)} = -\frac{\beta}{1+at}.$$

Now x = a when t = 0; hence $C = -\beta$, and therefore

$$aa \cos^{-1} \frac{x}{a(1+at)} = \frac{a\beta t}{1+at},$$

$$x = a(1 + at) \cos \frac{\beta t}{a(1 + at)},$$

and therefore, from (1),

$$y = a(1 + at) \sin \frac{\beta t}{a(1 + at)};$$

which give the absolute position of the particle at any assigned time.

II. We proceed now to the consideration of the motion of a particle along a surface from which it is unable to detach itself, while the surface itself changes its position or its form, or both, according to any assigned law. • To fix the ideas, we suppose the particle to move between two surfaces indefinitely close together, so as to be expressed by the same equation.

Let x, y, z, be the co-ordinates of the particle at any time t; and let δx , δy , δz , be the increments of x, y, z, in an indefinitely small time δt ; also let dx, dy, dz, denote the increments of x, y, z, in passing from the point x, y, z, to any point near to it within the surface as it exists at the time t. Also let X, Y, Z, denote the resolved parts of the accelerating forces on the particle at the time t parallel to the axes of x, y, z; then, obtaining that the action of the surface on the particle is always in the direction of the normal at each point, we have, by D'Alembert's Principle combined with the Principle of Virtual Velocities,

$$\left(\frac{\delta^2 x}{\delta \epsilon^2} - X\right) dx + \left(\frac{\delta^2 y}{\delta \epsilon^2} - Y\right) dy + \left(\frac{\delta^2 z}{\delta \epsilon^2} - Z\right) dz = 0....(A').$$

Again, since the position and form of the surface vary according to an assigned law, its equation must evidently be known at any given time, and therefore we must have, from the nature of each particular problem, certain conditions between the quantities x, y, z, t, equivalent to a single equation

$$F = f(x, y, z, t) = 0 \dots (B')$$

Taking the total differential of (B'), we have

$$\frac{dF}{dx}\,dx + \frac{dF}{dy}\,dy + \frac{dF}{dz}\,dz = 0;$$

eliminating dz between this equation and (A'), we get

$$\left(\frac{\delta^{a}x}{\delta t^{a}} - X\right) \frac{dF}{dz} dx + \left(\frac{\delta^{a}y}{\delta t^{a}} - Y\right) \frac{dF}{dz} dy$$

$$= \left(\frac{\delta^{a}z}{\delta t^{a}} - Z\right) \left(\frac{dF}{dx} dx + \frac{dF}{dy} dy\right);$$

but dx and dy are independent quantities; we have therefore, by equating separately their coefficients on each side of the equation,

$$\left(\frac{\delta^{a}x}{\delta t^{a}} - X\right) \frac{dF}{dz} = \left(\frac{\delta^{a}z}{\delta t^{a}} - Z\right) \frac{dF}{dx},$$

$$\left(\frac{\delta^{a}y}{\delta t^{a}} - Y\right) \frac{dF}{dz} = \left(\frac{\delta^{a}z}{\delta t^{a}} - Z\right) \frac{dF}{dy},$$

and therefore also

$$\left(\frac{\delta^2 x}{\delta t^2} - X\right) \frac{dF}{dy} = \left(\frac{\delta^2 y}{\delta t^2} - Y\right) \frac{dF}{dx}$$

any two of these three relations, together with the equation (B), will give us three equations in x, y, z, t, whence x, y, z, are to be determined in terms of t.

The following example will serve to illustrate the use of these equations. We have taken a case where the form of the surface remains invariable, its position alone being liable to change. The analysis, however, in the solution of problems of the class which we are considering, receives its general character solely in consequence of the presence of t in the equation (B), and therefore the example which we have chosen is sufficient for the general object we have in view.

A particle descends by the action of gravity down a plane which revolves uniformly about a vertical axis through which it passes; to determine the motion of the particle.

Let the plane of x, y, be taken horizontal, the axis of x coinciding with the initial intersection of the revolving plane with the horizontal plane through the origin, and let the axis of z be taken vertically downwards; then, ω denoting the angular velocity of the plane, its equation at any time t will be

$$F = y \cos \omega t - x \sin \omega t = 0 \dots (1);$$

whence

$$\frac{dF}{dx} = -\sin \omega t, \quad \frac{dF}{dy} = \cos \omega t, \quad \frac{dF}{dz} = 0;$$

also X = 0, Y = 0, Z = g; and therefore, from either of the two first of the three general relations,

$$\frac{\delta^2 z}{\delta t^2} = g \ldots (2),$$

and from the third,

$$\frac{\delta^2 x}{\delta t^2} \cos \omega t + \frac{\delta^2 y}{\delta t^2} \sin \omega t = 0 \dots (3).$$

Let r denote the distance of the particle at any time from the axis of z; then

$$x = r \cos \omega t$$
, $y = r \sin \omega t$,

whence
$$\frac{\delta x}{\delta t} = \frac{\delta r}{\delta t} \cos \omega t - \omega r \sin \omega t$$
,

$$\frac{\delta^2 x}{\delta t^2} = \frac{\delta^2 r}{\delta t^2} \cos \omega t - 2\omega \frac{\delta r}{\delta t} \sin \omega t - \omega^2 r \cos \omega t,$$

$$\frac{\partial y}{\partial t} = \frac{\partial r}{\partial t} \sin \omega t + \omega r \cos \omega t,$$

$$\frac{\partial^2 y}{\partial t} = \frac{\partial^2 r}{\partial t} \sin \omega t + 2\omega \frac{\partial r}{\partial t} \cos \omega t - \omega^2 r \sin \omega t;$$

and therefore, from (8),

Let the initial values of z, $\frac{\partial z}{\partial t}$, be 0, β ; and these of r, $\frac{\partial r}{\partial t}$, be a, a; then, from the equations (2) and (4), after executing obvious operations, we shall obtain

$$z = \frac{1}{2}gt^{2} + \beta t,$$

$$2\omega r = (\omega a + a) e^{\omega t} + (\omega a - a) e^{-\omega t},$$
and
$$\log \frac{(\omega^{2}r^{2} + a^{2} - \omega^{2}a^{2})^{\frac{1}{2}} + \omega r}{a + \omega a} = \frac{\omega}{g} \left\{ (2gz + \beta^{2})^{\frac{1}{2}} - \beta \right\};$$

the two first of these equations give the position of the particle on the revolving plane, and therefore, by virtue of the equation (1), the absolute position of the particle at any time; while the third is the equation to the path which the particle describes on the plane.

SECT. 2. Systems of Particles.

(1) Two heavy particles P, P, (fig. 157), are attached to a rigid imponderable rod APP which is oscillating in a vertical plane about a fixed point in its extremity A; to determine the motion.

Let m, m', be the masses of the two particles; AP = a, AP' = a'. Draw AB vertically downwards and let $\angle PAB = \theta$. Let ds, ds', denote the elements of the circular paths described by P, P', in a small time dt, estimated in a direction corresponding to an increase of θ . Then the effective moving forces of the two particles will be $m \frac{d^2s}{dt^2}$, $m' \frac{d^2s'}{dt^2}$, the moments of which about the point A will be $ma \frac{d^2s}{dt^2}$, $m'a' \frac{d^2s'}{dt^2}$. Also the moments

of the impressed forces will be $-mag \sin \theta$, $-m'a'g \sin \theta$. Hence, for the equilibrium of the impressed forces, and the effective forces applied in directions opposite to their own, we have

 $ma \frac{d^2s}{dt^2} + m'a' \frac{d^2s'}{dt^2} + (ma + m'a') g \sin \theta = 0.$

But $ds = ad\theta$, $ds' = a'd\theta$; hence

$$(ma^2 + m'a'^2)\frac{d^2\theta}{dt^2} + (ma + m'a')g \sin \theta = 0;$$

a result which shews that the rod will oscillate isochronously with a perfect pendulum of which the length is

$$\frac{ma^2 + m'a'^2}{ma + m'a'}.$$

(2) Two particles, attached to the extremities of a fine inextensible thread, are placed upon two inclined planes with a common summit; to determine the motion of the particles and the tension of the thread at any time.

Let m, m', be the masses of the particles; a, a', the inclinations of the planes to the horizon; x, x', the distances of the particles from the common summit of the planes at any time. Then the impressed accelerating forces on the particles m, m', estimated down the two planes, will be $g \sin a$, $g \sin a'$, respectively, and the effective accelerating forces, estimated in the same directions, will be $g \frac{d^2x}{dt^2}$, $g \frac{d^2x'}{dt^2}$. Hence, for the equilibrium of the impressed forces, and the effective forces applied in directions

opposite to their own, we have $g(m \sin \alpha - m' \sin \alpha') = m \frac{d^2x}{d^2} - m' \frac{d^2x'}{d^2} \dots \dots (1).$

But, if c denote the length of the thread,

$$x + x' = c$$
, $\frac{d^2x}{dt^2} + \frac{d^2x'}{dt^2} = 0$;

hence, from (1),

$$(m+m')\frac{d^2x}{dt^2} = g(m \sin \alpha - m' \sin \alpha') \dots (2);$$

which determines the common acceleration of the two particles estimated in accordance with an increase of x; should the expression for $\frac{d^2x}{dt^2}$ be a negative quantity, x will decrease and x' increase.

If T denote the tension of the thread, we shall have, for the equilibrium of the impressed moving forces T, $mg \sin a$, exerted on the particle m, and the effective moving force $m \frac{d^2x}{dt^2}$ applied in a direction opposite to its own,

$$T = m \left(g \sin \alpha - \frac{d^2x}{dt^2} \right)$$
$$= \frac{mm'g}{m+m'} (\sin \alpha + \sin \alpha'), \text{ by (2)};$$

which gives the value of T, which is therefore of invariable magnitude during the whole motion.

Poisson; Traité de Mécanique, tom. 11. p. 12.

(3) One body draws up another on the wheel and axle; to determine the motion of the weights and the tensions of the strings.

Let a, a', denote the radii of the wheel and axle; m, m', the masses of the bodies suspended from them; s the arc described, at the end of the time t, by a molecule μ of the mass of the wheel and axle, r the distance of the molecule from the axis of rotation; x, x', the vertical distances, below the horizontal plane through the axis, of the masses m, m'.

Then the moment of the impressed forces about the axis of rotation will be mag - m'a'g;

and the moments of the effective forces, estimated in the same direction, will be

$$ma \; rac{d^2x}{dt^2} - m'a' \; rac{d^2x'}{dt^2} + \sum \left(\mu r \; rac{d^2s}{dt^2}
ight).$$

Hence, by D'Alembert's Principle,

$$ma \frac{d^2x}{dt^2} - m'a' \frac{d^2x'}{dt^2} + \Sigma \left(\mu r \frac{d^2s}{dt^2}\right) = mag - m'a'g \dots (1).$$

Let θ represent the whole angle through which the wheel and axle have rotated at the end of the time t; then b, b', denoting the initial values of x, x', it is clear that

$$x = b + a\theta$$
, $x' = b' - a'\theta$,

and therefore

$$\frac{d^3x}{dt^2} = a \frac{d^3\theta}{dt^2}, \quad \frac{d^3x'}{dt^2} = -a' \frac{d^3\theta}{dt^2} \dots (2).$$

Also, it is manifest that $s = r\theta$, and therefore

$$\Sigma\left(\mu r \frac{d^3s}{dt^2}\right) = \Sigma\left(\mu r^2 \frac{d^3\theta}{dt^2}\right) = \frac{d^3\theta}{dt^2} \Sigma\left(\mu r^2\right) = Mk^2 \frac{d^3\theta}{dt^2}...(3),$$

where Mk² denotes the moment of inertia of the wheel and axle together about the axis of rotation.

From (1), (2), (3), we obtain

$$(ma^2 + m'a'^2 + Mk^2)\frac{d^2\theta}{dt^2} = mag - m'a'g,$$

and therefore, if the system be supposed to have no motion when t = 0, $m_0 - m_0' c'$

 $\theta = \frac{1}{2}gt^{2}\frac{ma - m'a'}{ma^{2} + m'a'^{2} + Mk^{2}} \cdot \cdot \cdot \cdot \cdot (4).$

Let T denote the tension of the string supporting m, then

$$T = m\left(g - \frac{d^{2}x}{dt^{2}}\right)$$

$$= m\left(g - a\frac{d^{2}\theta}{dt^{2}}\right)$$

$$= mg\left\{1 - \frac{a\left(ma - m'a'\right)}{ma^{2} + m'a'^{2} + Mk^{2}}\right\}$$

$$= mg\frac{m'a'\left(a + a'\right) + Mk^{2}}{ma^{2} + m'a'^{2} + Mk^{2}}.$$

Similarly, the tension of the other string being denoted by T',

$$T' = m'g \frac{ma(a' + a) + Mk^2}{m'a'^2 + ma^2 + Mk^2}.$$

(4) A uniform heavy rod OA, (fig. 158), which is at liberty to oscillate in a vertical plane about a horizontal axis through O, falls from a horizontal position; to determine the angle included between the direction of the rod and the direction of the pressure for any position of the rod.

Let Ox, Oy, be the axes of co-ordinates in the plane of oscillation, Ox being horizontal and Oy vertical; let Oz be at right angles to the plane xOy. Let U, V, represent the resolved parts of the reaction of the axis Oz upon the rod, estimated along xO,

y0. Let ρ = the density of the rod, κ = the area of a section of it taken at right angles to its length; let P be any point in OA, draw PM at right angles to Ox; let

$$OM = x$$
, $PM = y$, $OP = r$, $OA = a$, $\angle AOx = \theta$.

Then, by D'Alembert's Principle, resolving forces parallel to Ox,

$$U = -\int_0^x \left\{ \kappa \rho \, d\mathbf{r} \, \frac{d^2x}{dt^2} \right\} = -\kappa \rho \int_0^x \left(d\mathbf{r} \, \frac{d^2x}{dt^2} \right). \quad (1);$$

resolving forces parallel to Oy,

$$V = -\kappa\rho \int_0^{\pi} \left\{ d\mathbf{r} \left(\frac{d^3y}{dt^2} - g \right) \right\} \dots \dots (2);$$

and, taking moments about the axis Oz,

$$\int_{a}^{a} u \rho \, g \, dr \cdot x = \int_{a}^{a} \left\{ u \rho \, dr \left(x \, \frac{d^{3}y}{dt^{2}} - y \, \frac{d^{3}x}{dt^{2}} \right) \right\},$$

$$g \int_{a}^{a} x \, dr = \int_{a}^{a} \left\{ dr \left(x \, \frac{d^{3}y}{dt^{2}} - y \, \frac{d^{3}x}{dt^{2}} \right) \right\} \dots (3).$$

But, from the geometry, we see that

 $x = r \cos \theta$, $\frac{dx}{dt} = -r \sin \theta \frac{d\theta}{dt}$, $\frac{d^2x}{dt^2} = -r \cos \theta \frac{d\theta^2}{dt^2} - r \sin \theta \frac{d^2\theta}{dt^2}$, and similarly

$$\frac{d^3y}{dt^2} = -r \sin \theta \frac{d\theta^2}{dt^2} + r \cos \theta \frac{d^3\theta}{dt^2};$$

hence, from (1), we have

$$U = \kappa \rho \int_0^a r dr \left(\cos \theta \, \frac{d\theta^2}{dt^2} + \sin \theta \, \frac{d^3\theta}{dt^2} \right)$$
$$= \frac{1}{2} a^3 \kappa \rho \left(\cos \theta \, \frac{d\theta^3}{dt^3} + \sin \theta \, \frac{d^3\theta}{dt^3} \right) \dots \dots (4);$$

and, from (2),

$$V = \kappa \rho g a + \frac{1}{2} a^3 \kappa \rho \left(\sin \theta \frac{d\theta^3}{dt^3} - \cos \theta \frac{d^3 \theta}{dt^3} \right) \dots \dots (5).$$

Again, from (3), substituting for x and y their values in terms of r and θ , we get

$$g\int_0^a \cos\theta \, rdr = \int_0^a r^2 dr \, \frac{d^2\theta}{dt^2},$$

and therefore

$$\frac{1}{2}ga^2\cos\theta = \frac{1}{3}a^3\frac{d^3\theta}{d^2}, \quad \frac{d^3\theta}{d^2\theta} = \frac{3g}{2a}\cos\theta;$$

multiplying by $2\frac{d\theta}{dt}$, integrating, and bearing in mind that $\theta = 0$ when $\frac{d\theta}{dt} = 0$, we have

$$\frac{d\theta^2}{dt^2} = \frac{3g}{a} \sin \theta.$$

Hence, substituting for $\frac{d\theta}{dt}$ and $\frac{d^2\theta}{dt^2}$ their values in (4) and (5), we obtain $U = \frac{2}{3} \cos q \sin \theta \cos \theta$

$$U = \frac{9}{4} \kappa \rho \ ag \sin \theta \cos \theta,$$

$$V = \frac{1}{4} \kappa \rho \ ag (10 - 9 \cos^2 \theta).$$

From these equations we get

$$U\cos\theta + V\sin\theta = \frac{5}{2}\kappa\rho ag\sin\theta$$
,
 $V\cos\theta - U\sin\theta = \frac{1}{2}\kappa\rho ag\cos\theta$.

But $U\cos\theta + V\sin\theta$ and $V\cos\theta - U\sin\theta$ are the expressions for the resolved parts of the reaction of the fixed axis, estimated along AO and at right angles to AO; hence, if ϕ denote the inclination of the resultant reaction to the line AO produced, or of the resultant pressure on the axis to the line OA, we shall have

$$\tan \phi = \frac{V \cos \theta - U \sin \theta}{U \cos \theta + V \sin \theta} = \frac{1}{10} \cot \theta,$$
$$\tan \theta \tan \phi = \frac{1}{10}.$$

CHAPTER VII.

MOTION OF RIGID BODIES ABOUT FIXED AXES.

Various Problems.

Let F denote the resolved part of any one of a system of forces acting on a rigid body, at right angles to a fixed axis, r being the perpendicular distance between the fixed axis and the direction of F. Then Fr will be the moment of this force about the axis, and if $\Sigma(Fr)$ denote the sum of the moments of all the forces affected by their appropriate signs, we shall have, for the determination of the motion of the body, the general formula

$$\frac{d\omega}{dt} = \frac{\Sigma(Fr)}{Mk^2},$$

where $\omega =$ the angular velocity of the body after a time t, and Mk^2 = its moment of inertia about the fixed axis.

(1) An isosceles right-angled triangle ABC (fig. 159) is suspended at the right angle A, and its side AB is kept vertical by a ring at B; an angular velocity ω being communicated to the triangle round AB, to determine the magnitude of ω that there may be no pressure at B.

Bisect BC in L, join AL, and take $AG = \frac{2}{3}AL$; then G will be the centre of gravity of the triangle; draw GH at right angles to AC. Take P any point in the area of the triangle, and draw PM at right angles to AB. Let AM = x, PM = y, AC = a=AB; m = the mass of a unit of area of the triangle. Then

 $AH = AG \cos \frac{\pi}{4} = \frac{2}{3} AL \cos \frac{\pi}{4} = \frac{2}{3} a \left(\cos \frac{\pi}{4}\right)^2 = \frac{1}{3} a.$

Also the area of the triangle is equal to $\frac{1}{2}a^2$, and therefore its mass to $\frac{1}{2}ma^2$; hence the moment of the triangle about an axis through A at right angles to its plane at any instant, in consequence of gravity, is

 $\frac{1}{5}ma^2q \cdot \frac{1}{3}a = \frac{1}{6}ma^3q$.

Again, the moment about the same axis due to centrifugal force is equal to

$$\iint m\omega^2 y dx dy \cdot x = \frac{1}{2} m\omega^2 \int y^2 x dx$$
$$= \frac{1}{2} m\omega^2 \int_0^a (a - x)^2 x dx = \frac{1}{24} m\omega^2 a^4.$$

Now, since there is no pressure on the ring at B, the moments of gravity and of centrifugal force about the axis through A must be equal; hence we have

$$\tfrac{1}{6} ma^3 g = \tfrac{1}{24} m\omega^2 a^4,$$

and therefore

$$\omega^2 = \frac{4g}{a}$$
, $\omega = 2\left(\frac{g}{a}\right)^{\frac{1}{4}}$.

(2) Supposing the force which acts on the crank of a steamengine to be vertical, and to vary as the sine of the angle through which the crank has revolved at any time from a vertical position; to find the angular velocity of the crank in any position, the moment of the resistance being always equal to half the greatest moment of the force, and the moment of the weight of the crank being considered negligible.

Let AO (fig. 160) be the crank, O being the fixed extremity; draw Ox vertical; let $\angle AOx = \theta$ at any time t; F = the force acting at the extremity A; OA = a: assume $F = \mu \sin \theta$; let mk^2 denote the moment of inertia of the crank about O.

Then, the moment of the resistance about O being $\frac{1}{2}\mu a$, we have, for the motion of the crank,

$$mk^{2} \frac{d^{2}\theta}{dt^{2}} = \mu \sin \theta \cdot a \sin \theta - \frac{1}{2} \mu a$$
$$= -\frac{1}{2} \mu a \cos 2\theta :$$

multiplying by 2 $\frac{d\theta}{dt}$ and integrating, we obtain

$$mk^2 \frac{d\theta^2}{dt^2} = C - \frac{1}{2} \mu a \sin 2\theta:$$

let ω denote the angular velocity of the crank when $\theta = 0$; then

$$mk^2\omega^2=C$$
;

hence

$$\frac{d\theta^2}{dt^2} = \omega^2 - \frac{\mu a \sin 2\theta}{2mk^2},$$

which gives the angular velocity of the crank in any position: from this result we see that the angular velocity is always when the crank is in either a horizontal or a vertical position.

SECT. 2. Centre of Oscillation.

Conceive a body of any figure acted on by gravity to be oscillating about a fixed horizontal axis AB (fig. 161); let G be the centre of gravity of the body; draw GO at right angles to AB. Produce OG to a point C such that

$$OC = \frac{k^2 + k^2}{h},$$

where h = OG and k = the radius of gyration of the body about an axis through G parallel to AB; then if the whole mass of the body be collected at the point C, the period of its oscillations about AB will be the same as before. The point C is called the Centre of Oscillation or of Agitation.

The theory of the Centre of Oscillation of bodies originated in questions addressed, about the year 1646, by Mersenne to the mathematicians of his day, who were called upon by him to exert their ingenuity to discover the time of oscillation of bodies moveable about horizontal axes. It is rather singular that all those who first attempted the solution of this celebrated problem, among whom Mersenne¹ himself is to be numbered, together with Descartes,² Roberval,³ Wallis,⁴ and Fabri,⁵ tacitly supposed the Centre of Oscillation to be coincident with the Centre of Percussion; a supposition which, although true, is by no means obvious without a rigorous demonstration. On the strength of this assumption, however, the Centre of Oscillation was correctly determined in the case of certain figures. Descartes gave a true solution of the case where a plane area oscillates in planum, but failed in the case of solid bodies and of plane

¹ Mersenni Reflexiones Physico-Mathematica, cap. x1. et x11.

Lettres de Descartes, tom. III. p. 487, &c.

³ Lettres de Descartes, ib.

⁴ Mechanica, sive De Motu.

^{*} Tract. de Motu, Append. Physico Math. De Centro Percussionis.

areas oscillating in latus. Roberval assigned correctly the position of the Centre of Oscillation, not only of plane areas oscillating in planum, but also in certain instances of oscillation in latus, while together with Descartes he failed to give a correct solution of the problem in the case of solid figures. The labours of Huyghens, who in his earlier efforts to obtain a solution of Mersenne's problem had been utterly baffled, were at length crowned with success, and accordingly in the fourth part of his Horologium Oscillatorium, which appeared in the year 1673, was given the first rigorous and general investigation of the Centre of Oscillation. The two following axioms constitute the basis of his researches: first, that the centre of gravity of a system of heavy bodies cannot of itself rise to an altitude greater than that from which it has fallen, whatever change be made in the mutual disposition of the bodies; and secondly, that a compound pendulum will always ascend to the same height as that from which it has descended freely. Some years after the publication of the Horologium Oscillatorium, the truth of these fundamental axioms, which although true, it must be admitted, are not sufficiently elementary, was called in question by the Abbé Catelan,1 who substituted certain frail theories of his own in place of the valuable researches of Huyghens. The attention of the mathematicians of the day having been more closely directed to the subject by the controversy which arose between Huyghens and Catelan, the views of Huyghens received ample corroboration from the more elementary investigations of L' Hôpital, James Bernoulli, and other mathematicians. For information respecting the subsequent history of Mersenne's problem, the reader is referred to the Chapter on D'Alembert's Principle.

(1) To find at what point of the rod of a perfect pendulum must be fixed a given weight of indefinitely small volume, so as to have the greatest effect in accelerating the pendulum.

Let m be the mass of the bob of the perfect pendulum, and a its length; m' the mass of the given weight, and a' the distance of its point of attachment from the centre of suspension; l the distance between the centre of suspension and the centre of

Journal des Scavans, 1682 et 1684.

oscillation of the complex pendulum. Then we shall have, m and m' being both of indefinitely small volume,

$$l=\frac{ma^2+m'a'^2}{ma+m'a'}.$$

Now the shorter the rod of a perfect pendulum, the shorter will be the time of its oscillations; hence we must have l a minimum; differentiating then with respect to a' we get

$$\frac{dl}{da'} = \frac{2m'a'(ma + m'a') - m'(ma^2 + m'a'^2)}{(ma + m'a')^2} = 0 ;$$

hence

$$m'a'^{4} + 2maa' = ma^{2},$$
 $m'^{2}a'^{2} + 2ma \ m'a' + m^{2}a^{2} = (m^{2} + mm')a^{2},$
 $m'a' + ma = (m^{2} + mm')^{\frac{1}{4}}a,$
 $a' = \frac{a}{m'} \left\{ (m^{2} + mm')^{\frac{1}{4}} - m \right\},$

which determines the required point of attachment.

Lady's and Gentleman's Diary, 1742. Diarian Repository, p. 394. Euler; De Motu Corp. Solid., Prob. 48. Cor. 1. p. 216.

(2) To compare the times in which a circular plate will vibrate round a horizontal tangent and round a horizontal axis, through the point of contact, at right angles to the tangent.

Let l, l', denote the lengths of the isochronous pendulums in the former and latter case respectively; a the radius of the plate; k, k', the radii of gyration about axes through the centre of the plate parallel in each case to the axis of oscillation. Then

$$l = \frac{a^2 + k^2}{a}$$
, $l' = \frac{a^2 + k'^2}{a}$.

Let A denote the area of the plate; r the distance of a point within it from its centre, and θ the inclination of this distance to the horizon when the plate is hanging at rest. Then

$$Ak^2 = \iint r d\theta dr \cdot r^2 \sin^2 \theta, \text{ between the proper limits,}$$
$$= \int_0^{2\pi} \int_0^a r^3 \sin^2 \theta \ d\theta dr = \frac{1}{4} a^4 \int_0^{2\pi} \sin^2 \theta d\theta = \frac{1}{4} \pi a^4.$$

Also

$$Ak^{2} = \iint rd\theta dr \cdot r^{2} = \int_{0}^{2\pi} \int_{0}^{a} r^{3}d\theta dr = \frac{1}{4} a^{4} \int_{0}^{2\pi} d\theta = \frac{1}{2} \pi a^{4}.$$

But $A = \pi a^2$; hence $k^2 = \frac{1}{4} a^2$ and $k'^2 = \frac{1}{2} a^2$; and therefore $l = a + \frac{1}{4} a = \frac{6}{4} a$, $l' = a + \frac{1}{2} a = \frac{6}{4} a$.

Hence, if t, t', denote the times of vibration,

$$\frac{t}{t'} = \left(\frac{l}{l'}\right)^{\frac{1}{4}} = \left(\frac{5}{6}\right)^{\frac{1}{4}}.$$

(3) If l and h be the distances of the centres of oscillation and gravity of a mercurial pendulum of which the weight is m, from the axis of suspension, and h' be the distance of the centre of gravity of a small quantity of mercury μ by the addition of which the pendulum is made to vibrate seconds exactly, to determine the approximate ratio of μ to m, L being the length of the seconds pendulum, and r the radius of the cylinder containing the mercury.

The moment of inertia of the mercury μ , which may be regarded approximately as a circular lamina of fluid, about any diameter, and therefore about a diameter parallel to the axis from which the pendulum is suspended will be $\frac{1}{4} \mu r^2$, and therefore its moment of inertia about the axis of suspension will be

$$\mu (h^2 + \frac{1}{4} r^2).$$

Also the radius of gyration of the mercury m about a line through its centre of gravity parallel to the axis of suspension being k, the moment of inertia about the axis of suspension will be $m(h^2 + k^2)$. Hence, by the formula for the Centre of Oscillation, we have approximately

$$(\mu h' + mh) L = \mu (h'^2 + \frac{1}{4} r^2) + m (h^2 + k^2).$$

But also we shall have

hence
$$\begin{aligned} hl &= h^2 + h^2; \\ (\mu h' + mh) \ L &= \mu \left(h'^2 + \frac{1}{4} r^2 \right) + mhl, \\ \frac{\mu}{m} \left\{ h' \left(L - h' \right) - \frac{1}{4} r^2 \right\} &= h (l - L), \\ \frac{\mu}{m} &= \frac{4h (l - L)}{4h' \left(L - h' \right) - r^2}. \end{aligned}$$

(4) A bent lever, of which the arms are of lengths a and b, and the angle between them θ , makes small oscillations in its own plane about the angular point; to find the length of the isochronous simple pendulum.

Required length = $\frac{\frac{2}{3}(a^3 + b^3)}{(a^4 + 2a^2b^2 \cos \theta + b^4)^{\frac{1}{4}}}.$

(5) A uniform rod of given length is bent into the form of a cycloid, and oscillates about a horizontal line joining its extremities; to find the length of the isochronous pendulum.

If a be the length of the rod, the length of the isochronous pendulum will be $\frac{1}{6}a$.

(6) A pendulum consists of an indefinitely thin rigid rod OA, and a globe of which the centre is A; to determine the point A', in the line OA, at which the centre of another globe must be fixed in order that the oscillations of the system of the two globes may be executed in the smallest time possible.

Let OA = a, OA' = a'; also let r, r', be the radii, and m, m', the masses of the globes A, A'. Then

$$a' = \frac{1}{m'} \left\{ m \left(m + m' \right) a^2 + \frac{2}{5} m' \left(mr^2 + m'r'^2 \right) \right\}^{\frac{1}{4}} - \frac{ma}{m'}.$$

Euler; Theoria Motus Corporum Solidorum, p. 215.

CHAPTER VIII.

MOTION OF RIGID BODIES. FREE AXES. SMOOTH SURFACES.

IF a body be in motion about a Principal Axis, and be acted on by forces which do not tend to perturb the direction of this axis; then, the motion of the centre of gravity of the body remaining the same as if all the forces were impressed on the mass condensed at this point, the Principal Axis will always remain parallel to itself as an axis of permanent rotation, and the angular acceleration about this axis will be the same as if it were a fixed axis. The discovery of the existence of three principal axes in every body as axes of permanent rotation is due to Professor Segner of Gottingen, by whom it was communicated to the world in a memoir entitled Specimen Theoriæ Turbinum, published at Halle in the year 1755. For the complete development of the theory of rotation about permanent axes, the student is referred to Euler's Theoria Motus Corporum Solidorum, cap. VIII., a work of the greatest value for those who wish to acquire profound views on the subject of the motion of rigid bodies.

If a body be revolving at any instant of time about an axis which is not a principal one, this axis will not be one of permanent rotation; the body will revolve successively about a series of instantaneous axes, the positions of which both in relation to the body and to absolute space are different. The solution of the great physical problem of the Precession of the Equinoxes, published by D'Alembert² in the year 1749, unfolded a complete method for the investigation of the general problem of

¹ The following is the definition of Principal Axes given by Euler, Theoria Motus Corporum Solidorum, p. 175: "Axes principales cujusque corporis sunt tres illi axes per ejus centrum inertiæ transcuntes, quorum respectu momenta inertiæ sunt vel maxima vel minima."

^{*} Recherches sur la Précession des Equinoxes, 1749.

In the following year was published by Euler a memoir entitled Découverte d'un nouveau principe de Méchanique, the object of which was to investigate general formulæ for the motion of a body under the most general circumstances of motion and force. The equations, however, expressing under the most simple form the general conditions of rotation, were first given by Euler² in the year 1758, who availed himself of the principles of simplification afforded by the recent discoveries of Segner respecting the existence of the three Principal Axes of material bodies. The consideration of the general problem of rotation was resumed by D'Alembert, and presented under its most general aspect in the first volume of his Opuscules Mathematiques, published in 1761, where he expresses disapprobation of the title prefixed by Euler to his memoir of 1749, in consideration of his own investigations on the Precession of the Equinoxes. The subject of rotation was thoroughly investigated and exemplified by Euler in his Theoria Motus Corporum Solidorum et Rigidorum, which appeared in the year 1767. The same subject was afterwards investigated by Lagrange on more general principles of analysis. In the year 1777 appeared a memoir entitled 'A new Theory of the Rotatory Motion of Bodies affected by Forces disturbing such motion,' by Landen, a celebrated English mathematician, in which he expresses himself dissatisfied with the conclusions of the great continental philosophers on the subject of rotation. The subject was again resumed by Landen, a few years afterwards, when he developes more fully his own views, and persists in his opposition to the doctrines of his predecessors. There is a memoir by Wildbore in the Philosophical Transactions for the year 1790, in which the subject is investigated under a new light: the conclusions of the author are unfavourable to the cause of Landen, whose views are in fact now generally

¹ Mémoires de l'Académie des Sciences de Berlin, 1750.

^{*} Ibid. 1758.

¹ Specimen Theoriæ Turbinum, 1755.

Mémoires de l'Académie des Sciences de Berlin, 1773; Mécanique Analytique, Seconde Partie, ection 1x.

b Philosophical Transactions, 1777.

⁴ Ibid. 1785.

exploded. For further information on the history of the theory of rotation and Landen's controversy, the student is referred to a memoir by Mr. Whewell, in the second volume of the Cambridge Philosophical Transactions, 1827. The investigation of Euler's general equations of rotatory motion has been effected with great elegance and simplicity by Mr. O'Brien, in the fifth chapter of his Mathematical Tracts, part 1.

SECT. 1. Single Body.

(1) A rod PQ (fig 162) of uniform thickness and density, having been placed in a given position with one end upon a smooth horizontal plane OA, and the other leaning against a smooth vertical plane OB, descends in a vertical plane AOB by the action of gravity; to determine where the rod will detach itself from the vertical plane.

Let PG = a = GQ, G being the centre of gravity of the rod; let GH be vertical and equal to y at any time t of the motion; OH = x, $\angle QPO = \phi$; k = the radius of gyration of the rod about G; R = the reaction of the vertical plane, which will be horizontal, and S = that of the horizontal plane, which will be vertical; m = the mass of the rod.

Then, for the motion of the rod, we have, resolving forces horizontally, d^2x

$$m \frac{d^3x}{dt^2} = R \dots (1);$$

resolving vertically, $m \frac{d^3y}{dt^2} = S - mg \dots (2);$

and, taking moments about G,

$$mk^2 \frac{d^2 \phi}{dt^2} = Ra \sin \phi - Sa \cos \phi \dots (3).$$

Eliminating R and S between the three equations (1), (2), (3), we have

$$k^2 \frac{d^2 \phi}{dt^2} = a \sin \phi \frac{d^2 x}{dt^2} - a \cos \phi \frac{d^2 y}{dt^2} - ag \cos \phi \dots (4).$$

Now, from the geometry, it is clear that

$$x = a \cos \phi$$
, $y = a \sin \phi$.

and therefore

$$\frac{dx}{dt} = -a \sin \phi \frac{d\phi}{dt}, \quad \frac{d^3x}{dt} = -a \cos \phi \frac{d\phi^3}{dt^3} - a \sin \phi \frac{d^3\phi}{dt^3}, \quad \dots \quad (5)_s$$

$$\frac{dy}{dt} = a \cos \phi \, \frac{d\phi}{dt} \,, \quad \frac{d^3y}{dt^2} = -a \sin \phi \, \frac{d\phi^2}{dt^2} + a \cos \phi \, \frac{d^3\phi}{dt^2} \,;$$

hence we have

$$a \sin \phi \frac{d^3x}{dt^3} - a \cos \phi \frac{d^3y}{dt^3} = -a^2 \frac{d^3\phi}{dt^3};$$

and therefore, from (4),

$$(a^3+k^3)\frac{d^3\phi}{dt^2}=-ag\cos\phi\ldots\ldots(6);$$

multiplying by 2 $\frac{d\phi}{dt}$, and integrating, we have

$$(a^3+k^3)\frac{d\phi^3}{d\ell^3}=C-2ag\sin\phi;$$

but, if a be the initial value of ϕ , we have, since $\frac{d\phi}{dt} = 0$ initially, $0 = C - 2aq \sin a$;

and therefore
$$(a^3 + k^2) \frac{d\phi^2}{dt^2} = 2ag (\sin a - \sin \phi) \dots (7)$$
.

Now, at the instant when the rod detaches itself from the vertical plane, R=0; hence, by (1) and the value of $\frac{d^3x}{dt^2}$ in (5), $\cos\phi \frac{d\phi^2}{dt^2} + \sin\phi \frac{d^2\phi}{dt^2} = 0$;

and therefore, by (6) and (7),

$$2ag (\sin a - \sin \phi) \cos \phi = ag \cos \phi \sin \phi$$
;

whence, since φ cannot be equal to $\frac{1}{2}\pi$, we have

$$2 \sin \alpha - 2 \sin \phi = \sin \phi, \sin \phi = \frac{2}{3} \sin \alpha;$$

which gives the position of the rod at the moment of its separation from the vertical plane.

This problem was proposed by Weston, a disciple of Landen's, in the *Lady's and Gentleman's Diary* for the year 1757; and solved by Peter Walton, a contributor to the *Diary*. See *Diarian Repository*, p. 467.

(2) A uniform rod of given length hangs horizontally by two equal vertical strings attached to its ends; if it be twisted horizontally through a very small angle, so that its centre of gravity remains in the same vertical line, to find the time of an oscillation, the inertia of the strings being neglected.

Let P, Q, (fig. 163), be the points from which the strings PA, QB, are suspended, AB being the position in which the rod will rest; let ab be the position of the rod at any instant after disturbance; G the centre of gravity of AB, and therefore approximately of ab. Let AG = a = BG, AP = b = BQ, $\triangle AGa = 0$, m = 1 the mass of the rod.

Then, for small oscillations, the tension of each string may be considered equal to $\frac{1}{2}$ mg and Aa equal to $a\theta$. Also the resolved part of the tension of aP along aA will be nearly equal to

$$\frac{1}{2} mg \frac{a\theta}{b} = \frac{mag\theta}{2b},$$

and its moment about G will be nearly equal to

$$\frac{ma^{3}g\theta}{2h}$$
;

similarly for the tension of the string bQ: hence for the angular motion of ab about G, taking into account the tensions of both the strings,

$$mk^3\frac{d^2\theta}{dt^2}=-\frac{ma^2g\theta}{b};$$

but $k^2 = \frac{1}{3}a^2$; hence

$$\frac{d^2\theta}{d\ell^2}=-\frac{3g}{h}\theta;$$

multiplying by 2 $\frac{d\theta}{dt}$, and integrating, we have

$$\frac{d\theta^2}{dt^2} = C - \frac{3g}{h} \theta^2 = \frac{3g}{h} (a^2 - \theta^2),$$

supposing a to be the greatest value of θ .

Hence the time of an oscillation will be equal to

$$-\left(\frac{b}{3g}\right)^{\frac{1}{4}}\int_{-\infty}^{-\infty}\frac{d\theta}{(\alpha^2-\theta^2)^{\frac{1}{4}}}=\left(\frac{b}{3g}\right)^{\frac{1}{4}}\pi;$$

which is that of a simple pendulum, of which the length is $\frac{1}{3}b$, and is independent of the length of the rod.

Lady's and Gentleman's Diary, 1842, p. 51.

(3) A heterogeneous sphere is placed upon a perfectly smooth horizontal plane, its centre of gravity being slightly distant from the vertical through its geometrical centre; to find the time of the small oscillation of the centre of gravity about the geometrical centre.

Let ABS (fig. 164) be a vertical section of the sphere passing through C its geometrical centre and G its centre of gravity. Draw GQ horizontal, intersecting the vertical line SCK, through the point of contact of the sphere and the horizontal plane, in the point Q; draw GM vertical to cut the horizontal plane in M; and let CGA be the radius through G. Let $\angle AGM = \phi = \angle ACS$; CG = c, MG = y, m = the mass of the sphere, k = the radius of gyration about G, R = the reaction of the plane at S upon the sphere, which will exert itself vertically.

Then for the motion of the sphere we have, resolving forces vertically,

$$m\frac{d^3y}{dt^3}=R-mg\ldots(1),$$

and, taking moments about G,

$$mk^2 \frac{d^2\phi}{dt^2} = -Rc \sin \phi \dots (2);$$

but $y = a - c \cos \phi$, and therefore, from (1),

$$R = mg - mc \frac{d^2 \cos \varphi}{dt^2};$$

hence from (2) we have

$$mk^2 \frac{d^2\phi}{dt^2} = -mcg \sin \phi + mc^2 \sin \phi \frac{d^2 \cos \phi}{dt^2},$$

and therefore

$$(k^2 + c^2 \sin^2 \phi) \frac{d^2 \phi}{dt^2} + c^2 \sin \phi \cos \phi \frac{d\phi^2}{dt^2} = -cg \sin \phi;$$

multiplying this equation by $2\frac{d\phi}{dt}$, and integrating, we get

$$(k^2 + c^2 \sin^2 \phi) \frac{d\phi^2}{dt^2} = 2cg \cos \phi + C;$$

1

suppose a to be the initial value of ϕ , then, $\frac{d\phi}{dt}$ being supposed to be initially zero, we have

$$0 = 2cg \cos a + C,$$

and therefore

$$(k^2 + c^2 \sin^2 \phi) \frac{d\phi^2}{dt^2} = 2cg (\cos \phi - \cos \alpha), \dots (3),$$

whence, $\frac{dt}{d\phi}$ being considered negative because as t increases $d\phi$ is negative from the beginning to the end of every complete oscillation,

 $\frac{dt}{d\phi} = -\frac{1}{(2cq)^{\frac{1}{4}}} \frac{(k^2 + c^3 \sin^2 \phi)^{\frac{1}{4}}}{(\cos \phi - \cos \alpha)^{\frac{1}{4}}};$

now from (3) we see that when $\frac{d\phi}{dt} = 0$, $\cos \phi = \cos a$, and therefore $\phi = \pm a$, the positive value of ϕ corresponding to the beginning and the negative value to the end of a complete oscillation; hence, if T denote the time of a complete oscillation,

$$T = -\frac{1}{(2cg)^{\frac{1}{4}}} \int_{-a}^{-a} d\phi \frac{(k^{2} + c^{2} \sin^{2} \phi)^{\frac{1}{4}}}{(\cos \phi - \cos a)^{\frac{1}{4}}}$$
$$= \frac{1}{(2cg)^{\frac{1}{4}}} \int_{-a}^{-a} d\phi \frac{(k^{2} + c^{2} \sin^{2} \phi)^{\frac{1}{4}}}{(\cos \phi - \cos a)^{\frac{1}{4}}}.$$

Assume

$$\sin\frac{\phi}{2}=s, \quad \sin\frac{a}{2}=b;$$

then

$$\frac{1}{2}\cos\frac{\phi}{2}\,d\phi=ds,\quad d\phi=\frac{2ds}{(1-s^2)^{\frac{1}{4}}},$$

 $\cos \phi - \cos \alpha = 2(b^2 - s^2), \sin^2 \phi = 4s^2(1 - s^2);$

hence we have

$$T = \frac{1}{(cg)^{\frac{1}{4}}} \int_{-s}^{-s} ds \, \frac{(1-s^2)^{-\frac{1}{4}} \left\{ k^2 + 4c^3s^2 (1-s^2) \right\}^{\frac{1}{4}}}{(b^2-s^2)^{\frac{1}{4}}},$$

or, considering powers of the small quantity s beyond the second to be negligible,

$$= \frac{k}{(cg)^{\frac{1}{4}}} \int_{-\delta}^{4\delta} ds \frac{1 + \frac{1}{2} s^2 + \frac{2c^2}{k^2} s^3}{(b^3 - s^2)^{\frac{1}{4}}}$$

$$= \frac{k}{(cg)^{\frac{1}{4}}} \int_{-\delta}^{4\delta} ds \left\{ \frac{1}{(b^3 - s^2)^{\frac{1}{4}}} + \frac{4c^3 + k^3}{2k^3} \frac{s^3}{(b^3 - s^2)^{\frac{1}{4}}} \right\};$$

$$\int_{-\delta}^{4\delta} \frac{ds}{(b^3 - s^2)^{\frac{1}{4}}} = \pi, \int_{-\delta}^{4\delta} \frac{s^3 ds}{(b^3 - s^2)^{\frac{1}{4}}} = \frac{1}{2}\pi b^3;$$

but

hence we have, for the time of a complete oscillation,

$$T = \frac{\pi k}{(cg)^{\frac{1}{4}}} \left(1 + \frac{4c^3 + k^3}{4k^3} b^3 \right) = \frac{\pi k}{(cg)^{\frac{1}{4}}} \left(1 + \frac{4c^3 + k^3}{4k^3} \sin^2 \frac{a}{2} \right).$$

Euler; Nova Acta Acad. Petrop. 1783; p. 119.

(4) A beam ΔB (fig. 165) is placed with one end Δ upon a smooth inclined plane EF; to find the motion of the beam and its pressure on the plane at any time.

Let G be the position of the centre of gravity of the beam at any time t from the commencement of the motion, R = the reaction of the plane upon the extremity A, $\angle BAF = \phi$; let E be the initial position of A, and β the initial value of ϕ ; m = the mass of the beam, k = its radius of gyration about G, EA = z, EH = x, GH = y, a = the inclination of FE to the horizon.

Then for the motion of the beam we have, resolving forces parallel to the plane,

$$m\frac{d^2x}{dt^2}=mg\sin a\cdot\ldots(1);$$

resolving forces at right angles to the plane,

$$m \frac{d^3y}{dt^2} = R - mg \cos a \dots (2);$$

and, taking moments about G,

$$mk^2\frac{d^2\phi}{dt^2}=-Ra\cos\phi....(3).$$

From (1) we get

$$\frac{dx}{dt} = gt \sin a + C;$$

but $\frac{dx}{dt} = 0$ when t = 0; and therefore C = 0; hence

$$\frac{dx}{dt} = gt \sin a;$$

integrating, and observing that $x = a \cos \beta$ when t = 0, we have $x = \frac{1}{2}gt^2 \sin a + a \cos \beta \dots (4)$,

which gives the position of the point H at any assigned time from the commencement of the motion.

Again, from (2) and (3), by the elimination of R,

$$a \cos \phi \frac{d^3y}{dt^3} = -k^3 \frac{d^3\phi}{dt^6} - ag \cos a \cos \phi;$$

but, by the geometry, we see that $y = a \sin \phi$; hence

$$a^2\cos\phi\,\frac{d^2\sin\phi}{dt^2}=-\,k^3\,\frac{d^2\phi}{dt^2}-ag\,\cos\alpha\cos\phi\,;$$

and therefore, multiplying by $2 \frac{d\phi}{dt}$, and integrating,

$$a^{2}\left(\frac{d\sin\phi}{dt}\right)^{2}=C-k^{2}\frac{d\phi^{2}}{dt^{2}}-2ag\cos\alpha\sin\phi;$$

but, initially, $\frac{d\phi}{dt} = 0$ and $\phi = \beta$; hence there is

$$0 = C - 2ag \cos a \sin \beta,$$

and therefore

and

$$(a^2\cos^2\phi+k^2)\frac{d\phi^2}{dt^2}=2ag\cos\alpha\left(\sin\beta-\sin\phi\right)\ldots..(5),$$

which gives the angular velocity of the beam for every position which it can assume during its descent.

From the geometry it is evident that

$$z = x - a \cos \phi$$

$$= \frac{1}{2} g t^{2} \sin a + a (\cos \beta - \cos \phi), \text{ by (4)},$$

$$y = a \sin \phi;$$

if therefore from (5) we could obtain ϕ in terms of t, we might determine the values of y and z at any time from the beginning of the motion.

Again, for the pressure on the plane at any time, we have, from (3),

$$R = -\frac{mk^3}{a\cos\phi} \frac{d^3\phi}{dt^2};$$

but, from (5),

$$\frac{d\phi^2}{dt^2} = 2ag \cos \alpha \frac{\sin \beta - \sin \phi}{a^2 \cos^2 \phi + k^2},$$

and therefore, differentiating with respect to t, and dividing by $2 \frac{d\phi}{dt} \cos \phi$,

$$\frac{1}{\cos \phi} \frac{d^3 \phi}{dt^2} = \frac{2a^3 g \cos a \sin \phi (\sin \beta - \sin \phi)}{(a^3 \cos^3 \phi + k^3)^3} - \frac{ga \cos a}{a^3 \cos^3 \phi + k^2};$$
hence
$$R = \frac{mk^2 g \cos a}{a^3 \cos^2 \phi + k^3} - \frac{2ma^3 k^2 g \cos a \sin \phi (\sin \beta - \sin \phi)}{(a^3 \cos^3 \phi + k^3)^3}$$

$$= \frac{mk^2 g \cos a}{(a^3 \cos^2 \phi + k^3)^3} \left\{ k^2 + a^3 (1 + \sin^2 \phi - 2 \sin \beta \sin \phi) \right\},$$

which gives the pressure on the plane for any of the successive positions of the beam.

Fuss; Nova Acta Petrop. 1795; p. 70.

(5) A cylinder KLM, (fig. 166), is placed with its axis horizontal upon a smooth inclined plane; a string EPMKL, one end E of which is attached to a fixed point at a distance EA from the plane equal to the radius of the cylinder, having been wound about the cylinder in a vertical plane through the centre of gravity O of the cylinder at right angles to its axis; to find the tension of the string and the velocity of decrease of its angle of inclination to the plane corresponding to any position of the cylinder in its descent; the length of the free string being initially equal to zero.

Let M be the point of contact of the section KLM of the cylinder, about which the string is wound, with the inclined plane; and P the point in which the free string EP touches the cylinder. Produce EP to meet the inclined plane in S; join OP, OM; at any time t from the commencement of the motion let AM = x, T = the tension of the string, $\angle ESA = \angle POM = \theta$, $\phi =$ the whole angle through which the cylinder has revolved

about its centre of gravity; also, let m = the mass of the cylinder, k = its radius of gyration about its axis, a = the inclination of the plane to the horizon, and AE = MO = a.

Then for the motion of the cylinder we have, resolving forces parallel to the plane,

$$m\frac{d^2x}{dt^2}=mg\sin\alpha-T\cos\theta\ldots(1);$$

and taking moments about O, the centre of gravity,

$$mk^2 \frac{d^2\phi}{dt^2} = Ta \dots (2):$$

by the elimination of T between these two equations, we get

$$a\frac{d^2x}{dt^2}=ag\sin a-k^2\cos\theta\frac{d^2\phi}{dt^2}\ldots(3).$$

Take along EA, produced if necessary, Ep equal to EP: then, if the cylinder were made to roll from E to p, and then Ep were made to revolve about E into the position EP, the cylinder would clearly on the whole have revolved about its centre of gravity through the very angle which actually belongs to its real motion in setting free the length EP of the string. Now, in the first stage of the hypothetical motion, the cylinder will obviously move through an angle equal to $\frac{Ep}{a}$ or $\frac{EP}{a}$, which is equal to $\cot \theta$; and, in the second stage, through an angle $pES = \frac{1}{2}\pi - \theta$, in an opposite direction. Hence clearly we have

$$\phi = \cot \theta - (\frac{1}{2}\pi - \theta) = \cot \theta + \theta - \frac{1}{2}\pi \dots (4).$$

Also, from the geometry, it is obvious that

From (4) and (5) we have

$$d\phi = -\frac{d\theta}{\sin^2\theta} + d\theta = -\frac{\cos^2\theta}{\sin^2\theta} d\theta \dots (6),$$

$$dx = -\frac{a\cos\theta}{\sin^2\theta} d\theta. \dots (7).$$

Multiplying (8) by $2\frac{dx}{dt}$, we get

$$2a\frac{dx}{dt}\frac{d^3x}{dt^2} = 2ag\sin a\frac{dx}{dt} - 2k^2\frac{dx}{dt}\cos\theta\frac{d^3\phi}{dt^2};$$

but from (6) and (7) it is clear that

$$\cos\theta\,\frac{dx}{dt}=a\,\frac{d\phi}{dt}\ldots\ldots(8);$$

hence we obtain

$$2\frac{dx}{dt}\frac{d^3x}{dt^2} = 2g\sin a\frac{dx}{dt} - 2k^3\frac{d\phi}{dt}\frac{d^3\phi}{dt^2};$$

integrating, and adding the arbitrary constant C,

$$\frac{dx^3}{dt^2} + k^3 \frac{d\phi^2}{dt^3} = 2gx \sin \alpha + C;$$

but, initially, $\frac{dx}{dt} = 0$, $\frac{d\phi}{dt} = 0$, x = a; hence

$$0 = 2ga \sin a + C,$$

and therefore

$$\frac{dx^2}{dt^2} + k^2 \frac{d\phi^2}{dt^2} = 2g \sin \alpha (x - \alpha);$$

substituting in this equation the values of x, $d\phi$, dx, given in (5), (6), (7), we have

$$\left(\frac{a^2\cos^2\theta}{\sin^4\theta} + \frac{k^2\cos^4\theta}{\sin^4\theta}\right)\frac{d\theta^2}{dt^2} = 2g\sin\alpha\left(\frac{a}{\sin\theta} - a\right),$$

and therefore

$$\frac{d\theta^2}{dt^2} = \frac{2ga \sin a (1 - \sin \theta) \sin^3 \theta}{\cos^2 \theta (a^2 + k^2 \cos^2 \theta)} \dots (9);$$

which gives the angular velocity of the string about E in terms of its inclination to the plane, or for any position of the cylinder.

Again, from (8), we have

$$a \frac{d^2 \phi}{dt^2} = \cos \theta \frac{d^2 x}{dt^2} - \sin \theta \frac{dx}{dt} \frac{d\theta}{dt},$$

and therefore

$$\frac{d^2x}{dt^2} = \frac{a}{\cos\theta} \frac{d^2\phi}{dt^2} + \tan\theta \frac{dx}{dt} \frac{d\theta}{dt}$$
$$= \frac{a}{\cos\theta} \frac{d^2\phi}{dt^2} - \frac{a}{\sin\theta} \frac{d\theta^2}{dt^2}, \text{ by (7)};$$

substituting this value of $\frac{d^3x}{dt^2}$ in (3), we obtain

$$\left(\frac{a}{\cos\theta} + \frac{k^2}{a}\cos\theta\right)\frac{d^2\phi}{dt^2} = g\sin\alpha + \frac{a}{\sin\theta}\frac{d\theta^2}{dt^2},$$

and therefore, by (9),

$$\frac{a^{2} + k^{2} \cos^{2} \theta}{a \cos \theta} \frac{d^{2} \phi}{dt^{2}} = g \sin a + \frac{2ga^{2} \sin a \sin^{2} \theta (1 - \sin \theta)}{\cos^{2} \theta (a^{2} + k^{2} \cos^{2} \theta)}$$

$$= g \sin a \frac{a^{2} (1 + \sin^{2} \theta - 2 \sin^{3} \theta) + k^{2} \cos^{4} \theta}{\cos^{2} \theta (a^{2} + k^{2} \cos^{2} \theta)};$$

hence, by (2), we have for the tension of the string for any position of the cylinder,

$$T = \frac{mk^3}{a} \frac{d^2\phi}{dt^2} = mk^2g \sin a \frac{a^2(1 + \sin^2\theta - 2\sin^3\theta) + k^2\cos^4\theta}{\cos\theta (a^2 + k^2\cos^2\theta)^3}.$$

Euler; Nova Acta Acad. Petrop. 1795; p. 64.

(6) A uniform heavy rod OA, (fig. 167), which is at liberty to oscillate in a vertical plane about a horizontal axis through O, falls from a horizontal position; to determine the angle included between the direction of the rod and the direction of the pressure upon the fixed axis, for any position of the rod.

From O draw Om at right angles to OA and to the fixed axis; and produce AO indefinitely to a point n. Let R, S, denote the resolved parts of the reaction of the fixed axis along Om, On, for any position of the rod. Draw Ox horizontal and at right angles to the fixed axis. Let OA = a; m = the mass of the rod; $\angle AOx = \theta$, at any time t. Then, for the motion of the rod about its centre of gravity G, the moment of inertia about G being $\frac{1}{12}ma^2$,

$$\frac{1}{12}ma^3\frac{d^2\theta}{dt^2} = \frac{1}{2}aR,$$

$$ma\frac{d^2\theta}{dt^2} = 6R \dots (1).$$

Also, for the motion about O, the moment of inertia about O being $\frac{1}{3}ma^2$, $\frac{1}{3}ma^2\frac{d^2\theta}{dt^2} = mg \cdot \frac{1}{2}a\cos\theta,$ $2a\frac{d^2\theta}{dt^2} = 3g\cos\theta \cdot \dots (2).$

Again, equating S to the resolved part of the weight along OA and the centrifugal force,

$$S = mg \sin \theta + \int_{0}^{a} \left(\frac{m \, dr}{a} \, r \, \frac{d\theta^{2}}{dt^{2}} \right)$$

$$= mg \sin \theta + \frac{1}{2} ma \, \frac{d\theta^{2}}{dt^{2}} \dots (4).$$

Again, multiplying (2) by $\frac{d\theta}{dt}$, and integrating,

$$a \frac{d\theta^2}{dt^2} = C + 3g \sin \theta;$$

but $\frac{d\theta}{dt} = 0$, when $\theta = 0$; hence C = 0, and therefore

$$a\,\frac{d\theta^2}{dt^2}=3g\,\sin\,\theta\,;$$

hence, from (4), we get

$$S = mg \sin \theta + \frac{3}{2} mg \sin \theta = \frac{5}{2} mg \sin \theta \dots (5).$$

Let ϕ be the angle which the whole reaction of the fixed axis makes with the line On; then

$$\tan \phi = \frac{R}{S}.$$

and therefore, by the equations (4) and (5),

$$\tan \theta \tan \phi = \frac{1}{10}$$
,

which gives the value of ϕ for any position of the rod; ϕ is evidently the angle between the direction of the whole pressure on the fixed axis and the length OA of the rod.

A solution of this problem was given in Chap. vi., by the direct application of D'Alembert's Principle.

(7) A uniform rod, acted on by gravity, is oscillating in a vertical plane about one extremity; to find the tendency of the vis inertia in any position to bend the rod at any point, and to ascertain the point at which this tendency is a maximum.

Let OA (fig. 168) be the position of the rod at any time t; Ox an indefinite horizontal line through O, the fixed extremity of the rod, in the vertical plane through OA. Take C any

point in OA, P any point in CA. Let OA = 2a, OC = c, OP = r, $\angle AOx = \theta$, k =the radius of gyration about O; m =the mass of the rod.

Then the force gained by an element dr of the rod at the point P, resolved at right angles to OP, will be equal to

$$m\,\frac{dr}{2a}\bigg(r\,\frac{d^3\theta}{dt^2}-g\,\cos\,\theta\bigg);$$

and the moment of this about C will be equal to

$$\frac{m}{2a}\left(r\,\frac{d^2\theta}{d\ell^2}-g\,\cos\,\theta\right)(r-c)\,dr;$$

hence the whole moment to produce bending at C will be equal to $\frac{m}{2a} \int_{c}^{2a} \left(r \frac{d^{2}\theta}{dt^{2}} - g \cos \theta \right) (r - c) dr \dots (1).$

But, for the motion of the rod, we have

$$mk^3 \frac{d^2\theta}{dt^2} = mga \cos \theta,$$

and therefore, $\frac{4}{3}a^2$ being the value of k^2 ,

$$\frac{d^2\theta}{dt^2} = \frac{3g}{4a}\cos\theta.$$

Hence the expression (1) becomes

$$\frac{mg \cos \theta}{8a^3} \int_{c}^{2a} (3r - 4a) (r - c) dr$$

$$= \frac{mg \cos \theta}{8a^3} \int_{c}^{2a} \left\{ 3 (r - c) - (4a - 3c) \right\} (r - c) dr$$

$$= \frac{mg \cos \theta}{8a^3} \left\{ (2a - c)^3 - \frac{1}{2} (4a - 3c) (2a - c)^2 \right\}$$

$$= \frac{mg \cos \theta}{8a^3} \left(2a - c \right)^2 \left\{ 2a - c - \frac{1}{2} (4a - 3c) \right\}$$

$$= \frac{mg \cos \theta}{16a^3} c (2a - c)^2.$$

When this expression is a maximum, we have

$$(2a-c)^2-2c(2a-c)=0,$$

 $2a-3c=0, c=\frac{2}{3}a,$
 $OC=\frac{1}{3}OA.$

or

(8) An angular velocity having been impressed apon a heterogeneous sphere, about an axis, perpendicular to the vertical plane which contains its centre of gravity G and its geometrical centre C, and passing through G (fig. 164), it is then placed upon a smooth horizontal plane; to determine the magnitude of the impressed angular velocity that G may rise into a point in the vertical line SCK through C, and there rest; the initial magnitude of the angle between CG and the vertical radius CS being given.

Let CG = c, k = the radius of gyration about G, a = the initial value of the angle GCS, and w = the required angular velocity: then will be determined by the equation

$$(k^2 + c^2 \sin^2 a) \omega^2 = 2cg (1 + \cos a).$$

Euler; Nova Acta Acad. Petrop. 1783; p. 119.

(9) A chain, ten yards long, consisting of indefinitely small equal links, being laid straight on a perfectly smooth horizontal plane, except one part, a yard in length, which hangs down perpendicularly below the plane; in what time will the chain entirely quit the plane?

Time = 2.890663 seconds nearly. Lady's and Gentleman's Diary, 1758, Diarian Repository, p. 683.

SECT. 2. Several Bodies.

(1) A wheel and axle is loaded with given weights P and Q, (fig. 169), which are not in equilibrium; to determine their motion and the tension of the strings by which the weights are suspended.

Through C, the centre of the wheel and axle, draw the horizontal line ACB meeting the strings in A and B; let AC = a, BC = a'; m =the mass of P, m' =that of Q, $\mu =$ that of the wheel and axle together; k = the radius of gyration of the wheel and axle about their common axis; AP = x, BQ = x', T = the tension of AP, T = the tension of BQ; θ = the angle through which the wheel and axle have revolved at the end of the time t about their common axis.

Then, for the motion of P, we have

$$m\frac{d^2x}{dt^2}=mg-T....(1),$$

for the motion of Q,

$$m'\frac{d^3x'}{dt^2}=m'g-T'\ldots(2);$$

and, for the rotation of the wheel and axle,

$$\mu k^2 \frac{d^2 \theta}{dt^2} = Ta - Ta' \dots (3).$$

But, from the geometry, it is clear that

$$\frac{dx}{dt} = a \frac{d\theta}{dt}, \quad \frac{dx'}{dt} = -a' \frac{d\theta}{dt};$$

hence, from (1) and (2),

$$ma\frac{d^2\theta}{dt^2}=mg-T....(4),$$

$$-m'a'\frac{d^3\theta}{dt^3}=m'g-T'\ldots\ldots\ldots\ldots\ldots\ldots(5).$$

Substituting the values of T and T from (4) and (5) in the equation (3), we get

$$\mu k^2 \frac{d^2\theta}{dt^2} = ma \left(g - a \frac{d^2\theta}{dt^2} \right) - m'a' \left(g + a' \frac{d^2\theta}{dt^2} \right),$$

$$\left(ma^2 + m'a'^2 + \mu k^2 \right) \frac{d^2\theta}{dt^2} = g \left(ma - m'a' \right) \dots (6);$$

whence θ is immediately obtained in terms of t, the initial values of θ and $\frac{d\theta}{dt}$ being supposed to be known.

From (4) and (6) we have

$$T = mg - \frac{mag (ma - m'a')}{ma^2 + m'a'^2 + \mu k^2};$$

and from (5), (6),

$$T' = m'g + \frac{m'a'g (ma - m'a')}{ma^2 + m'a'^2 + \mu k^2}.$$

(2) Two equal uniform rods AC, BC, (fig. 170), having a compass joint at C, are laid in a line upon a horizontal plane. A string CDP having a given weight P at one end passes over a smooth pin D above the plane, and has its other end fastened to C which is vertically beneath the pin; to determine the motion when P descends.

Let AC = 2a = BC, $\angle CAB = \theta$; R = the vertical reaction of the plane at each of the points A and B; T = the tension of the string; S = the mutual action of the two rods at the joint, which will evidently take place in a horizontal line parallel to AB; m = the mass of each of the rods, μ = the mass of the weight P. Let G be the centre of gravity of the rod AC; draw GH, CE, at right angles to AB; let EH = x, GH = y, k = the radius of gyration of AC about G.

Then, for the motion of the rod AC, we have, resolving forces horizontally,

$$m\frac{d^3x}{dt^2}=S\ldots\ldots(1);$$

resolving vertically, $\frac{1}{2}T$ being the force exerted by the string on each rod, $m\frac{d^2y}{dt^2} = R + \frac{1}{2}T - mg \dots (2);$

and, taking moments about G,

$$mk^2 \frac{d^2\theta}{dt^2} = Sa \sin \theta + \frac{1}{2} Ta \cos \theta - Ra \cos \theta \dots (3).$$

Also, for the motion of P, the increment of DP being double that of GH,

$$2\mu \frac{d^2y}{dt^2} = \mu g - T \dots (4).$$

Multiplying the equation (2) by $a \cos \theta$, we have

ma cos
$$\theta\left(\frac{d^2y}{dt^2}+g\right)=(R+\frac{1}{2}T)$$
 a cos θ ,

and therefore, adding this equation to the equation (3),

$$mk^2 \frac{d^2\theta}{dt^2} + ma \cos \theta \left(\frac{d^2y}{dt^2} + g\right) = Sa \sin \theta + Ta \cos \theta$$
;

hence, from (1) and (4),

$$mk^2\frac{d^2\theta}{dt^2} + ma\cos\theta\left(\frac{d^2y}{dt^2} + g\right) = ma\sin\theta\frac{d^2x}{dt^2} + \mu a\cos\theta\left(g - 2\frac{d^2y}{dt^2}\right);$$

and therefore, since $x = a \cos \theta$, and $y = a \sin \theta$,

$$mk^{2} \frac{d^{2}\theta}{dt^{2}} + ma^{2} \left(\cos\theta \frac{d^{2}\sin\theta}{dt^{2}} - \sin\theta \frac{d^{2}\cos\theta}{dt^{2}}\right) + 2\mu a^{2}\cos\theta \frac{d^{2}\sin\theta}{dt^{2}}$$
$$= ag(\mu - m)\cos\theta,$$

$$(ma^2 + mk^2)\frac{d^2\theta}{dt^2} + 2\mu a^2 \cos\theta \frac{d^2 \sin\theta}{dt^2} = ag(\mu - m)\cos\theta.$$

Multiplying both sides of this equation by $2 \frac{d\theta}{dt}$, and integrating,

we have, since
$$\frac{d\theta}{dt} = 0$$
 when $\theta = 0$,

$$(ma^2 + mk^2 + 2\mu a^2 \cos^2 \theta) \frac{d\theta^2}{dt^2} = 2ag(\mu - m) \sin \theta,$$

which determines the angular velocity of the rods for any position.

The value of $\frac{d\theta}{dt}$ and therefore of $\frac{d^2\theta}{dt^2}$ being known in terms of θ , we may readily obtain the values of R, S, and T, from the equations (1), (2), (4), in terms of the same angle.

(3) A tube, moveable in a horizontal plane about a vertical axis, is charged with any number of balls at assigned intervals; supposing a given angular velocity to be communicated to the tube, it is required to determine the motion of the tube and of the balls.

Let a, a', a', \dots be the initial distances of the balls from the fixed axis, and r, r', r', \dots their distances at any time t from the commencement of the motion. Let m, m', m', \dots be the masses of the balls, μ of the tube; and let θ be the angle through which the tube has revolved at the end of the time t. Let R, R', R', \dots denote the mutual actions and reactions of the balls and the tube. Then, μk^2 denoting the moment of inertia of the tube about the vertical axis, we shall have, for the motion of the tube,

Also, for the motion of the balls, m, m', m', \dots we have

$$m \frac{d^{2}x}{dt^{2}} = R \frac{y}{r}, \qquad m \frac{d^{2}y}{dt^{2}} = -R \frac{x}{r} \dots (2),$$

$$m' \frac{d^{2}x'}{dt^{2}} = R' \frac{y'}{r'}, \qquad m' \frac{d^{2}y'}{dt^{2}} = -R' \frac{x'}{r'} \dots (3),$$

$$m' \frac{d^{2}x''}{dt^{2}} = R' \frac{y'}{r'}, \qquad m' \frac{d^{2}y''}{dt^{2}} = -R' \frac{x'}{r} \dots (4),$$

$$\vdots \qquad \vdots \qquad \vdots \qquad \vdots$$

where (x, y), (x', y'), (x', y'), ... are the rectangular co-ordinates of the balls at the time t.

Multiplying the former and the latter of the equations (2) by y and x respectively, and subtracting the latter from the former of the resulting equations, we get

$$m\left(y\,\frac{d^2x}{dt^2}-x\,\frac{d^2y}{dt^2}\right)=Rr;$$

and therefore, since $x = r \cos \theta$, $y = r \sin \theta$, the axis of x being supposed to coincide with the initial position of the tube, we may readily obtain, by substitution,

$$m \frac{d}{dt} \left(r^2 \frac{d\theta}{dt} \right) = -Rr.$$

In like manner, from the equations of (3), (4),...we may get

$$m' \frac{d}{dt} \left(r'^2 \frac{d\theta}{dt} \right) = -R'r',$$

$$m' \frac{d}{dt} \left(r'^2 \frac{d\theta}{dt} \right) = -R'r',$$

Hence from (1) we have

$$\mu k^2 \frac{d^2\theta}{dt^2} = -\frac{d}{dt} \left(mr^2 \frac{d\theta}{dt} + m'r'^2 \frac{d\theta}{dt} + m'r''^2 \frac{d\theta}{dt} + \dots \right);$$

integrating, we get

$$\mu k^{2} \frac{d\theta}{dt} = C - (mr^{2} + m'r'^{2} + m''r''^{2} + \cdots) \frac{d\theta}{dt};$$

but, supposing ω to be the initial value of $\frac{d\theta}{dt}$, we have

$$\mu k^2 \omega = C - (ma^2 + m'a'^2 + m''a''^2 + ...) \omega$$
;

hence

$$\frac{d\theta}{dt} = \frac{\mu k^2 + ma^2 + m'a'^2 + m'a'^2 + \dots}{\mu k^2 + mr^2 + m'r'^2 + m'r'^2 + \dots} \omega \dots (5).$$

Again, from the equations (2), we have

$$x\frac{d^2x}{dt^2}+y\frac{d^2y}{dt^2}=0,$$

and thence, substituting for x and y their values in r and θ ,

$$\frac{d^2r}{dt^2}=r\frac{d\theta^2}{dt^2}\ldots\ldots(6).$$

In the same way, from (3), we may get

$$\frac{d^3r'}{dt^2}=r'\,\frac{d\theta^2}{dt^2};$$

and therefore, eliminating $\frac{d\theta}{dt}$ between these two equations,

$$r'\frac{d^3r}{dt^2}=r\frac{d^3r'}{dt^2};$$

integrating and bearing in mind that both $\frac{dr}{dt}$ and $\frac{dr'}{dt}$ are initially equal to zero,

$$r'\frac{dr}{dt} = r\frac{dr'}{dt},$$

$$\frac{dr}{dt} = \frac{dr'}{r'};$$

and therefore

integrating again we have, a, a', being the initial values of r, r',

$$\frac{r}{a} = \frac{r'}{a'}, \quad r' = \frac{a'}{a} r.$$

In precisely the same way it may be shown that

$$r' = \frac{a''}{a} r, \quad r''' = \frac{a'''}{a} r, \quad \dots$$

Hence from (5) we have

$$\frac{d\theta}{dt} = \frac{\mu k^2 + ma^2 + m'a'^2 + m'a'^2 + \dots}{\mu k^2 + \left(ma^2 + m'a'^2 + m'a'^2 + \dots\right)\frac{r^2}{a^2}} \omega \dots (7).$$

From (6) and (7) we obtain

$$\frac{d^2r}{dt^2} = \omega^2 r \left\{ \frac{\mu k^2 + ma^2 + m'a'^2 + m'a'^2 + \dots + \frac{r^2}{a^2}}{\mu k^2 + \left(ma^2 + m'a'^2 + m'a'^2 + \dots\right)\frac{r^2}{a^2}} \right\}.$$

Multiplying both sides of this equation by $2\frac{dr}{dt}$, integrating, and bearing in mind that $\frac{dr}{dt} = 0$ when r = a, we shall easily see that

$$\frac{dr^2}{dt^2} = \omega^3 (r^2 - a^3) \frac{\mu k^3 + ma^3 + m'a'^2 + m''a''^2 + \dots}{\mu k^3 + (ma^3 + m'a'' + m''a''^2 + \dots) \frac{r^2}{a^3}} \dots (8).$$

The equations (7) and (8) will give us, for any assigned distance of the ball m from the axis of rotation, the angular velocity of the tube and the velocity of the ball m within it. If between (7) and (8) we eliminate dt, we shall obtain the differential equation in polar co-ordinates to the path of m in the horizontal plane passing through the axis of the tube. Similar results may evidently be obtained for the other balls with which the tube is charged.

Cor. If $\mu = 0$, the equations (7) and (8) become

$$\frac{d\theta}{dt} = \frac{a^3\omega}{r^3}, \quad \frac{dr^3}{dt^3} = (r^3 - a^3) \frac{a^3\omega^3}{r^3},$$

and therefore, eliminating dt,

$$\frac{dr^3}{d\theta^2} = (r^3 - a^3) \frac{r^3}{a^3}, \quad d\theta = \frac{adr}{r (r^3 - a^3)^{\frac{1}{2}}} = -\frac{d^2 \frac{a}{r}}{\left(1 - \frac{a^3}{r^3}\right)^{\frac{1}{2}}}$$

integrating, and remembering that $\theta = 0$ when r = a,

$$\theta = \cos^{-1}\frac{a}{r}, \quad \frac{a}{r} = \cos \theta.$$

Again, to determine the relation between r and t, we have

$$dt = \frac{1}{\omega a} \frac{rdr}{(r^2 - a^2)^{\frac{1}{4}}}, \quad t = \frac{1}{\omega a} (r^2 - a^2)^{\frac{1}{4}},$$

$$e \frac{r^2}{a} = 1 + \omega^2 t^2.$$

and therefore

Similar relations holding good for the other balls, we have, for the equations to their paths,

$$\frac{a}{r} = \frac{a'}{r'} = \frac{a''}{r''} = \dots = \cos \theta,$$

which shew that they all move in straight lines at right angles to the initial position of the tube; and for their distances from the axis of rotation at any time,

$$\frac{r^2}{a^2} = \frac{r'^2}{a'^3} = \frac{r'^2}{a'^2} = \ldots = 1 + \omega^2 t^2.$$

Clairaut; Mém. de l'Acad. des Sciences de Paris, 1742, p. 48. Daniel Bernoulli; Mém. de l'Acad. des Sciences de Berlin, 1745, p. 54. Euler; Opuscula, de motu corporum tubis mobilibus inclusorum, p. 71.

(4) A heavy particle P descends down a smooth inclined plane BA, (fig. 171), forming the upper surface of a solid BAC, which is capable of sliding freely along a smooth horizontal plane OAx; to determine the motion of the particle and of the body, both of which are supposed to have initially no motion.

Let PM be at right angles to Ox, and let B be the point in the inclined plane which the particle occupies initially; let A be supposed to coincide with O at the commencement of the motion. Let OM = x, PM = y, OA = s, AB = a, BP = s', $\triangle BAC = a$; and R = the action and reaction of the plane and the particle. Then, m denoting the mass of the particle and m' of the body, we shall have

$$m \frac{d^3y}{dt^3} = R \cos a - mg \dots (1),$$

$$m \frac{d^3x}{dt^3} = -R \sin a \dots (2),$$

$$m' \frac{d^3s}{dt^3} = R \sin a \dots (3).$$

But $y = (a - s') \sin a$, $x = s + (a - s') \cos a$; hence, from (1),

$$m \sin a \frac{d^3s'}{dt'} = mg - R \cos a \dots (4),$$

and, from (2),

$$m\frac{d^2s}{dt^2}-m\cos a\frac{d^2s'}{dt^2}=-R\sin a\ldots (5).$$

Adding together (3) and (5), we have

$$(m+m')\frac{d^2s}{dt^2}-m\cos\alpha\frac{d^2s'}{dt^2}=0....(6).$$

Multiplying (3) by cos a and (4) by sin a, we have, adding together the resulting equations,

$$m'\cos\alpha\frac{d^2s}{dt^2}+m\sin^2\alpha\frac{d^2s'}{dt^2}=mg\sin\alpha\ldots$$
 (7).

Multiplying (6) by $\sin^2 \alpha$, (7) by $\cos \alpha$, and adding together the resulting equations,

$$(m \sin^2 a + m') \frac{d^2s}{dt^2} = mg \sin a \cos a;$$

integrating twice with respect to t, and bearing in mind that s = 0 and $\frac{ds}{dt} = 0$ when t = 0, we obtain

$$s = \frac{1}{2}gt^2 \frac{m \sin a \cos a}{m \sin^2 a + m'} \cdot \dots (8).$$

Again, multiplying (7) by m + m', (6) by $m' \cos a$, and subtracting the latter of the resulting equations from the former, we have

$$m (m \sin^2 \alpha + m') \frac{d^2 s'}{dt^2} = mg \sin \alpha (m + m'),$$

$$\frac{d^2s'}{dt^2} = \frac{g\sin \alpha (m+m')}{m\sin^2 \alpha + m'};$$

integrating twice, and recollecting that s' = 0 and $\frac{ds'}{dt} = 0$ when t = 0, we get

$$s' = \frac{1}{2}gt^2 \frac{(m + m')\sin a}{m\sin^2 a + m'} \dots (9).$$

The equation (8) gives the position of the moveable inclined plane, and (9) the place of the particle on the plane at any time. Again, by (3),

$$R = \frac{m'}{\sin \alpha} \frac{d^2s}{dt^2} = \frac{mm'g \cos \alpha}{m \sin^2 \alpha + m'},$$

which gives the value of the mutual pressure of the particle and the plane; the value of which, therefore, is invariable.

John Bernoulli; Comment. Acad. Petrop. 1730, p. 11. Opera, tom. 111. p. 365. Euler; Opuscula, de motu corporum tubis mobilibus inclusorum, p. 28.

(5) A heavy particle is placed within a thin tube APB, (fig. 172), situated in a vertical plane, which passes through a horizontal line OE; the tube is attached rigidly to a body ABC, the lower surface of which is flat, and in contact with a smooth horizontal plane, along which it is able to slide freely;

supposing the particle and the body to be initially at rest, to find their subsequent motions.

Let A be the point of the tube which the particle occupies initially, and O the initial position of the point B of the body; let OB = s, length AP of the tube = s'; OM = x, PM = y, where PM is vertical; ϕ the inclination to the horizon of an element of the tube at P; m = the mass of the particle and m' = the mass of the body; R = the action and reaction of the tube and the particle. Then, for the motion of the particle, we have

$$m \frac{d^2x}{dt^2} = -R \sin \phi \dots (1),$$

$$m \frac{d^2y}{dt^2} = R\cos\phi - mg \dots (2);$$

and, for the motion of the body,

$$m'\frac{d^2s}{dt^2}=R\sin\phi.....(3).$$

Again, from the geometry it is evident that

$$dx = ds - \cos \phi \ ds' \dots (4),$$

and

From (1) and (3) we have

$$m \frac{d^2x}{dt^2} + m' \frac{d^2s}{dt^2} = 0 ;$$

integrating, we get

$$m\frac{dx}{dt}+m'\frac{ds}{dt}=C,$$

where C is an arbitrary constant; and therefore, by (4),

$$(m+m')\frac{ds}{dt}-m\cos\phi\frac{ds'}{dt}=C;$$

but, by the conditions of the problem, $\frac{ds}{dt} = 0$ and $\frac{ds'}{dt} = 0$ simultaneously, and therefore C = 0; hence

$$(m+m')\frac{ds}{dt}-m\cos\varphi\frac{ds'}{dt}=0.....(6).$$

Again, from (2) and (3),

$$m'\cos\phi\frac{d^2s}{dt^2}-m\sin\phi\frac{d^2y}{dt^2}=my\sin\phi,$$

and therefore, by (5),

$$m'\cos\phi \frac{d^2s}{dt^2} + m\sin\phi \frac{d}{dt}\left(\sin\phi \frac{ds'}{dt}\right) = mg\sin\phi;$$

but, from (6),
$$\frac{d^2s}{dt^2} = \frac{m}{m+m'} \frac{d}{dt} \left(\cos \phi \frac{ds'}{dt}\right);$$

hence we have

$$\frac{m'}{m+m'}\cos\phi\,\frac{d}{dt}\left(\cos\phi\,\frac{ds'}{dt}\right)+\sin\phi\,\frac{d}{dt}\left(\sin\phi\,\frac{ds'}{dt}\right)=g\,\sin\phi.$$

Multiply both sides of this equation by $2\frac{ds'}{dt}$ and integrate; then

$$\frac{m'}{m+m'}\left(\cos\phi\,\frac{ds'}{dt}\right)^2 + \left(\sin\phi\,\frac{ds'}{dt}\right)^2 = 2g\int\sin\phi\,ds',$$

$$(m\sin^2\phi + m')\,\frac{ds'^2}{dt^2} = 2g\left(m+m'\right)\int\sin\phi\,ds',$$

$$\frac{ds'}{dt} = \left\{2g\left(m+m'\right)\right\}^{\frac{1}{2}}\left\{\frac{\int\sin\phi\,ds'}{m\sin^2\phi + m'}\right\}^{\frac{1}{2}}.....(7);$$

from this equation, when the form of the tube and therefore the relation between ϕ and s' is known, we may determine the relation between s' and t.

Again, for the determination of the relation between s and t, we have from (6) and (7),

$$\frac{ds}{dt} = \left(\frac{2g}{m+m'}\right)^{\frac{1}{4}} m \cos \phi \left\{\frac{\int \sin \phi \ ds'}{m \sin^2 \phi + m'}\right\}^{\frac{1}{4}}:$$

the integral $\int \sin \phi \, ds'$, which enters into these formulæ, must evidently be so taken as to vanish when s' = 0.

Clairaut; Mémoires de l'Académie des Sciences de Paris, 1742, p. 41. Euler; Opuscula, de motu corporum tubis mobilibus inclusorum, p. 48.

(6) A smooth groove ALA', (fig. 173), is carved in a vertical plane in the body ACBA', which is placed upon a smooth horizontal plane, along which it is able to slide freely; to find the form of the groove that a heavy particle may oscillate in it tautochronously, the time of an oscillation being given.

Let L be the lowest point of the groove, LM a vertical line through L, meeting a horizontal line PM through the place P of the particle at any time t; let PM = x', ML = y', arc LP = s'.

Then by the equation (7) of the preceding problem, putting -ds' in place of ds', and retaining in other respects the same notation, we have

$$\frac{ds'}{dt} = -\left\{2g(m+m')\right\}^{\frac{1}{4}} \left\{\frac{-\int \sin \phi \ ds'}{m \sin^2 \phi + m'}\right\}^{\frac{1}{4}}; \\ -\int \sin \phi \ ds' = k - y',$$

but

if k be the initial value of y'; hence

$$\frac{ds'}{dt} = -\left\{2g(m + m')\right\}^{\frac{1}{4}} \left\{ \frac{k - y'}{m \frac{dy'^2}{ds'^2} + m'} \right\},\,$$

and therefore, if r be the time of half an oscillation,

$$\tau = -\frac{1}{\left\{2g(m+m')\right\}^{\frac{1}{4}}}\int_{a}^{0} \frac{ds'}{(k-y')^{\frac{1}{4}}} \left(m\frac{dy'^{2}}{ds'^{2}} + m'\right)^{\frac{1}{4}}dy' \dots (1),$$

a quantity which, by the nature of the problem, must be independent of k: hence

$$\int \frac{ds'}{(k-y')^{\frac{1}{4}}} \left(m \frac{dy'^{2}}{ds'^{2}} + m' \right)^{\frac{1}{4}} dy'$$

must be of zero dimensions in y' and k; and therefore

$$\frac{\frac{ds'}{dy'}}{(k-y')^{\frac{1}{4}}}\left(m\frac{dy'^2}{ds'^2}+m'\right)^{\frac{1}{4}}$$

must evidently be of -1 dimensions in k and y': but it is clear that $\frac{ds'}{ds'} \left(m \frac{dy'^2}{ds'^2} + m' \right)^{\frac{1}{2}}$

does not involve k; hence this expression must be of $-\frac{1}{2}$ dimensions in y', and therefore, a being some constant quantity,

$$\frac{ds'}{dy'} \left(m \frac{dy'^2}{ds'^2} + m' \right)_{1}^{1} = \frac{a}{y'^{\frac{1}{4}}} \dots (2),$$

$$m + m' \frac{ds'^2}{dy'^2} = \frac{a^2}{y'},$$

$$m' \frac{dx'^2}{dy'^2} = \frac{a^2 - (m + m') y'}{y'},$$

$$\frac{dx'}{dy'} = \left(\frac{m+m'}{m'}\right)^{\frac{1}{2}} \left(\frac{2a-y'}{y'}\right)^{\frac{1}{2}},$$

where $2a = \frac{a^3}{m + m'}$; integrating, we get for the equation to the curve,

 $z' = \left(\frac{m+m'}{m'}\right)^{\frac{1}{2}} \left\{ (2ay'-y'^{\frac{1}{2}})^{\frac{1}{2}} + a \text{ vers}^{-1} \frac{y'}{a} \right\} \dots (3),$

which is an elongated cycloid, which may be constructed from the ordinary cycloid by increasing the distance of each point of the curve from the axis in the ratio of $(m + m')^{\frac{1}{2}}$ to $m^{\frac{1}{2}}$.

Again, from (1) and (2), we have

$$\tau = -\frac{\alpha}{\{2g(m+m')\}^{\frac{1}{4}}} \int_{a}^{0} \frac{dy'}{(ky'-y'^2)^{\frac{1}{4}}}$$

$$= \frac{\pi\alpha}{\{2g(m+m')\}^{\frac{1}{4}}};$$

$$\alpha^{2} = \frac{2g\tau^{2}}{\pi^{2}}(m+m'),$$

whence

and therefore

$$2a = \frac{a^3}{m + m'} = \frac{2g\tau^3}{\pi^3}, \quad a = \frac{g\tau^3}{\pi^3}.$$

Hence (3) may be written

$$x' = \left(\frac{m + m'}{m'}\right)^{\frac{1}{4}} \left\{ \left(\frac{2g\tau^2}{\pi^2} y' - y'^2\right)^{\frac{1}{4}} + \frac{g\tau^2}{\pi^2} \operatorname{vers}^{-1} \frac{\pi^2 y'}{g\tau^2} \right\}.$$

Euler; Opuscula, de motu corporum tubis mobilibus inclusorum. p. 51.

(7) A heavy particle P descends down a smooth inclined plane CA, (fig. 174), forming the upper surface of a body CAE which is capable of sliding freely down a smooth inclined plane OB, with which its lower flat surface is in contact; to determine the motion of the particle and of the body, both of which are supposed to be initially at rest, and the pressure of the particle on the body.

Let C be the initial position of the particle on the body, and O the initial position of the point A of the body; $\angle BAC = a$; $\angle OBF = \beta$, BF being horizontal; OA = s, CP = s', R =the

required pressure; m = the mass of the particle, m' = that of the body; t = the time from the commencement of the motion. Then

$$s = \frac{1}{2}gt^2 \frac{(m+m')\sin\beta + m\cos\alpha\sin(\alpha-\beta)}{m\sin^2\alpha + m'},$$

$$s' = \frac{1}{2}gt^2 \frac{(m+m')\sin\alpha^2\cos\beta}{m\sin^2\alpha + m'},$$

$$R = \frac{mm'g\cos\alpha\cos\beta}{m\sin^2\alpha + m}$$

Euler; Ibid. p. 35.

(8) Any number of heavy particles, P, P', P', ..., (fig. 175), are descending down a smooth inclined plane BA, forming the upper surface of a body BAC capable of sliding freely along a smooth horizontal plane OE, with which its lower flat surface is in contact; to determine the motion of the particles and of the body, and the pressures which they exert upon the body.

Let R, R', R', be the pressures, and m, m', m'', the masses of the particles P, P', P'',; let OA = S, BP = s, BP' = s', BP'' = s'',, O being a fixed point in the line OE, and B on the inclined plane BA; $\angle BAC = a$; M =the mass of the body. Then

$$S = \frac{1}{2}gt^{2} \frac{(m + m' + m' + ...) \sin a \cos a}{M + (m + m' + m' +) \sin^{2} a} + At + B,$$

$$s = \frac{1}{2}gt^{2} \frac{(M + m + m' + m' +) \sin a}{M + (m + m' + m' +) \sin^{2} a} + at + b,$$

$$s' = \frac{1}{2}gt^{2} \frac{(M + m + m' + m' +) \sin a}{M + (m + m' + m' +) \sin^{2} a} + a't + b',$$

$$\vdots \qquad \vdots \qquad \vdots \qquad \vdots \qquad \vdots$$

$$R = \frac{R'}{m'} = \frac{R'}{m'} = = \frac{M \cos a}{M + (m + m' + m' + m' +) \sin^{2} a}.$$

The quantities A, B, a, b, a', b', are arbitrary constants, to be determined from the initial circumstances of the motion.

Euler; *Ibid.* p. 40.

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CHAPTER IX.

MOTION OF RIGID BODIES.

SECT. 1. Single Body.

(1) A cylinder descends down a perfectly rough inclined plane by the action of gravity, its axis being horizontal; to determine the motion of the cylinder and the friction at any time of its descent.

Let G (fig. 176) be the centre of gravity of the cylinder at any instant of its descent; OA the course of the point of contact H of the circular section of the cylinder through G down the inclined plane; OH = x, a = the angle of inclination of OA to the horizon, $\theta =$ the whole angle through which the cylinder has revolved about its centre of gravity in moving from O to H; a = the radius of the cylinder, k = its radius of gyration about its axis. Let F denote the friction of the plane on the cylinder, which, from the signification of perfect roughness, is supposed to be sufficient to prevent sliding, and m the mass of the cylinder.

Then for the motion of the cylinder we have, resolving forces parallel to OA,

$$m\frac{d^2x}{dt^2} = mg \sin \alpha - F \dots (1);$$

and, taking moments about G,

$$mk^2\frac{d^2\theta}{dt^2}=Fa\ldots(2).$$

But, since F is sufficiently great to secure perfect rolling, we must evidently have $x = a\theta$; and therefore by (2),

$$mk^2\frac{d^2x}{dt^2}=Fa^2;$$

hence from (1) we get

$$ma^{2} \frac{d^{2}x}{dt^{2}} = ma^{2}g \sin a - mk^{2} \frac{d^{2}x}{dt^{2}},$$

$$(a^{2} + k^{2}) \frac{d^{2}x}{dt^{2}} = a^{2}g \sin a,$$

$$d^{2}x$$

or, since
$$k^3 = \frac{1}{2}a^2$$
, $3\frac{d^2x}{dt^2} = 2g \sin a$;

and therefore, integrating twice, and supposing $\frac{dx}{dt} = 0$ when II is at O, we have

$$x = \frac{1}{3}g \sin \alpha t^{2},$$

$$\theta = \frac{g \sin \alpha t^{2}}{3a} \dots (3);$$

and therefore

hence we have also, from (2) and (3),

$$F = \frac{mk^3}{a} \frac{d^3\theta}{dt^2} = \frac{2mgk^3 \sin a}{3a^3} = \frac{1}{3} mg \sin a.$$

(2) A globe descends from instantaneous rest down the surface of a perfectly rough hemispherical bowl, the centre of the globe always remaining in the same vertical plane; to determine the velocity of the globe at any position of its descent.

Let ABA' (fig. 177) be the vertical section of the bowl made by the plane in which the centre C of the globe is always situated, O being the centre of the bowl and OA a horizontal radius. Let M be the point in which the globe touches the bowl at any time of its motion, B being the initial position of M. Draw the radii OB, OCM; and let C'C be the circular arc described by the centre of gravity of the globe. Let $\angle AOM = \theta$, $\angle AOB = \alpha$, C'C=s, $\alpha=$ the radius of the globe, r= the radius of the bowl, $\phi=$ the angle which the globe has described about its centre of gravity in the motion from B to M, m= the mass of the globe; F= the friction of the bowl upon the globe at the point M, which is supposed to be sufficiently great to prevent all sliding.

Then for the motion of the centre of gravity of the globe, which will not be affected by our supposing all the impressed forces to be applied at C in their proper directions,

$$m\frac{d^3s}{dt^2}=-F+mg\cos\theta\ldots(1);$$

and for the motion of the globe about its centre of gravity,

$$mk^2\frac{d^3\phi}{dt^2}=Fa\ldots(2).$$

From the points B and M, draw two indefinite straight lines BkB' and MkT, tangents to the section of the hemispherical bowl; along BB measure a length Bm equal to the circular arc BM; then, if we were to conceive the globe to roll from B along the length Bm, and then Bm to be applied along BM so as to coincide with it, mB' being, as soon as m coincides with M, a tangent both to the circle AMA' and to the globe; it is evident that the globe would have revolved about its centre through the same angle as by its actual motion of rolling down the arc BM. Now by rolling along Bm it would have revolved about its centre through an angle $\frac{Bm}{a} = \frac{BM}{a} = \frac{r}{a}(\theta - a)$; and by the transference of m to M, it would have revolved through an angle equal to $\angle BkM = \angle BOM = \theta - \alpha$, in an opposite direction. we see that the whole actual angle through which the globe revolves about its centre in its actual motion from B to M, is equal to $\frac{r-a}{a}(\theta-a)=\phi$.

Hence, putting for ϕ its value in (2), we have

$$\frac{mk^2}{a}(r-a)\frac{d^2\theta}{dt^2}=Fa....(3).$$

Again, it is clear from the geometry that $s = (r - a)(\theta - a)$, and therefore, from (1),

$$m(r-a)\frac{d^2\theta}{dt^2}=-F+mg\cos\theta\ldots(4).$$

Eliminating F between (3) and (4), we obtain

$$m(r-a)\frac{d^2\theta}{dt^2} + \frac{mk^2}{a^2}(r-a)\frac{d^2\theta}{dt^2} = mg\cos\theta,$$

$$\left(1 + \frac{k^2}{a^2}\right)(r-a)\frac{d^2\theta}{dt^2} = g\cos\theta,$$

or, since k^2 is equal to $\frac{2}{5}a^2$,

$$\frac{d^2\theta}{dt^2} = \frac{5g\cos\theta}{7(r-a)}.....(5);$$

multiplying by $2\frac{d\theta}{dt}$, and integrating,

$$\frac{d\theta^a}{dt^a} = C + \frac{10g\sin\theta}{7(r-a)};$$

but, when $\theta = a$, $\frac{d\theta}{dt}$ is equal to zero; hence

$$0 = C + \frac{10g \sin a}{7(r-a)}$$

and therefore

$$\frac{d\theta^a}{dt^a} = \frac{10g\left(\sin\theta - \sin\alpha\right)}{7\left(r - a\right)};$$

whence

$$\frac{ds^2}{dt^2} = \frac{10g}{7} (r - a) (\sin \theta - \sin a);$$

and therefore also, if s' =the arc BM,

$$\frac{ds^{2}}{dt^{2}} = r^{2} \frac{d\theta^{2}}{dt^{2}} = \frac{10r^{2}g (\sin \theta - \sin a)}{7 (r - a)}.$$

For the magnitude of the friction at any time, we have, from (3),

$$F = \frac{mk^2}{a^3} (r - a) \frac{d^2\theta}{dt^2}$$
$$= \frac{5mgk^2 \cos \theta}{7a^3}, \text{ by the equation (5)}.$$

(3) A heterogeneous sphere rolls along a perfectly rough horizontal plane, its rotatory motion taking place always about an instantaneous axis normal to the vertical plane which passes through it geometrical centre and its centre of gravity; to determine its angular velocity for any position in its path.

Let C (fig. 178) be the geometrical centre and G the centre of gravity of the sphere at any time; S the point of contact of the vertical section of the sphere containing C and G with the horizontal plane; OSE the rectilinear locus of the points of contact; CGA a radius of the sphere; GM, CS, perpendiculars upon the plane.

Let F denote the friction of the plane at any time upon the sphere, estimated in the direction EO, and R the vertical reaction of the plane; m = the mass of the sphere; k = the radius of gyration about an axis through G at right angles to the vertical

section containing C and G; OM=x, GM=y, CS=CA=a, $\angle AGM=\angle ACS=\phi$, CG=c.

Then for the motion of the sphere we have, resolving forces parallel to OE,

$$m\frac{d^3x}{dt^3}=-F.\ldots.(1);$$

resolving forces vertically,

$$m \frac{d^2y}{dt^2} = R - mg \dots (2);$$

and, taking moments about the centre of gravity,

$$mk^2\frac{d^2\phi}{d\ell^2} = Fy - Rc\sin\phi \dots (3).$$

Now, since the friction is supposed to be sufficiently rough to prevent all sliding, we have from the geometry,

$$x + c \sin \phi = b + a\phi,$$

b being the value of x when $\phi = 0$; and therefore

$$\frac{dx}{dt} = a \frac{d\phi}{dt} - c \cos \phi \frac{d\phi}{dt},$$

$$\frac{d^2x}{dt^2} = (a - c \cos \phi) \frac{d^2\phi}{dt^2} + c \sin \phi \frac{d\phi^2}{dt^2}.$$

Again, from the geometry,

$$y = a - c \cos \phi$$

and therefore

$$\frac{dy}{dt} = c \sin \phi \frac{d\phi}{dt}, \quad \frac{d^3y}{dt^2} = c \sin \phi \frac{d^3\phi}{dt^2} + c \cos \phi \frac{d\phi^2}{dt^2}.$$

Hence, from (1),

$$F = -m \left\{ (a - c \cos \phi) \frac{d^2 \phi}{dt^2} + c \sin \phi \frac{d\phi^2}{dt^2} \right\},\,$$

and, from (2),

$$R = m\left(\frac{d^3y}{dt^2} + g\right) = m\left(c\sin\phi\frac{d^3\phi}{dt^2} + c\cos\phi\frac{d\phi^3}{dt^2} + g\right).$$

Substituting these expressions for R and F in (3), we get

$$k^{2} \frac{d^{2} \phi}{dt^{2}} = -\left\{ (a - c \cos \phi) \frac{d^{2} \phi}{dt^{2}} + c \sin \phi \frac{d\phi^{2}}{dt^{2}} \right\} (a - c \cos \phi)$$
$$- c \sin \phi \left(c \sin \phi \frac{d^{2} \phi}{dt^{2}} + c \cos \phi \frac{d\phi^{2}}{dt^{2}} + g \right),$$

and, by simplification,

$$(a^2 + k^2 + c^2 - 2ac \cos \phi) \frac{d^2\phi}{dt^2} + ac \sin \phi \frac{d\phi^2}{dt^2} = -cg \sin \phi;$$

multiplying by $2 \frac{d\phi}{dt}$, and integrating,

$$(a^2 + k^2 + c^2 - 2ac \cos \phi) \frac{d\phi^2}{dt^2} = C + 2cg \cos \phi.$$

Let t=0 when $\phi=0$, and let ω be the initial angular velocity about G; then $(\alpha^2 + k^2 + c^2 - 2ac) \omega^2 = C + 2cg,$

$$(a^{2}+c^{2}+k^{2}-2ac\cos\phi)\frac{d\phi^{2}}{dt^{2}}=\{(a-c)^{2}+k^{2}\}\omega^{2}-2cg(1-\cos\phi),$$

which gives the angular velocity about G at any time in terms of the whole angle described.

(4) A pendulum of any figure is firmly attached to a solid circular cylinder as an axis; this axis is placed horizontally within a hollow circular horizontal cylinder of larger diameter, and of which the surface is perfectly rough; in the hollow cylinder there is a slit, through which the pendulum hangs; supposing the initial position of the pendulum to be very nearly a position of equilibrium, to find the length of an isochronous simple pendulum.

Let g (fig. 179) denote the position of the centre of gravity of the pendulum and its axis, regarded as one mass, at any instant of the small motions; let O, c, denote the centres of the circular sections of the hollow and the solid cylinders made by a vertical plane through g; let gp meet at right angles the vertical line OAp, which cuts in A the circular section MA N of the hollow cylinder; join Oc and produce the line to a, which will be the point of contact of the sections of the solid and the hollow cylinders; let e be the point of contact of the section of the solid cylinder in its lowest position with that of the hollow one, so that the farc Aa will be equal to the arc ea. Let Op = x, gp = y, ce = ca = b, AO = Oa = a; $\phi = \angle cfO$ made by pro-

ducing cg to meet Op produced; $\angle AOa = \theta$, cg = c, $\angle ccf = \beta$; m = the mass of the pendulum and its axis together, k = the radius of gyration about g of the pendulum and axis regarded as one mass, R = the pressure of the hollow upon the solid cylinder, F = the friction of the hollow cylinder upon the solid one in the direction of a tangent to the arc aN.

Then for the motion of the system we have, resolving forces vertically,

$$m\frac{d^3x}{dt^2} = mg - R\cos\theta - F\sin\theta;$$

resolving forces horizontally,

$$m\frac{d^3y}{dt^3}=-R\sin\theta+F\cos\theta,$$

and, taking moments about g,

$$mk^{2}\frac{d^{2}\phi}{dt^{2}}=-Rc\sin\left(\theta+\phi\right)+F\left\{ c\cos\left(\theta+\phi\right)-b\right\} .$$

Now, θ and ϕ being by the hypothesis small angles, we may neglect their second and higher powers in the equations, and we get d^2x

$$m\frac{d^2x}{dt^2}=mg-R-F\theta,$$

$$m\frac{d^3y}{dt^2}=-R\theta+F,$$

$$mk^2\frac{d^2\phi}{dt^2}=-Rc(\theta+\phi)+F(c-b).$$

Eliminating R and F between these three equations, we shall finally obtain, as far as the first order of small quantities,

$$k^2 \frac{d^2 \phi}{dt^2} + cg \left(\theta + \phi\right) = \left(\frac{d^2 y}{dt^2} + g\theta\right) (c - b). \dots (1).$$

But, from the geometry, it is clear that

-
$$y = (a - b) \sin \theta - c \sin \phi = (a - b) \theta - c \phi$$
 nearly

and therefore, putting for brevity a - b = e,

$$\frac{d^2y}{dt^2}=e\frac{d^2\theta}{dt^2}-c\frac{d^2\phi}{dt^2};$$

hence the equation (1) becomes

$$(k^{2}+c^{3}-cb)\frac{d^{2}\phi}{dt^{2}}-e(c-b)\frac{d^{2}\theta}{dt^{2}}+cg\phi+bg\theta=0....(2).$$

Now, since there is no sliding, we may shew, by precisely the same method as in the case of problem (2), that

$$\phi + \beta = \frac{a-b}{b} \theta = \frac{e}{b} \theta,$$

 $\phi + \beta$ being evidently the whole angle described by the solid cylinder about its axis in rolling from a to A. Hence, from (2), we have

$$\{k^{3} + (c - b)^{3}\} \frac{d^{3}\phi}{dt^{6}} + \frac{g(b^{3} + ce)}{e}\phi + \frac{b^{2}\beta g}{e} = 0;$$

$$\frac{g(b^{3} + ce)}{e}\phi + \frac{b^{3}\beta g}{e} = \frac{g(b^{3} + ce)}{e}\psi;$$

then $\frac{d^2\phi}{dt^2} = \frac{d^2\psi}{dt^2}$, and the equation becomes

$$\{k^2 + (c-b)^3\} \frac{d^2\psi}{d\ell^2} + \frac{g(b^2 + ce)}{e} \psi = 0,$$

$$\frac{e\{k^2 + (c-b)^3\}}{b^2 + ce} \frac{d^2\psi}{d\ell^2} + g\psi = 0.$$

or

let

Hence, if l denote the length of a perfect pendulum isochronous with the period of ϕ , and therefore of ψ , we shall have

$$l = \frac{e\left\{k^{2} + (c - b)^{2}\right\}}{b^{2} + ce} = \frac{(a - b)\left\{k^{2} + (c - b)^{2}\right\}}{b^{2} + c(a - b)}.$$

Euler; Acta Acad. Petrop. 1780; P. 2; p. 164.

(5) At the extremities A and B of a uniform beam AB, (fig. 180), are two small rings, capable of sliding along the horizontal and vertical rods Ox, Oy; the friction between the ends of the beam and the rods is equal to the normal pressure on each; to determine the motion of the beam.

Let G be the centre of gravity of the beam; draw GH vertical; let OH = x, GH = y, AG = BG = a, $\angle BAO = \theta$; R, S, the normal reactions of the rods Ox, Oy, and therefore R, S, the frictions along xO, Oy; m = the mass of the beam, k = the radius of gyration about G.

Then, for the motion of the beam,

$$m\frac{d^3x}{dt^2} = S - R. \qquad (1),$$

$$m\frac{d^3y}{dt^3} = S + R - mg \qquad (2),$$

$$mk^2\frac{d^3\theta}{dt^2} = (S - R) a \cos \theta + (S + R) a \sin \theta \qquad (3).$$

Substituting the values of S-R and S+R from (1) and (2) in the equation (3), we get

But
$$z = a \cos \theta \frac{d^{2}x}{dt^{2}} + \left(g + \frac{d^{2}y}{dt^{2}}\right) a \sin \theta \dots (4).$$

$$z = a \cos \theta, \quad \frac{dx}{dt} = -a \sin \theta \frac{d\theta}{dt},$$

$$\frac{d^{2}x}{dt^{2}} = -a \cos \theta \frac{d\theta^{2}}{dt^{2}} - a \sin \theta \frac{d^{2}\theta}{dt^{2}};$$
and
$$y = a \sin \theta, \quad \frac{dy}{dt} = a \cos \theta \frac{d\theta}{dt},$$

$$\frac{d^{2}y}{dt^{2}} = -a \sin \theta \frac{d\theta^{2}}{dt^{2}} + a \cos \theta \frac{d^{2}\theta}{dt^{2}}.$$

Substituting these values of $\frac{d^2x}{dt^2}$, $\frac{d^2y}{dt^2}$, in the equation (4),

we have

$$k^2 \frac{d^2 \theta}{dt^2} + a^2 \frac{d\theta^2}{dt^2} = ag \sin \theta;$$

or, changing the independent variable from t to θ ,

$$-k^2\frac{\frac{d^2t}{d\theta^2}}{\frac{dt^3}{d\theta^3}} + \frac{a^2}{\frac{dt^2}{d\theta^2}} = ag \sin \theta:$$

multiplying both sides of the equation by $2e^{\frac{2a^2}{k^2}\theta}$, we obtain

$$\frac{d}{d\theta} \left\{ \frac{k^2 \epsilon^{\frac{2a^2}{k^2} \theta}}{\frac{dt^2}{d\theta^2}} \right\} = 2ag \epsilon^{\frac{2a^2}{k^2} \theta} \sin \theta;$$

integrating,

Now, integrating by parts, we shall get

$$\int_{\varepsilon^{\frac{2a^2}{k^2}}}^{\frac{2a^2}{k^2}} \sin \theta \, d\theta = \frac{k^4}{k^4 + 4a^4} \left\{ \frac{2a^2}{k^2} \sin \theta - \cos \theta \right\} \varepsilon^{\frac{2a^2}{k^2}};$$

hence, from (5),

$$k^{2} \frac{\epsilon^{\frac{2a^{2}}{k^{2}}\theta}}{dt^{2}} \frac{d\theta^{2}}{dt^{2}} = \frac{2agk^{4}}{k^{4} + 4a^{4}} \left\{ C' + \left(\frac{2a^{2}}{k^{2}} \sin \theta - \cos \theta \right) \frac{\epsilon^{\frac{2a^{2}}{k^{2}}\theta}}{k^{4} + 4a^{4}} \right\},$$

$$\frac{\epsilon^{\frac{2a^{2}}{k^{2}}\theta}}{dt^{2}} \frac{d\theta^{4}}{dt^{2}} = \frac{2agk^{2}}{k^{4} + 4a^{4}} \left\{ C' + \left(\frac{2a^{2}}{k^{2}} \sin \theta - \cos \theta \right) \frac{\epsilon^{\frac{2a^{2}}{k^{2}}\theta}}{k^{4} + 4a^{4}} \right\}.$$

Since $a^2 = 3k^2$, we have

$$\varepsilon^{\bullet \theta} \frac{d\theta^{\theta}}{dt^{\theta}} = \frac{6g}{37a} \left\{ C' + (6 \sin \theta - \cos \theta) \varepsilon^{\bullet \theta} \right\}.$$

Suppose that a, ω , are simultaneous values of θ , $\frac{d\theta}{dt}$; then

$$\varepsilon^{\mathbf{e}_{x}}\omega^{2}=\frac{6g}{37a}\left\{C'+\left(6\sin a-\cos a\right)\varepsilon^{\mathbf{e}_{x}}\right\};$$

hence

$$\varepsilon^{\omega\theta} \frac{d\theta^a}{dt^a} - \varepsilon^{6x}\omega^2 = \frac{6g}{37a} \left\{ (6 \sin \theta - \cos \theta) \varepsilon^{6\theta} - (6 \sin a - \cos a) \varepsilon^{6\tau} \right\},\,$$

which determines the angular velocity of the beam at every position in its descent.

(6) To determine the motion of a cylinder upon a perfectly rough plane of indefinite extent, which, having been initially horizontal, has a uniform motion downwards about a fixed horizontal axis within itself, to which the line of contact of the cylinder and plane is always parallel, and with which it initially coincides.

Let a vertical plane at right angles to the axis of revolution, and passing through the centre of gravity C of the cylinder, cut the revolving plane at any instant of the motion in the line OA, (fig. 181); let Ox be the initial position of OA; draw CM, CN, at right angles to Ox, OA; let OM = x, CM = y, ON = r, CN = a, $\omega =$ the angular velocity of OA about O, $\triangle AOx = \omega t$, k = the radius of gyration of the cylinder about its axis, R = the normal reaction of the plane in the direction NC, T = the tangential reaction along NO, $\theta =$ the angle through which the

cylinder has revolved about its centre of gravity at the end of the time t, m = the mass of the cylinder.

Then for the motion of the cylinder, we have

$$m \frac{d^3x}{dt^3} = R \sin \omega t - T \cos \omega t \dots (1),$$

$$m \frac{d^3y}{dt^3} = mg - R \cos \omega t - T \sin \omega t \dots (2),$$

$$mk^2 \frac{d^3\theta}{dt^3} = Ta \dots (3).$$

Again, the angle through which the cylinder would revolve about C, by rolling along Ox through a space r, would be $\frac{r}{a}$, and that due to the motion of Ox into the position OA would be et; hence $\frac{r}{a} + et$ is the angle through which it has actually revolved at the end of the time t; or

$$\theta = \frac{r}{a} + \omega t \quad \quad (4).$$

From the equations (1) and (2), we have

$$\sin \omega t \frac{d^2x}{dt^2} - \cos \omega t \frac{d^2y}{dt^2} = \frac{R}{m} - g \cos \omega t \dots (5);$$

and from (1), (2), (3),

$$\cos \omega t \frac{d^3x}{dt^2} + \sin \omega t \frac{d^3y}{dt^2} = g \sin \omega t - \frac{k^3}{a} \frac{d^3\theta}{dt^2};$$

but, from (4), we get $a \frac{d^2\theta}{dt^2} = \frac{d^2r}{dt^2}$;

hence
$$\cos \omega t \frac{d^2x}{dt^2} + \sin \omega t \frac{d^2y}{dt^2} = g \sin \omega t - \frac{k^2}{a^2} \frac{d^2r}{dt^2} \dots (6)$$

From the geometry it is clear that

 $x = r \cos \omega t + a \sin \omega t$, $y = r \sin \omega t - a \cos \omega t$; differentiating these expressions twice with respect to t, we shall get

$$\frac{d^2x}{dt^2} = \cos \omega t \frac{d^2r}{dt^2} - 2\omega \sin \omega t \frac{dr}{dt} - \omega^2 r \cos \omega t - a\omega^2 \sin \omega t,$$

$$\frac{d^2y}{dt^2} = \sin \omega t \frac{d^2r}{dt^2} + 2\omega \cos \omega t \frac{dr}{dt} - \omega^2 r \sin \omega t + a\omega^2 \cos \omega t$$

substituting these expressions for $\frac{d^2x}{dt^2}$, $\frac{d^3y}{dt^2}$, in the equations (5) and (6), we obtain

$$2\omega \frac{dr}{dt} + a\omega^2 = g \cos \omega t - \frac{R}{m} \dots (7),$$

and

$$\frac{a^3+k^3}{a^3}\frac{d^3r}{dt^2}-\omega^2r=g\sin \omega t\ldots\ldots(8).$$

Since $a^2 = 2k^2$, the equation (8) becomes

$$\frac{d^2r}{dt^2} - \frac{2}{3}\omega^2r = \frac{2}{3}g\sin\omega t;$$

the integral of this equation is

$$r = -\frac{2g}{5\omega^2}\sin \omega t + C\varepsilon^{\omega't} + C'\varepsilon^{-\omega't},$$

where $\omega' = (3)^{\frac{1}{2}}\omega$, and C, C', are arbitrary constants. If we determine C and C' from the conditions that r = 0, $\frac{dr}{dt} = 0$, initially, we shall have

$$r = -\frac{2g}{5\omega^2}\sin \omega t + \frac{3^{\frac{3}{2}}g}{5\omega^2 2^{\frac{1}{2}}}(\varepsilon^{\omega't} - \varepsilon^{-\omega't}),$$

which determines the position of the cylinder at any time before it detaches itself from the revolving plane: differentiating with respect to t,

$$\frac{dr}{dt} = -\frac{2g}{5\omega}\cos\omega t + \frac{g}{5\omega}\left(\varepsilon^{\omega't} + \varepsilon^{-\omega't}\right);$$

hence, from (7), we obtain

$$\frac{R}{m} = g \cos \omega t - a\omega^2 + \frac{4}{5}g \cos \omega t - \frac{2g}{5} \left(\varepsilon_{\cdot}^{\omega't} + \varepsilon_{\cdot}^{-\omega't} \right)$$
$$= \frac{9}{5}g \cos \omega t - a\omega^2 - \frac{2g}{5} \left(\varepsilon_{\cdot}^{\omega't} + \varepsilon_{\cdot}^{-\omega't} \right),$$

which gives the value of R for any time of the motion of the cylinder upon the plane: when R=0, or when the cylinder leaves the plane,

$$9q \cos \omega t = 5a\omega^2 + 2q \left(\varepsilon^{\omega't} + \varepsilon^{-\omega't}\right),$$

an equation which fixes the epoch of the separation.

*

(7) A sphere is projected directly down an inclined plane with a motion both of translation and of rotation; the motion of rotation is the same in point of direction as that which would correspond to perfect downward rolling, but greater in magnitude: to determine the motion of the sphere, having given the coefficients both of statical and of dynamical friction between the sphere and the inclined plane.

Let OA (fig. 182) be the inclined plane, C the position of the sphere's centre, and M its point of contact with OA at the end of a time t from the beginning of the motion. Let μ = the coefficient of dynamical friction between the sphere and the plane, a = the radius of the sphere, OM = s, ϕ = the angle through which the sphere has revolved at the end of the time t about its centre of gravity, R = the normal reaction of the plane, m = the mass of the sphere, a = the inclination of the plane to the horizon.

Then, for the motion of the sphere, we have

$$m \frac{d^3s}{dt^3} = mg \sin a + \mu R \dots (1),$$

$$mk^2\frac{d^2\phi}{dt^2}=-\mu aR \dots (2):$$

but, since C has no motion at right angles to the plane, we have also $R = mg \cos a$; hence, from (1),

$$\frac{d^2s}{dt^2} = g \left(\sin \alpha + \mu \cos \alpha \right) \dots (3),$$

and, from (2),

$$k^2 \frac{d^2 \phi}{dt^2} = - \mu ag \cos \alpha \dots (4).$$

Integrating the equation (3) with respect to t, and denoting the initial value of $\frac{ds}{dt}$ by c, we get

$$\frac{ds}{dt} = gt \left(\sin \alpha + \mu \cos \alpha\right) + c \ldots (5);$$

similarly from (4), ω denoting the initial value of $\frac{d\phi}{dt}$,

$$\frac{d\phi}{dt} = \omega - \frac{\mu ag}{k^2} t \cos \alpha \dots (6).$$

As soon as, by the increase of t, $\frac{ds}{dt}$ becomes equal to $a\frac{d\phi}{dt}$, the motion will change its character, and our present equations will cease to be applicable. This event will take place when

$$gt \left(\sin \alpha + \mu \cos \alpha\right) + c = a\omega - \frac{\mu a^2 gt}{k^2} \cos \alpha,$$
or
$$t = \frac{k^2 \left(a\omega - c\right)}{\mu g \left(a^2 + k^2\right) \cos \alpha + k^2 g \sin \alpha} = t' \text{ suppose.}$$

For all values of t not greater than t, we have from (5) and (6), s and ϕ being considered to be initially zero,

$$s = \frac{1}{2} gt^2 \left(\sin \alpha + \mu \cos \alpha \right) + ct,$$

$$\phi = \omega t - \frac{\mu a gt^2}{2k^2} \cos \alpha;$$

the values of s and ϕ at the end of the first period of the motion will be obtained from these expressions by substituting t' in place of t.

When $\frac{ds}{dt}$ becomes equal to $a\frac{d\phi}{dt}$, there evidently exists at that instant no sliding between the plane and the sphere; and therefore, before dynamical friction can again come into play, the statical friction between the sphere and the plane must be overcome. First, let us suppose that the statical friction is sufficiently great to secure perfect rolling; and let F denote the tangential reaction of the plane against the descent of the sphere. The

$$m \frac{d^3s}{dt^2} = mg \sin \alpha - F \dots (1),$$

$$mk^2 \frac{d^3\phi}{dt^2} = aF \dots (2).$$

Also, there being no sliding, it is clear that

$$a\frac{d\phi}{dt}=\frac{ds}{dt}$$
, $a\frac{d^3\phi}{dt^3}=\frac{d^3s}{dt^3}$;

hence from (2) there is

equations of motion will be,

$$mk^2\frac{d^2s}{dt^2}=a^2F,$$

and therefore, from (1),

$$(a^3+k^2)\frac{d^3s}{dt^3}=a^3g\sin a \dots (3);$$

and therefore, also,

$$(a^2+k^2)\frac{d^3\phi}{dt^2}=ag\sin a \ldots (4).$$

Integrating, we get from (3) and (4),

$$\frac{ds}{dt} = \frac{a^2g\sin a}{a^2 + k^2}t + c',$$

$$\frac{d\phi}{dt} = \frac{ag\sin\alpha}{a^2 + k^2} t + \omega';$$

where c', ω' , are the values of $\frac{ds}{dt}$, $\frac{d\phi}{dt}$, at the end of the first stage of the motion, the time being now reckoned from the commencement of the second period: c' is clearly equal to $a\omega'$.

Integrating again, we get

$$s = \frac{\frac{1}{2} a^2 g t^2 \sin \alpha}{a^3 + k^3} + c't + s';$$

$$\phi = \frac{\frac{1}{2} agt^2 \sin \alpha}{a^2 + k^2} + \omega' t + \phi',$$

s', ϕ' , being the values of ϕ , s, at the end of the first period.

Also
$$F = \frac{mk^2}{a} \frac{d^2\phi}{dt^2} = \frac{mk^2g \sin \alpha}{a^2 + k^2},$$

which is the value of the statical friction necessary to secure perfect rolling in the second stage of the motion.

If the statical friction be less than this, dynamical friction will arise, and will evidently exert itself up the plane. Hence, for the motion,

$$\frac{d^2s}{dt^2}=g\,(\sin\,\alpha-\mu\,\cos\,\alpha)\,\ldots\,(\mathbf{A}),$$

$$k^2 \frac{d^2 \phi}{dt^2} = \mu ag \cos \alpha \dots (B).$$

It may be easily ascertained that the coefficient of g in the expression for $\frac{d^2s}{dt^2}$ is positive; for the coefficient of friction

necessary for perfect rolling

$$=\frac{F}{R}=\frac{F}{mq\cos a}=\frac{k^2\tan a}{a^2+k^2},$$

and, since μ is less than this by hypothesis, we have

$$\mu \angle \frac{k^2}{a^3 + k^3} \tan a$$
, and therefore $\mu \angle \tan a$, $\mu \cos a \angle \sin a$.

From (A) and (B) we have

$$\frac{ds}{dt} = gt \left(\sin \alpha - \mu \cos \alpha\right) + c',$$

$$\frac{d\phi}{dt} = \frac{\mu agt}{k^2} \cos a + \omega'.$$

It may be readily seen that $\frac{ds}{dt} - a \frac{d\phi}{dt}$ never becomes zero in the second stage of the motion, but is always positive; for, bearing in mind that $c' = a\omega'$,

$$\frac{ds}{dt} - a \frac{d\phi}{dt} = gt \left(\sin \alpha - \mu \cos \alpha - \mu \frac{a^2}{k^2} \cos \alpha \right)$$
$$= gt \cos \alpha \frac{a^2 + k^2}{k^2} \left(\frac{k^2}{a^2 + k^2} \tan \alpha - \mu \right);$$

hence the sphere will always rotate too slowly, in comparison with the velocity of translation, to correspond to perfect rolling.

If s' denote the space through which the sphere descends along the plane in the second stage of the motion, in consequence of sliding,

$$\frac{ds'}{dt} = \frac{ds}{dt} - a \frac{d\phi}{dt} = gt \left(\sin \alpha - \mu \cos \alpha \frac{a^3 + k^2}{k^3} \right),$$

$$s' = \frac{1}{2} gt^2 \left(\sin \alpha - \mu \cos \alpha \frac{a^3 + k^2}{k^3} \right).$$

Euler; Acta Acad. Petrop. P. 11. p. 131; 1781.

(8) A sphere, revolving about a horizontal axis, is placed upon an imperfectly rough plane inclined to the horizon at an angle $\tan^{-1}\mu$, where μ is the coefficient of dynamical friction, the direction of the sphere's rotation being opposite to that which would correspond to perfect downward rolling; to determine the motion of the sphere.

The notation remaining the same as in the preceding problem, we shall have, for the motion of the sphere,

$$m \frac{d^3s}{dt^2} = mg \sin \alpha - \mu R,$$

$$mk^3 \frac{d^3\phi}{dt^2} = -\mu aR;$$

but $R = mg \cos a$; hence we have

$$\frac{d^2s}{dt^2} = g \left(\sin \alpha - \mu \cos \alpha \right) = 0, \text{ by hypothesis };$$

and

$$\frac{d^2\phi}{d\ell^2} = -\frac{\mu ag \cos \alpha}{k^2}.$$

Integrating with respect to t, we get

$$\frac{ds}{dt} = C, \quad \frac{d\phi}{dt} = C' - \frac{\mu agt}{k^2} \cos \alpha,$$

where C, C', are arbitrary constants; but, initially, $\frac{ds}{dt} = 0$, $\frac{d\phi}{dt} = \omega$; and therefore C = 0, $C' = \omega$; hence

$$\frac{ds}{dt} = 0, \quad \frac{d\phi}{dt} = \omega - \frac{\mu agt}{k^2} \cos \alpha.$$

In the preceding investigation we have supposed the friction of the plane upon the sphere to act upwards; this will cease to be the case when $\frac{ds}{dt} + a \frac{d\phi}{dt}$, or, since $\frac{ds}{dt} = 0$, when $\frac{d\phi}{dt}$ becomes zero. Hence we see that, for a time equal to

$$\frac{k^2\omega}{\mu ag\cos a}$$
,

the centre of the sphere will remain stationary, and that at the end of this time the angular velocity of the sphere will become zero.

Before proceeding to investigate the nature of the motion after the end of the stationary period of the centre of the sphere, it will be necessary to determine the amount of upward friction requisite to cause the sphere subsequently to assume a motion of perfect rolling. The coefficient of friction requisite for this purpose, (see the preceding problem,) is equal to

$$\frac{k^3 \tan \alpha}{a^3 + k^3}.$$

But μ is equal to tan a, and therefore exceeds the requisite magnitude. The sphere will then proceed subsequently to roll down the plane without sliding; and the space described by its centre, at the end of a time t from the termination of its stationary interval, will be equal to

$$\frac{1}{2}\frac{a^2gt^2\sin \alpha}{a^2+k^2},$$

which, since $k^2 = \frac{2}{5}a^2$, is equal to $\frac{1}{14}gt^2 \sin a$. Also, putting for k^2 its value, the stationary interval will have for its value $\frac{2a\omega}{5g\sin a}$.

Euler; Acta Acad. Petrop. P. 11. p. 131; 1781.

(9) A homogeneous sphere attracted towards a given centre of force varying directly as the distance, is projected with a given velocity along a plane passing through that centre, friction being such as to prevent all sliding; to determine the path described by the sphere.

Let O, (fig. 183), the centre of force, be taken as the origin of co-ordinates, and let Ox, Oy, which are at right angles to each other and in the plane along which the sphere rolls, be taken as the co-ordinate axes. Let C be the centre of the sphere at any time of its motion, P its point of contact with the plane xOy; join CO, CP, and draw PM parallel to yO; draw also Oz at right angles to the plane xOy. Let OC = r, CP = c, OM = x, PM = y, m = the mass of the sphere, $\mu =$ the attraction of the central force upon a unit of mass collected at a point at a unit of distance; let X, Y, denote the friction of the plane on the sphere, estimated parallel to Ox, Oy; ω' , ω' , the angular velocities of the sphere at any time about diameters parallel to Ox, Oy, estimated in the directions indicated by the arrows in the planes yOz, zOx; k = the radius of gyration of the sphere about a diameter.

It may be readily ascertained that the attraction on the whole sphere will be equal to a force μmr in the direction CO, and the

resolved parts of this force parallel to Ox, Oy, are evidently $-\mu mx$, $-\mu my$. Hence, for the motion of the sphere, we have

$$m \frac{d^3x}{dt^2} = X - \mu mx. \qquad (1),$$

$$m \frac{d^3y}{dt^2} = Y - \mu my \qquad (2),$$

$$mk^2 \frac{d\omega}{dt} = cY. \qquad (3),$$

$$mk^2 \frac{d\omega}{dt} = -cX. \qquad (4).$$

But, since the friction is sufficiently great to prevent all sliding, it is clear that

$$\frac{dx}{dt} = c\omega', \quad \frac{dy}{dt} = -c\omega';$$

hence, from (3) and (4), we get

$$Y = -\frac{mk^3}{c^3} \frac{d^3y}{dt^2}, \quad X = -\frac{mk^3}{c^3} \frac{d^3x}{dt^3}$$

$$(\omega'''^2 - \omega''^2) \sum (\delta m.yz) + \left(\omega'''\omega' - \frac{d\omega''}{dt}\right) \sum (\delta m.xy) - \left(\omega''\omega' + \frac{d\omega''}{dt}\right) \sum (\delta m.xz) + \omega'''\omega'' \sum (\delta m.y^2) - \omega''\omega''' \sum (\delta m.z^2) + \frac{d\omega'}{dt} \sum \{\delta m.(y^2 + z^2)\} = L,$$

where L denotes the moment of the forces about the first of the three straight lines. Similar equations, mutatis mutandis, will hold in relation to the axes of y and z. If the body be a homogeneous sphere, as in the present problem,

 $\Sigma(\delta m.yz) = 0$, $\Sigma(\delta m.xy) = 0$, $\Sigma(\delta m.xz) = 0$, $\Sigma(\delta m.y^2) - \Sigma(\delta m.z^2) = 0$, and therefore, adopting the notation of the text,

$$mk^2\frac{d\omega'}{dt}=L,$$

L being equal to cY. Similarly

$$mk^2 \frac{d\omega''}{dt} = -cX.$$

[•] If ω' , ω'' , ω''' , denote the angular velocities of a rigid body about three straight lines through its centre of inertia, parallel to the axes of z, y, z, respectively, and δm denote an element of the mass, then, (See Pratt's Mechanical Philosophy, First Edit. p. 428),

and therefore, from (1), (2),

$$\frac{c^{2} + k^{2}}{c^{3}} \frac{d^{2}x}{dt^{2}} = -\mu x,$$

$$\frac{c^{2} + k^{2}}{c^{2}} \frac{d^{2}y}{dt^{2}} = -\mu y;$$

or, since
$$k^2 = \frac{2}{5}c^3$$
,
$$\frac{d^2x}{dt^2} = -\frac{5\mu}{7}x, \quad \frac{d^2y}{dt^2} = -\frac{5\mu}{7}y:$$

integrating, and adding arbitrary constants,

$$\frac{dx^2}{dt^2} = C - \frac{5\mu}{7} x^2$$
, $\frac{dy^2}{dt^2} = C' - \frac{5\mu}{7} y^2$;

let a, b, be the initial values of x, y, and a, β , of $\frac{dx}{dt}$, $\frac{dy}{dt}$; then

$$a^2 = C - \frac{5\mu}{7} a^2$$
, $\beta^2 = C' - \frac{5\mu}{7} b^2$,

and therefore

$$\frac{dx^3}{dt^3} - \alpha^2 = \frac{5\mu}{7}(a^2 - x^3), \quad \frac{dy^2}{dt^2} - \beta^2 = \frac{5\mu}{7}(b^2 - y^3):$$

eliminating dt, and putting

$$a^{\prime 2} = \frac{7a^{2}}{5\mu} + a^{3}, \quad b^{\prime 2} = \frac{7\beta^{2}}{5\mu} + b^{2},$$

$$\frac{dy^{2}}{b^{\prime 2} - y^{3}} = \frac{dx^{2}}{a^{\prime 2} - x^{3}}, \qquad \frac{dy}{(b^{\prime 2} - y^{3})^{\frac{1}{2}}} = \frac{dx}{(a^{\prime 2} - x^{2})^{\frac{1}{2}}};$$

integrating, we have

$$\sin^{-1}\frac{y}{h'} = \sin^{-1}\frac{x}{a'} + \cos^{-1}c' \dots (5),$$

where c' is an arbitrary constant; hence

$$\frac{y}{b'} = \frac{c'x}{a'} + (1 - c^2)^{\frac{1}{2}} \left(1 - \frac{x^2}{a'^2} \right)^{\frac{1}{2}},$$

$$\frac{c'^2x^2}{a'^2} - 2c' \frac{x}{a'} \frac{y}{b'} + \frac{y^2}{b'^2} = (1 - c'^2) \left(1 - \frac{x^2}{a'^2} \right),$$

$$\frac{x^2}{a'^2} - 2c' \frac{x}{a'} \frac{y}{b'} + \frac{y^2}{b'^2} = 1 - c'^2;$$

putting a, b, for x, y, we have for the determination of c',

$$\frac{a^2}{a'^2} - 2c'\frac{a}{a'}\frac{b}{b'} + \frac{b^2}{b'^2} = 1 - c'^2;$$

let m, n, be the two values of c given by this quadratic; then for the equation to the path of the sphere we shall have

$$\frac{x^2}{a'^2} - 2m \frac{x}{a'} \frac{y}{b'} + \frac{y^2}{b'^2} = 1 - m^2,$$

or

$$\frac{x^3}{a'^3} - 2n \frac{x}{a'} \frac{y}{b'} + \frac{y^3}{b'^3} = 1 - n^3.$$

Thus, m and n being neither of them greater than unity, we see that the sphere will describe one or other of two ellipses, the centres of which coincide with the origin of co-ordinates. The ellipse, which is to be taken, will depend upon the sign of the ratio α to β .

SECT. 2. Several Bodies.

(1) A cylinder rolls down a perfectly rough inclined plane, while a string coils round it which unwinds from an equal cylinder revolving about its axis which is fixed, the position of the latter cylinder being such that the string is parallel to the plane; to find the accelerative force of descent, the tension of the string, and the friction of the inclined plane.

Let O (fig. 184) be the centre of gravity of the descending cylinder at any time of its motion down the plane BA, M being its point of contact with the plane; let C be the centre of gravity of the other cylinder; join CO. Let CO = x at any time t; a = the radius of each of the cylinders, a = the inclination of the plane BA to the horizon, T = the tension of the uncoiled string, F = the friction of the inclined plane exerted upon the cylinder O at M in the direction MB; m = the mass of each of the cylinders, and k = the radius of gyration of each about its axis; let θ , θ , denote the angles through which the cylinders O, C, have revolved about their axes at the end of the time t.

Then, for the motion of the cylinder O, we have

$$m\frac{d^3x}{dt^2}=mg\sin\alpha-F-T.....(1),$$

$$mk^3\frac{d^3\theta}{dt^2}=(F-T)a\dots(9);$$

and, for the motion of the cylinder C,

$$mk^2\frac{d^2\theta}{dt^2}=Ta \ldots (8).$$

Multiplying the equations (1) and (8) by a and 2 respectively, and adding the resulting equations to (2), we get

$$k^2 \frac{d^2\theta}{d\ell^2} + 2k^2 \frac{d^2\theta'}{d\ell^2} + k^2 \frac{d^2x}{d\ell^2} = ag \sin \alpha;$$

but, from the geometry, it is evident that

$$a\frac{d^2\theta}{dt} = \frac{d^2x}{dt^2}, \quad a\frac{d^2\theta}{dt^2} = 2\frac{d^2x}{dt^2};$$

hence we have
$$(a^3 + 5k^2) \frac{d^3x}{dt^2} = a^3 g \sin a$$
,

or, since $2k^2 = a^2$,

$$\frac{d^2x}{dx} = \frac{2}{3}g \sin \alpha.$$

Again, from (8),

$$T = \frac{mk^2}{a} \frac{d^3\theta}{dt^2} = \frac{2mk^2}{a^3} \frac{d^3x}{dt^2} = m \frac{d^3x}{dt^2} = \frac{1}{2} mg \sin \alpha.$$

Lastly, from (2),
$$F = T + \frac{mk^2}{a} \frac{d^2\theta}{dt^2} = T + \frac{1}{2} ma \frac{d^2\theta}{dt^2}$$

$$= T + \frac{1}{2}m \frac{d^2x}{dt^2} - \frac{1}{2}mg \sin \alpha + \frac{1}{2}mg \sin \alpha$$
$$= \frac{1}{2}mg \sin \alpha.$$

(2) A string, the free end of which hangs through a ring and has a weight attached to it, is wound about a circular section of a cylinder, made by a vertical plane passing through the ring and through the centre of gravity of the cylinder; the cylinder rests upon a rough horizontal plane, and its diameter is equal to the altitude of the ring above this plane: to determine the motion of the weight and of the cylinder, which are supposed to be initially in a state of instantaneous rest.

Let R (fig. 185) be the position of the ring; NRP the free portion of the string meeting in the point O the locus OMA of the point M in which the circular section of the cylinder touches the plane at any time; C the position at any time of the centre of gravity of the cylinder. Let a = the radius of the cylinder, OP = y, OM = x, ϕ = the angle through which the cylinder has revolved about its axis at the end of the time t, F = the action of the plane on the cylinder estimated in the direction AO, m = the mass of the cylinder, k = its radius of gyration about its axis, m' = the mass of P, T = the tension of the string.

Then, for the motion of the weight, we have

$$m'\frac{d^3y}{dt^2}=m'g-T....(1);$$

and for the motion of the cylinder, both in respect to translation and to rotation,

$$m \frac{d^3x}{dt^2} = -T - F \dots (2),$$

$$mk^2 \frac{d^2\phi}{dt^2} = Ta - Fa \dots (3).$$

Now, since the horizontal plane along which the cylinder moves may be either perfectly or imperfectly rough, we shall have to consider two cases of the motion. We will first suppose the plane to be perfectly rough, that is, to be sufficiently rough to prevent all sliding.

Since the cylinder rolls without sliding, we must evidently have $dx = -ad\phi$; and therefore, by (3),

$$-mk^2\frac{d^2x}{dt^2}=Ta^2-Fa^2.....(4);$$

also, from (2),

and therefore, adding together these two last equations,

$$-m(a^2+k^2)\frac{d^2x}{dt^2}=2Ta^2.....(6);$$

and therefore, by (1),

$$2m'a^{2}\frac{d^{2}y}{dt^{2}}-m(a^{2}+k^{2})\frac{d^{2}x}{dt^{2}}=2m'a^{2}g:$$

but, if *l* denote the original length of the free string, it is clear that

$$x + y = l + a\phi$$
, $\frac{d^3x}{dt^2} + \frac{d^3y}{dt^2} = a\frac{d^3\phi}{dt^2} = -\frac{d^3x}{dt^2}$, $\frac{d^3y}{dt^2} = -2\frac{d^3x}{dt^2}$;

hence we have

$$\left\{4m'a^2+m(a^2+k^3)\right\}\frac{d^3x}{dt^2}=-2m'a^2g\ldots(7);$$

integrating, and bearing in mind that $\frac{dx}{dt} = 0$ when t = 0,

$${4m'a^2 + m(a^2 + k^2)}\frac{dx}{dt} = -2m'a^2gt$$
:

integrating again, and taking c for the initial value of x,

$$x = c - \frac{m'a^2gt^2}{4m'a^2 + m(a^2 + k^2)}$$

We may easily get also

$$y = l - c + \frac{2m'a^2gt^2}{4m'a^2 + m(a^2 + k^2)};$$

and therefore, since $a\phi = x + y - l$

$$\phi = \frac{m'agt^2}{4m'a^2 + m(a^2 + k^2)}.$$

Also, from (4) and (5),

$$2Fa^2 = -m(a^2 - k^2)\frac{d^2x}{dt^2},$$

and consequently, from (7),

$$F = \frac{mm'(a^2 - k^2)g}{4m'a^2 + m(a^2 + k^2)}.$$

Also, from (6), we have
$$T = -\frac{m(a^2 + k^2)}{2a^2} \frac{d^2x}{dt^2}$$
$$= \frac{mm'g(a^2 + k^2)}{4m'a^2 + m(a^2 + k^2)}.$$

We will now proceed to the consideration of the case when the friction of the plane is not sufficient to prevent sliding: and since the value which we obtained for F, the friction necessary to prevent sliding, is positive, therefore this force would act during the whole motion in the direction AO, which shews that, if the action of the plane on the cylinder be not sufficient to secure perfect rolling, the dynamical friction will be exerted in the direction AO. Let μ denote the coefficient of dynamical friction; then, putting μmg instead of F in the equations (2) and (3), we have

$$m \frac{d^3x}{d\ell^3} = -T - \mu mg,$$

$$m k^3 \frac{d^3\phi}{d\ell^3} = Ta - \mu mag;$$

from these two equations, together with the equation (1), and the appropriate geometrical relations, we may easily show that

$$\phi = \frac{1}{2} \frac{(1 - 2\mu) m' - \mu m}{mk^3 + m' (a^3 + k^3)} agt^3,$$

$$x = c - \frac{1}{2} gt^3 \frac{m' (2\mu a^3 + k^3) + m\mu k^3}{mk^3 + m' (a^3 + k^3)},$$

$$y = l - c + \frac{1}{2} gt^3 \frac{m' (a^2 + k^3) - \mu m (a^3 - k^3)}{mk^4 + m' (a^3 + k^3)},$$

$$T = \frac{\mu a^2 - (1 - \mu) k^2}{mk^2 + m' (a^2 + k^2)} mm'g.$$

In a memoir by Fuss, from which this problem has been extracted, is discussed the more general case when the cylinder descends down an inclined plane, and the ring is replaced by a pulley of considerable inertia.

Fuss; Nova Acta Acad. Petrop. 1787; p. 176.

(3) A cylinder rolls, without sliding, down a moveable inclined plane, which rests on a perfectly smooth horizontal surface; to determine the motion of the plane and of the cylinder.

The axis of the cylinder is supposed to be horizontal, and a vertical plane, at right angles to the axis, to contain the centre of gravity C(fig. 186) of the cylinder and the centre of gravity of the inclined plane: let Ox be the intersection of this vertical plane with the smooth surface which supports the inclined plane AB. Draw CM at right angles to Ox. Let R be the mutual normal action and reaction of the cylinder and inclined plane,

F the action of the plane on the cylinder along AB; m = the mass of the cylinder, $mk^2 =$ its moment of inertia about its axis; m' = the mass of the inclined plane; OM = x, CM = y, OA = x'; $\theta =$ the angle through which the cylinder has revolved about its axis at the end of the time t; a = the inclination of AB to Ox, a = the radius of the cylinder.

Then, for the motion of the cylinder, we have

$$m\frac{d^3x}{dt^2}=-R\sin\alpha+F\cos\alpha.....(1),$$

$$m\frac{d^2y}{dt^2}=R\cos\alpha+F\sin\alpha-mg.....(2),$$

$$mk^3\frac{d^3\theta}{dt^3}=Fa....(3);$$

and, for the motion of the inclined plane,

$$m'\frac{d^2x'}{dt^2}=R\sin a-F\cos a....(4).$$

From the equations (1) and (4), we get

Again, from the geometry, we see that

$$y\cos\alpha=a+(x-x')\sin\alpha;$$

hence

$$\cos \alpha \frac{d^3y}{dt^3} = \sin \alpha \left(\frac{d^3x}{dt^2} - \frac{d^3x'}{dt^3} \right),$$

and therefore, by (5),

$$m'\cos\alpha\frac{d^3y}{dt^2}=(m+m')\sin\alpha\frac{d^3x}{dt^2}\dots(6).$$

Also, since no sliding takes place between the cylinder and the plane, it is clear that

$$\frac{dx}{dt} = \frac{dx'}{dt} - a \cos a \frac{d\theta}{dt},$$

and therefore

$$a\cos a \frac{d^2\theta}{dt^2} = \frac{d^2x'}{dt^2} - \frac{d^2x}{dt^2};$$

hence, by the aid of (5), we have

$$m'a \cos \alpha \frac{d^3\theta}{dt^6} = -(m+m')\frac{d^3x}{dt^6}\dots (7).$$

Again, from (1) and (2),

$$m \cos a \frac{d^3x}{dt^2} + m \sin a \frac{d^3y}{dt^2} = F - mg \sin \alpha$$

and therefore, by (8),

$$a \cos a \frac{d^2x}{dt^2} + a \sin a \frac{d^2y}{dt^2} = k^2 \frac{d^2\theta}{dt^2} - ag \sin a$$
:

substituting in this equation the expressions for $\frac{d^2y}{dt^2}$ and $\frac{d^2\theta}{dt^2}$, given in (6) and (7), we obtain

 $\{m'a^2\cos^2 a + (m+m')(a^2\sin^2 a + k^2)\}\frac{d^2x}{dt^2} = -m'a^2g\sin a\cos a$:

which gives the value of $\frac{d^3x}{dt^3}$, which it appears therefore is constant during the whole motion; the values of $\frac{d^3x}{dt^3}$, $\frac{d^3y}{dt^3}$, $\frac{d^3y}{dt^3}$,

may now be readily obtained by the aid of the equations (5), (6), (7), and will be constant during the whole motion.

Knowing the values of $\frac{d^3x}{dt^2}$, $\frac{d^3y}{dt^2}$, $\frac{d^3\theta}{dt^2}$, $\frac{d^3x'}{dt^2}$, we may immediately obtain the values of x, y, θ , x', in terms of t, if we have given the initial values of x, x', $\frac{dx}{dt}$, $\frac{dx'}{dt}$. The values of R and

F may also be readily obtained from the equations (1), (2), (3), (4).

(4) A bullet is fired with a given velocity into a body in a direction passing through the centre of gravity of the body; the body is initially at rest and is capable of free motion, not being under the action of any forces; to determine the velocities of the bullet and of the body when the bullet has traversed any space within the body; the resistance of the body to the motion of the bullet being supposed to be a constant force.

Let k denote the constant retarding force, m the mass of the bullet, μ of the body, β the initial velocity of the bullet; then if

u and v denote the velocities of the bullet and of the body when the bullet has traversed a space x within the body,

$$u = \frac{m\beta}{m+\mu} + \frac{\mu}{m+\mu} \left\{ \beta^a - \frac{2k}{m\mu} (m+\mu) x \right\}^{\frac{1}{4}},$$

$$v = \frac{m\beta}{m+\mu} - \frac{m}{m+\mu} \left\{ \beta^a - \frac{2k}{m\mu} (m+\mu) x \right\}^{\frac{1}{4}}.$$

Camus; Mém. de l'Acad. des Sciences de Paris, 1738, p. 147.

CHAPTER X.

DYNAMICAL PRINCIPLES.

SECT. 1. Vis Viva.

THE term Vis Viva was first introduced into the language of Mechanics by Leibnitz, in a memoir published in the Acta Eruditorum for the year 1695, entitled Specimen dynamicum pro admirandis naturæ legibus circa corporum vires et mutuas actiones detegendis et ad suas causas revocandis: it was intended by its author to signify the force of a body in actual motion, called otherwise its Vis Motrix or Moving Force, as distinguished from the statical pressure of a body, which has merely a tendency to motion, against a fixed obstacle; the statical force of a body he designated by the appellation of Vis Mortua. Leibnitz contended, in opposition to the received doctrine of the Cartesians, that the proper measure of the Vis Viva or Moving Force of a body, is the product of its mass into the square of its velocity, the measure adopted by the disciples of Descartes having been the same as that of the Quantity of Motion, namely, the product of the mass and the first power of the velocity. This contrariety of opinion in respect to the estimation of Moving Force, gave rise to one of the most memorable controversies in the annals of philosophy; almost all the mathematicians of Europe ultimately arranging themselves as partizans, either of the Cartesian or of the Leibnitzian doctrine. Among the adherents of Leibnitz may be mentioned John and Daniel Bernoulli, Poleni, Wolff, 's Gravesande, Camus, Muschenbroek, Papin, Hermann, Bulfinger, Kenig, and eventually Madame du Châtelet; while in the opposite ranks may be named Maclaurin, Clarke, Stirling, Desaguliers, Catalan, Robins, Mairan, and Voltaire. The Vis Motrix, or, as Leibnitz expressed it, the Vis Viva of a moving body was regarded as

a power inherent in the body, by which it is able to encounter a certain amount of resistance before losing the whole of its velocity: the question reduced itself, therefore, to the determination of an appropriate measure of this amount of resistance, to which the Moving Force was supposed to be proportional. Leibnitz regarded the product of the mass of the body and the space through which it must move, under the action of a given retarding force, to lose the whole of its velocity, as the correct measure of the whole resistance expended in the destruction of its motion, and therefore as a proper representative of the Vis Motrix or Vis Viva of the body. Now, by the theory of uniform acceleration, $mv^2 = 2mfs$, m being the mass of the body, and s the space which it must describe, under the action of a constant retarding force f, to lose the whole of its velocity v: hence it is evident that, according to the doctrine of Leibnitz, mv² will represent the body's Vis Viva. On the other hand, the Cartesians estimated the whole resistance necessary for the destruction of the body's velocity by the product of the mass of the body and the whole time of the action of the given retarding force; and therefore, by the formula mv = mft, it would follow that mv is the proper measure of the Vis Motrix, or in the language of Leibnitz, of the Vis Viva of the body. The memorable controversy of the Vis Viva, after raging for the space of about thirty years, was finally set to rest by the luminous observations of D'Alembert in the preface to his Dynamique, who declared the whole dispute to be a mere question of terms, and as having no possible connection with the fundamental principles of Mechanics. Since the publication of D'Alembert's work, the term Vis Viva has been used to signify merely the algebraical product of the mass of a moving body and the square of its velocity, while the words Moving Force have been universally employed, agreeably to the definition given by Newton in the Principia, in the signification of the product of the mass of a body and the accelerating force to which it is conceived to be subject, no physical theory whatever in regard to the absolute nature of force being supposed to be involved in these definitions. For additional information respecting the controversy of the Vis Viva, the reader is referred to Montucla's Histoire des Mathematiques, tom. III.; Hutton's

Mathematical Dictionary under the word Force; and Whewell's History of the Inductive Sciences.

The Principle of the Conservation of Vis Viva is comprehended in the following proposition: If a system of particles, any number of which are rigidly connected together, move from one position to another, either with or without constraint, under the action of finite accelerating forces, external or internal; the change of the vis viva of the whole system will be independent of the actions of the particles arising from their mutual connections, and will be equal to the sum of the changes which would be experienced by the vis viva of each particle, were it constrained to move unconnectedly from its original to its new position, under the action of the very accelerating forces to which it is subject in the actual state of the motion. This Principle immediately furnishes us with a first integral of the differential equations of motion, which is frequently of great use; especially if the co-ordinates of the position of the moving system involve only one independent variable, as in the problem of the Centre of Oscillation, when the Principle is sufficient for the complete determination of the motion.

The Principle employed by Huyghens' as the basis of his investigations on the problem of the Centre of Oscillation, constitutes under an indirect form a particular instance of the Principle of the Conservation of Vis Viva. John Bernoulli', however, was the first who enunciated the theory of the Conservation of Vis Viva, a name which he gave to the Principle, as a general law of nature, from which he deduced that of Huyghens as a particular case. Daniel Bernoulli' afterwards extended the application of the Principle to the motion of bodies subject to mutual attraction, or solicited towards fixed centres by forces varying as any functions of the distances. A demonstration of

¹ Si pendulum è pluribus ponderibus compositum, atque è quiete dimissum, partem quamcunque oscillationis integræ confecerit, atque inde porro intelligantur pondera ejus singula, relicto communi vinculo, celeritates acquisitas aursum convertere, ac quousque possunt ascendere; hoc facto, centrum gravitatis ex omnibus compositæ, ad eandem altitudinem reversum erit, quam ante inceptam oscillationem obtinebat. Horolog. Oscillator. p. 126.

¹ Opera, passim.

⁸ Mémoires de l'Académie des Sciences de Berlin, 1748.

the Principle in particular cases was first given by D'Alembert' by the aid of his general Principle of Dynamics, the same method of proof being, it was evident, of general application.

(1) A uniform rod AB (fig. 187) moves in a vertical plane, within a hemisphere; to determine its angular velocity in any of its positions, its initial position being one of instantaneous rest.

Let O be the centre of the sphere; G the middle point of AB, which will be its centre of gravity; GH a perpendicular from G upon the horizontal radius through O, which is in the plane of the rod's motion; let OG = c, AG = BG = a, k = theradius of gyration about G; OH = x, GH = y, and $\theta =$ the angle of inclination of AB to the horizon at any time t. Then, by the Principle of the Conservation of Vis Viva, m being the mass of the rod,

$$m\left(\frac{dx^3}{dt^4} + \frac{dy^3}{dt^2} + k^3 \frac{d\theta^2}{dt^4}\right) = C + 2mgy;$$

let k be the initial value of y; then, since $\frac{dx}{dt}$, $\frac{dy}{dt}$, $\frac{d\theta}{dt}$, are initially zero, we have 0 = C + 2mgh;

hence
$$\frac{dx^3}{dt^3} + \frac{dy^3}{dt^2} + R^2 \frac{d\theta^3}{dt^2} = 2g(y - h).$$

But from the geometry it is plain that

$$x = c \sin \theta$$
, $y = c \cos \theta$;

whence

$$\frac{dx}{dt} = c \cos \theta \, \frac{d\theta}{dt}, \quad \frac{dy}{dt} = -c \sin \theta \, \frac{d\theta}{dt};$$

we have, therefore,

$$(c^3+k^3)\frac{d\theta^6}{dt^2}=2og(\cos\theta-\cos\alpha),$$

a being the initial value of θ ; hence, putting for k^2 its value $\frac{1}{2}a^2$, we have, for the angular velocity of the rod in any of its positions,

$$\frac{d\theta^a}{dt^a} = \frac{6cg}{8c^3 + a^a} (\cos \theta - \cos \alpha).$$

(2) A rod PQ (fig. 188) is kept in a vertical position by means of two small rings A and A'; its lower end P is supported on an inclined plane BC, which is at liberty to move

¹ Traité de Dynamique, Seconde Partie, chap IV. p. 252.

freely on a horizontal plane; to determine the motion of the rod and the plane.

Produce QP to meet the horizontal plane in the point O; let OP = y, OB = x, at any time of the motion; h = the initial value of y, a = the inclination of the inclined plane to the vertical, m = the mass of the rod, m' = the mass of the inclined plane.

Then, by the Principle of the Conservation of Vis Viva,

$$m'\frac{dx^2}{dt^2}+m\frac{dy^2}{dt^2}=C-2gmy;$$

but, supposing the rod and the plane to be initially in a state of instantaneous rest, 0 = C - 2gmh;

hence

$$m'\frac{dx^2}{dt^2}+m\frac{dy^2}{dt^2}=2gm(h-y);$$

but, from the geometry,

$$x = y \tan a$$
, $\frac{dx}{dt} = \tan a \frac{dy}{dt}$;

hence we have

$$(m' \tan^2 a + m) \frac{dy^2}{dt^2} = 2mg (h - y),$$

$$(m' \tan^2 a + m)^{\frac{1}{4}} \frac{dy}{(h-y)^{\frac{1}{4}}} = -(2mg)^{\frac{1}{4}} dt,$$

the negative sign being taken, because y decreases as t increases: therefore, by integration,

$$2(m'\tan^2 a + m)^{\frac{1}{4}}(h - y)^{\frac{1}{4}} = C + (2mq)^{\frac{1}{4}}t:$$

but y = h when t = 0; and therefore C = 0; hence

$$2(m'\tan^2 a + m)(h - y) = mgt^2,$$

and therefore, for the value of y at any instant of the motion,

$$y = h - \frac{\frac{1}{2} mgt^2}{m + m' \tan^2 a};$$

and therefore, for the value of x,

$$x = h \tan \alpha - \frac{\frac{1}{2} mgt^2 \tan \alpha}{m + m' \tan^2 \alpha}$$

(3) AB (fig. 189) is a uniform beam, capable of moving freely about a hinge A; the extremity B rests upon an inclined

plane CE, which forms the upper surface of a body ECD; the body rests with a flat base upon a smooth horizontal plane passing through A, the vertical plane which contains AB being supposed to cut the plane surface of the body CED at right angles, and to pass through its centre of gravity; to determine the motion of the beam and the body.

Let G be the centre of gravity of AB; draw GH at right angles to the straight line ACD; let m, m', denote the masses of the beam and of the body; AH = x, GH = y, $\angle BAC = \theta$, $\angle ECD = a$, AC = x', k = the radius of gyration of AB about G. Then, by the Principle of the Conservation of Vis Viva,

$$m\left(\frac{dx^{2}}{dt^{2}}+\frac{dy^{2}}{dt^{2}}\right)+mk^{2}\frac{d\theta^{2}}{dt^{2}}+m'\frac{dx^{2}}{dt^{2}}=C-2mgy;$$

but from the geometry we see that

$$x = a \cos \theta$$
, $y = a \sin \theta$, $x' = \frac{2a}{\sin a} \sin (a - \theta)$...(1);

hence we have

$$m (a^{2} + k^{2}) \frac{d\theta^{2}}{dt^{2}} + \frac{4a^{2}}{\sin^{2} a} m' \cos^{2} (a - \theta) \frac{d\theta^{2}}{dt^{2}} = C - 2mga \sin \theta,$$

$$\{m (a^{2} + k^{2}) \sin^{2} a + 4m'a^{2} \cos^{2} (a - \theta)\} \frac{d\overline{\theta}^{2}}{dt^{2}} = \sin^{2} a (C - 2mga \sin \theta);$$
let β be the value of θ when $\frac{d\theta}{dt} = 0$; then

$$0 = \sin^2 \alpha \ (C - 2mga \sin \beta),$$

and therefore we get

$$\{m (a^2+k^2)\sin^2 a + 4m'a^2\cos^2(a-\theta)\}\frac{d\theta^2}{dt^2} = 2mag\sin^2 a (\sin\beta-\sin\theta),$$
 which gives the value of $\frac{d\theta^2}{dt}$ for any assigned value of θ ; whence, by the aid of the equations (1), we may obtain the values of $\frac{dx}{dt}$, $\frac{dy}{dt}$, $\frac{dx'}{dt}$, for any position of the beam.

(4) A uniform lever ACB, (fig. 190), of which the arms AC and BC are at right angles to each other, rests in equilibrium when AC is inclined at a given angle to the horizon; if AC be raised to a horizontal position, C being fixed, to find the angle through which it will fall.

Let CA = 2a, CB = 2a', m = the mass of AC, m' = the mass of BC; let θ , θ' , be the inclinations of CA, CB, to the horizon, at any time of the motion.

Then the vis viva of the lever will be equal to

$$2mag \sin \theta + 2m'a'g \sin \theta' + C;$$

but, when $\theta = 0$ and therefore $\theta' = \frac{1}{2}\pi$, the vis viva is equal to zero; hence 0 = 2m'a'g + C;

hence the vis viva for any position of the lever is equal to

$$2mag \sin \theta + 2m'a'g \sin \theta' - 2m'a'g.$$

Now, when the value of θ is a maximum, the vis viva will again become zero; hence, for the required value of θ ,

$$ma \sin \theta + m'a' \sin \theta' = m'a' \dots (1).$$

Let β , β' , be the values of θ , θ' , for the equilibrium of the lever; then $ma \cos \beta = m'a' \cos \beta'$;

hence from (1) there is

$$\cos \beta' \sin \theta + \cos \beta \sin \theta' = \cos \beta$$
,

or, since
$$\beta' = \frac{1}{2}\pi - \beta$$
, $\theta' = \frac{1}{2}\pi - \theta$,

 $\sin \beta \sin \theta + \cos \beta \cos \theta = \cos \beta$, $\cos (\theta - \beta) = \cos \beta$; and therefore $\theta = 2\beta$, the angle through which CA falls.

(5) To determine the motion of a pendulum, the axis of which is a cylinder resting upon two perfectly rough planes which coincide with the same horizontal plane, the cylindrical axis being thus capable of rolling along the planes.

Let C (fig. 191) be the centre of a vertical section of the cylindrical axis made by a plane containing the centre of gravity of the pendulum; C may be regarded as the centre of gravity of the axis. Let G be the centre of gravity of the pendulum and cylinder together, and mk^2 their moment of inertia about a horizontal line through G parallel to the axis, m denoting the sum of their masses. Let GH be drawn at right angles to the horizontal plane along which the axis rolls; let G be the point of contact of the section G of the axis with this plane at any time of the motion, G being the position of G corresponding to the equilibrium of the system. Let GG = c, GG = a, $GGK = \phi$, GK being a vertical line, GH = x, GH = y.

Then the vis viva of the system at the time t due to the motion of G will be $m\left(\frac{dx^2}{dt^2} + \frac{dy^2}{dt^2}\right)$, and the vis viva due to rotation about G will be $mk^2\frac{d\phi^2}{dt^2}$; hence the whole vis viva of the system will be equal to

$$m\left(k^2\frac{d\phi^2}{dt^2}+\frac{dx^2}{dt^2}+\frac{dy^2}{dt^2}\right):$$

also the sum of the products of the mass of each molecule of the system into the vertical space through which it has descended, will be equal to my together with some constant quantity depending upon the initial circumstances of the system. Hence, by the Principle of Conservation of Vis Viva,

$$m\left(k^2\frac{d\varphi^2}{dt^2}+\frac{dx^3}{dt^2}+\frac{dy^2}{dt^3}\right)=C+2gmy;$$

but from the geometry it is evident that

$$x = a \sin \phi - c\phi$$
, $y = a \cos \phi - c$,

and therefore

$$\frac{dx}{dt} = (a\cos\phi - c)\frac{d\phi}{dt}, \quad \frac{dy}{dt} = -a\sin\phi\frac{d\phi}{dt};$$

hence we have

$$(a^2 + c^2 + k^2 - 2ac \cos \phi) \frac{d\phi^2}{dt^2} = C' + 2g(a \cos \phi - c);$$

let a be the maximum value of ϕ , then

$$0 = C' + 2g (a \cos a - c),$$

and therefore

$$(a^2 + c^2 + k^2 - 2ac \cos \phi) \frac{d\phi^2}{dt^2} = 2ga (\cos \phi - \cos \alpha),$$

which gives the angular velocity of the pendulum for every position which it assumes during its motion.

For the period of a semi-oscillation we have

$$T = \frac{1}{(2ag)^{\frac{1}{4}}} \int_{0}^{x} \frac{(a^{2} + c^{2} + k^{2} - 2ac \cos \phi)^{\frac{1}{4}}}{(\cos \phi - \cos a)^{\frac{1}{4}}} d\phi \dots (1).$$

This integration cannot be effected except, which we will suppose to be the case, the amplitude of the pendulum's oscillation is very small.

Assume then
$$\sin \frac{\phi}{2} = s$$
, $\sin \frac{\alpha}{2} = b$;

whence $\cos \phi = 1 - 2 \sin^2 \frac{\phi}{2} = 1 - 2s^2$,

$$\cos \alpha = 1 - 2\sin^2\frac{\alpha}{2} = 1 - 2b^2,$$

and

$$d\phi = \frac{4sds}{\sin \phi} = \frac{4sds}{(1 - \cos^2 \phi)^{\frac{1}{4}}} = \frac{2ds}{(1 - s^2)^{\frac{1}{4}}}.$$

Hence, from (1),

$$T = \frac{1}{(2ag)^{\frac{1}{4}}} \int_{0}^{s} \frac{(a^{2} + c^{2} + k^{2} - 2ac + 4ac s^{2})^{\frac{1}{4}}}{(2b^{2} - 2s^{2})^{\frac{1}{4}}} \cdot \frac{2ds}{(1 - s^{2})^{\frac{1}{4}}},$$

and therefore, putting $(a - c)^2 + k^2 = h^2$,

$$T = \frac{1}{(ag)^{\frac{1}{4}}} \int_{0}^{b} \frac{(h^{2} + 4ac \ s^{2})^{\frac{1}{4}} ds}{(1 - s^{2})^{\frac{1}{4}} (b^{2} - s^{2})^{\frac{1}{4}}};$$

but s being a small quantity, we have, neglecting small quantities of orders higher than the second,

$$\frac{1}{(1-s^2)^{\frac{1}{4}}} = (1-s^2)^{-\frac{1}{4}} = 1 + \frac{1}{2}s^2,$$

$$(h^2 + 4ac s^2)^{\frac{1}{2}} = h\left(1 + \frac{2ac}{h^2}s^2\right),$$

and therefore

$$\frac{\left(h^2 + 4ac\ s^2\right)^{\frac{1}{4}}}{\left(1 - s^2\right)^{\frac{1}{4}}} = h\left(1 + \frac{4ac + h^2}{2h^2}s^2\right) \qquad \text{nearly}:$$

hence

$$T = \frac{h}{(ag)^{\frac{1}{4}}} \int_{0}^{b} ds \left\{ \frac{1}{(b^{2} - s^{2})^{\frac{1}{4}}} + \frac{4ac + h^{2}}{2h^{2}} \frac{s^{2}}{(b^{2} - s^{2})^{\frac{1}{4}}} \right\}$$

$$= \frac{h}{(ag)^{\frac{1}{4}}} \left\{ \frac{\pi}{2} + \frac{4ac + h^{2}}{2h^{2}} \frac{\pi b^{2}}{4} \right\}$$

$$= \frac{\pi h}{2(ag)^{\frac{1}{4}}} + \frac{\pi b^{2} (4ac + h^{2})}{8h (ag)^{\frac{1}{4}}};$$

and therefore the period of a complete oscillation is equal to

$$\frac{\pi h}{(ag)^{\frac{1}{4}}} + \frac{\pi b^{2} (4ac + h^{2})}{4h (ag)^{\frac{1}{4}}}.$$

Euler has discussed this problem, starting with the general equations of motion, and investigated the pressure on the plane at any time, as well as the horizontal action of the plane upon the cylinder which shall be sufficient to prevent sliding.

Euler; Nova Acta Acad. Petrop. 1788; p. 145.

(6) Two equal particles are attached to the extremities A, A', (fig. 192), of a straight lever ACA' having equal arms without weight, and are each attracted to a centre of force in O which varies inversely as the cube of the distance; CO is vertical and equal to CA or CA'; supposing the lever to be placed originally in a given position, to find the time of its becoming vertical.

Let CO = CA = a; OA = r and OA' = r' at any time of the motion; $\angle ACO = \theta$; m = the mass of each of the particles, $\mu =$ the attraction of the force in O upon a unit of matter collected in a point at a unit of distance; $\alpha =$ the initial value of θ .

Then the vis viva of the two particles together will be, at any time t, $2ma^2 \frac{d\theta^2}{dt^3}$; hence

$$2ma^{3}\frac{d\theta^{2}}{d\ell^{3}}=\int\left(-\frac{\mu m}{r^{3}}dr-\frac{\mu m}{r'^{3}}dr'\right)+C,$$

and therefore

$$2a^3\frac{d\theta^3}{dt^3}=\frac{\mu}{r^3}+\frac{\mu}{r'^3}+C''.$$

Now from the geometry we have

$$r^{2} = 2a^{2}(1 - \cos \theta), \quad r^{2} = 2a^{2}(1 + \cos \theta);$$

hence

$$2a^{2} \frac{d\theta^{3}}{dt^{3}} = \frac{\mu}{2a^{3}} \left(\frac{1}{1 - \cos \theta} + \frac{1}{1 + \cos \theta} \right) + C'$$
$$= \frac{\mu}{a^{3}} \frac{1}{\sin^{2} \theta} + C'';$$

but when $\theta = a$, $\frac{d\theta}{dt} = 0$; hence

$$0=\frac{\mu}{a^2}\,\frac{1}{\sin^2 a}+C',$$

and therefore

$$2a^2\frac{d\theta^3}{dt^2} = \frac{\mu}{a^2\sin^2\alpha} \frac{\sin^2\alpha - \sin^2\theta}{\sin^2\theta};$$

hence, $\frac{d\theta}{dt}$ being negative because θ decreases with the increase of t, $\left(\frac{2}{\mu}\right)^{\frac{1}{2}} a^{2} \frac{\sin \theta \ d\theta}{\left(\sin^{2} a - \sin^{2} \theta\right)^{\frac{1}{2}}} = -\frac{dt}{\sin a}$

$$\left(\frac{2}{\mu}\right)^{\frac{1}{2}} a^{2} \frac{d \cos \theta}{\left(\cos^{2} \theta - \cos^{2} a\right)^{\frac{1}{2}}} = \frac{dt}{\sin a};$$

integrating we get

$$\left(\frac{2}{\mu}\right)^{\frac{1}{4}} a^{2} \log \left\{\cos \theta + \left(\cos^{2} \theta - \cos^{2} a\right)^{\frac{1}{4}}\right\} = \frac{t}{\sin a} + C;$$

but $\theta = a$ when t = 0; hence $C = \left(\frac{2}{\mu}\right)^{\frac{1}{2}} a^2 \log \cos a$; and therefore

$$\left(\frac{2}{\mu}\right)^{\frac{1}{4}} a^{2} \log \frac{\cos \theta + \left(\cos^{2} \theta - \cos^{4} \alpha\right)^{\frac{1}{4}}}{\cos \alpha} = \frac{t}{\sin \alpha}.$$

When ACA' becomes vertical, $\theta = 0$, and we have for the required value of t,

$$t = \left(\frac{2}{\mu}\right)^{\frac{1}{2}} a^{2} \sin a \log \frac{1 + \sin a}{\cos a} = \left(\frac{2}{\mu}\right)^{\frac{1}{2}} a^{2} \sin a \log \tan \frac{\pi + 2a}{4}.$$

(7) BFG (fig. 193) is a heavy body of any form, of which C is the centre of gravity; an inextensible string attached to a fixed point E is wound about the circumference of a circle ALH, having C for its centre, and representing an axis; EA is vertical; to determine the velocity of C when the body has descended from rest through a given altitude, under the action of gravity, by the uncoiling of the string.

Let a be the radius of the axis, k the radius of gyration of the body about C; v the velocity acquired by C, after descending through a space x. Then

$$v^2 = \frac{2ga^2x}{a^2+k^2}.$$

This problem is one of the 'Theoremata Selecta,' given by John Bernoulli, 'pro conservatione virium vivarum demonstranda et experimentis confirmanda.

Comment. Acad. Petrop. 1727, p. 200. Opera, tom. 111. p. 127.

(8) A particle A (fig. 194) descends down the curve CKA, drawing a particle B up the curve CLB by means of a string passing over the point C; to determine the velocities of the particles after moving from rest through any corresponding spaces.

Let m, m', be the masses of A, B, respectively; v, v', their velocities after moving through vertical spaces equal to y, y'; then, ds, ds', denoting elements of the two curves,

$$v^2 = 2g(my - m'y') \frac{ds^2}{mds^2 + m'ds'^2}, \ v'^2 = 2g(my - m'y') \frac{ds'^2}{mds^2 + m'ds'^2}.$$

John Bernoulli; Act. Erudit. Lips. 1735. Mai. p. 210; Opera, tom. 111. p. 257. Hermann; Mémoires de St. Pétersbourg, tom. 11. D'Alembert; Traité de Dynamique, p. 123; Seconde Edition.

SECT. 2. Vis Viva and the Conservation of the Motion of the Centre of Gravity.

The Principle of the Conservation of the Motion of the Centre of Gravity, under its most general form, asserts that, the motion of the centre of gravity of a free system of bodies disposed relatively to each other in any conceivable manner, is always the same as if the bodies were all united in the centre of gravity, and at the same time each of them were animated by the same accelerating forces as in their actual state. The discovery of the Principle is due to Newton, by whom it received a demonstration in the particular case where the system is subject to no external force, when the centre of gravity will either remain at rest or move in a straight line with a uniform velocity. D'Alembert' afterwards extended the Principle to the case where each body is supposed to be solicited by a constant accelerating force acting in parallel lines, or directed towards a fixed point and varying as the dis-Finally, Lagrange' expressed the Principle under its most general form for every law of force to which the bodies can be subject.

¹ Principia; Axiomata sive Leges Motus, Cor. 4.

³ Traité de Dynamique, Seconde Partie, Chap. 11.

Mécanique Analytique, tom. i. p. 257, &c.

(1) A smooth groove KAL (fig. 195) is carved in a vertical plane in the body KBCL, which is placed upon a smooth horisontal plane, along which it is able to slide freely; to find the form of the groove that a heavy particle, placed within it, may oscillate in it tautochronously, the time of an oscillation being given.

Let P be the place of the particle in the groove at any time; draw PN vertically to meet the horizontal plane in N, which will lie in the line OE formed by the intersection of a vertical plane through the groove with the horizontal plane. Let A be the lowest point of the groove, draw AM horizontally, AA' vertically. Let O be a fixed point in OE; OA' = x', $ON = x_1$, $PN = y_1$, AM = x, PM = y; let k_1 , k, be the initial values of y_1 , y; m = the mass of the particle, m' = the mass of the body.

Then, by the Principle of the Conservation of the Motion of the Centre of Gravity, since no forces act upon the particle and body parallel to OE,

Also, by the Principle of the Conservation of Vis Viva,

$$m' \frac{dx'^2}{dt'^2} + m \left(\frac{dx_1^2}{dt'^2} + \frac{dy_1^2}{dt'^2}\right) = 2mg(k_1 - y_1)....(2).$$

But, from the geometry, it is evident that

$$\frac{dx_1}{dt} = \frac{dx'}{dt} + \frac{dx}{dt} \dots (3),$$

and

From (1) and (3) we have

$$\frac{dx_1}{dt} = \frac{m'}{m+m'} \frac{dx}{dt}, \quad \frac{dx'}{dt} = -\frac{m}{m+m'} \frac{dx}{dt} \cdot \dots (5).$$

Hence, from (2), (4), (5), we see that

$$\frac{m'}{m+m'} \frac{dx^2}{dt^2} + \frac{dy^3}{dt^2} = 2g(k-y);$$

and therefore, if r denote the time of a semi-oscillation,

$$\tau = -\frac{1}{(2g)^{\frac{1}{4}}} \int_{k}^{c} \left(\frac{m'}{m+m'} \frac{dx^{2}}{dy^{3}} + 1 \right)^{\frac{1}{4}} dy \dots (6).$$

This value of τ must be independent of k in order that the particle may oscillate tautochronously, and therefore we must have, it being necessary that the coefficient of dy be of -1 dimensions in y and k,

$$\left(\frac{m'}{m+m'} \frac{dx^2}{dy^2} + 1\right)^{\frac{1}{4}} = \frac{a}{v^{\frac{1}{4}}} \dots \dots (7),$$

where a is a constant quantity; hence

$$\frac{dx}{dy} = \left(\frac{m+m'}{m'}\right)^{\frac{1}{4}} \left(\frac{a^2-y}{y}\right)^{\frac{1}{4}},$$

and therefore, by integration,

$$x = \left(\frac{m + m'}{m'}\right)^{\frac{1}{4}} \left\{ \left(\alpha^{2}y - y^{2}\right)^{\frac{1}{4}} + \frac{1}{2}\alpha^{2} \text{ vers}^{-1} \frac{2y}{\alpha^{2}} \right\} \dots (8).$$

But, from (6) and (7),

$$\tau = -\frac{1}{(2g)^{\frac{1}{4}}} \int_{a}^{b} \frac{a \, dy}{(ky - y^2)^{\frac{1}{4}}} = \frac{\pi a}{(2g)^{\frac{1}{4}}}, \quad a^2 = \frac{2g\tau^2}{\pi^2},$$

and therefore from (8) we get, for the equation to the groove,

$$x = \left(\frac{m + m'}{m'}\right)^{\frac{1}{4}} \left\{ \left(\frac{2g\tau^{2}}{\pi^{2}} y - y^{2}\right)^{\frac{1}{4}} + \frac{g\tau^{2}}{\pi^{2}} \text{ vers}^{-1} \frac{\pi^{2}y}{g\tau^{2}} \right\}.$$

Clairaut; Mémoires de l'Académie des Sciences de Paris, 1742, p. 41. Euler; Opuscula, de motu corporum tubis mobilibus inclusorum, p. 48.

SECT. 3. Vis Viva and the Conservation of Areas.

The Principle of the Conservation of Areas asserts, that if a system of particles be subject only to mutual actions, the sum of the products of the mass of each particle into the projection (on any proposed plane) of the area described by its radius vector round any assigned point, is proportional to the time. The same

principle holds good also if the system be subject to external forces, provided that they be such that the algebraical sum of their moments about a line through the assigned point at right angles to the proposed plane be zero. This principle, which is in fact a generalization of Newton's theorem respecting the areas described by a single body about a centre of force, was discovered, about the same time, by Euler, Daniel Bernoulli, and D'Arcy; the enunciation of the Principle given by Euler and Bernoulli being expressed under a form somewhat different from that given by D'Arcy, under which it is now generally expressed. The discovery of the Principle was suggested to these three mathematicians by the consideration of the problem of the motion of several bodies within a tube of given form, moving about a fixed point.

(1) P, II, (fig. 196), are two material particles attached to an inflexible straight line POII, moveable in a horizontal plane about a fixed point O; the particle II is fixed to the inflexible line, while the particle P is capable of aliding along it; to determine the path described by P, corresponding to any initial velocities of the particles.

Let OE be an immoveable straight line passing through O; PO = r, $\Pi O = a$, m = the mass of P, $\mu =$ the mass of Π , $\angle POE = \theta$. Then, by the Principle of the Conservation of Areas, since the only force to which the moving system is subject is the reaction of the fixed point O, we have

$$(mr^2 + \mu a^2) \frac{d\theta}{dt} = C \dots (1),$$

where C is some constant quantity.

Again, by the Principle of the Conservation of Vis Viva,

$$m\left(\frac{dr^2}{dt^2}+r^2\frac{d\theta^2}{dt^2}\right)+\mu a^2\frac{d\theta^2}{dt^2}=C',$$

$$m\frac{dr^2}{dt^2}+(mr^2+\mu a^2)\frac{d\theta^2}{dt^2}=C'....(2),$$

C' being a constant quantity.

- 1 Opuscula, de motu corporum tubis mobilibus inclusorum, p. 48, 1746.
- ¹ Mémoires de l'Académie des Sciences de Berlin, 1745, p. 54.
- ³ Mémoires de l'Académie des Sciences de Paris, 1747, p. 348.

Eliminating dt between (1) and (2), we obtain

$$m \frac{dr^3}{d\theta^3} + mr^2 + \mu a^2 = \frac{C'}{C^3} (mr^2 + \mu a^2)^3$$

$$m\frac{dr^{2}}{d\theta^{2}} = \left\{\frac{C'}{C^{2}}(mr^{2} + \mu a^{2}) - 1\right\}^{\bullet}(mr^{2} + \mu a^{2}),$$

which is the differential equation to P's path.

In order to determine C and C', suppose that b, ω , u, are the initial values of r, $\frac{d\theta}{dt}$, $\frac{dr}{dt}$, respectively. Then, from (1),

$$(mb^2 + \mu a^2) \omega = C,$$

which determines C; and, from (2),

$$mu^2 + (mb^2 + \mu a^3) \omega^2 = C',$$

which determines C.

Clairaut; Mém. de l'Acad. des Sciences de Paris, 1742, p. 22. D'Arcy; Mém. de l'Acad. des Sciences de Paris, 1747, p. 351. D'Alembert; Traité de Dynamique, p. 104, seconde edit.

(2) A straight rod PQ, (fig. 197), subject to the condition of always passing through a small fixed ring at O, is in motion on a horizontal plane; to determine the path of its centre of gravity G.

At any time t of the motion let OG = r, $\angle GOE = \theta$, OE being a fixed line in the plane. Let m be the mass of an element of the rod at any distance ρ from O, and let μ be the mass of the whole rod.

Then, by the Principle of the Conservation of Areas, the only force which acts on the rod being the reaction of the ring,

$$C = \Sigma \left(m\rho^2 \frac{d\theta}{dt} \right) = \Sigma \left(m\rho^2 \right) \frac{d\theta}{dt} = \mu \left(r^2 + k^2 \right) \frac{d\theta}{dt} \dots (1),$$

k being the radius of gyration of the rod about its centre of gravity, and C a constant quantity.

Again, by the Principle of the Conservation of Vis Viva, the ring being considered perfectly smooth,

$$C' = \mu \left(\frac{dr^2}{dt^2} + r^2 \frac{d\theta^2}{dt^2} \right) + \mu k^2 \frac{d\theta^2}{dt^2}$$

$$= \mu \frac{dr^2}{dt^2} + \mu (r^2 + k^2) \frac{d\theta^2}{dt^2} \dots (2),$$

C' being a constant quantity.

Eliminating dt between (1) and (2), we have

$$\frac{dr^2}{d\theta^2} + r^2 + k^2 = \frac{\mu C'}{C^2} (r^2 + k^2)^2,$$

$$\frac{dr^2}{d\theta^2} = \left\{ \frac{\mu C'}{C^2} (r^2 + k^2) - 1 \right\} (r^2 + k^2) \dots (3).$$

Suppose that a, u, ω , are the initial values of r, $\frac{dr}{dt}$, $\frac{d\theta}{dt}$, respectively; then, by (1) and (2),

$$C = \mu (a^2 + k^2) \omega$$
, $C' = \mu u^2 + \mu (a^2 + k^2) \omega^2$:

hence the equation (3) becomes

$$\begin{split} \frac{dr^2}{d\theta^2} &= \left\{ \frac{u^2 + (a^2 + k^2)}{(a^2 + k^2)^2} \frac{\omega^2}{\omega^2} (r^2 + k^2) - 1 \right\} (r^2 + k^2) \\ &= \frac{r^2 + k^2}{\omega^2 (a^2 + k^2)^2} \left\{ (r^2 + k^2) u^2 + \omega^2 (a^2 + k^2) (r^2 - a^2) \right\}, \end{split}$$

which is the differential equation to the path of G.

Clairaut; Mémoires de l'Acad. des Sciences de Paris.
1742, p. 38-41.

(3) Two equal particles P, P, (fig. 198), are attached to the extremities of a rod PP; the middle point O of the rod is fixed; the rod is able to move in every direction about O; to determine the motion of the particles corresponding to any initial circumstances, the weight of the rod being neglected.

Through the point O draw a straight line AOB; with O as a centre and radius equal to OP, describe the two indefinite circular arcs APk, Al, the latter of which is supposed to lie within an assigned plane. Let OP=a, $\angle AOP=\phi$, $\angle kAl=\theta$; m= the mass of each of the particles. Then, t denoting the corresponding time, we shall have, by the Principle of the Conservation of Vis Viva, whether the particles be subject to the action of gravity or not,

$$2ma^{3}\left(\frac{d\phi^{3}}{dt^{3}} + \sin^{2}\phi \frac{d\theta^{3}}{dt^{2}}\right) = C,$$

$$\frac{d\phi^{3}}{dt^{2}} + \sin^{2}\phi \frac{d\theta^{3}}{dt^{2}} = c \dots (1),$$

or

C, c, being constant quantities.

Again, by the Principle of the Conservation of Areas, we have

$$2ma^2\sin^2\phi\;\frac{d\theta}{dt}=C_1,$$

or

$$\sin^2 \phi \, \frac{d\theta}{dt} = c_1 \, \dots \, (2),$$

 C_1 , c_1 , being constants.

Eliminating $\frac{d\theta}{dt}$ between the equations (1) and (2) we get

$$\frac{d\phi^2}{dt^2} + \frac{c_1^2}{\sin^2\phi} = c,$$

and therefore

$$\sin \phi \ d\phi = \left(c \sin^2 \phi - c_1^2\right)^{\frac{1}{2}} \ dt,$$

$$(c-c_1^2-c\cos^2\phi)^{\frac{1}{4}}dt=-d\cos\phi;$$

integrating, and adding an arbitrary constant c_2 ,

$$t + c_2 = \frac{1}{c^{\frac{1}{4}}} \cos^{-1} \frac{c^{\frac{1}{4}} \cos \phi}{(c - c_1^2)^{\frac{1}{4}}} \dots (3),$$

$$\cos \phi = \frac{(c-c_1^2)^{\frac{1}{4}}}{c^{\frac{1}{4}}}\cos \{c^{\frac{1}{4}}(t+c_2)\}...(4).$$

Suppose that when t = 0, $\phi = \beta$, $\frac{d\theta}{dt} = \omega$, $\frac{d\phi}{dt} = \omega'$; then

$$c = \omega^2 + \omega^2 \sin^2 \beta$$
, from (1);

$$c_1 = \omega \sin^2 \beta$$
, from (2)

and therefore, from (3),

$$c_2 = \frac{1}{\left(\omega^2 + \omega^2 \sin^2 \beta\right)^{\frac{1}{2}}} \cos^{-1} \frac{\left(\omega'^2 + \omega^2 \sin^2 \beta\right)^{\frac{1}{2}} \cos \beta}{\left(\omega'^2 + \omega^2 \sin^2 \beta \cos^2 \beta\right)^{\frac{1}{2}}}.$$

The value of $\cos \phi$ being given by (4), we may then, by the aid of (2), get an expression for θ in terms of t.

(4) A spherical shell, the interior radius of which is the nth of the exterior, is filled with fluid of the same density with itself; to compare the space through which it would roll from rest in a given time: the appropriate plane, with that which would be described by a solid sphere of the same size and weight rolling down the same plane.

Let m, m', denote the masses of the shell and fluid respectively; a, a', the exterior and interior radii; k, k', the radii of gyration of the shell and fluid about a diameter of the sphere; a the inclination of the plane to the horizon; θ , θ' , the angles through which the shell and fluid have revolved about their common centre of gravity from the beginning of the motion; x the space described by the centre of the sphere.

Then, by the Principle of the Conservation of Vis Viva,

$$m\left(\frac{dx^3}{dt^3}+k^2\frac{d\theta^3}{dt^3}\right)+m'\left(\frac{dx^3}{dt^3}+k'''\frac{d\theta''^3}{dt^3}\right)=C+2\left(m+m'\right)gx\sin\alpha;$$

but, since the resultant of all the forces which act on the fluidsphere passes through its centre of gravity, we have, by the Principle of the Conservation of Areas,

$$m'k'^2\frac{d\theta'}{dt}=C';$$

hence, from these two equations,

$$m\left(\frac{dx^{2}}{dt^{2}}+k^{2}\frac{d\theta^{2}}{dt^{2}}\right)+m'\frac{dx^{2}}{dt^{2}}=C''+2(m+m')gx\sin a;$$

but, since the sphere rolls, we have $x = a\theta$; hence, putting $m + m' = \mu$,

$$(\mu a^2 + mk^2) \frac{dx^2}{dt^2} = C''a^2 + 2\mu a^2gx \sin a;$$

differentiating with respect to t, and dividing by $2 \frac{dx}{dt}$,

$$(\mu a^2 + mk^2) \frac{d^2x}{dt^2} = \mu a^2g \sin a \dots (1).$$

Now,
$$mk^2 = \frac{2}{5} \mu a^2 - \frac{2}{5} m'a'^2 = \frac{2}{5} (\mu a^2 - m'a'^2);$$

but
$$m' = \frac{\mu}{n^3}$$
 and $a' = \frac{a}{n}$; hence

$$mk^3 = \frac{2}{5} \mu a^3 \frac{n^5 - 1}{n^5} .$$

Substituting this value of mk² in (1), we obtain

$$\left(1 + \frac{2}{5} \frac{n^5 - 1}{n^5}\right) \frac{d^2x}{dt^2} = g \sin a,$$

$$(7n^5 - 2) \frac{d^2x}{dt^2} = 5n^5 g \sin a;$$

integrating, and bearing in mind that $\frac{dx}{dt} = 0$, x = 0, when t = 0, we get

$$(7n^5-2) x = \frac{5}{2} n^5 g t^2 \sin a$$
.

In the case of the solid sphere, we shall obtain in a similar manner, $7x' = \frac{5}{2}gt^2 \sin a$,

x' denoting the space through which it has rolled along the plane at the end of the time t.

Hence finally we get
$$\frac{x}{x'} = \frac{7n^6}{7n^6 - 2}$$
.
If $n = 2$, we have $\frac{x}{x'} = \frac{112}{111}$.

Lady's and Gentleman's Diary, 1842, p. 51.

(5) The particle m (fig. 199) is connected with the particles m' and m' by means of two fine inextensible strings mOm', mOm', passing through a small smooth ring at O; m, m', m', all lie on a smooth horizontal plane passing through O; to determine the tensions of the two strings and the motions of the particles, supposing the particles to have received any initial impulses such as, at least at the commencement of the motion, to keep the strings at full stretch.

Draw through O a straight line OA in the plane of the motions; let Om=r, Om'=r', Om'=r', $\angle mOA=\theta$, $\angle m'OA=\theta'$, $\angle m'OA=\theta'$, at any time t; T'= the tension of the string mOm', and T'' of the string mOm'.

Then, since the only forces which act upon the particles pass through O, we have, by a formula in the theory of Central Forces,

$$\frac{d^{2}r'}{dt'} = r' \frac{d\theta'^{2}}{dt'} - \frac{T'}{m'},$$

$$\frac{d^{2}r'}{dt'} = r' \frac{d\theta'^{2}}{dt'^{2}} - \frac{T'}{m},$$

$$\frac{d^{2}r}{dt''} = r \frac{d\theta^{2}}{dt'^{2}} - \frac{T' + T'}{m}.$$
(1)

But, the strings being supposed to be kept at full stretch, we have

$$r + r' = c', \quad r + r'' = c', \ldots (2);$$

where d', c', denote the lengths of the strings mOm', mOm'; and therefore

$$\frac{d^3r}{dt^2} + \frac{d^3r'}{dt^2} = 0, \quad \frac{d^2r}{dt^2} + \frac{d^2r'}{dt^2} = 0;$$

hence, by the first and third of the formulæ (1),

$$\frac{T'}{m'} + \frac{T' + T'}{m} = r \frac{d\theta^2}{dt^2} + r' \frac{d\theta'^2}{dt^3} \dots (3);$$

and, by the second and third,

$$\frac{T}{m^a} + \frac{T' + T''}{m} = r \frac{d\theta^2}{dt^2} + r'' \frac{d\theta''^2}{dt^2} \dots (4).$$

Again, since the only forces which act upon the three particles pass through the point O, we have, by the Principle of the Conservation of Areas,

$$r^2\frac{d\theta}{dt}=e, \quad r'^2\frac{d\theta'}{dt}=e', \quad r''^2\frac{d\theta'}{dt}=e'', \ldots (5);$$

where e, e', e'', are invariable quantities: hence, from (3) and (4), we have

$$\frac{T'}{m'} + \frac{T' + T''}{m} = \frac{e^3}{r^3} + \frac{e'^2}{r'^3},$$

$$\frac{T''}{m'} + \frac{T' + T''}{m} = \frac{e^2}{r^3} + \frac{e'^2}{r'^3};$$

from these two equations we may readily ascertain that

$$(m + m' + m'') \frac{T'}{m'} = \frac{me^2}{r^3} + \frac{(m + m'') e'^2}{r'^3} - \frac{m''e'^3}{r'^3},$$

$$(m + m' + m'') \frac{T''}{m''} = \frac{me^2}{r^3} + \frac{(m + m'') e'^2}{r'^3} - \frac{m'e'^2}{r'^3},$$

$$(m + m' + m'') \frac{T' + T''}{m} = \frac{(m' + m'') e^2}{r^3} + \frac{m'e'^2}{r'^3} + \frac{m''e'^3}{r'^3},$$

$$(6),$$

which give the values of the tensions of the two strings mOm', mOm', and of the double string Om. It is important to observe that these values for the tensions hold good only so long as both the strings are at full stretch; if either of the strings become slack at any epoch of the motion, these formulæ will no longer apply; this will be evident when it is considered that in obtaining them we made use of the equations (2) which are grounded on the supposition that the strings are at full stretch. The formulæ themselves will indicate the epoch at which their inapplicability may commence by giving a zero value for either T' or T.

Again, by the Principle of Vis Viva,

$$m\left(r^{2}\frac{d\theta^{2}}{dt^{2}}+\frac{dr^{2}}{dt^{2}}\right)+m'\left(r'^{2}\frac{d\theta'^{2}}{dt^{2}}+\frac{dr'^{2}}{dt^{2}}\right)+m'\left(r''^{2}\frac{d\theta''^{2}}{dt^{2}}+\frac{dr''^{2}}{dt^{2}}\right)=C,$$

where C is some constant quantity: hence, observing that, by the equations (2), $\frac{dr'}{dt} = -\frac{dr}{dt} = \frac{dr'}{dt}$, we get

$$mr^{2}\frac{d\theta^{2}}{dt^{2}} + m'r'^{2}\frac{d\theta'^{2}}{dt^{3}} + m'r'^{2}\frac{d\theta'^{3}}{dt^{3}} + (m + m' + m'')\frac{dr^{3}}{dt^{2}} = C,$$

and therefore, by the equations (5),

$$\frac{me^2}{r^2} + \frac{m'e'^2}{r'^2} + \frac{m''e''^2}{r''^2} + (m+m'+m'')\frac{dr^2}{d\theta^2}\frac{e^3}{r^4} = C....(7),$$

and thence, by the equations (2), putting $m + m' + m'' = \mu$,

$$\frac{me^2}{r^2} + \frac{m'e'^2}{(c'-r)^2} + \frac{m'e''^2}{(c''-r)^2} + \frac{\mu e^2}{r^4} \frac{dr^2}{d\theta^2} = C \dots (8),$$

which is the differential equation to the path of m. Similarly may be obtained the differential equations to the paths of m' and m'. These equations will evidently cease to define the paths of the particles if at any time either of the strings become slack, or either T' or T' become zero. If either of the strings become slack at any time, then we shall have to investigate the motions of the two particles whose connecting string is not slack, the particle which belongs to the loose string moving along for a time without constraint. From the equation (7) it is evident that none of the quantities r, r', r', can ever become zero; or that the particles, so long as the strings are tight, will none of them arrive at the point O.

Suppose that ω , ω , σ , a, β , are the initial values of $\frac{d\theta}{dt}$, $\frac{d\theta}{dt}$,

 $\frac{d\theta'}{dt}$, r, $\frac{dr}{dt}$; then, from the equations (5),

$$e = a^2 \omega$$
, $e' = (e' - a)^2 \omega'$, $e' = (e'' - a)^2 \omega''$;

which give the values of e, e', e": and then, from (8),

$$C = ma^3\omega^3 + m'(c'-a)^2\omega'^2 + m''(c''-a)^2\omega''^2 + \mu\beta^2$$

If instead of attaching two particles m', m'', to m, we had attached any number of them, the problem would have been essentially of no greater difficulty.

Riccati; Comment. Bonon. tom. v. P. 1. p. 150; anno 1767.

(6) Two particles P, P', (fig. 200), are connected together by a rigid rod without inertia which passes through a small smooth ring at O; the rod rests upon a horizontal plane: supposing any impulse whatever to have been communicated to the particles, to find the paths which they will describe.

Let OE be a fixed line in the plane of the motion; OP = r, PP' = l; a, a', the initial values of OP, OP'; m, m', the masses of P, P'; $\angle POE = \theta$; ω , β , the initial values of $\frac{d\theta}{dt}$, $\frac{dr}{dt}$. Then the differential equation to P's path will be

$$\{mr^2 + m'(l-r)^2\} \{A [mr^2 + m'(l-r)^2] - 1\} = (m+m') \frac{dr^2}{d\theta^2},$$
where
$$A = \frac{(m+m')\beta^2 + (ma^2 + m'a'^2)\omega^2}{(ma^2 + m'a'^2)\omega};$$

and similarly for the path of P.

Clairaut; Mém. Acad. Paris, 1742, p. 38. D'Arcy; Ib. 1747, p. 352.

CHAPTER XI.

COEXISTENCE OF SMALL OSCILLATIONS.

Conceive that a particle or a system of particles, subject to certain fixed laws of geometrical connection or constraint, be slightly but generally deranged from a position of stable equilibrium, the invariable elements of the geometry being supposed to be free from particular relations. Then, if in the geometrical equations there be n independent variables, the motion of each member of the system may be represented by the composition of n primary oscillations of different periods, the periods of the n oscillations of any two members of the system being coexistent, while their amplitudes will generally be different. When the periods of the n elementary oscillations are commensurable, the whole system will return to its original state after an interval equal to the least common multiple of these periods; as in the case of vibrating cords and vibrating surfaces. This general property of sympathetic vibrations has been entitled the Principle of the Coexistence of small Oscillations or Vibrations.

Should the original derangement of the system from its position of equilibrium, instead of being perfectly general, be effected by peculiar adaptation, we may reduce the *n* elementary oscilla-

tions to any smaller number we may please.

If the fixed geometrical elements of the system be not, as we have supposed, free from particular relations, and if it receive a perfectly general derangement, there will as before arise in the system altogether n classes of oscillations; under these circumstances however a peculiarity occasionally presents itself, viz. that, although as we have supposed the original derangement be quite general, yet into the motion of no single member of the system will all the elementary oscillations enter; this case will then constitute a failure of the Principle of the Coexistence of small Oscillations.

The Principle of Coexistent Oscillations was first laid down

by Daniel Bernoulli, who has written several memoirs on the subject in the St. Petersburgh Transactions. See particularly Nov. Comment. Petrop. vol. XIX. p. 281. The student is referred also to Lagrange, Mécanique Analytique, tom. 1. p. 347, and to Poisson, Traité de Mécanique, tom. 11. p. 426, where he will find investigations of the Principle based on the first principles of Mechanics.

(1) To determine the nature of the oscillations of a material particle within the surface of an ellipsoid, in the neighbourhood of the lower extremity of a vertical; s.

Let 2a, 2b, denote the lengths of the two horizontal axes of the ellipsoid, 2c representing the length of the vertical one; and let the co-ordinate axes be so chosen that a, b, may be parallel to the axes of x, y, and that c may coincide with the axis of z.

Then, by D'Alembert's Principle of abined with the Principle of Virtual Velocities, we have for the motion of the particle,

$$\frac{d^2x}{dt^2}\,\delta x + \frac{d^2y}{dt^2}\,\delta y + \left(\frac{d^2z}{dt^2} + g\right)\delta z = 0\,\ldots\,(1),$$

where x, y, z, denote the co-ordinates of the particle at any time t, and δx , δy , δz , the increments of x, y, z, in passing to any point of the surface very near to the position of the particle.

Again, by the equation to the ellipsoid, we have

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} + \frac{(c-z)^2}{c^2} = 1$$
;

and therefore, neglecting powers of the small quantities beyond the second,

$$c - z = c \left(1 - \frac{x^2}{a^2} - \frac{y^2}{b^2} \right)^{\frac{1}{4}} = c \left(1 - \frac{x^2}{2a^2} - \frac{y^2}{2b^2} \right),$$

$$z = \frac{1}{2} c \left(\frac{x^2}{a^2} + \frac{y^2}{b^2} \right),$$

$$\delta z = c \left(\frac{x \delta x}{a^2} + \frac{y \delta y}{b^2} \right);$$

hence, from (1), neglecting the products and powers, beyond the first, of small quantities in the coefficients of δx , δy , we get

$$\frac{d^2x}{dt^2} \delta x + \frac{d^2y}{dt^2} \delta y + \frac{cg}{a^2} x \delta x + \frac{cg}{b^2} y \delta y = 0,$$

and therefore $\left(\frac{d^2x}{dt^2} + \frac{cg}{a^2} x\right) \delta x + \left(\frac{d^2y}{dt^2} + \frac{cg}{b^2} y\right) \delta y = 0.$

Equating to zero the coefficients of δx , δy , which are independent of each other, we get

$$\frac{d^2x}{dt^2} + \frac{cg}{a^2} x = 0 \dots (2),$$

$$\frac{d^2y}{dt^2} + \frac{cg}{b^2} y = 0 \dots (3).$$

The integral of the equation (2) is

$$x = \beta \sin \left\{ \frac{(cg)^{\frac{1}{4}}}{a} t + \epsilon \right\},\,$$

and that of (3) is

$$y = \gamma \sin \left\{ \frac{(cg)^{\frac{1}{4}}}{b} t + \zeta \right\};$$

where β , γ , ε , are arbitrary constants, which may be determined from the initial values of x, y, $\frac{dx}{dt}$, $\frac{dy}{dt}$. It may be observed that the oscillation of the particle depends upon two simple oscillations of which $\frac{\pi a}{(eg)^{\frac{1}{4}}}$, are the periods; the number of inde-

pendent simple oscillations being the same as the number of independent variables in the geometrical equation to which the position of the particle is subject.

Poisson; Traité de Mécanique, tom. 11. p. 439.

(2) A uniform rod AB, (fig. 201), which is connected by a string OA with a fixed point O, having been slightly displaced from its position of equilibrium; to investigate the nature of its small oscillations.

Draw vertically the indefinite straight line Ox; take P any point in AB, draw PM at right angles to Ox, and produce BA to meet Ox in C. Let AB = 2a, OM = x, PM = y, AP = s, OA = l, $\angle AOx = \theta$, $\angle BCx = \phi$.

Then for the motion of the rod we have, by D'Alembert's Principle combined with the Principle of Virtual Velocities,

$$\int_0^{2a} \left\{ ds \left(\frac{d^2x}{dt^2} - g \right) \delta x \right\} + \int_0^{2a} \left\{ ds \frac{d^2y}{dt^2} \delta y \right\} = 0 \dots (1),$$

where dx, dy, denote the small spaces described by the element ds of the rod in the time dt, parallel to the co-ordinate axes; δx , δy , denoting the resolved parts of its virtual velocity.

Now, from the geometry, we have

$$x = l \cos \theta + s \cos \phi$$
, $y = l \sin \theta + s \sin \phi$;

and therefore, our object being to transform the equation (1) into an equation involving θ , ϕ , instead of x, y, and to retain small quantities only as far as the first order in the coefficients $\delta\theta$, $\delta\phi$, of the new equation, we have approximately

$$x = l\left(1 - \frac{1}{2}\theta^{2}\right) + s\left(1 - \frac{1}{2}\phi^{2}\right), \quad y = l\theta + s\phi,$$

$$\delta x = -l\theta \,\delta \theta - s\phi \,\delta \phi, \qquad \delta y = l\delta \theta + s\delta \phi,$$

$$\frac{d^{3}x}{dt^{2}} = 0, \quad \frac{d^{3}y}{dt^{2}} = l\frac{d^{3}\theta}{dt^{2}} + s\frac{d^{3}\phi}{dt^{2}}:$$

hence, substituting these values of x, y, \ldots in the equation (1), we have

$$\int_0^{2a} \left\{ g ds \left(l\theta \, \delta\theta + s\phi \, \delta\phi \right) \right\} + \int_0^{2a} \left\{ ds \left(l \, \frac{d^2\theta}{dt^2} + s \, \frac{d^2\phi}{dt^2} \right) \left(l \, \delta\theta + s \, \delta\phi \right) \right\} = 0.$$

Equating to zero the coefficient of $\delta\theta$, we get

$$\left(g\theta+l\frac{d^3\theta}{dt^3}\right)\int_0^{2a}ds+\frac{d^3\phi}{dt^2}\int_0^{2a}s\,ds=0,$$

and therefore

$$l \frac{d^2\theta}{dt^2} + a \frac{d^2\phi}{dt^2} + g\theta = 0 \dots (2);$$

and, equating to zero the coefficient of $\delta \phi$, we obtain

$$\left(g\phi + l\frac{d^{2}\theta}{dt^{2}}\right)\int_{0}^{2a} s\,ds + \frac{d^{2}\phi}{dt^{2}}\int_{0}^{2a} s^{2}ds = 0,$$

and therefore

$$l\frac{d^{2}\theta}{dt^{2}} + \frac{4}{3}a\frac{d^{2}\phi}{dt^{2}} + g\phi = 0 \dots (3).$$

In order to integrate the equations (2) and (3), assume

$$\theta = a \sin \left\{ \left(\frac{g}{\rho} \right)^{\frac{1}{4}} t + \epsilon \right\}, \quad \phi = \beta \sin \left\{ \left(\frac{g}{\rho} \right)^{\frac{1}{4}} t + \epsilon \right\};$$

substituting these expressions for θ in (2), and dividing by $\sin \left\{ \left(\frac{g}{\rho} \right)^{\frac{1}{2}} t + \epsilon \right\}$, we get

$$-\frac{la}{\rho}-\frac{a\beta}{\rho}+a=0, \text{ or } a\beta=a(\rho-l)...(4);$$

and substituting in (3), we get, in the same way,

$$-\frac{la}{\rho}-\frac{4a\beta}{3\rho}+\beta=0, \quad \text{or} \quad \beta(3\rho-4a)=3la....(5).$$

Eliminating a and β between the two equations (4) and (5), we obtain

$$\frac{3\rho - 4a}{a} = \frac{3l}{\rho - l}, \quad \text{or} \quad 3\rho^2 - (4a + 3l)\rho + al = 0.$$

Let the two values of ρ deducible from this quadratic be denoted by m, m': then the motion of the rod will be completely determined by the equations

$$\theta = a \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{4}} t + \epsilon \right\} + a' \sin \left\{ \left(\frac{g}{m'} \right)^{\frac{1}{4}} t + \epsilon' \right\} \dots (6),$$

$$\phi = \beta \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{4}} t + \epsilon \right\} + \beta' \sin \left\{ \left(\frac{g}{m'} \right)^{\frac{1}{4}} t + \epsilon' \right\} \dots (7).$$

In these two equations there are six arbitrary constants, a, a', β , β' , ϵ , ϵ' ; they are not however all of them independent of each other; in fact, by (4), since a and a' correspond respectively to the values m and m' of the quantity ρ , we have

$$\beta = \frac{a}{a}(m-l), \quad \beta' = \frac{a'}{a}(m'-l);$$

hence, from (7), we see that

$$\phi = \frac{a}{a}(m-l)\sin\left\{\left(\frac{g}{m}\right)^{\frac{1}{4}}t + \varepsilon\right\} + \frac{a'}{a}(m'-l)\sin\left\{\left(\frac{g}{m'}\right)^{\frac{1}{4}}t + \varepsilon'\right\}..(8).$$

The four constants a, a', ϵ , ϵ' , involved in the two equations (7) and (8), may be determined if we have given the initial circumstances of the rod, or the initial values of θ , $\frac{d\theta}{dt}$, ϕ , $\frac{d\phi}{dt}$.

If a' = 0, $\beta' = 0$, then we have

$$\theta = a \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{2}} t + \epsilon \right\}, \quad \phi = \beta \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{2}} t + \epsilon \right\},$$

and the oscillations of θ and ϕ will evidently be regular and isochronous, the time of vibration being equal to $\pi \left(\frac{m}{g}\right)^{l}$.

If α' , β' , be not equal to zero, the oscillations of θ and ϕ will be compounded of two simple and isochronous vibrations.

Suppose that at two different times t', t', the values of θ and of $\frac{d\theta}{dt}$ are the same. This will manifestly be the case if

$$\left(\frac{g}{m}\right)^{\frac{1}{2}}t' + \varepsilon = \left(\frac{g}{m}\right)^{\frac{1}{2}} \varepsilon + 2\lambda\pi,$$

$$\left(\frac{g}{m'}\right)^{\frac{1}{2}}t' + \varepsilon' = \left(\frac{g}{m'}\right)^{\frac{1}{2}}t' + \varepsilon' + 2\lambda'\pi,$$

and

λ, λ', being any integers; hence

$$2\lambda\pi\left(\frac{m}{g}\right)^{\frac{1}{2}}=t'-t'=:\ \ \forall\pi\left(\frac{m'}{g}\right)^{\frac{1}{2}},$$

and therefore

$$\lambda m^{\frac{1}{4}} = \lambda' m'^{\frac{1}{4}},$$

or m, m', must be to each other as two square numbers.

It will be observed that, in agreement with the general theory of the Coexistence of small Oscillations, the number of independent oscillations of θ and ϕ is two, which is the same as the number of the independent geometrical variables.

The following is another method of solving this problem.

Let G (fig. 202) be the position of the centre of gravity of the rod at any time t; draw GH at right angles to the vertical line Ox; m = the mass of the rod, $mk^3 =$ its moment of inertia about G, T = the tension of the string AO, OH = x, GH = y. Then, the rest of the notation being the same as before, we have, for the motion of the rod,

$$m \frac{d^2x}{dt^2} = mg - T\cos\theta \dots (1),$$

$$m \frac{d^2y}{dt^2} = -T\sin\theta \dots (2),$$

$$mk^2 \frac{d^2\phi}{dt^2} = -a\sin(\phi - \theta)T \dots (3).$$

Eliminating T between (1) and (2), and omitting small quantities of higher orders than the first, we have

$$\frac{d^2y}{dt^2}+g\theta=0\ldots(4);$$

and, eliminating T between (1) and (3), we get in the same manner $d^2 a$

$$k^2 \frac{d^2 \phi}{dt^2} + ag (\phi - \theta) = 0 \dots (5).$$

But

$$y = a \sin \phi + l \sin \theta = a\phi + l\theta$$
, nearly;

hence (4) becomes

$$l\frac{d^2\theta}{dt^2} + a\frac{d^2\phi}{dt^2} + g\theta = 0;$$

and, putting for k^2 its value $\frac{1}{3}a^2$ in (5), we have

$$\frac{1}{3}a\frac{d^3\phi}{dt^2}+g(\phi-\theta)=0.$$

These two last equations are equivalent to the equations (2) and (3) in the former investigation.

Daniel Bernoulli; Novi Comment. Petrop. 1773, tom. xvIII. p. 247. Euler; Ib. p. 268.

(3) A pendulum of any figure is firmly attached to a solid circular cylinder as an axis; this axis is supported in a horizontal position at its two extremities, which rest within two hollow circular cylinders placed horizontally and of the same dimensions; to investigate the small oscillations of the pendulum corresponding to any initial state of displacement and motion, the surfaces in contact being considered perfectly smooth.

Let G (fig. 208) be the centre of gravity of the pendulum and its axis, regarded as one mass, at any time of the motion; let the plane of the paper represent the vertical plane through G, which cuts the axis of the solid cylinder at right angles in the point C. Let the circular arc MAN be the common intersection of the two concave cylinders with the plane of the paper, when produced to meet it. From O, the centre of the arc MAN, draw OAx vertically; GH at right angles to Ox; produce GC to meet Ox in K; join OC, and produce it to a, which will be the

point of contact between MAN and the circular section of the solid cylinder made by the plane of the paper. Let OH = x, GH = y, AO = a, Ca = b, $\angle AKC = \phi$, $\angle COx = \theta$, CG = c; m =the mass of the pendulum and its axis together; k =their radius of gyration about G; R =the reaction of the hollow cylinders against the axis of the pendulum.

Then, for the motion of the pendulum, we have

$$m \frac{d^2x}{dt^2} = mg - R \cos \theta \dots (1),$$

$$m \frac{d^3y}{dt^2} = -R \sin \theta \dots (2),$$

$$mk^2 \frac{d^2\phi}{dt^2} = -Rc \sin (\phi - \theta) \dots (3).$$

From (1) and (2) we get, as far as small quantities of the first order, $\frac{d^3y}{dt^2} + g\theta = 0 \dots (4);$

and, from (1) and (8), to the same degree of approximation,

$$k^{2}\frac{d^{3}\phi}{dt^{2}}+cg(\phi-\theta)=0....(5).$$

Now, from the geometry,

$$y = (a - b) \sin \theta + c \sin \phi$$

= $(a - b) \theta + c\phi$, nearly:

hence from (4) we obtain

Assume
$$\theta = \alpha \sin \left\{ \left(\frac{g}{r} \right)^{\frac{1}{4}} t + \epsilon \right\}, \quad \phi = \beta \sin \left\{ \left(\frac{g}{r} \right)^{\frac{1}{4}} t + \epsilon \right\}$$
:

then from (5) we may get

$$\beta (cr - k^2) = acr \dots (7),$$

and, from (6), $c\beta$

$$c\beta = \alpha \{r - (a - b)\},\$$

and therefore, eliminating α and β ,

$$(cr-k^2)(r-a+b)=c^3r.$$

Let the two roots of this quadratic in r be denoted by m and m'; then, for the general values of θ and ϕ , we have

$$\theta = a \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{4}} t + \epsilon \right\} + a' \sin \left\{ \left(\frac{g}{m'} \right)^{\frac{1}{4}} t + \epsilon' \right\} \dots (8),$$

$$\phi = \beta \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{4}} t + \epsilon \right\} + \beta' \sin \left\{ \left(\frac{g}{m'} \right)^{\frac{1}{4}} t + \epsilon' \right\} \dots (9).$$

From (7) we have, β , β' , being the values of β , and α , α' , those of α , corresponding to the values m, m', of r,

$$\beta = \frac{acm}{cm - k^2}, \quad \beta' = \frac{a'cm'}{cm' - k^2};$$

hence, from (9), we have

$$\phi = \frac{acm}{cm - k^2} \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{2}} t + \epsilon \right\} + \frac{a'cm'}{cm' - k^2} \sin \left\{ \left(\frac{g}{m'} \right)^{\frac{1}{2}} t + \epsilon' \right\} \dots (10).$$

In the equations (8) and (10) there are four arbitrary constants, a, a', ϵ , ϵ' , which may be determined if we have given the initial values of θ , ϕ , $\frac{d\theta}{dt}$, $\frac{d\phi}{dt}$.

If a' = 0, $\beta' = 0$, we have

$$\theta = a \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{2}} t + \epsilon \right\}, \quad \phi = \beta \sin \left\{ \left(\frac{g}{m} \right)^{\frac{1}{2}} t + \epsilon \right\};$$

and the oscillations of θ and ϕ will be regular and isochronous, the time of vibration being $\pi \left(\frac{m}{g}\right)^{\frac{1}{2}}$.

If a' and β' have finite values, the oscillations of θ and ϕ will be compounded of two simple isochronous oscillations.

Euler; Acta Acad. Petrop. 1780, P. II. p. 133.

(4) A string AEFB (fig. 204) is attached to two fixed points A, B, in the same horizontal line. From E, F, points so chosen that AE, EF, FB, are all equal, two masses are suspended by strings EM, FN, of different lengths, the masses being equal. Supposing the system to be slightly deranged from its position of equilibrium, to investigate the nature of its small oscillations.

At any time t let EM, FN, make angles ϕ , ϕ' , with the vertical. Let AE, EF, BF, make angles $a + \omega$, ω' , $a - \omega'$, with the horizon, the values of these angles being a, 0, a, for the position of equilibrium. Draw Mm, Nn, horizontally to meet the vertical line Amn in the points m, n. Let $AE = EF = l^{2}B \cdot a$, EM = k, FN = k', Am = x, Mm = y, An = x', Nn = y'.

By D'Alembert's Principle and the Principle of Virtual Velocities, we have, for the motion of the system,

$$\left(\frac{d^2x}{dt^2}-g\right)\delta x+\left(\frac{d^2x'}{dt^2}-g\right)\delta x'+\frac{d^2y}{dt^2}\delta y+\frac{d^2y'}{dt^2}\delta y'=0\ldots(1).$$

Our object is now to express x, y, x', y', in terms of ω, ϕ, ϕ' , and to substitute their values in this equation. This computation must be effected as far as small quantities of the second order.

By the geometry it is plain that

$$a\cos(a+\omega)+a\cos\omega'+a\cos(a-\omega')=2a\cos\alpha+a$$
,

and therefore

$$\cos \alpha (1 - \frac{1}{2}\omega^2) - \sin \alpha \cdot \omega + 1 - \frac{1}{2}\omega'^2 + \cos \alpha (1 - \frac{1}{2}\omega''^2) + \sin \alpha \cdot \omega'' = 2\cos \alpha + 1;$$

whence

$$\cos \alpha \cdot \omega^2 + 2 \sin \alpha \cdot \omega + \omega'^2 + \cos \alpha \cdot \omega'^2 - 2 \sin \alpha \cdot \omega' = 0...(2).$$

Again, by the geometry,

$$a \sin (\alpha + \omega) = a \sin \omega' + a \sin (\alpha - \omega'),$$

and therefore

$$\sin\alpha\left(1-\tfrac{1}{2}\omega^2\right)+\cos\alpha\cdot\omega=\omega'+\sin\alpha\left(1-\tfrac{1}{2}\omega''^2\right)-\cos\alpha\cdot\omega\ ;$$
 whence

$$2 \cos \alpha \cdot \omega - \sin \alpha \cdot \omega^2 = 2\omega' - \sin \alpha \cdot \omega^2 - 2 \cos \alpha \cdot \omega' \cdot \ldots \cdot (3).$$

Now, as far as the first order of small quantities, we have. from (2), $2 \sin \alpha \cdot \omega = 2 \sin \alpha \cdot \omega$,

and therefore

$$\boldsymbol{\omega}' = \boldsymbol{\omega}$$
;

and from (3) we have

$$2\cos\alpha.\omega = 2\omega' - 2\cos\alpha.\omega'' = 2\omega' - 2\cos\alpha.\omega,$$

and therefore $\omega' = 2 \cos \alpha \cdot \omega$.

Substituting these values of ω' , ω'' , in the terms of the second order in (2) and (3), we get

$$(2 \cos \alpha + 4 \cos^2 \alpha) \omega^2 + 2 \sin \alpha \cdot \omega - 2 \sin \alpha \cdot \omega'' = 0,$$
and
$$\cos \alpha \cdot \omega = \omega' - \cos \alpha \cdot \omega''$$
:

from these two last equations we see that

$$(2\cos^2\alpha + 4\cos^3\alpha)\omega^2 + 2\sin\alpha\cos\alpha.\omega - 2\sin\alpha.\omega' + 2\sin\alpha\cos\alpha.\omega = 0$$

and therefore

$$\omega' = 2 \cos \alpha \cdot \omega + \frac{\cos^2 \alpha + 2 \cos^3 \alpha}{\sin \alpha} \omega^2 \cdot \dots (4).$$

Again, as far as our approximation requires,

$$x = a \sin(a + \omega) + k \cos \phi = a \sin \alpha \left(1 - \frac{1}{2}\omega^2\right) + a \cos \alpha \cdot \omega + k \left(1 - \frac{1}{2}\varphi^2\right),$$

$$\delta x = -a \sin a \omega \delta \omega + a \cos a \delta \omega - k \phi \delta \phi,$$

$$\frac{d^2x}{dt^2} = a \cos a \frac{d^2\omega}{dt^2};$$

$$y = a \cos(\alpha + \omega) + k \sin \phi = a \cos \alpha \left(1 - \frac{1}{2}\omega^2\right) - a \sin \alpha \cdot \omega + k\phi,$$

$$\delta y = -a \sin a \, \delta \omega + k \, \delta \phi,$$

$$\frac{d^3y}{dt^2} = -a \sin a \frac{d^3\omega}{dt^2} + k \frac{d^3\phi}{dt^2};$$

$$x' = a \sin (a + \omega) - a \sin \omega' + k' \cos \phi'$$

=
$$a \sin a (1 - \frac{1}{2}\omega^2) + a \cos a \cdot \omega - a\omega' + k' (1 - \frac{1}{2}\phi'^2)$$

$$= a \sin a - a \cos a.\omega - a \frac{\sin^3 a + 2 \cos^2 a + 4 \cos^3 a}{2 \sin a} \omega^2 + k'(1 - \frac{1}{2}\phi'^2),$$

by (4);

$$\delta x' = -a \cos a \delta \omega - a \frac{\sin^2 \alpha + 2 \cos^2 \alpha + 4 \cos^3 \alpha}{\sin \alpha} \omega \delta \omega - k' \phi' \delta \phi',$$

$$\frac{d^2x'}{dt^2}=-a\cos\alpha\,\frac{d^2\omega}{dt^2};$$

$$y' = a \cos(\alpha + \omega) + a \cos \omega' + k' \sin \varphi'$$

$$= a \cos a \left(1 - \frac{1}{2} \omega^{2}\right) - a \sin a \cdot \omega + a \left(1 - \frac{1}{2} \omega^{2}\right) + k' \phi',$$

$$\delta y' = k' \delta \phi' - a \sin a \delta \omega,$$

$$\frac{d^2y'}{dt^2} = k' \frac{d^2\phi'}{dt^2} - a \sin \alpha \frac{d^2\omega}{dt^2}.$$

Hence, by the equation (1), there is

$$2a^{2}\cos^{2}a\frac{d^{2}\omega}{dt^{2}}\delta\omega + kg\phi\delta\phi + ga\frac{2+4\cos^{3}a}{\sin a}\omega\delta\omega$$

$$+ gk'\phi'\delta\phi' + \left(a\sin a\frac{d^{2}\omega}{dt^{2}} - k\frac{d^{2}\phi}{dt^{3}}\right)(a\sin a\delta\omega - k\delta\phi)$$

$$+ \left(k'\frac{d^{2}\phi'}{dt^{2}} - a\sin a\frac{d^{2}\omega}{dt^{2}}\right)(k'\delta\phi' - a\sin a\delta\omega)$$

Hence, equating to zero the coefficients of the independent quantities $\delta \phi$, $\delta \phi'$, $\delta \omega$, we get

$$k \frac{d^2 \phi}{dt^2} \quad a \sin \alpha \frac{d^2 \omega}{dt^2} + g \phi = 0 \dots (5),$$

$$k' \frac{d^2 \phi'}{dt^2} - a \sin \alpha \frac{d^2 \omega}{dt^2} + g \phi' = 0 \dots (6),$$

$$2a \frac{d^3\omega}{d\ell^2} - k \sin \alpha \frac{d^3\phi}{d\ell^2} - k' \sin \alpha \frac{d^3\phi'}{d\ell^2} + g \frac{2 + 4\cos^3\alpha}{\sin\alpha} \omega = 0....(7).$$

Eliminating $\frac{d^3\phi}{dt^3}$ and $\frac{d^3\phi'}{dt^2}$ between (5), (6), (7), we get

$$2a \sin a \cos^2 a \frac{d^2\omega}{dt^2} + g \{(2+4 \cos^3 a) \omega + \sin^2 a \cdot \phi + \sin^2 a \cdot \phi'\} = 0 \dots (8)$$

Let r denote the length of a pendulum, isochronous with one of the elementary oscillations, and assume accordingly

$$\omega = \Omega \sin \left\{ \begin{pmatrix} g \\ r \end{pmatrix} \right\} t + \varepsilon ,$$

$$\phi = F \sin \left\{ \right\} ,$$

$$\phi' = F' \sin \left\{ \right\} .$$

Then, from (5), (6), (8), we have

$$(k-r) F = a \sin a.\Omega,$$

$$(k'-r) F' = a \sin a.\Omega,$$

$$-2a \sin a \cos^2 a \frac{1}{r} \Omega + (2 + 4 \cos^3 a) \Omega + \sin^2 a F' + \sin^2 a F' = 0;$$

and, by eliminating the constants F, F', Ω , we have a cubic equation in r,

$$\frac{2 \sin a \cos^3 a}{r} + \frac{\sin^3 a}{r - k} + \frac{\sin^3 a}{r - k'} - \frac{2 + 4 \cos^3 a}{a} = 0 \dots (9).$$

Let l, l', l', be the three roots of this equation; then, for the complete solution of the problem, we have

$$\begin{split} \omega'' &= \omega = \Omega \sin \left\{ \left(\frac{g}{\bar{l}} \right)^{\frac{1}{4}} t + \varepsilon \right\} + \Omega' \sin \left\{ \left(\frac{g}{\bar{l'}} \right)^{\frac{1}{4}} t + \varepsilon' \right\} + \Omega'' \sin \left\{ \left(\frac{g}{\bar{l'}} \right)^{\frac{1}{4}} t + \varepsilon' \right\}, \\ \omega' &= 2 \cos \alpha \cdot \omega, \end{split}$$

$$\begin{split} \phi &= F_1 \sin \left\{ \left(\frac{g}{l} \right)^{\frac{1}{4}} t + \epsilon \right\} + F_2 \sin \left\{ \left(\frac{g}{l} \right)^{\frac{1}{4}} t + \epsilon' \right\} + F_3 \sin \left\{ \left(\frac{g}{l''} \right)^{\frac{1}{4}} t + \epsilon' \right\}, \\ \phi' &= F_1' \sin \left\{ \right\} + F_2' \sin \left\{ \right\} + F_3' \sin \left\{ \right\}. \end{split}$$

This problem may be solved also in the following manner, which is Euler's method of considering it.

Let P, Q, be the tensions of the strings EM, FN, and m the mass of each of the bodies.

Then, for the motion of the bodies, we have, approximately,

$$m \frac{d^{2}x}{dt^{2}} = mg - P \cos \phi = mg - P \dots (1),$$

$$m \frac{d^{3}y}{dt^{3}} = -P \sin \phi = -mg\phi \dots (2),$$

$$m \frac{d^{3}x'}{dt^{2}} = mg - Q \cos \phi' = mg - Q \dots (3),$$

$$m \frac{d^{3}y'}{dt^{3}} = -Q \sin \phi' = -mg\phi' \dots (4);$$

these four equations being true as far as the first order of small quantities.

Let T denote the tension of the string EF; then, since the three tensions acting upon the point E must be in equilibrium, there is

$$\frac{T}{P} = \frac{\sin\left\{\phi + \frac{1}{2}\pi + a + \omega\right\}}{\sin\left\{\frac{1}{2}\pi - (a + \omega) + \frac{1}{2}\pi - \omega'\right\}} = \frac{\cos\left(a + \omega + \phi\right)}{\sin\left(a + \omega + \omega'\right)}.$$

Similarly, for the tensions at F, we have

$$\frac{Q}{T} = \frac{\sin (a - \omega'' - \omega')}{\cos (a - \omega'' - \phi')};$$

hence

$$\frac{Q}{P} = \frac{\sin (a - \omega'' - \omega') \cos (a + \omega + \phi)}{\cos (a - \omega'' - \phi') \sin (a + \omega + \omega')}.$$

Hence, as far as small quantities of the first order,

$$Q \left\{ \sin \alpha + \cos \alpha \left(\omega + \omega' \right) \right\} \left\{ \cos \alpha + \sin \alpha \left(\omega' + \phi' \right) \right\}$$

$$= P \left\{ \sin \alpha - \cos \alpha \left(\omega' + \omega'' \right) \right\} \left\{ \cos \alpha - \sin \alpha \left(\omega + \phi \right) \right\},$$

and therefore

$$Q\left\{\sin \alpha \cos \alpha + \sin^2 \alpha \left(\omega'' + \phi'\right) + \cos^2 \alpha \left(\omega + \omega'\right)\right\}$$

$$= P\left\{\sin \alpha \cos \alpha - \sin^2 \alpha \left(\omega + \phi\right) - \cos^2 \alpha \left(\omega' + \omega'\right)\right\}...(5).$$

Now, by the geometry,

$$\cos (\alpha + \omega) + \cos \omega' + \cos (\alpha - \omega'') = 2 \cos \alpha + 1,$$

and therefore, as far as the first order of small quantities,

$$-\sin\alpha \cdot \omega + \sin\alpha \cdot \omega'' = 0,$$

$$\omega'' = \omega \qquad (6)$$

Also, by the geometry,

$$\sin (\alpha + \omega) = \sin \omega' + \sin (\alpha - \omega''),$$

$$\cos \alpha \cdot \omega = \omega' - \omega'' \cos \alpha = \omega' - \omega \cos \alpha,$$

$$\omega' = 2\omega \cos \alpha \cdot \dots (7).$$

Hence by (5), (6), (7), we have

$$Q \left\{ \sin \alpha \cos \alpha + \sin^2 \alpha \left(\omega + \phi' \right) + \cos^2 \alpha \left(1 + 2 \cos \alpha \right) \omega \right\}$$

$$= P \left\{ \sin \alpha \cos \alpha - \sin^2 \alpha \left(\omega + \phi \right) - \cos^2 \alpha \left(1 + 2 \cos \alpha \right) \omega \right\} \cdots (8).$$

Eliminating P and Q between (1), (3), and (8), we have, as far as small quantities of the first order,

$$\left(\frac{d^2x}{dt^2} - \frac{d^2x'}{dt^2}\right) \sin \alpha \cos \alpha = -g\left\{(2 + 4\cos^3\alpha)\omega + \sin^2\alpha \cdot \phi + \sin^2\alpha \cdot \phi'\right\} \dots (9).$$

But
$$x = a \sin (a + \omega) + k \cos \phi = a \cos a \cdot \omega + ...,$$

 $y = a \cos (a + \omega) + k \sin \phi = -a \sin a \cdot \omega + k\phi + ...,$
 $x' = a \sin (a + \omega) - a \sin \omega' + k' \cos \phi' = a \cos a \cdot \omega - a\omega' + ...$
 $= -a \cos a \cdot \omega + ...,$

$$y' = a\cos(\alpha + \omega) + a\cos\omega' + k'\sin\varphi' = -a\sin\alpha.\omega + k'\varphi' + ...;$$
 hence, by (9),

$$2a \sin a \cos^2 a \frac{d^2 \omega}{dt^2} + g \left\{ (2 + 4 \cos^3 a) \omega + \phi \sin^2 \alpha + \phi' \sin^2 \alpha \right\} = 0$$
:

and, by (2) and (4),

$$k\frac{d^2\phi}{dt^2}-a\sin\alpha\frac{d^2\omega}{dt^2}+g\phi=0,$$

$$k'\frac{d^2\varphi'}{dt^2}-a\sin\alpha\frac{d^2\omega}{dt^2}+g\varphi'=0;$$

which are the same three linear equations as (5), (6), (8), in the former solution.

If k be equal to k', the cubic equation (9) of the former solution will degenerate into a quadratic, and the variations of ω , ϕ , ϕ' , will no longer be expressible by the composition of the same elementary vibrations. This will be an instance of the failure of the Principle of the Coexistence of small Oscillations.

Euler; Act. Acad. Petrop. 1779, P. H. p. 95.

(5) One of the scales of a common balance having been slightly displaced from its position of rest, in a vertical plane passing through the beam; to investigate the nature of the oscillatory motions of the two scales and of the beam, to which the displacement will give rise.

Let O (fig. 205) be the point of suspension of the whole balance, G its centre of gravity, AB the beam, P and Q the scales, which are here supposed to be material points. Draw aOb horizontal, aAa, $bB\beta$, vertical. Let AC = a = BC, OC = b, OG = c, AP = l = BP, Mk^2 = the moment of inertia of the beam about O, m = the mass of P and of Q supposed to be equal.

Let ϕ be the angle which, at any time t, the beam makes with the horizon; $\angle PAa = \eta$, $\angle QB\beta = \theta$. Also put

$$\frac{g}{l}=n^2, \quad \frac{b}{l}=h, \quad \frac{Mc+2mb}{Mk^2} g=p^3, \quad -\frac{mbg}{Mk^2}=q,$$

and let $-\mu_1^2$, $-\mu_2^2$, represent the two roots of the quadratic

$$z^{2} + (n^{2} + p^{2} - 2hq)z + n^{2}p^{2} = 0.$$

Then, bearing in mind that, initially,

$$\phi = 0$$
, $\eta = \varepsilon$, $\theta = 0$,
 $\frac{d\phi}{dt} = 0$, $\frac{d\eta}{dt} = 0$, $\frac{d\theta}{dt} = 0$,

where ε is a known constant, we shall obtain for the complete expression of the motions

$$\begin{split} \varphi &= \frac{2\epsilon h p^2 q^2}{\mu_1^2 - \mu_1^2} (\cos \mu_1 t - \cos \mu_2 t), \\ 2\eta &= \frac{2\epsilon h p^2 q}{\mu_2^2 - \mu_1^2} \left(\frac{\cos \mu_1 t}{\mu_1^2 - p^2} - \frac{\cos \mu_2 t}{\mu_2^2 - p^2} \right) + \epsilon \cos nt, \\ 2\theta &= \frac{2\epsilon h p^2 q}{\mu_2^2 - \mu_1^2} \left(\frac{\cos \mu_1 t}{\mu_1^2 - p^2} - \frac{\cos \mu_2 t}{\mu_2^2 - p^2} \right) - \epsilon \cos nt. \end{split}$$

(6) One of the scales of a common balance having been slightly displaced from its position of rest, in a vertical plane at right angles to the beam; to investigate the nature of the oscillatory motions of the two scales and of the beam.

Let AB (fig. 206) be the original position of the beam, PQ its position at any time t; p, q, the projections of the positions of

the scales considered as material points at the same time. Let AC = a = BC, AP = z = BQ, Pp = x, Qq = y, $Mk^2 =$ the moment of inertia of the beam round C, m = the mass of each scale, l = the length of the string by which each scale is suspended. If we put, for simplicity,

$$\frac{g}{l}=n^2, \quad \frac{g}{l}\left(1+2\frac{ma^2}{Mk^2}\right)=n^2,$$

we shall have, for the complete expression of the motions, the initial value of x being c, while those of y, $\frac{dx}{dt}$, $\frac{dy}{dt}$, are all zero;

$$x = c \cos\left(\frac{n'-n}{2}t\right) \cos\left(\frac{n'+n}{2}t\right),$$

$$y = -c \sin\left(\frac{n'-n}{2}t\right) \sin\left(\frac{n'+n}{2}t\right),$$

$$z = \frac{2g}{l} \frac{ma^2}{Mk^2} \frac{c}{n'^2} \sin^2\frac{n't}{2}.$$

Investigations of the two last problems are given in a paper on the Sympathy of Pendulums, in the Cambridge Mathematical Journal, Vol. 11. p. 120.

CHAPTER XII.

IMPULSIVE FORCES.

If two rigid bodies impinge against each other, their motions both of translation and of rotation will generally experience modification, the determination of the nature of which, in the case of bodies of which the positions and motions are assigned at the instant before impact, constitutes the general problem of collision. The process of collision may be divided into two stages of indefinitely small duration: in the former stage, by the force of compression, which we will denote by R, the two points in which the bodies touch each other are constrained to assume equal resolved velocities in the direction of the common normal to their surfaces; in the latter stage, by the force of restitution, if the bodies be not inelastic, an additional reaction eR takes place between them, where e denotes their common elasticity. Let ω_1 , ω_2 , ω_3 , denote the angular velocities of one of the bodies about its principal axes and v_1 , v_2 , v_3 , the components of the velocity of its centre of gravity, at the conclusion of the former stage of the collision; let ω_1' , ω_2' , ω_3' , v_1' , v_2' , v_3' , denote the analogous quantities in relation to the other body. Then, for the expression of the motion of the former body, as modified by the force of compression, we shall have six equations involving, together with known quantities, the symbols ω_1 , ω_2 , ω_3 , v_1 , v_2 , v_3 , R; and in like manner for the latter body we shall have six equations involving ω_1' , ω_2' , ω_3' , v_1' , v_2' , v_3' , R. Thus we shall have in all twelve equations involving thirteen variables. Another equation is supplied by the condition that the points of the two bodies in which their contact takes place shall have an equal resolved velocity in the direction of the common normal. Thus we shall be able to determine completely the modification of the motions of the two bodies due to the force of compression as well as the magnitude

of this force. An additional modification must be applied, in the case of elastic bodies, in consequence of the force of restitution eR, which, from the investigation for the former stage of the collision, has become a known force. If one of the bodies be immoveable, the simplification of the method of investigation which we have described is obvious, the thirteen equations of which we made mention being reduced in this case to seven, and the common normal velocity of the two points of contact being zero. For ample information on this subject the student is referred to Poisson's Traité de Mécanique, tom. II. p. 254, seconde édition.

SECT. 1. Single Body. Smooth Surfaces.

(1) A beam of imperfect elasticity, moving anyhow in a vertical plane, impinges upon a smooth horizontal plane; to determine the initial motion of the beam after impact.

We will commence with supposing the beam to be inelastic; in this case the extremity of the beam which strikes the horizontal plane will continue after impact to slide along it without detaching itself. Let PQ (fig. 207) represent the beam at any time after collision; KL being the section of the horizontal plane made by the vertical plane through PQ; G the centre of gravity of PQ; draw GH at right angles to KL. Let GH = y. QG = a, $\angle GQH = \theta$; k = the radius of gyration about G. m = the mass of the beam; ω , ω' , the angular velocities of the beam about G estimated in the direction of the arrows in the figure, just before and just after impact; u, v, the vertical velocities of G estimated downwards just before and just after impact; B the blow of collision.

Then, $\omega' - \omega$ being the angular velocity communicated by the blow, we shall have, if β be the value of θ at the instant of $\omega' - \omega = \frac{Ba \cos \beta}{mk^2} \dots (1);$ impact,

and, u - v being the velocity of G which is destroyed by the blow, $u-r=\frac{B}{m}$(2).

Again, by the geometry, we get

$$y = a \sin \theta$$
;

and therefore, t denoting the interval between the instant of collision and the arrival of the beam at the position represented in the figure,

 $\frac{dy}{dt} = a \cos \theta \, \frac{d\theta}{dt};$

From (1), (2), (3), we get

$$u - \frac{B}{m} = a \cos \beta \left(\omega + \frac{Ba \cos \beta}{mk^2}\right),$$

$$\frac{B}{m} \left(1 + \frac{a^2}{k^2} \cos^2 \beta\right) = u - a\omega \cos \beta,$$

$$B = mk^2 \frac{u - a\omega \cos \beta}{a^2 \cos^2 \beta + k^2}.................(4).$$

Hence, from (1),

$$\omega' = \omega + a \cos \beta \frac{u - a\omega \cos \beta}{a^2 \cos^2 \beta + k^2} = \frac{au \cos \beta + k^2\omega}{a^2 \cos^2 \beta + k^2};$$

and, from (2),

$$v = u - k^2 \frac{u - a\omega \cos \beta}{a^2 \cos^2 \beta + k^2} = a \cos \beta \frac{au \cos \beta + k^2\omega}{a^2 \cos^2 \beta + k^2}$$

Next let us suppose the beam to be imperfectly elastic, its clasticity being denoted by e; in this case the value of B given in (4) must be increased in the ratio of 1 + e to 1; and therefore, instead of the equation (4), we have

$$B = (1 + e) mk^2 \frac{u - a\omega \cos \beta}{a^2 \cos^2 \beta + k^2},$$

which determines the magnitude of the blow of impact: substituting this value of B in (1), we get

$$\omega' = \omega + (1 + e) a \cos \beta \frac{u - a\omega \cos \beta}{a^2 \cos^2 \beta + k^2}$$

$$= \frac{(k^2 - a^2 e \cos^2 \beta) \omega + (1 + e) au \cos \beta}{a^2 \cos^2 \beta + k^2};$$

and, substituting in (2),

$$v = u - (1 + e) k^{2} \frac{u - a_{\omega} \cos \beta}{a^{2} \cos^{2} \beta + k^{2}}$$

$$= \frac{a^{2} u \cos^{2} \beta - ek^{2} u + (1 + e) k^{2} a_{\omega} \cos \beta}{a^{2} \cos^{2} \beta + k^{2}}.$$

The velocity of G parallel to the plane KL will be the same before and after impact. The end B of the beam will evidently after collision detach itself from the horizontal plane, since v is less and ω' greater when e has a finite value than when it is equal to zero.

(2) A body AB, (fig. 208), after sliding from a given altitude down an inclined plane Oy, impinges against a small obstacle at C; to determine the impulsive reaction of the obstacle and the motion of the body immediately after collision.

Let G be the centre of gravity of the body; draw GH at right angles to Oy; Ox parallel to HG. Let GH = a, CH = b, m = the mass of the body, k = the radius of gyration about G; c = the velocity of G immediately before impact. We will commence with supposing the body to be perfectly inelastic; in this case the point C of the body will remain during collision in contact with the obstacle, the body rotating about this point. Let R, S, denote the impulsive reactions of the obstacle parallel to Ox, Oy, and let u, v, denote the velocities of G parallel to Ox, Oy, on the completion of the impact; also let w represent the angular velocity of rotation about G at the same instant.

Then we have, for the motion of translation,

$$mu = R \dots (1),$$

$$mv = mc - S \dots (2);$$

and, for the motion of rotation,

$$mk^2\omega = Sa - Rb \dots (3).$$

Again, the velocity of the point C of the body, estimated parallel to Ox, will be

$$u - \omega CG \cos \angle GCH \text{ or } u - b\omega$$
,

the former term of this expression arising from the motion of G, and the latter from the rotation of the body about G.

Also the velocity of the point C, parallel to Oy, will be

$$v - \omega . CG \sin \angle GCH$$
 or $v - a\omega$,

the former term being due to the motion of G, and the latter to the rotation about G. But the point C of the body, which is perfectly inelastic, remains at rest during the collision; hence, evidently,

$$u - b\omega = 0$$
.....(4); $v - a\omega = 0$(5).

From (1) and (4) we have

$$R = mb\omega \dots (6),$$

and from (2), (5),

$$S = m (c - a\omega), \ldots (7);$$

substituting these values of R and S in (3), we obtain

$$k^2\omega = ac - a^2\omega - b^2\omega,$$

and therefore,

$$\omega = \frac{ac}{a^2 + b^2 + k^2}, \quad u = \frac{abc}{a^2 + b^2 + k^2}, \quad v = \frac{a^2c}{a^2 + b^2 + k^2};$$

hence also, from (6),

$$R=\frac{mabc}{a^2+b^2+k^2};$$

and, from (7),

$$S = m\left(c - \frac{a^2c}{a^2 + b^2 + k^2}\right) = \frac{mc\left(b^2 + k^2\right)}{a^2 + b^2 + k^2}.$$

If the body be supposed to be elastic, we must increase these values of R and S in the ratio of 1 + e to 1, e denoting the elasticity. Hence

$$R = \frac{m(1+e) abc}{a^2 + b^2 + k^2}, \quad S = \frac{mc(1+e)(b^2 + k^2)}{a^2 + b^2 + k^2};$$

and therefore from (1), (2), (3),

$$u = \frac{(1+e)abc}{a^2+b^2+k^2}, \quad v = c \frac{a^3-e(b^2+k^2)}{a^3+b^2+k^2}, \quad \omega = \frac{(1+e)ac}{a^2+b^2+k^2}.$$

(3) A beam AB (fig. 209) is originally in a vertical position, hanging from the point O, along the line Oy; supposing the extremity A of the beam to be projected from O with a given velocity along a smooth horizontal groove Ox, to determine the motion of the beam.

Let AB be the position of the beam after a time t from the projection of A, G its centre of gravity; draw GH at right angles to Ox; let OH = x, GH = y, $\angle OAG = \theta$, AG = a; m =the mass of the beam, k =its radius of gyration about G.

Then, for the motion of the beam at any time after the projection, we have, by the Principle of the Conservation of Vis Viva,

 $m\left(\frac{dx^2}{dt^2} + \frac{dy^2}{dt^2} + k^2\frac{d\theta^2}{dt^2}\right) = C + 2mgy \dots (1);$

and, by the Principle of the Conservation of the Motion of the Centre of Gravity,

$$\frac{dx}{dt} = C' \dots (2)$$

From (1) and (2), we have

$$m\left(\frac{dy^2}{dt^3}+k^2\frac{d\theta^2}{dt^2}\right)=C''+2mgy;$$

but, from the geometry, we see that $y = a \sin \theta$; hence

$$m \left(a^{3} \cos^{3} \theta + k^{2}\right) \frac{d\theta^{2}}{dt^{2}} = C'' + 2mga \sin \theta \dots (3).$$

Let B denote the blow of projection which is impressed upon the end A of the beam; u the velocity of A's projection, and u the angular velocity of the beam about G immediately after the blow. Also let v be the velocity communicated to G by the blow.

Then we shall have

$$mv = B, \qquad mk^3\omega = Ba, \ldots (4).$$

Again, the velocity of A along Ox will be equal to

$$v + a\omega$$

the former term being due to the motion of G, and the latter to the rotation about G; but the velocity of A is also u by the hypothesis; hence v + au = u;

but, from the equations (4), we have $k^2\omega = av$; we obtain, therefore, a^2 k^2u au

$$u = v + \frac{a^2}{k^2}v$$
, $v = \frac{k^2u}{a^2 + k^2}$, $\omega = \frac{au}{a^2 + k^2}$.

Now,
$$\theta = \frac{1}{2}\pi$$
, $\frac{d\theta}{dt} = \omega$, simultaneously; hence, from (3),

$$mk^2\omega^2=C''+2mga;$$

and therefore

$$(a^{2} \cos^{2} \theta + k^{2}) \frac{d\theta^{2}}{dt^{2}} = k^{2} \omega^{2} - 2ga (1 - \sin \theta)$$

$$= \frac{k^{2} a^{2} u^{2}}{(a^{2} + k^{2})^{2}} - 2ga (1 - \sin \theta) \dots (5).$$

Also, the value of $\frac{dx}{dt}$ being constant, as is shewn by the equation (2),

$$\frac{dx}{dt} = v = \frac{k^2 u}{a^2 + k^2}, \quad x = \frac{k^2 u t}{a^2 + k^2};$$

which gives the velocity of G parallel to Ox, and the value of x at any time of the motion; the angular velocity of the beam for every position is given by (5), whence θ is to be ascertained in terms of t.

(4) An inelastic beam AB, (fig. 210), capable of moving in a vertical plane about a fixed horizontal axis through A, falls from a given position, and impinges against an immoveable obstacle at C; to determine the shock on the axis.

Let G be the centre of gravity of the beam; AM a horizontal line through A; m = the mass of the beam; $\angle GAM = \theta$ at any time t of the descent; a = the initial value of θ ; k = the radius of gyration about G; AG = a.

Then, for the motion of the beam in its fall,

$$m(a^2+k^2)\frac{d^2\theta}{dt^2}=mga\cos\theta;$$

multiplying by 2 $\frac{d\theta}{dt}$ and integrating

$$m(a^2+k^2)\frac{d\theta^2}{dt^2}=2mag\sin\theta+C;$$

but $\theta = a$ when $\frac{d\theta}{dt} = 0$; hence

$$0 = 2mag \sin a + C,$$

and therefore
$$(a^2 + k^2) \frac{d\theta^2}{dt^2} = 2ag (\sin \theta - \sin a)$$
.

Let $\angle CAM = \beta$ and, at the instant before collision, let $\frac{d\theta}{dt} = \omega$; then, clearly,

$$(a^2 + k^2) \omega^2 = 2ag (\sin \beta - \sin \alpha)$$
....(1).

Let R, R', denote the impulsive reactions of the obstacle C and the axis A, at the instant of impact; both of which will evidently be at right angles to the length of the beam. Now the effect of the reaction R is to destroy the whole of the angular velocity of the beam about A, by impressing upon it an equal and opposite angular velocity; hence, putting CA = c,

$$m_{\omega}(a^2+k^2)=Rc\ldots(2)$$

Again, the difference of the moments of R and R' about the centre of gravity of the beam being

$$R(c-a)-R'a$$

we must have

$$R(c-a)-R'a=mk^2\omega\ldots\ldots(3).$$

From (2) and (3) we obtain

$$m_{\omega}(a^2 + k^2)(c - a) - R'ac = mk^2c_{\omega},$$

 $R'ac = m_{\omega}\{(c - a)(a^2 + k^2) - ck^2\},$
 $R' = m_{\omega}\{a - \frac{a^2 + k^2}{c}\};$

and therefore, from (1),

$$R' = m (2ag)^{\frac{1}{2}} \left(\frac{\sin \beta - \sin \alpha}{a^2 + k^2} \right)^{\frac{1}{2}} \left(\alpha - \frac{\alpha^2 + k^2}{c} \right)$$

If R' = 0, we must have

$$a-\frac{a^2+k^2}{c}=0, \ c=\frac{a^2+k^2}{a};$$

and therefore C must be the centre of oscillation of the beam at the moment of collision.

If the beam be elastic we must increase the value of R given by (2) in the ratio of 1 + e to 1, e denoting the elasticity; we shall then have, from (3),

$$R' = \frac{m\omega}{ac} \left\{ (1+c)(c-a)(a^2+k^2) - ck^2 \right\}.$$

(5) An inelastic beam, which is moving without rotation along a smooth horizontal plane, impinges upon a fixed rod at right angles to the plane; to determine the impulsive reaction of the rod and the motion of the beam subsequent to the impact.

Let AB (fig. 211) be the position of the beam at the instant of impact; O the place of the obstacle; G the centre of gravity of the beam; G'G the line of G's motion before collision. Produce OB indefinitely to x, and draw the indefinite line yOy' at right angles to Ox and meeting GG' in G'. Let R = the impulsive reaction of O, which will be exerted along the line Oy'; w = the velocity of G before impact; LOG'G = a; LOG = c; LOG'G = a; LO

Then, by the equations of impulsive motion,

Again, the velocity of the point O of the beam in the direction Oy, the instant after impact, must be $v_y - c\omega$, v_y being its velocity due to the velocity of G, and $-c\omega$ its velocity due to the rotation of the beam about G; but, the beam being inelastic, the effect of the impact is to destroy the resolved part of O's velocity at right angles to AB; hence v_y must be equal to $c\omega$.

We have, then, from (2),

$$mc_{\omega} = mu \cos \alpha - R$$
,

and therefore, by the aid of (3),

$$mc^2\omega = mcu \cos \alpha - mk^2\omega$$
,

or

$$\omega = \frac{cu \cos a}{c^2 + k^2}.$$

Hence, from (3),

$$R=\frac{mk^3u\,\cos\,a}{c^2+k^2}\,;$$

and consequently, from (2),

$$v_y = u \cos a - \frac{k^3 u \cos a}{c^3 + k^2} = \frac{c^2 u \cos a}{c^2 + k^2}$$
.

IMPULSIVE FORCES.

Also, from (1), $v_z = u \sin a$.

Thus we have determined completely the instantaneous motions of the beam after the impact, and the impulsive reaction of the rod at O.

y be ascertained that, if the original motion be precisely as our particular figure represents it, on the consummation to impact, the beam will detach itself from the obstacle and then move along freely with the velocities v_x , v_y , ω , which we have obtained above. In fact we should find, if we were to assume the beam:

obstacle, that the obstacle would have to exert a section instead of a reaction.

SECT. 2. Seve

Smooth Surfaces.

(1) A heavy s altitude upon a 1 smooth inclined plane; the 1 smooth horizontal plane, same vertical plane contains.

falls down from a given oper surface of which is a capable of sliding along a arface being flat; also the res of gravity both of the

sphere and of the body: to determine the initial motions of the sphere and of the body, both of which are supposed to be perfectly inelastic.

Let ABH denote the section of the body made by a plane passing through its centre of gravity, AH being a line in the horizontal plane.

Let V be the velocity of the sphere just before impact; u, v, the resolved parts of its velocity after impact, perpendicular and parallel to the hypotenuse BA of the triangle BAH; u' the velocity of the body parallel to AH after the impact; m, m', the respective masses of the sphere and body, and B the blow of collision; $\angle BAH = a$.

Then, for the motion of the sphere, the blow being at right angles to BA, $mu = mV \cos a - B \dots (1)$,

 $v = V \sin \alpha \dots (2);$

and, for the motion of the body,

 $m'u' = B \sin \alpha \dots (3)$

These three equations involve four unknown quantities, u, t,

 \mathbf{z}' , B: for the solution of the problem, then, another equation will be necessary. This will be obtained by the consideration that, the sphere and the body being both perfectly inelastic, the effect of their collision is merely to prevent the penetration of the one into the interior of the other, without causing any recoil, which could result only from the existence of elasticity: hence the velocity of the ball, after collision, at right angles to the line BA must be equal to the velocity of any fixed point in this line estimated in the same direction.

Now the velocity of any fixed point in BA at right angles to this line is evidently $u' \sin a$; and therefore we have

$$u'\sin\alpha=u$$
.....(4).

From the equations (1), (3), (4), we obtain

$$mB \sin^2 a = mm' V \cos a - m'B$$
,

and therefore

which gives the magnitude of the blow.

From (3) and (5), we get

$$u' = \frac{mV \sin a \cos a}{m \sin^2 a + m'},$$

which determines the motion of the body; and therefore, by (4),

$$u=\frac{mV\sin^2\alpha\cos\alpha}{m\sin^2\alpha+m'}.$$

D'Arcy; Mémoires de l'Académie des Sciences de Paris, 1747, p. 344.

(2) Two billiard balls B and C (fig. 213) are lying in contact on the table; to find the direction in which the ball B must be struck by a third ball A so as to go off in a given direction BD; the balls being of equal volume and weight, and perfectly smooth.

Let the direction AB, which joins the centres of A and B at the instant of their collision, make an angle θ with the straight line CBE, and let $\angle CBD = a$. We will first suppose the balls to be inelastic. Let a, a', denote the resolved parts of the velocity of A in the direction AB, before and after collision respectively; and b, c, the velocities of B, C. Let m represent

the mass of each of the balls, R the blow between A, B, and S that between B, C. Then, for the motion of B, resolving forces at right angles to EC. $mb \sin a = R \sin \theta \dots (1),$ and, resolving parallel to EC. $mb \cos a = R \cos \theta - S \dots (2)$. Also, for the motion of C, we have $mc = S \dots (3)$ Again, since after collision the velocities of B and C in the direction EC must be equal, there is b cos a = c..... From (8) and (4) we get $mb \cos a = S_1$ and therefore, from (2), $mb \cos a = R \cos \theta - mb \cos a$, $R \cos \theta = 2mb \cos \alpha \dots$. From (1) and (5) there is mb sin a cos $\theta = 2mb$ cos a sin θ , and therefore $\tan \theta = \frac{1}{2} \tan a,$

which determines the point in which A must come into collision with B.

If we introduce the consideration of elasticity, the magnitudes of R and S will each have to be increased in the ratio of 1 + eto 1. Now the direction of B's motion will evidently not be affected by any alteration in the absolute magnitudes of R and S, provided that the ratio between their intensities be not changed. Thus we see that the consideration of elasticity will not modify the solution of the problem.

(3) A billiard ball impinges simultaneously upon two other billiard balls which are resting in contact; to determine the motions of the three balls after collision.

Let A (fig. 214) denote the centre of the impinging ball at the moment of impact; A', A'', of the other two balls. AA', AA'', and produce them indefinitely to points a', a'; draw Aa a common tangent to the two balls A', A''. evidently after collision A's motion will be confined to the straight line Aa; while A', A'', will proceed to move along A'a', A''a''. Let u, v, be the velocities of A before and after impact; v' the velocity after impact of each of the balls A', A''. Since AA'A'' is evidently an equilateral triangle, $\angle aAa' = \frac{1}{6}\pi = \angle aAa''$; let B be the blow of collision between A, A', and A, A''; and m the mass of each of the three balls. Then

$$mv = mu - 2B \cos \frac{1}{6}\pi \dots (1),$$

 $mv' = B \dots (2).$

Let us first suppose the three balls to be perfectly inelastic; then the instant after impact the balls A, A', will move in contact, as well as the balls A, A''; hence we must evidently have $v' = v \cos \frac{1}{8} \pi$;

we have, therefore, from (2),

$$B=mv\,\cos\frac{\pi}{a};$$

and consequently, from (1),

$$mv = mu - 2mv \cos^2 \frac{\pi}{6}$$
, $v = \frac{u}{1 + 2 \cos^2 \frac{\pi}{6}} = \frac{2u}{5}$;

and therefore

$$v'=v\,\cos\frac{\pi}{6}=\frac{3^4}{5}\,u,$$

and

$$B=\frac{3^{\frac{1}{4}}mu}{5}\ldots\ldots(3).$$

If the balls be elastic, and e denote their elasticity, we must increase the value of B in (3) in the ratio of 1 + e to 1; hence we have

$$B = (1 + e) \frac{3^{\frac{1}{4}} mu}{5} \dots (4),$$

and therefore, from (1),

$$mv = mu - \frac{3}{5}(1 + e) mu,$$

 $v = \frac{1}{5}(2 - 3e) u;$

and from (2), (4), $mv' = (1 + e) \frac{3^{\frac{1}{4}} mu}{5}$,

$$v' = \frac{3^{\frac{1}{4}}}{5}(1+e) u.$$

Maclaurin; Treatise of Fluxions. D'Alembert; Traité de Dynamique, p. 227.

(4) A ball C (fig. 215) impinges upon an inelastic beam AB with a given velocity, at right angles to its length; to determine the magnitude of the blow and the initial circumstances of the motion of the beam and ball.

Let G be the centre of gravity of the beam; m' its mass, k the radius of gyration about G; EG = a, u = the velocity of the ball before impact, v = its velocity immediately afterwards; R = the magnitude of the mutual impulse: let v' be the velocity of G and ω the angular velocity of the beam about G just after collision.

Then, for the initial motion of the ball after collision, m denoting its mass, $mv = mu - R \dots (1)$;

and, for the initial motion of the beam,

$$m'v' = R \dots (2),$$

 $m'k^2\omega = Ra \dots (3).$

Again, the velocity of the point E of the beam will be equal to $v' + a\omega$,

the former term of the expression being due to the motion of G, and the latter to the rotation about G; but, the beam and ball being inelastic, the velocity of the point E of the beam after collision must be equal to that of the point E of the ball, and therefore of the point C: hence we have

$$v=v'+a \ldots (4).$$

From (1), (2), (3), (4), we obtain

$$u-\frac{R}{m}=\frac{R}{m'}+\frac{Ra^2}{m'k^2},$$

and therefore

$$R = \frac{u}{\frac{1}{m} + \frac{1}{m'} + \frac{a^2}{m'k^2}} = \frac{mm'k^3u}{(m+m') \ k^2 + ma^2}.$$

Hence, from (1), we get

$$v = u - \frac{m'k^2u}{(m+m')k^2 + ma^2} = \frac{m(k^2 + a^2)u}{(m+m')k^2 + ma^2},$$
from (2),
$$v' = \frac{mk^2u}{(m+m')k^2 + ma^2},$$

and, from (3),
$$\omega = \frac{mau}{(m+m') k^2 + ma^2}.$$

(5) A cylinder is revolving with a given angular velocity round an axis which is horizontal, when it suddenly begins to draw up a weight, consisting of inelastic materials, by means of an inextensible string wound round the cylinder; to determine the time the system will continue in motion, and the original distance of the weight from the cylinder, that at the instant the motion ceases the weight may just touch the cylinder.

Let a = the radius of the cylinder, m = its mass, k^2 = the radius of gyration about its axis; m' = the mass of the weight; let ω , ω' , denote the angular velocities of the cylinder just before and just after beginning to draw up the weight; u = the velocity of the weight at the commencement of its motion, B = the impulsive force exerted initially by the string on the weight.

Then we shall have

$$\omega' = \omega - \frac{Ba}{mk^3}, \quad u = \frac{B}{m'};$$

but $u = a\omega'$; hence

$$\frac{B}{m'} = a\omega - \frac{Ba^2}{mk^2},$$

and therefore

$$B = \frac{mm'a\omega k^2}{m'a^2 + mk^2}, \qquad u = \frac{ma\omega k^2}{m'a^2 + mk^2} \cdot \dots (1).$$

Let θ denote the angle through which the cylinder has revolved about its axis at the end of the time t from the commencement of the raising of the weight; and let x be the corresponding distance of the weight below the horizontal plane through the axis of the cylinder. Then, by the Principle of the Conservation of Vis Viva,

$$m'\frac{dx^3}{dt^3} + mk^3\frac{d\theta^3}{dt^3} = 2m'gx + C \dots (2)$$
:

but if b denote the value of x at the commencement of the raising of the weight, it is clear from the geometry that

$$x + a\theta = b$$
, and therefore $a \frac{d\theta}{dt} = -\frac{dx}{dt}$;

hence, from (2),

$$\frac{m'a^2+mk^2}{a^2}\frac{dx^2}{dt^2}=2m'gx+C;$$

differentiating with respect to t, and dividing by $2\frac{dx}{dt}$,

$$(m'a^2 + mk^2)\frac{d^2x}{dt^2} = m'a^2g;$$

integrating with respect to t, we have

$$(m'a^2 + mk^2)\frac{dx}{dt} = C + m'a^2gt;$$

, but $\frac{dx}{dt} = -u$ when t = 0; hence $C = -(m'a^2 + mk^2)u$, or, by (1),

 $C = -ma_{\omega}k^2$, and therefore

$$(m'a^2 + mk^2) \frac{dx}{dt} = m'a^2gt - mawk^2 \dots (3);$$

integrating again with respect to t, we get

$$(m'a^2 + mk^2)x = C' + \frac{1}{2}m'a^2gt^2 - ma\omega k^2t;$$

but x=b when t=0; hence $C'=(m'a^2+mk^2)b$, and therefore

$$(m'a^2 + mk^3) x = (m'a^3 + mk^3) b + \frac{1}{2} m'a^3gt^3 - mawk^2t \dots (4).$$

Let t' denote the time when the motion ceases for an instant; then, from (3), since $\frac{dx}{dt} = 0$ when t = t',

$$0 = m'a^2gt' - ma\omega k^2, \qquad t' = \frac{m\omega k^2}{m'ag}.$$

Hence also, from (4), since x = 0 when t = t',

$$(m'a^2 + mk^2) b = ma\omega k^2 t' - \frac{1}{2} m'a^2 g t'^2 = \frac{m^2 \omega^2 k^4}{2 m' j},$$

which gives the required value of b.

(6) Two inelastic spheres, of which A and a (fig. 216) are the centres, are attached to rigid rods CA and ca, which are capable of motion in a single plane about axes through C and c at right angles to the plane; supposing the spheres to impinge against each other with given velocities, it is required to determine their initial velocities after impact.

Join Aa, and produce it indefinitely both ways to points a, β ; from C, c, draw CG, cg, at right angles to $\alpha\beta$ at the

moment of collision. Let C, c, represent the lines CG, cg, Ω , ω , the angular velocities of CA, ca, about C, c, respectively immediately before, and Ω' , ω' , immediately after collision, the angular motions being estimated in the directions indicated by the arrows in the figure; B the blow of collision; I the moment of inertia of the sphere A with its rod AC about the axis through C, i the moment of inertia of the other sphere and rod about c.

Then, $\Omega' - \Omega$ being the angular velocity which CA gains, and $\omega - \omega'$ that which ca loses by the shock, we shall have

$$I(\Omega'-\Omega)=B.C, \quad i(\omega-\omega')=B.c...(1).$$

But, the spheres being inelastic, the point P of the sphere a will, the instant after collision, have the same velocity along $a\beta$ as the point P of the sphere A: hence

$$\Omega'.CP. \sin \angle CPG = \omega'.cP. \sin \angle cPg,$$

 $\Omega'C = \omega'c....(2).$

Now, from the equations (1),

$$cI(\Omega'-\Omega)+Ci(\omega'-\omega)=0,$$

and therefore, by (2), we get

$$cI(c\omega'-C\Omega)+C^2i(\omega'-\omega)=0,$$

$$(c^2I + C^2i) \omega' = C(cI\Omega + Ci\omega),$$

and therefore $(c^2I + C^2i)\Omega' = c(cI\Omega + Ci\omega);$

which two equations give the values of Ω' , ω' .

The solution of a particular case of this problem was unsuccessfully attempted by John Bernoulli, son of the celebrated John Bernoulli, in the *Mémoires de St. Pétersbourg*, tom. VII.: the correct solution of the problem in its most general form, was given by D'Alembert, *Traité de Dynamique*, p. 221; seconde edition.

SECT. 3. Rough Surfaces.

(1) An inelastic cylinder O (fig. 217) having rolled down a perfectly rough plane CA, impinges upon a perfectly rough plane C'A, the axis to the cylinder being parallel to the intersection of the two planes; to find the velocity with which the cylinder will commence its ascent up the second

plane, and the limiting angle of inclination of the two planes for which the ascent is possible.

Let $\angle CAC' = a$, k = the radius of gyration of the cylinder about its axis, a = the radius of the cylinder, m = its mass; u =the velocity of the centre O of a circular section of the cylinder parallel to the plane CA just before impact, and v = the velocity after impact up the plane AC'; R = the impulsive force of friction exerted by the plane AC' upon the cylinder at the moment of impact to secure perfect rolling.

Then, for the motion of the centre of gravity of the cylinder parallel to AC, we have

$$mv = R - mu \cos a \dots (1);$$

and, for the value of the decrement of the angular velocity of the cylinder about its axis owing to the impulse R, we have the expression Ra

mk2 ,

which, by (1), is equal to

$$\frac{a\left(v+u\cos u\right)}{k^2};$$

but, the planes being both perfectly rough, it is evident that $\frac{u}{a}$ is the angular velocity before, and $\frac{v}{a}$ after the impact; hence

$$\frac{u}{a} - \frac{v}{a} = \frac{a (v + u \cos a)}{k^2};$$

$$u - v = \frac{a^2}{k^2} (v + u \cos a);$$

but $a^2 = 2k^2$; hence

$$u-v=2v+2u \cos a;$$

and therefore,

$$v = \frac{1}{3} \left(1 - 2 \cos a \right) u,$$

which gives the velocity with which the cylinder ascends the plane AC'.

Since v can evidently, from the nature of the case, have neither a zero nor a negative value, it is clear that the smallest possible value of a for the ascent is given by the equation

$$1-2\cos\alpha=0,\quad\cos\alpha=\frac{1}{2};$$

whence $\frac{1}{2}\pi$ is the limiting value of a.

(2) A ball sliding without rotation along a smooth horizontal plane impinges obliquely against a perfectly rough vertical plane; to determine the subsequent motion of the ball.

Let Ox, Oy, Oz, (fig. 218), be three rectangular axes, the plane xOy being horizontal and passing through the centre C of the ball, and the plane xOz being the rough vertical plane against which the ball impinges. Let E be the point in which the ball strikes against the vertical plane; CF the direction of the motion of C before collision. Let u be the velocity of C before impact, a the inclination of CF to Ox; v_a , v_y , the resolved parts of the velocity of C parallel to Ox, Oy, after the impact; w the angular velocity of the ball about C after impact; X, Y, the impulsive reactions of the rough plane along xO, and parallel to Oy, during collision; m the mass of the ball; a the radius of the ball, k the radius of gyration about C.

First we will suppose the ball to be inelastic. For the motion of the ball after impact, we have

$$mv_{a} = mu \cos a - X \dots (1),$$

 $mv_{y} = Y - mu \sin a \dots (2),$
 $mk^{2}\omega = Xa \dots (3).$

Now, the ball being perfectly inelastic, the velocity of C at right angles to the vertical plane will be destroyed by the impact, or $v_{\bullet} = 0$; hence, from (2),

$$Y = mu \sin a \dots (4)$$
.

Also, the vertical plane being perfectly rough, the ball must roll without sliding after impact; hence we must have $a\omega = v_s$, and therefore from (1), (3), we get

Next, let us suppose the ball to be elastic, e denoting the clasticity; then v_a' , v_y' , ω' , denoting on the new supposition what v_a , v_a , ω , were taken to denote on the old one, we shall have

$$mv'_{\bullet} = mu \cos a - (1 + e) X \dots (6),$$

 $mv'_{\bullet} = mu \sin a - (1 + e) Y \dots (7),$
 $mk''_{\bullet} = (1 + e) Xa \dots (8),$

From the equations (5) and (6), we obtain

$$v'_{a} = u \cos \alpha - \frac{(1+e)k^{2}u \cos \alpha}{a^{2}+k^{2}} = \frac{a^{2}-ek^{2}}{a^{2}+k^{2}}u \cos \alpha;$$

from (4) and (7),

$$v' = u \sin \alpha - (1 + e) u \sin \alpha = -eu \sin \alpha$$
;

and from (5), (8),

$$\omega' = \frac{(1+e) \ au \cos \alpha}{a^2 + k^2};$$

which values of v_{s}' , v_{s}' , ω' , completely determine the subsequent motion of the ball.

(3) A perfectly rough sphere is placed upon a perfectly rough horizontal plane which is made to rotate with a uniform angular velocity about a vertical axis; to determine the path described by the sphere in space.

Let Ox (fig. 219) be the vertical axis about which the plane revolves; let Ox, Oy, be any two horizontal lines fixed in space and at right angles to each other; P the point of contact of the sphere with the revolving plane at any time t; draw PM parallel to yO. Let OM = x, PM = y; a = the radius of the sphere, m = its mass, $mk^2 =$ its moment of inertia about a diameter; $\omega =$ the angular velocity of the revolving plane about Ox, the motion being supposed to take place in the direction indicated by the arrow in the plane xOy in the figure; let X, Y, denote the resolved parts of the friction exerted by the plane on the sphere, estimated parallel to Ox, Oy, respectively; ω' , ω' , the angular velocities of the sphere about diameters parallel to Ox, Oy, the directions of these velocities being estimated in the manner indicated by the arrows in the planes yOz, zOx.

For the motion of the centre of gravity of the sphere we have

$$m \frac{d^3x}{dt^3} = X \dots (1),$$

$$m \frac{d^2y}{dt^2} = Y \dots (2);$$

and, for the rotation of the sphere,

$$mk^2\frac{d\omega'}{dt}=Ya....(3),$$

$$mk^2\frac{d\omega''}{dt}=-Xa.....(4).$$

From (1) and (4) we have, eliminating X,

and from (2), (3), eliminating Y,

$$a \frac{d^3y}{dt^3} = k^3 \frac{d\omega'}{dt} \dots (6).$$

Integrating the equations (5), (6), and adding arbitrary constants, we have

$$a\frac{dx}{dt}=C-k^2\omega'....(7),$$

$$a\frac{dy}{dt}=C'+k^2\omega',\ldots(8).$$

Now the linear velocity of the centre of gravity of the sphere relatively to the rough plane, in consequence of the rolling of the sphere, will be

 $a\omega'$ parallel to Ox, $-a\omega'$ parallel to Oy;

and the linear velocity due to the rotation of the rough plane will be $-\omega y$ parallel to Ox, ωx parallel to Oy;

hence, $\frac{dx}{dt}$, $\frac{dy}{dt}$, being the whole linear velocity of the centre of the sphere parallel to Ox, Oy, respectively, we have

$$\frac{dx}{dt} = -\omega y + a\omega', \quad \frac{dy}{dt} = \omega x - a\omega';$$

and therefore, by the aid of (7) and (8), eliminating ω' and ω' ,

$$a\frac{dx}{dt} = C - \frac{k^2}{a} \left(\omega y + \frac{dx}{dt} \right), \quad \left(1 + \frac{a^2}{k^2} \right) \frac{dx}{dt} = \frac{aC}{k^2} - \omega y,$$

$$a\frac{dy}{dt}=C'+\frac{k^2}{a}\left(\omega x-\frac{dy}{dt}\right), \left(1+\frac{a^2}{k^2}\right)\frac{dy}{dt}=\frac{a\,C'}{k^2}+\,\omega x\,;$$

eliminating t between these two equations, we get

or

$$(aC - k^2\omega y) dy = (aC' + k^2\omega x) dx:$$

integrating and adding an arbitrary constant C', we obtain

$$2aCy - k^3\omega y^2 = 2aC'x + k^2\omega x^3 + C'',$$

$$x^3 + y^2 + \frac{2aC'}{k^2\omega}x - \frac{2aC}{k^2\omega}y + \frac{C'}{k^2\omega} = 0.....(9).$$

We proceed now to the determination of the arbitrary constants. Let the initial distance of the centre of the sphere from the axis of x be δ , and let the axis of x be so chosen as to pass through the initial position of the point of contact of the sphere with the rough plane. Then, since the initial impulse of the friction of the revolving plane upon the sphere is at right angles to the axis of x, we shall have initially $\frac{dx}{dt} = 0$, $\omega^* = 0$; hence from (7) we see that C = 0. Again, F denoting the impulse of the friction when the sphere is just placed upon the revolving plane, and $\left(\frac{dy}{dt}\right)$, (ω), denoting the values of $\frac{dy}{dt}$, ω , just after the impulse, we shall have

$$m\left(\frac{dy}{dt}\right) = F \dots (10),$$

 $mk^2(\omega) = Fa \dots (11).$

But, since there is no sliding between the sphere and the plane, it is clear that

$$\left(\frac{dy}{dt}\right) = b\omega - a(\omega') \dots (12),$$

where b_{ω} is the velocity of the centre of the sphere parallel to Ox due to the rotation of the plane, and $-a(\omega')$ the velocity estimated in the same direction due to the rolling of the sphere along the plane: hence from (10) we have

$$m \{b\omega - a(\omega')\} = F,$$

and therefore, by (11),

$$a\{b\omega-a(\omega')\}=k^2(\omega'), \quad (\omega')=\frac{ab\omega}{a^2+k^2};$$

and then, by (12),

$$\left(\frac{dy}{dt}\right) = b\omega - \frac{a^2b\omega}{a^2 + k^2} = \frac{k^2b\omega}{a^3 + k^3}.$$

But from (8) we have

$$a\left(\frac{dy}{dt}\right)=C'+k^2(\omega');$$

hence, putting for $\left(\frac{dy}{dt}\right)$ and (ω') their values,

$$\frac{abk^2\omega}{a^2+k^2}=C'+\frac{abk^2\omega}{a^2+k^2},\quad C'=0.$$

Since C = 0 and C' = 0, we have, from (9),

$$x^{2}+y^{2}+\frac{C''}{k^{2}\omega}=0;$$

but x = b when y = 0, and therefore

$$b^2 + \frac{C''}{k^2\omega} = 0,$$

and we get for the equation to the path of the centre of the sphere in space $x^2 + y^2 = b^2$,

the equation to a circle having O for its centre.

(4) An inelastic homogeneous cylinder rolls down a perfectly rough inclined plane which terminates in a perfectly rough horizontal plane; to find the velocity of the cylinder along the horizontal plane, and the blow which it receives when it first impinges upon it.

Let a = inclination of the inclined plane to the horizon, m = the mass of the cylinder, u = the velocity of its axis the instant before and u' the instant after impact, F = the initial impulse of the friction of the horizontal plane, estimated in the direction of the sphere's motion along it, and B the normal impulse of the horizontal plane; then

$$u' = \frac{1}{3}(1 + 2\cos a)u$$
, $F = \frac{1}{3}mu(1 - \cos a)$, $B = mu\sin a$.

(5) A homogeneous cylinder slides, without rolling, down an inclined plane which is for a certain space quite smooth, and after acquiring a given velocity is suddenly caused by the roughness of the surface to roll without sliding: to determine the velocity of the axis of the cylinder the instant rolling commences, and to find the initial impulse of friction.

If u be the velocity of the cylinder the instant before and u the instant after the commencement of perfect rolling, m the mass of the cylinder, F the initial impulse of friction; then

$$u' = \frac{2}{3}u, \quad F = \frac{1}{3}mu.$$

APPENDIX.

A cory of Bernoulli's programme, which had been received by the celebrated David Gregory, is at the present time in the possession of Mr. Gregory, Fellow of Trinity College. The following reprint of the challenge will probably be acceptable to those who take an interest in the antiquities of science.

Acutissimis qui toto orbe florent Mathematicis.

S. P. D.

JOHANNES BERNOULLI, Math. P. P.

"Cum compertum habeamus, vix quicquam esse quod magis excitet generosa ingenia ad moliendum quod conducit augendis scientiis, quam difficilium pariter et utilium quæstionum propositionem, quarum enodatione tanquam singulari si qua alia via ad nominis claritatem perveniant sibique apud posteritatem æterna extruant monumenta: Sic me nihil gratius Orbi Mathematico facturum speravi quam si imitando exemplum tantorum Virorum Mersenni, Pascalii, Fermatii, præsertim recentis illius Anonymi Ænigmatistæ Florentini aliorumque qui idem ante me fecerunt, præstantissimis hujus ævi Analystis proponerem aliquod problema, quo quasi Lapide Lydio suas methodos examinare, vires intendere et si quid invenirent nobiscum communicare possent, ut quisque suas exinde promeritas laudes à nobis publicè id profitentibus consequeretur.

"Factum autem illud est ante semestre in Actis Lips. m. Jun. pag. 269. Ubi tale problema proposui cujus utilitatem cum jucunditate conjunctam videbunt omnes qui cum successu ei se applicabunt. Sex mensium spatium à prima publicationis die Geometris concessum est, intra quod si nulla solutio prodiret in lucem, me meam exhibiturum promisi: Sed ecce elapsus est

terminus et nihil solutionis comparuit; nisi quod Celeb. Leibnitius de profundiore Geometriâ præclarè meritus me per literas certiorem fecerit, se jam feliciter dissolvisse nodum pulcherrimi hujus uti vocabat et inauditi antea problematis, insimulque humaniter rogavit, ut præstitutum limitem ad proximum pascha extendi paterer, quo interea apud Gallos Italosque idem illud publicari posset nullusque adeo superesset locus ulli de angustiâ termini querelæ; Quam honestam petitionem non solum indulsi, sed ipse hanc prorogationem promulgare decrevi, visurus num qui sint qui nobilem hanc et arduam quæstionem aggressuri, post longum temporis intervallum tandem Enodationis compotes fierent. Illorum interim in gratiam ad quorum manus Acta Lipsiensia non perveniunt, propositionem hic repeto.

PROBLEMA MECHANICO-GEOMETRICUM DE LINEA CELERRIMI DESCENSUS.

"Determinare lineam curvam data duo puncta in diversis ab horizonte distantiis et non in eadem rectà verticali posita connectentem, super qua mobile proprià gravitate decurrens et à superiori puncto moveri incipiens citissime descendat ad punctum inferius.

"Sensus problematis hic est, ex infinitis lineis quæ duo illa data puncta conjungunt, vel ab uno ad alterum duci possunt eligatur illa, juxta quam si incurvetur lamina tubi canalisve formam habens, ut ipsi impositus globulus et liberè dimissus iter suum ab uno puncto ad alterum emetiatur tempore brevissimo.

"Ut vero omnem ambiguitatis ansam præcaveamus, scire B.L. volumus, nos hîc admittere Galilæi hypothesin de cujus veritate sepositâ resistentiâ jam nemo est saniorum Geometrarum qui ambigat, Velocitates scilicet acquisitas gravium cadentium esse in subduplicata ratione altitudinum emensarum, quanquam aliàs nostra solvendi methodus universaliter ad quamvis aliam hypothesin sese extendat.

"Cum itaque nihil obscuritatis supersit, obnixè rogamus omnes et singulos hujus ævi Geometras, accingant se promtè, tentent, discutiant quicquid in extremo suarum methodorum recessu absconditum tenent; Rapiat qui potest præmium quod Solutori paravimus, non quidem auri non argenti summam quo

abjecta tantum et mercenaria conducuntur ingenia, à quibus ùt nihil laudabile sic nihil quod scientiis fructuosum expectamus, sed cùm virtus sibi ipsi sit merces pulcherrima, atque gloria immensum habeat calcar, offerimus præmium quale convenit ingenui sanguinis Viro, consertum ex honore, laude et plausu, quibus magni nostri Apollinis perspicacitatem publicè et privatim, scriptis et dictis coronabimus, condecorabimus et celebrabimus.

"Quod si verò festum paschatis præterierit nemine deprehenso qui quæsitum nostrum solverit, nos quæ ipsi invenimus publica non invidebimus; Incomparabilis enim Leibnitius solutiones tum suam tum nostram ipsi jam pridem commissam protinus ut spero in lucem emittet, quas si Geometræ ex penitiori quodam fonte petitas perspexerint, nulli dubitamus quin angustos vulgaris Geometriæ limites agnoscant, nostraque proin inventa tanto pluris faciant, quanto pauciores eximiam nostram quæstionem soluturi extiterint etiam inter illos ipsos qui per singulares quas tantopere commendant methodos, interioris Geometriæ latibula non solum intimè penetrâsse, sed etiam ejus pomæria Theorematis suis aureis, nemini ut putabant cognitis, ab aliis tamen jam longè priùs editis mirum in modum extendisse gloriantur.

PROBLEMA ALTERUM PURE GEOMETRICUM, QUOD PRIORI SUB-NECTIMUS ET STRENÆ LOCO ERUDITIS PROPONIMUS.

"Ab Euclidis tempore vel Tyronibus notum est; Ductam utcunque à puncto dato rectam lineam, à circuli peripherià ita secari ut rectangulum duorum segmentorum inter punctum datum et utramque peripheriæ partem interceptorum sit eidem constanti perpetuo æquale. Primus ego ostendi in eod. Actor. Jun. pag. 265. hanc proprietatem infinitis aliis curvis convenire, illamque adeo circulo non esse essentialem. Arrepta hinc occasione, proposui Geometris determinandam curvam vel curvas, in quibus non rectangulum sed solidum sub uno et quadrato alterius segmentorum æquetur semper eidem; sed à nemine hactenus solvendi modus prodiit; exhibebimus eum quandocunque desiderabitur: Quoniam autem non nisi per curvas transcendentes quæsito satisfacimus, en aliud cujus solutio per merè algebricas in nostra est potestate. Quæritur Curva, ejus proprietatis, ut

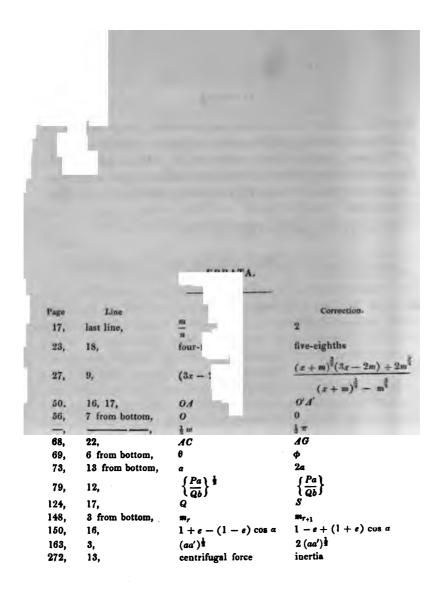
duo illa segmenta ad quamcunque potentiam datam elevata ct simul sumta faciant ubique unam eandemque summam.

"Casum simplicissimum existente sc. numero potentiæ 1, ibidem in actis pag. 266. jam solutum dedimus, generalem verò solutionem quam etiamnum premimus, Analystis eruendam relinquimus.

Dabam Groningæ ipsis Cal. Jan. 1697.

Groningæ, Typis Catharinæ Zandt, Provincialis Academiæ Typographæ, 1697.

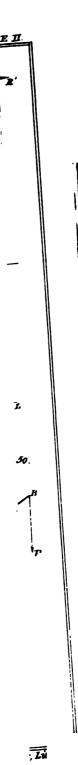
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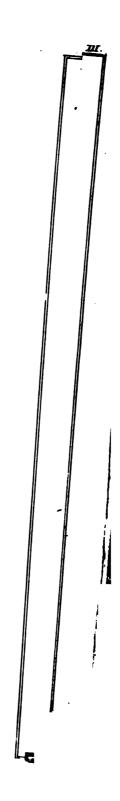
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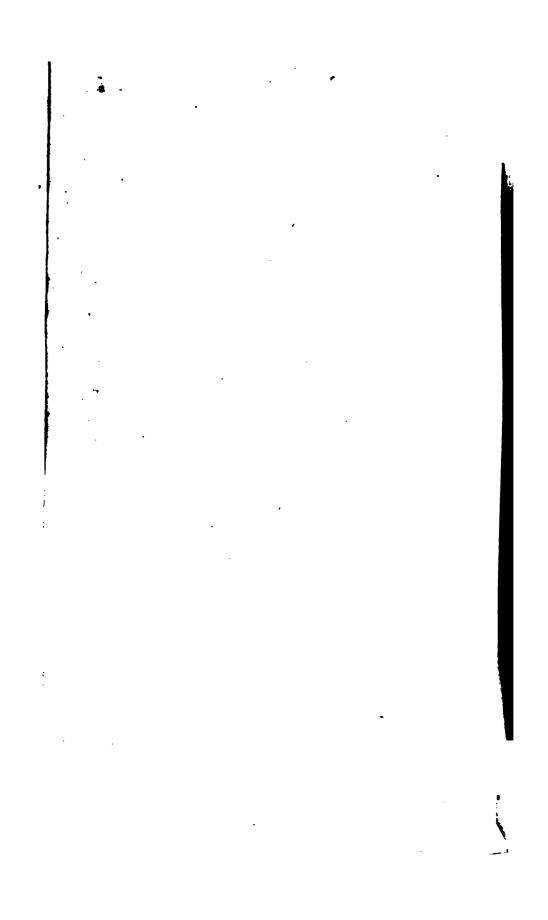




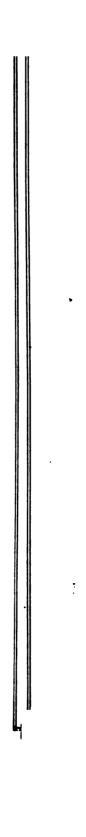


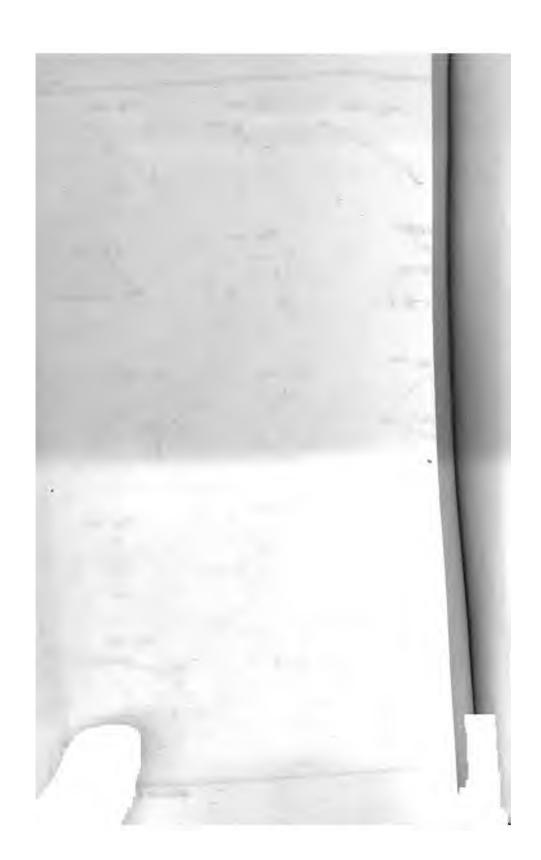


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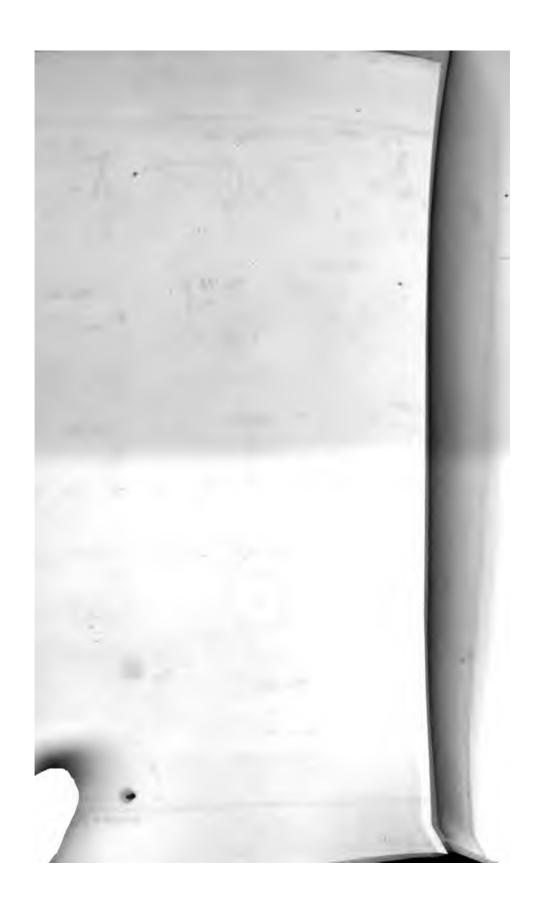


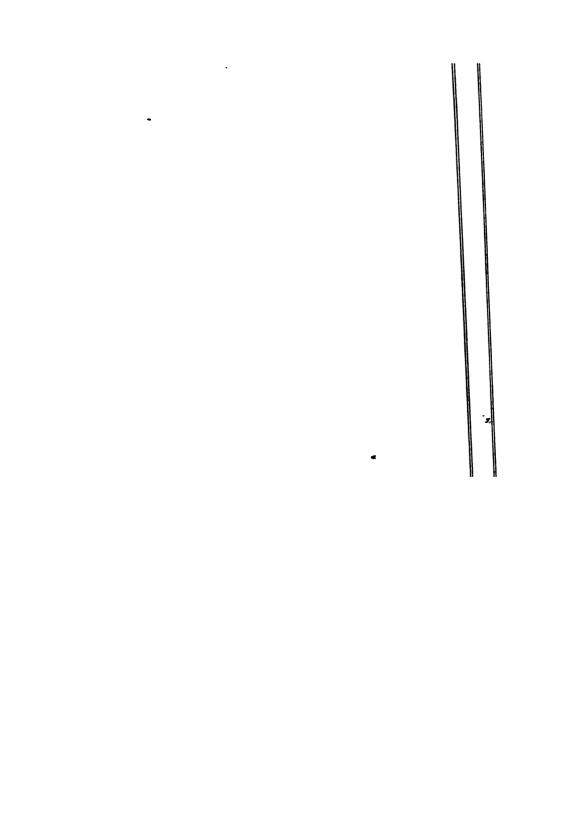


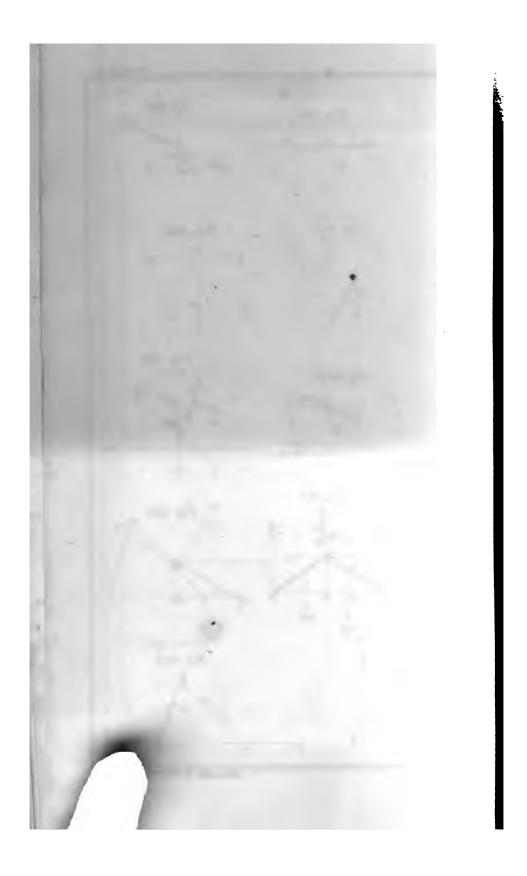




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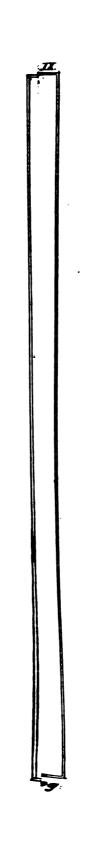












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